

$\omega(1650)$   
was  $\omega(1600)$

$I^G(J^{PC}) = 0^-(1^{--})$

### $\omega(1650)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1649 ± 24 OUR AVERAGE</b>		Error includes scale factor of 2.3.			
1609 ± 20	315	1 ANTONELLI	92 DM2	1.34–2.4e <sup>+</sup> e <sup>-</sup> → $\rho\pi$	
1663 ± 12	435	2 ANTONELLI	92 DM2	1.34–2.4e <sup>+</sup> e <sup>-</sup> → $\omega\pi\pi$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1700 ± 20		EUGENIO	01 SPEC	18 π <sup>-</sup> p → $\omega\eta n$	
1652 ± 35		3 AKHMETSHIN	00D CMD2	e <sup>+</sup> e <sup>-</sup> →	
1705 ± 26		4 AKHMETSHIN	00D CMD2	e <sup>+</sup> e <sup>-</sup> →	
1643 ± 14		5 ACHASOV	99E RVUE	0.75–1.80 e <sup>+</sup> e <sup>-</sup> → $\pi^+\pi^-\pi^0$	
1820 <sup>+190</sup> -150		6 ACHASOV	98H RVUE	e <sup>+</sup> e <sup>-</sup> → $\pi^+\pi^-\pi^0$	
1840 <sup>+100</sup> -70		7 ACHASOV	98H RVUE	e <sup>+</sup> e <sup>-</sup> → $\omega\pi^+\pi^-$	
1780 <sup>+170</sup> -300		8 ACHASOV	98H RVUE	e <sup>+</sup> e <sup>-</sup> → $K^+K^-$	
~2100		9 ACHASOV	98H RVUE	e <sup>+</sup> e <sup>-</sup> → $K_S^0 K^\pm \pi^\mp$	
1600 ± 30		1 CLEGG	94 RVUE	e <sup>+</sup> e <sup>-</sup> → $\rho\pi$	
1607 ± 10		2 CLEGG	94 RVUE	e <sup>+</sup> e <sup>-</sup> → $\omega\pi\pi$	
1635 ± 35		10 CLEGG	94 RVUE	e <sup>+</sup> e <sup>-</sup> → $\rho\pi$	
1625 ± 21		10 CLEGG	94 RVUE	e <sup>+</sup> e <sup>-</sup> → $\omega\pi\pi$	
1670 ± 20		ATKINSON	83B OMEG	20–70 γp → $3\pi X$	
1657 ± 13		CORDIER	81 DM1	e <sup>+</sup> e <sup>-</sup> → $\omega 2\pi$	
1679 ± 34	21	ESPOSITO	80 FRAM	e <sup>+</sup> e <sup>-</sup> → 3π	
1652 ± 17		COSME	79 OSPK 0	e <sup>+</sup> e <sup>-</sup> → 3π	

<sup>1</sup> From a two Breit-Wigner fit.

<sup>2</sup> From a single Breit-Wigner plus background fit.

<sup>3</sup> Using the data of AKHMETSHIN 00D and ANTONELLI 92. The energy-independent width of the  $\omega(1420)$  and  $\omega(1650)$  mesons assumed.

<sup>4</sup> Using the data of AKHMETSHIN 00D and ANTONELLI 92. The  $\rho\pi$  dominance for the energy dependence of the  $\omega(1420)$  and  $\omega(1650)$  width assumed.

<sup>5</sup> Using the data of DOLINSKY 91, ANTONELLI 92, AKHMETSHIN 98, and ACHASOV 99E. From a fit to two Breit-Wigner functions interfering between them and with the  $\omega, \phi$  tails with fixed (+, -, +) phases.

<sup>6</sup> Using data from BARKOV 87, DOLINSKY 91, and ANTONELLI 92.

<sup>7</sup> Using the data from ANTONELLI 92.

<sup>8</sup> Using the data from IVANOV 81 and BISELLO 88B.

<sup>9</sup> Using the data from BISELLO 91C.

<sup>10</sup> From a single Breit-Wigner fit.

## $\omega(1650)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>220±35 OUR AVERAGE</b>	Error includes scale factor of 1.6.				
159±43	315	11 ANTONELLI	92 DM2		1.34–2.4e <sup>+</sup> e <sup>-</sup> → $\rho\pi$
240±25	435	12 ANTONELLI	92 DM2		1.34–2.4e <sup>+</sup> e <sup>-</sup> → $\omega\pi\pi$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
250±50		EUGENIO	01 SPEC		18 π <sup>-</sup> p → $\omega\eta n$
284±31		13 AKHMETSHIN 00D	CMD2		e <sup>+</sup> e <sup>-</sup> → $\omega\pi^+\pi^-$
370±25		14 AKHMETSHIN 00D	CMD2		e <sup>+</sup> e <sup>-</sup> → $\omega\pi^+\pi^-$
272±29		15 ACHASOV	99E RVUE		0.75–1.80 e <sup>+</sup> e <sup>-</sup> → $\pi^+\pi^-\pi^0$
140±50		11 CLEGG	94 RVUE		e <sup>+</sup> e <sup>-</sup> → $\rho\pi$
86±20		12 CLEGG	94 RVUE		e <sup>+</sup> e <sup>-</sup> → $\omega\pi\pi$
350±80		16 CLEGG	94 RVUE		e <sup>+</sup> e <sup>-</sup> → $\rho\pi$
401±63		16 CLEGG	94 RVUE		e <sup>+</sup> e <sup>-</sup> → $\omega\pi\pi$
160±20		ATKINSON	83B OMEG		20–70 γp → $3\pi X$
136±46		CORDIER	81 DM1		e <sup>+</sup> e <sup>-</sup> → $\omega 2\pi$
99±49	21	ESPOSITO	80 FRAM		e <sup>+</sup> e <sup>-</sup> → $3\pi$
42±17		COSME	79 OSPK 0		e <sup>+</sup> e <sup>-</sup> → $3\pi$

11 From a two Breit-Wigner fit.

12 From a single Breit-Wigner plus background fit.

13 Using the data of AKHMETSHIN 00D and ANTONELLI 92. The energy-independent width of the  $\omega(1420)$  and  $\omega(1650)$  mesons assumed.

14 Using the data of AKHMETSHIN 00D and ANTONELLI 92. The  $\rho\pi$  dominance for the energy dependence of the  $\omega(1420)$  and  $\omega(1650)$  width assumed.

15 Using the data of DOLINSKY 91, ANTONELLI 92, AKHMETSHIN 98, and ACHASOV 99E. From a fit to two Breit-Wigner functions interfering between them and with the  $\omega,\phi$  tails with fixed (+,-,+,-) phases.

16 From a single Breit-Wigner fit.

## $\omega(1650)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1 \rho\pi$	seen
$\Gamma_2 \omega\pi\pi$	seen
$\Gamma_3 \omega\eta$	seen
$\Gamma_4 e^+e^-$	seen

## $\omega(1650) \Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$

$\Gamma(\rho\pi) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$	EVTS	DOCUMENT ID	TECN	$\Gamma_1\Gamma_4/\Gamma$
<b>134±14</b>	435	17 ANTONELLI	92 DM2	1.34–2.4e <sup>+</sup> e <sup>-</sup> → hadrons

• • • We do not use the following data for averages, fits, limits, etc. • • •

$27 \pm 7$	$^{18}$ ACHASOV	99E RVUE	$0.75-1.80 \pi^+ \pi^- \rightarrow \pi^+ \pi^- \pi^0$
$93 \pm 27$	315	ANTONELLI	92 DM2 $1.34-2.4 e^+ e^- \rightarrow \rho \pi$
$96 \pm 35$		DONNACHIE	89 RVUE $e^+ e^- \rightarrow \rho \pi$

$^{17}$  From a coupled fit of  $\rho \pi$  and  $\omega \pi \pi$  channels.

$^{18}$  Using the data of DOLINSKY 91, ANTONELLI 92, AKHMETSHIN 98, and ACHASOV 99E. From a fit to two Breit-Wigner functions interfering between them and with the  $\omega, \phi$  tails with fixed  $(+, -, +)$  phases.

### $\Gamma(\omega \pi \pi) \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	$\Gamma_2 \Gamma_4 / \Gamma$
$170 \pm 17$	435	$^{19}$ ANTONELLI	92 DM2	$1.34-2.4 e^+ e^- \rightarrow \text{hadrons}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$192 \pm 33$		$^{20}$ AKHMETSHIN 00D	CMD2	$e^+ e^- \rightarrow \omega \pi^+ \pi^-$
$200 \pm 36$		$^{21}$ AKHMETSHIN 00D	CMD2	$e^+ e^- \rightarrow \omega \pi^+ \pi^-$
$135 \pm 16$	435	$^{22}$ ANTONELLI	92 DM2	$1.34-2.4 e^+ e^- \rightarrow \omega \pi \pi$
$56 \pm 31$		DONNACHIE	89 RVUE	$e^+ e^- \rightarrow \omega 2\pi$

$^{19}$  From a coupled fit of  $\rho \pi$  and  $\omega \pi \pi$  channels.

$^{20}$  Using the data of AKHMETSHIN 00D and ANTONELLI 92. The energy-independent width of the  $\omega(1420)$  and  $\omega(1650)$  mesons assumed.

$^{21}$  Using the data of AKHMETSHIN 00D and ANTONELLI 92. The  $\rho \pi$  dominance for the energy dependence of the  $\omega(1420)$  and  $\omega(1650)$  width assumed.

$^{22}$  From a single Breit-Wigner fit.

### $\omega(1650)$ BRANCHING RATIOS

#### $\Gamma(\rho \pi)/\Gamma(\omega \pi \pi)$

#### $\Gamma_1/\Gamma_2$

VALUE	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$0.17 \pm 0.05$	$^{23}$ ACHASOV	99E RVUE	$0.75-1.80 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$^{23}$ Using the data of DOLINSKY 91, ANTONELLI 92, AKHMETSHIN 98, and ACHASOV 99E. From a fit to two Breit-Wigner functions interfering between them and with the $\omega, \phi$ tails with fixed $(+, -, +)$ phases.			

### $\omega(1650)$ REFERENCES

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AKHMETSHIN 00D	PL B489 125	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
ACHASOV 99E	PL B462 365	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
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AKHMETSHIN 98	PL B434 426	R.R. Akhmetshin <i>et al.</i>
CLEGG 94	ZPHY C62 455	A.B. Clegg, A. Donnachie (LANC, MCHS)
ANTONELLI 92	ZPHY C56 15	A. Antonelli <i>et al.</i> (DM2 Collab.)
BISELLO 91C	ZPHY C52 227	D. Bisello <i>et al.</i> (DM2 Collab.)
DOLINSKY 91	PRPL 202 99	S.I. Dolinsky <i>et al.</i> (NOVO)
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