

**D<sup>0</sup>**

$$I(J^P) = \frac{1}{2}(0^-)$$

## **D<sup>0</sup> MASS**

The fit includes  $D^\pm$ ,  $D^0$ ,  $D_s^\pm$ ,  $D^{*\pm}$ ,  $D^{*0}$ , and  $D_s^{*\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1864.7 ± 0.5 OUR NEW UNCHECKED FIT</b>				Error includes scale factor of 1.1. [1864.5 ± 0.5 MeV OUR 2002 FIT Scale factor = 1.1]
<b>1864.1 ± 1.0 OUR AVERAGE</b>				
1864.6 ± 0.3 ± 1.0	641	BARLAG	90C ACCM	$\pi^-$ Cu 230 GeV
1852 ± 7	16	ADAMOVICH	87 EMUL	Photoproduction
1861 ± 4		DERRICK	84 HRS	$e^+ e^-$ 29 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1856 ± 36	22	ADAMOVICH	84B EMUL	Photoproduction
1847 ± 7	1	FIORINO	81 EMUL	$\gamma N \rightarrow \bar{D}^0 +$
1863.8 ± 0.5		<sup>1</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.77 GeV
1864.7 ± 0.6		<sup>1</sup> TRILLING	81 RVUE	$e^+ e^-$ 3.77 GeV
1863.0 ± 2.5	238	ASTON	80E OMEG	$\gamma p \rightarrow \bar{D}^0$
1860 ± 2	143	<sup>2</sup> AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$
1869 ± 4	35	<sup>2</sup> AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$
1854 ± 6	94	<sup>2</sup> ATIYA	79 SPEC	$\gamma N \rightarrow D^0 \bar{D}^0$
1850 ± 15	64	BALTAY	78C HBC	$\nu N \rightarrow K^0 \pi\pi$
1863 ± 3		GOLDHABER	77 MRK1	$D^0, D^+$ recoil spectra
1863.3 ± 0.9		<sup>1</sup> PERUZZI	77 MRK1	$e^+ e^-$ 3.77 GeV
1868 ± 11		PICCOLO	77 MRK1	$e^+ e^-$ 4.03, 4.41 GeV
1865 ± 15	234	GOLDHABER	76 MRK1	$K\pi$ and $K3\pi$

<sup>1</sup> PERUZZI 77 and SCHINDLER 81 errors do not include the 0.13% uncertainty in the absolute SPEAR energy calibration. TRILLING 81 uses the high precision  $J/\psi(1S)$  and  $\psi(2S)$  measurements of ZHOLENTZ 80 to determine this uncertainty and combines the PERUZZI 77 and SCHINDLER 81 results to obtain the value quoted. TRILLING 81 enters the fit in the  $D^\pm$  mass, and PERUZZI 77 and SCHINDLER 81 enter in the  $m_{D^\pm} - m_{D^0}$ , below.

<sup>2</sup> Error does not include possible systematic mass scale shift, estimated to be less than 5 MeV.

## **$m_{D^\pm} - m_{D^0}$**

The fit includes  $D^\pm$ ,  $D^0$ ,  $D_s^\pm$ ,  $D^{*\pm}$ ,  $D^{*0}$ , and  $D_s^{*\pm}$  mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4.78 ± 0.10 OUR FIT</b>			Error includes scale factor of 1.1.
<b>4.74 ± 0.28 OUR AVERAGE</b>			
4.7 ± 0.3		<sup>3</sup> SCHINDLER	81 MRK2 $e^+ e^-$ 3.77 GeV
5.0 ± 0.8		<sup>3</sup> PERUZZI	77 MRK1 $e^+ e^-$ 3.77 GeV

<sup>3</sup> See the footnote on TRILLING 81 in the  $D^0$  and  $D^\pm$  sections on the mass.

## **$D^0$ MEAN LIFE**

Measurements with an error  $> 20 \times 10^{-15}$  s have been omitted from the average.

<u>VALUE (<math>10^{-15}</math> s)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>410.3 \pm 1.5</math> OUR NEW AVERAGE</b>				$[(411.7 \pm 2.7) \times 10^{-15}$ s OUR 2002 AVERAGE]
409.6 $\pm$ 1.1 $\pm$ 1.5	210k	LINK	02F FOCS	$\gamma$ nucleus, $\approx 180$ GeV
407.9 $\pm$ 6.0 $\pm$ 4.3	10k	KUSHNIR...	01 SELX	$D^0 \rightarrow K^- \pi^+$ , $K^- \pi^+ \pi^+ \pi^-$
413 $\pm$ 3 $\pm$ 4	35k	AITALA	99E E791	$K^- \pi^+$
408.5 $\pm$ 4.1 $\pm$ 3.5	25k	BONVICINI	99 CLE2	$e^+ e^- \approx \gamma(4S)$
413 $\pm$ 4 $\pm$ 3	16k	FRABETTI	94D E687	$K^- \pi^+$ , $K^- \pi^+ \pi^+ \pi^-$
424 $\pm$ 11 $\pm$ 7	5118	FRABETTI	91 E687	$K^- \pi^+$ , $K^- \pi^+ \pi^+ \pi^-$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
417 $\pm$ 18 $\pm$ 15	890	ALVAREZ	90 NA14	$K^- \pi^+$ , $K^- \pi^+ \pi^+ \pi^-$
388 $\pm$ 23 $\pm$ 21	641	<sup>4</sup> BARLAG	90C ACCM	$\pi^-$ Cu 230 GeV
480 $\pm$ 40 $\pm$ 30	776	ALBRECHT	88I ARG	$e^+ e^-$ 10 GeV
422 $\pm$ 8 $\pm$ 10	4212	RAAB	88 E691	Photoproduction
420 $\pm$ 50	90	BARLAG	87B ACCM	$K^-$ and $\pi^-$ 200 GeV

<sup>4</sup> BARLAG 90C estimate systematic error to be negligible.

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$$|m_{D_1^0} - m_{D_2^0}|$$

The  $D_1^0$  and  $D_2^0$  are the mass eigenstates of the  $D^0$  meson, as described in the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<u>VALUE (<math>10^{10}</math> <math>\text{n s}^{-1}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
< 7	95	5 GODANG	00 CLE2	$e^+ e^-$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<32	90	6,7 AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
<24	90	8 AITALA	96C E791	$\pi^-$ nucleus, 500 GeV
<21	90	7,9 ANJOS	88C E691	Photoproduction

<sup>5</sup> This GODANG 00 limit is inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-$  (via  $D^0$ ))/ $\Gamma(K^- \pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0$ - $\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. The strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  is assumed to be small. If an arbitrary relative strong phase is allowed, the limit degrades by a factor of two.

<sup>6</sup> AITALA 98 allows interference between the doubly Cabibbo-suppressed and mixing amplitudes, and also allows  $CP$  violation in this term, but assumes that  $A_D = A_R = 0$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<sup>7</sup> This limit is inferred from  $R_M$  for  $f = K^+ \pi^-$  and  $f = K^+ \pi^- \pi^+ \pi^-$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. Decay-time information is used to distinguish doubly Cabibbo-suppressed decays from  $D^0$ - $\bar{D}^0$  mixing.

<sup>8</sup> This limit is inferred from  $R_M$  for  $f = K^+ \ell^- \bar{\nu}_\ell$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<sup>9</sup> ANJOS 88C assumes that  $y = 0$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. Without this assumption, the limit degrades by about a factor of two.

$$(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma = 2y$$

The  $D_1^0$  and  $D_2^0$  are the mass eigenstates of the  $D^0$  meson, as described in the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.003 ± 0.022 OUR AVERAGE</b>			Error includes scale factor of 1.4. See the ideogram below.		
-0.010 ± 0.020	+0.014 -0.016	18k	10 ABE	02I BELL	$e^+ e^- \approx \gamma(4S)$
-0.024 ± 0.050	± 0.028	3393	11 CSORNA	02 CLE2	$e^+ e^- \approx \gamma(4S)$
-0.050	+0.028 -0.032	± 0.006	12 GODANG	00 CLE2	$e^+ e^-$
0.0684 ± 0.0278 ± 0.0148	10k	10 LINK	00 FOCS	$\gamma$ nucleus	
+0.016 ± 0.058	± 0.021	10 AITALA	99E E791	$K^- \pi^+$ , $K^+ K^-$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$ \Delta\Gamma /\Gamma < 0.26$	90	13,14 AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
$ \Delta\Gamma /\Gamma < 0.20$	90	15 AITALA	96C E791	$\pi^-$ nucleus, 500 GeV
$ \Delta\Gamma /\Gamma < 0.17$	90	14,16 ANJOS	88C E691	Photoproduction

<sup>10</sup> LINK 00, AITALA 99E, and ABE 02I measure the lifetime difference between  $D^0 \rightarrow K^- K^+$  ( $CP$  even) decays and  $D^0 \rightarrow K^- \pi^+$  ( $CP$  mixed) decays, or  $y_{CP} = [\Gamma(CP+) - \Gamma(CP-)] / [\Gamma(CP+) + \Gamma(CP-)]$ . We list  $2y_{CP} = \Delta\Gamma/\Gamma$ .

<sup>11</sup> CSORNA 02 measures the lifetime difference between  $D^0 \rightarrow K^- K^+$  and  $\pi^- \pi^+$  ( $CP$  even) decays and  $D^0 \rightarrow K^- \pi^+$  ( $CP$  mixed) decays, or  $y_{CP} = [\Gamma(CP+) - \Gamma(CP-)] / [\Gamma(CP+) + \Gamma(CP-)]$ . We list  $2y_{CP} = \Delta\Gamma/\Gamma$ .

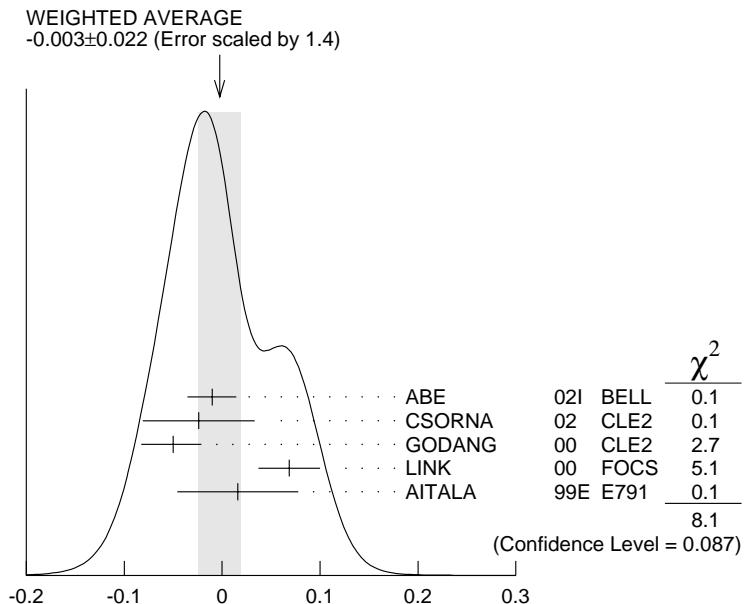
<sup>12</sup> This GODANG 00 limit is inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+ \pi^-)$  (via  $\bar{D}^0)) / \Gamma(K^- \pi^+$ ) given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0$ - $\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. The phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  is assumed to be small. This is a measurement of  $y'$  and is not the same as the  $y_{CP}$  of LINK 00.

<sup>13</sup> AITALA 98 allows interference between the doubly Cabibbo-suppressed and mixing amplitudes, and also allows  $CP$  violation in this term, but assumes that  $A_D = A_R = 0$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<sup>14</sup> This limit is inferred from  $R_M$  for  $f = K^+ \pi^-$  and  $f = K^+ \pi^- \pi^+ \pi^-$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. Decay-time information is used to distinguish doubly Cabibbo-suppressed decays from  $D^0$ - $\bar{D}^0$  mixing.

<sup>15</sup> This limit is inferred from  $R_M$  for  $f = K^+ \ell^- \bar{\nu}_\ell$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above.

<sup>16</sup> ANJOS 88C assumes that  $y = 0$ . See the note on “ $D^0$ - $\bar{D}^0$  Mixing,” above. Without this assumption, the limit degrades by about a factor of two.



$$(\Gamma_1 - \Gamma_2)/\Gamma = 2y$$

## $D^0$ DECAY MODES

$\bar{D}^0$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Inclusive modes</b>		
$\Gamma_1 e^+$ anything	[a] $(6.87 \pm 0.28) \%$	
$\Gamma_2 \mu^+$ anything	$(6.5 \pm 0.8) \%$	
$\Gamma_3 K^-$ anything	$(53 \pm 4) \%$	S=1.3
$\Gamma_4 \bar{K}^0$ anything + $K^0$ anything	$(42 \pm 5) \%$	
$\Gamma_5 K^+$ anything	$(3.4 \pm 0.6) \%$	
$\Gamma_6 \eta$ anything	[b] $< 13 \%$	CL=90%
$\Gamma_7 \phi$ anything	$(1.7 \pm 0.8) \%$	
<b>Semileptonic modes</b>		
$\Gamma_8 K^- \ell^+ \nu_\ell$	[c] $(3.43 \pm 0.15) \%$	S=1.2
$\Gamma_9 K^- e^+ \nu_e$	$(3.58 \pm 0.18) \%$	S=1.1
$\Gamma_{10} K^- \mu^+ \nu_\mu$	$(3.19 \pm 0.17) \%$	
$\Gamma_{11} K^- \pi^0 e^+ \nu_e$	$(1.1 \pm 0.8) \%$	S=1.6

$\Gamma_{12}$	$\overline{K}^0 \pi^- e^+ \nu_e$	( 1.8 $\pm$ 0.8 ) %	S=1.6
$\Gamma_{13}$	$\overline{K}^*(892)^- e^+ \nu_e$ × $B(K^{*-} \rightarrow \overline{K}^0 \pi^-)$	( 1.43 $\pm$ 0.23 ) %	
$\Gamma_{14}$	$K^*(892)^- \ell^+ \nu_\ell$		
$\Gamma_{15}$	$\overline{K}^*(892)^0 \pi^- e^+ \nu_e$		
$\Gamma_{16}$	$K^- \pi^+ \pi^- \mu^+ \nu_\mu$	< 1.2 $\times 10^{-3}$	CL=90%
$\Gamma_{17}$	$(\overline{K}^*(892)\pi)^- \mu^+ \nu_\mu$	< 1.4 $\times 10^{-3}$	CL=90%
$\Gamma_{18}$	$\pi^- e^+ \nu_e$	( 3.6 $\pm$ 0.6 ) $\times 10^{-3}$	

A fraction of the following resonance mode has already appeared above as a submode of a charged-particle mode.

$\Gamma_{19}$	$K^*(892)^- e^+ \nu_e$	( 2.15 $\pm$ 0.35 ) %
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### Hadronic modes with a $\overline{K}$ or $\overline{K}KK\overline{K}$

$\Gamma_{20}$	$K^- \pi^+$	( 3.80 $\pm$ 0.09 ) %	
$\Gamma_{21}$	$\overline{K}^0 \pi^0$	( 2.30 $\pm$ 0.22 ) %	
$\Gamma_{22}$	$\overline{K}^0 \pi^+ \pi^-$	[d] ( 5.97 $\pm$ 0.35 ) %	S=1.1
$\Gamma_{23}$	$\overline{K}^0 \rho^0$	( 1.55 $\pm$ 0.12 ) %	
$\Gamma_{24}$	$\overline{K}^0 f_0(980)$ × $B(f_0 \rightarrow \pi^+ \pi^-)$	( 2.8 $\pm$ 0.6 ) $\times 10^{-3}$	
$\Gamma_{25}$	$\overline{K}^0 f_2(1270)$ × $B(f_2 \rightarrow \pi^+ \pi^-)$	( 2.5 $\pm$ 1.0 ) $\times 10^{-3}$	
$\Gamma_{26}$	$\overline{K}^0 f_0(1370)$ × $B(f_0 \rightarrow \pi^+ \pi^-)$	( 5.1 $\pm$ 1.2 ) $\times 10^{-3}$	
$\Gamma_{27}$	$K^*(892)^- \pi^+$ × $B(K^{*-} \rightarrow \overline{K}^0 \pi^-)$	( 4.0 $\pm$ 0.4 ) %	
$\Gamma_{28}$	$K_0^*(1430)^- \pi^+$ × $B(K_0^*(1430)^- \rightarrow \overline{K}^0 \pi^-)$	( 7.3 $\pm$ 1.6 ) $\times 10^{-3}$	
$\Gamma_{29}$	$K^*(1680)^- \pi^+$ × $B(K^*(1680)^- \rightarrow \overline{K}^0 \pi^-)$		
$\Gamma_{30}$	$\overline{K}^0 \pi^+ \pi^-$ nonresonant	( 5.4 $\pm$ 12.0 ) $\times 10^{-4}$	
$\Gamma_{31}$	$K^- \pi^+ \pi^0$	[d] ( 13.0 $\pm$ 0.8 ) %	S=1.3
$\Gamma_{32}$	$K^- \rho^+$	( 10.1 $\pm$ 0.8 ) %	
$\Gamma_{33}$	$K^- \rho(1700)^+$ × $B(\rho(1700)^+ \rightarrow \pi^+ \pi^0)$	( 7.4 $\pm$ 1.6 ) $\times 10^{-3}$	
$\Gamma_{34}$	$K^*(892)^- \pi^+$ × $B(K^{*-} \rightarrow K^- \pi^0)$	( 2.0 $\pm$ 0.2 ) %	
$\Gamma_{35}$	$\overline{K}^*(892)^0 \pi^0$ × $B(\overline{K}^{*0} \rightarrow K^- \pi^+)$	( 1.87 $\pm$ 0.27 ) %	
$\Gamma_{36}$	$K_0^*(1430)^- \pi^+$ × $B(K_0^*(1430)^- \rightarrow K^- \pi^0)$	( 3.6 $\pm$ 0.8 ) $\times 10^{-3}$	

$\Gamma_{37}$	$\bar{K}_0^*(1430)^0 \pi^0$ $\times B(\bar{K}_0^*(1430)^0 \rightarrow K^- \pi^+)$	$(5.3 \pm 4.2) \times 10^{-3}$
$\Gamma_{38}$	$K^*(1680)^- \pi^+$ $\times B(K^*(1680)^- \rightarrow K^- \pi^0)$	$(1.7 \pm 0.6) \times 10^{-3}$
$\Gamma_{39}$	$K^- \pi^+ \pi^0$ nonresonant	$(1.04 \pm 0.50) \%$
$\Gamma_{40}$	$\bar{K}^0 \pi^0 \pi^0$	—
$\Gamma_{41}$	$\bar{K}^*(892)^0 \pi^0$ $\times B(\bar{K}^{*0} \rightarrow \bar{K}^0 \pi^0)$	$(9.3 \pm 1.3) \times 10^{-3}$
$\Gamma_{42}$	$\bar{K}^0 \pi^0 \pi^0$ nonresonant	$(8.5 \pm 2.2) \times 10^{-3}$
$\Gamma_{43}$	$K^- \pi^+ \pi^+ \pi^-$	[d] $(7.46 \pm 0.31) \%$
$\Gamma_{44}$	$K^- \pi^+ \rho^0$ total	$(6.2 \pm 0.4) \%$
$\Gamma_{45}$	$K^- \pi^+ \rho^0$ 3-body	$(4.7 \pm 2.1) \times 10^{-3}$
$\Gamma_{46}$	$\bar{K}^*(892)^0 \rho^0$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$	$(9.7 \pm 2.1) \times 10^{-3}$
$\Gamma_{47}$	$K^- a_1(1260)^+$ $\times B(a_1(1260)^+ \rightarrow \pi^+ \pi^+ \pi^-)$	$(3.6 \pm 0.6) \%$
$\Gamma_{48}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ total $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$	$(1.5 \pm 0.4) \%$
$\Gamma_{49}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ 3-body $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$	$(9.5 \pm 2.1) \times 10^{-3}$
$\Gamma_{50}$	$K_1(1270)^- \pi^+$ $\times B(K_1(1270)^- \rightarrow K^- \pi^+ \pi^-)$	[e] $(3.7 \pm 1.0) \times 10^{-3}$
$\Gamma_{51}$	$K^- \pi^+ \pi^+ \pi^-$ nonresonant	$(1.74 \pm 0.25) \%$
$\Gamma_{52}$	$\bar{K}^0 \pi^+ \pi^- \pi^0$	[d] $(10.9 \pm 1.3) \%$
$\Gamma_{53}$	$\bar{K}^0 \eta \times B(\eta \rightarrow \pi^+ \pi^- \pi^0)$	$(1.7 \pm 0.3) \times 10^{-3}$
$\Gamma_{54}$	$\bar{K}^0 \omega \times B(\omega \rightarrow \pi^+ \pi^- \pi^0)$	$(2.2 \pm 0.4) \%$
$\Gamma_{55}$	$K^*(892)^- \rho^+$ $\times B(K^{*-} \rightarrow \bar{K}^0 \pi^-)$	$(4.3 \pm 1.7) \%$
$\Gamma_{56}$	$\bar{K}^*(892)^0 \rho^0$ $\times B(\bar{K}^{*0} \rightarrow \bar{K}^0 \pi^0)$	$(4.8 \pm 1.1) \times 10^{-3}$
$\Gamma_{57}$	$K_1(1270)^- \pi^+$ $\times B(K_1(1270)^- \rightarrow \bar{K}^0 \pi^- \pi^0)$	[e] $(5.3 \pm 1.5) \times 10^{-3}$
$\Gamma_{58}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ 3-body $\times B(\bar{K}^{*0} \rightarrow \bar{K}^0 \pi^0)$	$(4.7 \pm 1.0) \times 10^{-3}$
$\Gamma_{59}$	$\bar{K}^0 \pi^+ \pi^- \pi^0$ nonresonant	$(2.3 \pm 2.3) \%$
$\Gamma_{60}$	$K^- \pi^+ \pi^0 \pi^0$	
$\Gamma_{61}$	$K^- \pi^+ \pi^+ \pi^- \pi^0$	$(4.0 \pm 0.4) \%$
$\Gamma_{62}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$	$(1.2 \pm 0.6) \%$
$\Gamma_{63}$	$\bar{K}^*(892)^0 \eta$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$ $\times B(\eta \rightarrow \pi^+ \pi^- \pi^0)$	$(2.8 \pm 0.6) \times 10^{-3}$
$\Gamma_{64}$	$K^- \pi^+ \omega \times B(\omega \rightarrow \pi^+ \pi^- \pi^0)$	$(2.7 \pm 0.5) \%$

$\Gamma_{65}$	$\bar{K}^*(892)^0 \omega$ $\times \text{B}(\bar{K}^{*0} \rightarrow K^- \pi^+)$ $\times \text{B}(\omega \rightarrow \pi^+ \pi^- \pi^0)$	$(6.5 \pm 2.4) \times 10^{-3}$
$\Gamma_{66}$	$\bar{K}^0 \pi^+ \pi^+ \pi^- \pi^-$	$(6.4 \pm 1.8) \times 10^{-3}$
$\Gamma_{67}$	$\bar{K}^0 \pi^+ \pi^- \pi^0 \pi^0 (\pi^0)$	
$\Gamma_{68}$	$\bar{K}^0 K^+ K^-$ In the fit as $\frac{1}{2}\Gamma_{80} + \Gamma_{70}$ , where $\frac{1}{2}\Gamma_{80} = \Gamma_{69}$ .	$(1.03 \pm 0.10) \%$
$\Gamma_{69}$	$\bar{K}^0 \phi \times \text{B}(\phi \rightarrow K^+ K^-)$	$(4.7 \pm 0.6) \times 10^{-3}$
$\Gamma_{70}$	$\bar{K}^0 K^+ K^- \text{non-}\phi$	$(5.6 \pm 0.9) \times 10^{-3}$
$\Gamma_{71}$	$K_S^0 K_S^0 K_S^0$	$(9.2 \pm 1.6) \times 10^{-4}$
$\Gamma_{72}$	$K^- \pi^+ \phi \times \text{B}(\phi \rightarrow K^+ K^-)$	
$\Gamma_{73}$	$K^+ K^- K^- \pi^+$	$(2.4 \pm 0.7) \times 10^{-4}$
$\Gamma_{74}$	$K^+ K^- \bar{K}^0 \pi^0$	

Fractions of many of the following modes with resonances have already appeared above as submodes of particular charged-particle modes. (Modes for which there are only upper limits and  $\bar{K}^*(892)\rho$  submodes only appear below.)

$\Gamma_{75}$	$\bar{K}^0 \eta$	$(7.7 \pm 1.1) \times 10^{-3}$
$\Gamma_{76}$	$\bar{K}^0 \rho^0$	$(1.47 \pm 0.29) \%$
$\Gamma_{77}$	$K^- \rho^+$	$(10.2 \pm 0.8) \%$
$\Gamma_{78}$	$\bar{K}^0 \omega$	$(2.31 \pm 0.35) \%$
$\Gamma_{79}$	$\bar{K}^0 \eta'(958)$	$(1.88 \pm 0.28) \%$
$\Gamma_{80}$	$\bar{K}^0 \phi$	$(9.4 \pm 1.1) \times 10^{-3}$
$\Gamma_{81}$	$K^- a_1(1260)^+$	$(7.2 \pm 1.1) \%$
$\Gamma_{82}$	$\bar{K}^0 a_1(1260)^0$	$< 1.9 \%$
$\Gamma_{83}$	$\bar{K}^0 f_2(1270)$	$(4.7 \pm 4.1) \times 10^{-4}$
$\Gamma_{84}$	$K^- a_2(1320)^+$	$< 2 \times 10^{-3}$
$\Gamma_{85}$	$K^*(892)^- \pi^+$	$(5.9 \pm 0.4) \%$
$\Gamma_{86}$	$\bar{K}^*(892)^0 \pi^0$	$(2.8 \pm 0.4) \%$
$\Gamma_{87}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \text{total}$	$(2.2 \pm 0.5) \%$
$\Gamma_{88}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \text{3-body}$	$(1.42 \pm 0.31) \%$
$\Gamma_{89}$	$K^- \pi^+ \rho^0 \text{total}$	$(6.2 \pm 0.4) \%$
$\Gamma_{90}$	$K^- \pi^+ \rho^0 \text{3-body}$	$(4.7 \pm 2.1) \times 10^{-3}$
$\Gamma_{91}$	$\bar{K}^*(892)^0 \rho^0$	$(1.45 \pm 0.32) \%$
$\Gamma_{92}$	$\bar{K}^*(892)^0 \rho^0 \text{transverse}$	$(1.5 \pm 0.5) \%$
$\Gamma_{93}$	$\bar{K}^*(892)^0 \rho^0 S\text{-wave}$	$(2.8 \pm 0.6) \%$
$\Gamma_{94}$	$\bar{K}^*(892)^0 \rho^0 S\text{-wave long.}$	$< 3 \times 10^{-3}$
$\Gamma_{95}$	$\bar{K}^*(892)^0 \rho^0 P\text{-wave}$	$< 3 \times 10^{-3}$
$\Gamma_{96}$	$\bar{K}^*(892)^0 \rho^0 D\text{-wave}$	$(1.9 \pm 0.6) \%$
$\Gamma_{97}$	$K^*(892)^- \rho^+$	$(6.6 \pm 2.6) \%$
$\Gamma_{98}$	$K^*(892)^- \rho^+ \text{longitudinal}$	$(3.2 \pm 1.3) \%$
$\Gamma_{99}$	$K^*(892)^- \rho^+ \text{transverse}$	$(3.4 \pm 2.0) \%$

$\Gamma_{100}$	$K^*(892)^-\rho^+ P\text{-wave}$	< 1.5	%	CL=90%
$\Gamma_{101}$	$K^-\pi^+ f_0(980)$			
$\Gamma_{102}$	$\bar{K}^*(892)^0 f_0(980)$			
$\Gamma_{103}$	$K_1(1270)^-\pi^+$	[e] ( 1.14 $\pm$ 0.31 ) %		
$\Gamma_{104}$	$K_1(1400)^-\pi^+$	< 1.2	%	CL=90%
$\Gamma_{105}$	$\bar{K}_1(1400)^0\pi^0$	< 3.7	%	CL=90%
$\Gamma_{106}$	$K^*(1410)^-\pi^+$			
$\Gamma_{107}$	$K_0^*(1430)^-\pi^+$	( 9.8 $\pm$ 2.0 ) $\times 10^{-3}$		
$\Gamma_{108}$	$\bar{K}_0^*(1430)^0\pi^0$	( 8.6 $\pm$ 6.8 ) $\times 10^{-3}$		
$\Gamma_{109}$	$K_2^*(1430)^-\pi^+$	( 2.0 $\pm$ 1.3 ) $\times 10^{-3}$		
$\Gamma_{110}$	$\bar{K}_2^*(1430)^0\pi^0$	< 3.3 $\times 10^{-3}$	CL=90%	
$\Gamma_{111}$	$K^*(1680)^-\pi^+$	( 8.2 $\pm$ 3.9 ) $\times 10^{-3}$	S=1.2	
$\Gamma_{112}$	$\bar{K}^*(892)^0\pi^+\pi^-\pi^0$	( 1.8 $\pm$ 0.9 ) %		
$\Gamma_{113}$	$\bar{K}^*(892)^0\eta$	( 1.8 $\pm$ 0.4 ) %		
$\Gamma_{114}$	$K^-\pi^+\omega$	( 3.0 $\pm$ 0.6 ) %		
$\Gamma_{115}$	$\bar{K}^*(892)^0\omega$	( 1.1 $\pm$ 0.4 ) %		
$\Gamma_{116}$	$K^-\pi^+\eta'(958)$	( 6.9 $\pm$ 1.8 ) $\times 10^{-3}$		
$\Gamma_{117}$	$\bar{K}^*(892)^0\eta'(958)$	< 1.0 $\times 10^{-3}$	CL=90%	
$\Gamma_{118}$	$K^-\pi^+\phi$	( 3.3 $\pm$ 1.7 ) $\times 10^{-4}$		

### Pionic modes

$\Gamma_{119}$	$\pi^+\pi^-$	( 1.43 $\pm$ 0.07 ) $\times 10^{-3}$		
$\Gamma_{120}$	$\pi^0\pi^0$	( 8.4 $\pm$ 2.2 ) $\times 10^{-4}$		
$\Gamma_{121}$	$\pi^+\pi^-\pi^0$	( 1.1 $\pm$ 0.4 ) %		
$\Gamma_{122}$	$\pi^+\pi^+\pi^-\pi^-$	( 7.3 $\pm$ 0.5 ) $\times 10^{-3}$		
$\Gamma_{123}$	$\pi^+\pi^+\pi^-\pi^-\pi^0$			
$\Gamma_{124}$	$\pi^+\pi^+\pi^+\pi^-\pi^-\pi^-$			

### Hadronic modes with a $K\bar{K}$ pair

$\Gamma_{125}$	$K^+K^-$	( 4.12 $\pm$ 0.14 ) $\times 10^{-3}$		
$\Gamma_{126}$	$K^0\bar{K}^0$	( 7.1 $\pm$ 1.9 ) $\times 10^{-4}$	S=1.2	
$\Gamma_{127}$	$K^0K^-\pi^+$	( 6.9 $\pm$ 1.0 ) $\times 10^{-3}$		
$\Gamma_{128}$	$\bar{K}^*(892)^0K^0$ $\times B(\bar{K}^{*0} \rightarrow K^-\pi^+)$	< 1.1 $\times 10^{-3}$	CL=90%	
$\Gamma_{129}$	$K^*(892)^+K^-$ $\times B(K^{*+} \rightarrow K^0\pi^+)$	( 2.5 $\pm$ 0.5 ) $\times 10^{-3}$		
$\Gamma_{130}$	$K^0K^-\pi^+$ nonresonant	( 2.3 $\pm$ 2.3 ) $\times 10^{-3}$		
$\Gamma_{131}$	$\bar{K}^0K^+\pi^-$	( 5.3 $\pm$ 1.0 ) $\times 10^{-3}$		
$\Gamma_{132}$	$K^*(892)^0\bar{K}^0$ $\times B(K^{*0} \rightarrow K^+\pi^-)$	< 6 $\times 10^{-4}$	CL=90%	
$\Gamma_{133}$	$K^*(892)^-K^+$ $\times B(K^{*-} \rightarrow \bar{K}^0\pi^-)$	( 1.3 $\pm$ 0.7 ) $\times 10^{-3}$		

$\Gamma_{134}$	$\bar{K}^0 K^+ \pi^-$ nonresonant	$(3.8 \pm 2.3) \times 10^{-3}$	
$\Gamma_{135}$	$K^+ K^- \pi^0$	$(1.24 \pm 0.35) \times 10^{-3}$	
$\Gamma_{136}$	$K_S^0 K_S^0 \pi^0$	$< 5.9 \times 10^{-4}$	
$\Gamma_{137}$	$K^+ K^- \pi^+ \pi^-$	$[f] (2.49 \pm 0.23) \times 10^{-3}$	
$\Gamma_{138}$	$\phi \pi^+ \pi^- \times B(\phi \rightarrow K^+ K^-)$	$(5.3 \pm 1.4) \times 10^{-4}$	
$\Gamma_{139}$	$\phi \rho^0 \times B(\phi \rightarrow K^+ K^-)$	$(2.9 \pm 1.5) \times 10^{-4}$	
$\Gamma_{140}$	$K^+ K^- \rho^0$ 3-body	$(9.0 \pm 2.3) \times 10^{-4}$	
$\Gamma_{141}$	$K^*(892)^0 K^- \pi^+ + c.c.$ $\times B(K^{*0} \rightarrow K^+ \pi^-)$	$[g] < 5 \times 10^{-4}$	
$\Gamma_{142}$	$K^*(892)^0 \bar{K}^*(892)^0$ $\times B^2(K^{*0} \rightarrow K^+ \pi^-)$	$(6 \pm 2) \times 10^{-4}$	
$\Gamma_{143}$	$K^+ K^- \pi^+ \pi^-$ non- $\phi$		
$\Gamma_{144}$	$K^+ K^- \pi^+ \pi^-$ nonresonant	$< 8 \times 10^{-4}$	CL=90%
$\Gamma_{145}$	$K^0 \bar{K}^0 \pi^+ \pi^-$	$(7.5 \pm 2.9) \times 10^{-3}$	
$\Gamma_{146}$	$K^+ K^- \pi^+ \pi^- \pi^0$	$(3.1 \pm 2.0) \times 10^{-3}$	

Fractions of most of the following modes with resonances have already appeared above as submodes of particular charged-particle modes.

$\Gamma_{147}$	$\bar{K}^*(892)^0 K^0$	$< 1.7 \times 10^{-3}$	CL=90%
$\Gamma_{148}$	$K^*(892)^+ K^-$	$(3.8 \pm 0.8) \times 10^{-3}$	
$\Gamma_{149}$	$K^*(892)^0 \bar{K}^0$	$< 9 \times 10^{-4}$	CL=90%
$\Gamma_{150}$	$K^*(892)^- K^+$	$(2.0 \pm 1.1) \times 10^{-3}$	
$\Gamma_{151}$	$\phi \pi^0$	$< 1.4 \times 10^{-3}$	CL=90%
$\Gamma_{152}$	$\phi \eta$	$< 2.8 \times 10^{-3}$	CL=90%
$\Gamma_{153}$	$\phi \omega$	$< 2.1 \times 10^{-3}$	CL=90%
$\Gamma_{154}$	$\phi \pi^+ \pi^-$	$(1.06 \pm 0.28) \times 10^{-3}$	
$\Gamma_{155}$	$\phi \rho^0$	$(5.7 \pm 3.0) \times 10^{-4}$	
$\Gamma_{156}$	$\phi \pi^+ \pi^-$ 3-body	$(7 \pm 5) \times 10^{-4}$	
$\Gamma_{157}$	$K^*(892)^0 K^- \pi^+ + c.c.$	$[g] < 7 \times 10^{-4}$	CL=90%
$\Gamma_{158}$	$K^*(892)^0 K^- \pi^+$		
$\Gamma_{159}$	$\bar{K}^*(892)^0 K^+ \pi^-$		
$\Gamma_{160}$	$K^*(892)^0 \bar{K}^*(892)^0$	$(1.4 \pm 0.5) \times 10^{-3}$	

### Radiative modes

$\Gamma_{161}$	$\rho^0 \gamma$	$< 2.4 \times 10^{-4}$	CL=90%
$\Gamma_{162}$	$\omega \gamma$	$< 2.4 \times 10^{-4}$	CL=90%
$\Gamma_{163}$	$\phi \gamma$	$< 1.9 \times 10^{-4}$	CL=90%
$\Gamma_{164}$	$\bar{K}^*(892)^0 \gamma$	$< 7.6 \times 10^{-4}$	CL=90%

**Doubly Cabibbo suppressed (DC) modes,**  
 **$\Delta C = 2$  forbidden via mixing (C2M) modes,**  
 **$\Delta C = 1$  weak neutral current (C1) modes,**  
**Lepton Family number (LF) violating modes, or**  
**Lepton number ( $L$ ) violating modes**

$\Gamma_{165}$	$K^+ \ell^- \bar{\nu}_\ell$ (via $\bar{D}^0$ )	C2M	$< 1.7 \times 10^{-4}$	CL=90%
$\Gamma_{166}$	$K^+ \pi^-$	DC	$(1.48 \pm 0.21) \times 10^{-4}$	
$\Gamma_{167}$	$K^+ \pi^-$ (via $\bar{D}^0$ )	C2M	$< 1.6 \times 10^{-5}$	CL=95%
$\Gamma_{168}$	$K^*(892)^+ \pi^-$		$(3.0 \pm 3.8) \times 10^{-4}$	
$\Gamma_{169}$	$K^+ \pi^- \pi^0$		$(5.6 \pm 1.7) \times 10^{-4}$	
$\Gamma_{170}$	$K^+ \pi^- \pi^+ \pi^-$	DC	$(3.1 \pm 1.0) \times 10^{-4}$	
$\Gamma_{171}$	$K^+ \pi^- \pi^+ \pi^-$ (via $\bar{D}^0$ )	C2M	$< 4 \times 10^{-4}$	CL=90%
$\Gamma_{172}$	$K^+ \pi^-$ or $K^+ \pi^- \pi^+ \pi^-$ (via $\bar{D}^0$ )		$< 1.0 \times 10^{-3}$	CL=90%
$\Gamma_{173}$	$\mu^-$ anything (via $\bar{D}^0$ )	C2M	$< 4 \times 10^{-4}$	CL=90%
$\Gamma_{174}$	$e^+ e^-$	C1	$< 6.2 \times 10^{-6}$	CL=90%
$\Gamma_{175}$	$\mu^+ \mu^-$	C1	$< 4.1 \times 10^{-6}$	CL=90%
$\Gamma_{176}$	$\pi^0 e^+ e^-$	C1	$< 4.5 \times 10^{-5}$	CL=90%
$\Gamma_{177}$	$\pi^0 \mu^+ \mu^-$	C1	$< 1.8 \times 10^{-4}$	CL=90%
$\Gamma_{178}$	$\eta e^+ e^-$	C1	$< 1.1 \times 10^{-4}$	CL=90%
$\Gamma_{179}$	$\eta \mu^+ \mu^-$	C1	$< 5.3 \times 10^{-4}$	CL=90%
$\Gamma_{180}$	$\pi^+ \pi^- e^+ e^-$	C1	$< 3.73 \times 10^{-4}$	CL=90%
$\Gamma_{181}$	$\rho^0 e^+ e^-$	C1	$< 1.0 \times 10^{-4}$	CL=90%
$\Gamma_{182}$	$\pi^+ \pi^- \mu^+ \mu^-$	C1	$< 3.0 \times 10^{-5}$	CL=90%
$\Gamma_{183}$	$\rho^0 \mu^+ \mu^-$	C1	$< 2.2 \times 10^{-5}$	CL=90%
$\Gamma_{184}$	$\omega e^+ e^-$	C1	$< 1.8 \times 10^{-4}$	CL=90%
$\Gamma_{185}$	$\omega \mu^+ \mu^-$	C1	$< 8.3 \times 10^{-4}$	CL=90%
$\Gamma_{186}$	$K^- K^+ e^+ e^-$	C1	$< 3.15 \times 10^{-4}$	CL=90%
$\Gamma_{187}$	$\phi e^+ e^-$	C1	$< 5.2 \times 10^{-5}$	CL=90%
$\Gamma_{188}$	$K^- K^+ \mu^+ \mu^-$	C1	$< 3.3 \times 10^{-5}$	CL=90%
$\Gamma_{189}$	$\phi \mu^+ \mu^-$	C1	$< 3.1 \times 10^{-5}$	CL=90%
$\Gamma_{190}$	$\bar{K}^0 e^+ e^-$	[h]	$< 1.1 \times 10^{-4}$	CL=90%
$\Gamma_{191}$	$\bar{K}^0 \mu^+ \mu^-$	[h]	$< 2.6 \times 10^{-4}$	CL=90%
$\Gamma_{192}$	$K^- \pi^+ e^+ e^-$	C1	$< 3.85 \times 10^{-4}$	CL=90%
$\Gamma_{193}$	$\bar{K}^*(892)^0 e^+ e^-$	[h]	$< 4.7 \times 10^{-5}$	CL=90%
$\Gamma_{194}$	$K^- \pi^+ \mu^+ \mu^-$	C1	$< 3.59 \times 10^{-4}$	CL=90%
$\Gamma_{195}$	$\bar{K}^*(892)^0 \mu^+ \mu^-$	[h]	$< 2.4 \times 10^{-5}$	CL=90%
$\Gamma_{196}$	$\pi^+ \pi^- \pi^0 \mu^+ \mu^-$	C1	$< 8.1 \times 10^{-4}$	CL=90%
$\Gamma_{197}$	$\mu^\pm e^\mp$	LF	$[i] < 8.1 \times 10^{-6}$	CL=90%
$\Gamma_{198}$	$\pi^0 e^\pm \mu^\mp$	LF	$[i] < 8.6 \times 10^{-5}$	CL=90%
$\Gamma_{199}$	$\eta e^\pm \mu^\mp$	LF	$[i] < 1.0 \times 10^{-4}$	CL=90%
$\Gamma_{200}$	$\pi^+ \pi^- e^\pm \mu^\mp$	LF	$[i] < 1.5 \times 10^{-5}$	CL=90%
$\Gamma_{201}$	$\rho^0 e^\pm \mu^\mp$	LF	$[i] < 4.9 \times 10^{-5}$	CL=90%

$\Gamma_{202}$	$\omega e^\pm \mu^\mp$	$LF$	$[i] < 1.2$	$\times 10^{-4}$	CL=90%
$\Gamma_{203}$	$K^- K^+ e^\pm \mu^\mp$	$LF$	$[i] < 1.8$	$\times 10^{-4}$	CL=90%
$\Gamma_{204}$	$\phi e^\pm \mu^\mp$	$LF$	$[i] < 3.4$	$\times 10^{-5}$	CL=90%
$\Gamma_{205}$	$\bar{K}^0 e^\pm \mu^\mp$	$LF$	$[i] < 1.0$	$\times 10^{-4}$	CL=90%
$\Gamma_{206}$	$K^- \pi^+ e^\pm \mu^\mp$	$LF$	$[i] < 5.53$	$\times 10^{-4}$	CL=90%
$\Gamma_{207}$	$\bar{K}^*(892)^0 e^\pm \mu^\mp$	$LF$	$[i] < 8.3$	$\times 10^{-5}$	CL=90%
$\Gamma_{208}$	$\pi^- \pi^- e^+ e^+ + \text{c.c.}$	$L$	$< 1.12$	$\times 10^{-4}$	CL=90%
$\Gamma_{209}$	$\pi^- \pi^- \mu^+ \mu^+ + \text{c.c.}$	$L$	$< 2.9$	$\times 10^{-5}$	CL=90%
$\Gamma_{210}$	$K^- \pi^- e^+ e^+ + \text{c.c.}$	$L$	$< 2.06$	$\times 10^{-4}$	CL=90%
$\Gamma_{211}$	$K^- \pi^- \mu^+ \mu^+ + \text{c.c.}$	$L$	$< 3.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{212}$	$K^- K^- e^+ e^+ + \text{c.c.}$	$L$	$< 1.52$	$\times 10^{-4}$	CL=90%
$\Gamma_{213}$	$K^- K^- \mu^+ \mu^+ + \text{c.c.}$	$L$	$< 9.4$	$\times 10^{-5}$	CL=90%
$\Gamma_{214}$	$\pi^- \pi^- e^+ \mu^+ + \text{c.c.}$	$L$	$< 7.9$	$\times 10^{-5}$	CL=90%
$\Gamma_{215}$	$K^- \pi^- e^+ \mu^+ + \text{c.c.}$	$L$	$< 2.18$	$\times 10^{-4}$	CL=90%
$\Gamma_{216}$	$K^- K^- e^+ \mu^+ + \text{c.c.}$	$L$	$< 5.7$	$\times 10^{-5}$	CL=90%

$\Gamma_{217}$  A dummy mode used by the fit.  $(10.7 \pm 3.4) \%$  S=1.1

- [a] The exclusive  $e^+$  modes  $K^- e^+ \nu_e$ ,  $K^- \pi^0 e^+ \nu_e$ ,  $\bar{K}^0 \pi^- e^+ \nu_e$  and  $\pi^- e^+ \nu_e$  are constrained to equal this (well-measured) inclusive fraction.
- [b] This is a weighted average of  $D^\pm$  (44%) and  $D^0$  (56%) branching fractions. See “ $D^+$  and  $D^0 \rightarrow (\eta \text{ anything}) / (\text{total } D^+ \text{ and } D^0)$ ” under “ $D^+$  Branching Ratios” in these Particle Listings.
- [c] This value averages the  $e^+$  and  $\mu^+$  branching fractions, after making a small phase-space adjustment to the  $\mu^+$  fraction to be able to use it as an  $e^+$  fraction; hence our  $\ell^+$  here is really an  $e^+$ .
- [d] The branching fraction for this mode may differ from the sum of the submodes that contribute to it, due to interference effects. See the relevant papers.
- [e] The two experiments measuring this fraction are in serious disagreement. See the Particle Listings.
- [f] The experiments on the division of this charge mode amongst its submodes disagree, and the submode branching fractions here add up to considerably more than the charged-mode fraction.
- [g] However, these upper limits are in serious disagreement with values obtained in another experiment.
- [h] This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.
- [i] The value is for the sum of the charge states or particle/antiparticle states indicated.

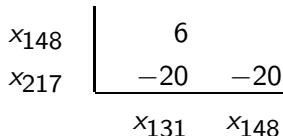
## CONSTRAINED FIT INFORMATION

An overall fit to 59 branching ratios uses 131 measurements and 1 constraint to determine 33 parameters. The overall fit has a  $\chi^2=63.3$  for  $99^\circ$  of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to 1.

$x_9$	6										
$x_{10}$	31	19									
$x_{11}$	0	-8	-1								
$x_{12}$	-1	-15	-3	-91							
$x_{18}$	1	24	5	-4	-8						
$x_{19}$	1	10	3	-1	-1	2					
$x_{20}$	13	46	42	-3	-7	11	8				
$x_{21}$	1	6	5	0	-1	1	16	11			
$x_{22}$	2	10	8	-1	-1	2	28	18	58		
$x_{30}$	0	1	0	0	0	0	2	1	4	6	
$x_{31}$	3	11	9	-1	-2	3	10	22	22	33	
$x_{43}$	5	18	17	-1	-3	4	3	40	5	8	
$x_{52}$	1	4	3	0	-1	1	12	8	25	42	
$x_{61}$	3	9	8	-1	-1	2	2	19	2	4	
$x_{70}$	1	4	3	0	-1	1	11	7	22	38	
$x_{75}$	1	4	3	0	-1	1	11	7	54	40	
$x_{78}$	1	3	3	0	-1	1	10	6	20	35	
$x_{80}$	1	5	4	0	-1	1	14	9	29	51	
$x_{85}$	2	9	7	-1	-1	2	25	17	51	88	
$x_{86}$	1	5	4	0	-1	1	6	10	18	19	
$x_{88}$	1	3	3	0	0	1	1	7	1	1	
$x_{92}$	1	2	2	0	0	0	1	4	2	4	
$x_{103}$	0	2	2	0	0	0	4	4	9	16	
$x_{107}$	1	3	2	0	0	1	8	5	16	28	
$x_{111}$	0	1	1	0	0	0	3	2	6	9	
$x_{113}$	1	3	3	0	0	1	2	6	5	8	
$x_{125}$	9	31	29	-2	-5	7	5	68	7	12	
$x_{126}$	0	2	2	0	0	0	6	4	12	20	
$x_{127}$	1	4	4	0	-1	1	9	9	19	33	
$x_{131}$	1	4	3	0	-1	1	7	7	13	23	
$x_{148}$	1	3	2	0	0	1	7	5	15	25	
$x_{217}$	-28	-17	-24	0	1	-4	-30	-33	-48	-70	
	$x_2$	$x_9$	$x_{10}$	$x_{11}$	$x_{12}$	$x_{18}$	$x_{19}$	$x_{20}$	$x_{21}$	$x_{22}$	

$x_{31}$	2									
$x_{43}$	0	9								
$x_{52}$	3	14	5							
$x_{61}$	0	4	28	2						
$x_{70}$	2	12	3	16	1					
$x_{75}$	2	14	3	17	1	15				
$x_{78}$	2	11	3	38	1	13	14			
$x_{80}$	3	17	4	21	2	-4	20	18		
$x_{85}$	5	36	7	37	3	33	35	31	45	
$x_{86}$	1	43	4	8	2	7	11	7	10	20
$x_{88}$	0	2	18	1	5	1	1	1	1	1
$x_{92}$	0	2	10	8	3	1	2	3	2	3
$x_{103}$	1	5	4	37	1	6	6	14	8	14
$x_{107}$	2	9	2	12	1	10	11	10	14	25
$x_{111}$	1	7	1	4	0	3	4	3	5	9
$x_{113}$	0	23	2	3	1	3	3	3	4	8
$x_{125}$	1	15	27	5	13	5	5	4	6	12
$x_{126}$	1	7	2	9	1	8	8	7	10	18
$x_{127}$	2	11	4	14	2	12	13	11	17	29
$x_{131}$	1	8	3	10	1	9	9	8	12	20
$x_{148}$	2	8	2	11	1	9	10	9	13	22
$x_{217}$	-6	-56	-27	-66	-20	-28	-33	-43	-37	-67
	$x_{30}$	$x_{31}$	$x_{43}$	$x_{52}$	$x_{61}$	$x_{70}$	$x_{75}$	$x_{78}$	$x_{80}$	$x_{85}$
$x_{88}$	1									
$x_{92}$	1	2								
$x_{103}$	3	1	3							
$x_{107}$	5	0	1	4						
$x_{111}$	4	0	0	1	3					
$x_{113}$	10	0	0	1	2	2				
$x_{125}$	7	5	3	3	3	2	4			
$x_{126}$	4	0	1	3	6	2	2	2		
$x_{127}$	6	1	1	5	9	3	3	6	7	
$x_{131}$	5	1	1	4	6	2	2	5	5	8
$x_{148}$	5	0	1	4	7	2	2	4	5	8
$x_{217}$	-36	-14	-22	-33	-25	-19	-25	-23	-15	-26
	$x_{86}$	$x_{88}$	$x_{92}$	$x_{103}$	$x_{107}$	$x_{111}$	$x_{113}$	$x_{125}$	$x_{126}$	$x_{127}$



## $D^0$ BRANCHING RATIOS

See the "Note on  $D$  Mesons" in the  $D^\pm$  Listings.

Some older now obsolete results have been omitted from these Listings.

### ———— Inclusive modes ——

#### $\Gamma(e^+ \text{anything})/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
-------	------	-------------	------	---------

**0.0687±0.0028 OUR FIT**

**0.0675±0.0029 OUR AVERAGE**

0.069 ± 0.003 ± 0.005	1670	ALBRECHT	96C ARG	$e^+ e^- \approx 10 \text{ GeV}$
0.0664±0.0018±0.0029	4609	17 KUBOTA	96B CLE2	$e^+ e^- \approx \gamma(4S)$
0.075 ± 0.011 ± 0.004	137	BALTRUSAIT	..85B MRK3	$e^+ e^- 3.77 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.15 ± 0.05		AGUILAR-...	87E HYBR	$\pi p, pp 360, 400$ GeV
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0.055 ± 0.037	12	SCHINDLER	81 MRK2	$e^+ e^- 3.771 \text{ GeV}$
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17 KUBOTA 96B uses  $D^{*+} \rightarrow D^0 \pi^+$  (and charge conjugate) events in which the  $D^0$  subsequently decays to  $X e^+ \nu_e$ .

#### $\Gamma(\mu^+ \text{anything})/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
-------	------	-------------	------	---------

**0.065±0.008 OUR FIT**

**0.060±0.007±0.012**      310      ALBRECHT      96C ARG       $e^+ e^- \approx 10 \text{ GeV}$

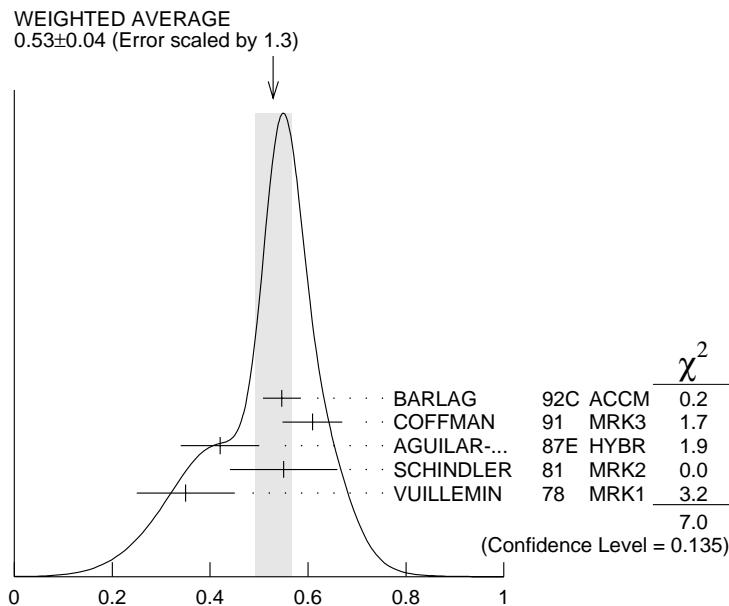
#### $\Gamma(K^- \text{anything})/\Gamma_{\text{total}}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
-------	------	-------------	------	---------

**0.53 ±0.04 OUR AVERAGE**      Error includes scale factor of 1.3. See the ideogram below.

0.546 <sup>+0.039</sup> <sub>-0.038</sub>	18 BARLAG	92C ACCM	$\pi^- \text{ Cu } 230 \text{ GeV}$
0.609±0.032±0.052	COFFMAN	91 MRK3	$e^+ e^- 3.77 \text{ GeV}$
0.42 ± 0.08	AGUILAR-...	87E HYBR	$\pi p, pp 360, 400 \text{ GeV}$
0.55 ± 0.11	121 SCHINDLER	81 MRK2	$e^+ e^- 3.771 \text{ GeV}$
0.35 ± 0.10	19 VUILLEMIN	78 MRK1	$e^+ e^- 3.772 \text{ GeV}$

18 BARLAG 92C computes the branching fraction using topological normalization.



$$\Gamma(K^- \text{anything})/\Gamma_{\text{total}}$$

$$[\Gamma(\bar{K}^0 \text{anything}) + \Gamma(K^0 \text{anything})]/\Gamma_{\text{total}} \quad \Gamma_4/\Gamma$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.42 ± 0.05 OUR AVERAGE</b>				
0.455 ± 0.050 ± 0.032		COFFMAN	91	MRK3 $e^+ e^-$ 3.77 GeV
0.29 ± 0.11	13	SCHINDLER	81	MRK2 $e^+ e^-$ 3.771 GeV
0.57 ± 0.26	6	VUILLEMIN	78	MRK1 $e^+ e^-$ 3.772 GeV

$$\Gamma(K^+ \text{anything})/\Gamma_{\text{total}} \quad \Gamma_5/\Gamma$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.034 ± 0.006 OUR AVERAGE</b>				
0.034 ± 0.007		19 BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV
0.028 ± 0.009 ± 0.004		COFFMAN	91	MRK3 $e^+ e^-$ 3.77 GeV
0.03 ± 0.05		AGUILAR-...	87E HYBR	$\pi p, pp$ 360, 400 GeV
0.08 ± 0.03	25	SCHINDLER	81	MRK2 $e^+ e^-$ 3.771 GeV

19 BARLAG 92C computes the branching fraction using topological normalization.

$$\Gamma(\phi \text{anything})/\Gamma_{\text{total}} \quad \Gamma_7/\Gamma$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0171 ± 0.0076 ± 0.0017</b>				
9	20 BAI	00C BES	$e^+ e^- \rightarrow D\bar{D}^*$ , $D^*\bar{D}^*$	

20 BAI 00C finds the average ( $\phi$  anything) branching fraction for the 4.03-GeV mix of  $D^+$  and  $D^0$  mesons to be  $(1.34 \pm 0.52 \pm 0.12)\%$ .

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Semileptonic modes

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 $\Gamma(K^-\ell^+\nu_\ell)/\Gamma_{\text{total}}$  $\Gamma_8/\Gamma$ 

We average our  $K^- e^+ \nu_e$  and  $K^- \mu^+ \nu_\mu$  branching fractions, after multiplying the latter by a phase-space factor of 1.03 to be able to use it with the  $K^- e^+ \nu_e$  fraction. Hence our  $\ell^+$  here is really an  $e^+$ .

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>COMMENT</u>
<b>0.0343±0.0015 OUR AVERAGE</b>		Error includes scale factor of 1.2.
0.0359±0.0018	PDG 02	Our $\Gamma(K^- e^+ \nu_e)/\Gamma_{\text{total}}$
0.0329±0.0017	PDG 02	$1.03 \times$ our $\Gamma(K^- \mu^+ \nu_\mu)/\Gamma_{\text{total}}$

 $\Gamma(K^- e^+ \nu_e)/\Gamma_{\text{total}}$  $\Gamma_9/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0358±0.0018 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.034 ±0.005 ±0.004</b>	55	ADLER	89	MRK3 $e^+ e^-$ 3.77 GeV

 $\Gamma(K^- e^+ \nu_e)/\Gamma(K^-\pi^+)$  $\Gamma_9/\Gamma_{20}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.94 ±0.04 OUR FIT</b>				
<b>0.95 ±0.04 OUR AVERAGE</b>				
0.978±0.027±0.044	2510	21 BEAN	93C CLE2	$e^+ e^- \approx \gamma(4S)$
0.90 ±0.06 ±0.06	584	22 CRAWFORD	91B CLEO	$e^+ e^- \approx 10.5$ GeV
0.91 ±0.07 ±0.11	250	23 ANJOS	89F E691	Photoproduction

- <sup>21</sup> BEAN 93C uses  $K^- \mu^+ \nu_\mu$  as well as  $K^- e^+ \nu_e$  events and makes a small phase-space adjustment to the number of the  $\mu^+$  events to use them as  $e^+$  events. A pole mass of  $2.00 \pm 0.12 \pm 0.18$  GeV/ $c^2$  is obtained from the  $q^2$  dependence of the decay rate.  
<sup>22</sup> CRAWFORD 91B uses  $K^- e^+ \nu_e$  and  $K^- \mu^+ \nu_\mu$  candidates to measure a pole mass of  $2.1^{+0.4+0.3}_{-0.2-0.2}$  GeV/ $c^2$  from the  $q^2$  dependence of the decay rate.  
<sup>23</sup> ANJOS 89F measures a pole mass of  $2.1^{+0.4}_{-0.2} \pm 0.2$  GeV/ $c^2$  from the  $q^2$  dependence of the decay rate.

 $\Gamma(K^-\mu^+\nu_\mu)/\Gamma(K^-\pi^+)$  $\Gamma_{10}/\Gamma_{20}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.84 ±0.04 OUR FIT</b>				
<b>0.84 ±0.04 OUR AVERAGE</b>				
0.852±0.034±0.028	1897	24 FRABETTI	95G E687	$\gamma Be \bar{E}_\gamma = 220$ GeV
0.82 ±0.13 ±0.13	338	25 FRABETTI	93I E687	$\gamma Be \bar{E}_\gamma = 221$ GeV
0.79 ±0.08 ±0.09	231	26 CRAWFORD	91B CLEO	$e^+ e^- \approx 10.5$ GeV

- <sup>24</sup> FRABETTI 95G extracts the ratio of form factors  $f_-(0)/f_+(0) = -1.3^{+3.6}_{-3.4} \pm 0.6$ , and measures a pole mass of  $1.87^{+0.11+0.07}_{-0.08-0.06}$  GeV/ $c^2$  from the  $q^2$  dependence of the decay rate.  
<sup>25</sup> FRABETTI 93I measures a pole mass of  $2.1^{+0.7+0.7}_{-0.3-0.3}$  GeV/ $c^2$  from the  $q^2$  dependence of the decay rate.  
<sup>26</sup> CRAWFORD 91B measures a pole mass of  $2.00 \pm 0.12 \pm 0.18$  GeV/ $c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(K^-\mu^+\nu_\mu)/\Gamma(\mu^+ \text{anything})$   $\Gamma_{10}/\Gamma_2$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.49 ± 0.06 OUR FIT</b>				
<b>0.472±0.051±0.040</b>	232	KODAMA	94	E653 $\pi^-$ emulsion 600 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.32 ± 0.05 ± 0.05	124	KODAMA	91	EMUL $pA$ 800 GeV

 $\Gamma(K^-\pi^0e^+\nu_e)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.011<sup>+0.008</sup><sub>-0.006</sub> OUR FIT</b>				Error includes scale factor of 1.6.
<b>0.016<sup>+0.013</sup><sub>-0.005</sub>±0.002</b>	4	27 BAI	91	MRK3 $e^+e^- \approx 3.77$ GeV
27 BAI 91 finds that a fraction $0.79^{+0.15+0.09}_{-0.17-0.03}$ of combined $D^+$ and $D^0$ decays to $\bar{K}\pi e^+\nu_e$ (24 events) are $\bar{K}^*(892)e^+\nu_e$ . BAI 91 uses 56 $K^-e^+\nu_e$ events to measure a pole mass of $1.8 \pm 0.3 \pm 0.2$ GeV/ $c^2$ from the $q^2$ dependence of the decay rate.				

 $\Gamma(\bar{K}^0\pi^-e^+\nu_e)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.018±0.008 OUR FIT</b>				Error includes scale factor of 1.6.
<b>0.028<sup>+0.017</sup><sub>-0.008</sub>±0.003</b>	6	28 BAI	91	MRK3 $e^+e^- \approx 3.77$ GeV
28 BAI 91 finds that a fraction $0.79^{+0.15+0.09}_{-0.17-0.03}$ of combined $D^+$ and $D^0$ decays to $\bar{K}\pi e^+\nu_e$ (24 events) are $\bar{K}^*(892)e^+\nu_e$ .				

 $\Gamma(K^*(892)^-e^+\nu_e)/\Gamma(K^-\bar{e}^+\nu_e)$   $\Gamma_{19}/\Gamma_9$ 

Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.60±0.10 OUR FIT</b>				
<b>0.51±0.18±0.06</b>		CRAWFORD	91B CLEO	$e^+e^- \approx 10.5$ GeV

 $\Gamma(K^*(892)^-e^+\nu_e)/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{19}/\Gamma_{22}$ 

Unseen decay modes of the  $\bar{K}^*(892)^-$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.36±0.06 OUR FIT</b>				
<b>0.38±0.06±0.03</b>	152	29 BEAN	93C CLE2	$e^+e^- \approx \gamma(4S)$

29 BEAN 93C uses  $K^*-\mu^+\nu_\mu$  as well as  $K^*-\bar{e}^+\nu_e$  events and makes a small phase-space adjustment to the number of the  $\mu^+$  events to use them as  $e^+$  events.

 $\Gamma(K^*(892)^-\ell^+\nu_\ell)/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{14}/\Gamma_{22}$ 

This is an average of the  $K^*(892)^-e^+\nu_e$  and  $K^*(892)^-\mu^+\nu_\mu$  ratios. Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.24±0.07±0.06	137	30 ALEXANDER	90B CLEO	$e^+e^- 10.5-11$ GeV

30 ALEXANDER 90B cannot exclude extra  $\pi^0$ 's in the final state. See nearby data blocks for more detailed results.

$\Gamma(\bar{K}^*(892)^0 \pi^- e^+ \nu_e)/\Gamma(K^*(892)^- e^+ \nu_e)$   $\Gamma_{15}/\Gamma_{19}$ Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
<0.64	90	31 CRAWFORD	91B CLEO	$e^+ e^- \approx 10.5$ GeV
31 The limit on $(\bar{K}^*(892)\pi)^- \mu^+ \nu_\mu$ below is much stronger.				

 $\Gamma(K^- \pi^+ \pi^- \mu^+ \nu_\mu)/\Gamma(K^- \mu^+ \nu_\mu)$   $\Gamma_{16}/\Gamma_{10}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.037</b>	90	KODAMA	93B E653	$\pi^-$ emulsion 600 GeV

 $\Gamma((\bar{K}^*(892)\pi)^- \mu^+ \nu_\mu)/\Gamma(K^- \mu^+ \nu_\mu)$   $\Gamma_{17}/\Gamma_{10}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.043</b>	90	32 KODAMA	93B E653	$\pi^-$ emulsion 600 GeV
32 KODAMA 93B searched in $K^- \pi^+ \pi^- \mu^+ \nu_\mu$ , but the limit includes other $(\bar{K}^*(892)\pi)^-$ charge states.				

 $\Gamma(\pi^- e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{18}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0036 ± 0.0006 OUR FIT</b>				
<b>0.0039 ± 0.0023 ± 0.0004</b>	7	33 ADLER	89 MRK3	$e^+ e^-$ 3.77 GeV

33 This result of ADLER 89 gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.057^{+0.038}_{-0.015} \pm 0.005$ . $\Gamma(\pi^- e^+ \nu_e)/\Gamma(K^- e^+ \nu_e)$   $\Gamma_{18}/\Gamma_9$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.101 ± 0.017 OUR FIT</b>				
<b>0.101 ± 0.018 OUR AVERAGE</b>				
0.101 ± 0.020 ± 0.003	91	34 FRABETTI	96B E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
0.103 ± 0.039 ± 0.013	87	35 BUTLER	95 CLE2	< 0.156 (90% CL)

34 FRABETTI 96B uses both  $e$  and  $\mu$  events, and makes a small correction to the  $\mu$  events to make them effectively  $e$  events. This result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.050 \pm 0.011 \pm 0.002$ .35 BUTLER 95 has  $87 \pm 33 \pi^- e^+ \nu_e$  events. The result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^\pi(0)}{f_+^K(0)}|^2 = 0.052 \pm 0.020 \pm 0.007$ .**———— Hadronic modes with a  $\bar{K}$  or  $\bar{K}K\bar{K}$  ——** $\Gamma(K^- \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{20}/\Gamma$ We list measurements *before* radiative corrections are made.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0380 ± 0.0009 OUR FIT</b>				
<b>0.0385 ± 0.0009 OUR AVERAGE</b>				
0.0382 ± 0.0007 ± 0.0012		36 ARTUSO	98 CLE2	CLEO average
0.0390 ± 0.0009 ± 0.0012	5392	37 BARATE	97C ALEP	From $Z$ decays
0.045 ± 0.006 ± 0.004		38 ALBRECHT	94 ARG	$e^+ e^- \approx \Upsilon(4S)$
0.0341 ± 0.0012 ± 0.0028	1173	37 ALBRECHT	94F ARG	$e^+ e^- \approx \Upsilon(4S)$

$0.0362 \pm 0.0034 \pm 0.0044$	<sup>37</sup> DECOMP	91J ALEP	From $Z$ decays
$0.045 \pm 0.008 \pm 0.005$	<sup>56</sup> ABACHI	88 HRS	$e^+ e^-$ 29 GeV
$0.042 \pm 0.004 \pm 0.004$	930 ADLER	88C MRK3	$e^+ e^-$ 3.77 GeV
$0.041 \pm 0.006$	263 SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV
$0.043 \pm 0.010$	130 PERUZZI	77 MRK1	$e^+ e^-$ 3.77 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.0381 \pm 0.0015 \pm 0.0016$	1165	<sup>41</sup> ARTUSO	98 CLE2	$e^+ e^-$ at $\gamma(4S)$
$0.0369 \pm 0.0011 \pm 0.0016$		<sup>42</sup> COAN	98 CLE2	
$0.0391 \pm 0.0008 \pm 0.0017$	4208	<sup>37,43</sup> AKERIB	93 CLE2	$e^+ e^- \approx \gamma(4S)$

<sup>36</sup> This combines the CLEO results of ARTUSO 98, COAN 98, and AKERIB 93.

<sup>37</sup> ABACHI 88, DECOMP 91J, AKERIB 93, ALBRECHT 94F, and BARATE 97C use  $D^*(2010)^+ \rightarrow D^0 \pi^+$  decays. The  $\pi^+$  is both slow and of low  $p_T$  with respect to the event thrust axis or nearest jet ( $\approx D^{*+}$  direction). The excess number of such  $\pi^+$ 's over background gives the number of  $D^*(2010)^+ \rightarrow D^0 \pi^+$  events, and the fraction with  $D^0 \rightarrow K^- \pi^+$  gives the  $D^0 \rightarrow K^- \pi^+$  branching fraction.

<sup>38</sup> ALBRECHT 94 uses  $D^0$  mesons from  $\bar{B}^0 \rightarrow D^{*+} \ell^- \bar{\nu}_\ell$  decays. This is a different set of events than used by ALBRECHT 94F.

<sup>39</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.24 \pm 0.02$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

<sup>40</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.25 \pm 0.05$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

<sup>41</sup> ARTUSO 98, following ALBRECHT 94, uses  $D^0$  mesons from  $\bar{B}^0 \rightarrow D^*(2010)^+ X \ell^- \bar{\nu}_\ell$  decays. Our average uses the CLEO average of this value with the values of COAN 98 and AKERIB 93.

<sup>42</sup> COAN 98 assumes that  $\Gamma(B \rightarrow \bar{D} X \ell^+ \nu)/\Gamma(B \rightarrow X \ell^+ \nu) = 1.0 - 3|V_{ub}/V_{cb}|^2 - 0.010 \pm 0.005$ , the last term accounting for  $\bar{B} \rightarrow D_s^+ K X \ell^- \bar{\nu}$ . COAN 98 is included in the CLEO average in ARTUSO 98.

<sup>43</sup> This AKERIB 93 value does not include radiative corrections; with them, the value is  $0.0395 \pm 0.0008 \pm 0.0017$ . AKERIB 93 is included in the CLEO average in ARTUSO 98.

### $\Gamma(\bar{K}^0 \pi^0)/\Gamma(K^- \pi^+)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.60 \pm 0.06</math> OUR FIT</b>				
<b><math>1.36 \pm 0.23 \pm 0.22</math></b>	119	ANJOS	92B E691	$\gamma$ Be 80–240 GeV

### $\Gamma(\bar{K}^0 \pi^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.385 \pm 0.031</math> OUR FIT</b>				
<b><math>0.378 \pm 0.033</math> OUR AVERAGE</b>				
$0.44 \pm 0.02 \pm 0.05$	1942	PROCARIO	93B CLE2	$e^+ e^-$ 10.36–10.7 GeV
$0.34 \pm 0.04 \pm 0.02$	92	<sup>44</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV
$0.36 \pm 0.04 \pm 0.08$	104	KINOSHITA	91 CLEO	$e^+ e^- \sim 10.7$ GeV

<sup>44</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^0 \pi^+ \pi^-)/\Gamma_{\text{total}}$  $\Gamma_{22}/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0597 ± 0.0035 OUR NEW UNCHECKED FIT</b>				Error includes scale factor of 1.1. [0.0592 ± 0.0035 OUR 2002 FIT Scale factor = 1.1]
<b>0.055 ± 0.005 OUR AVERAGE</b>				
0.0503 ± 0.0039 ± 0.0049	284	45 ALBRECHT	94F ARG	$e^+ e^- \approx \gamma(4S)$
0.064 ± 0.005 ± 0.010		ADLER	87 MRK3	$e^+ e^- 3.77 \text{ GeV}$
0.052 ± 0.016	32	46 SCHINDLER	81 MRK2	$e^+ e^- 3.771 \text{ GeV}$
0.079 ± 0.023	28	47 PERUZZI	77 MRK1	$e^+ e^- 3.77 \text{ GeV}$

<sup>45</sup> See the footnote on the ALBRECHT 94F measurement of  $\Gamma(K^-\pi^+)/\Gamma_{\text{total}}$  for the method used.

<sup>46</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times \text{branching fraction}$  to be  $0.30 \pm 0.08 \text{ nb}$ . We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6 \text{ nb}$ .

<sup>47</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times \text{branching fraction}$  to be  $0.46 \pm 0.12 \text{ nb}$ . We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6 \text{ nb}$ .

 $\Gamma(\bar{K}^0 \pi^+ \pi^-)/\Gamma(K^-\pi^+)$  $\Gamma_{22}/\Gamma_{20}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.57 ± 0.09 OUR NEW UNCHECKED FIT</b>				Error includes scale factor of 1.1. [1.56 ± 0.09 OUR 2002 FIT Scale factor = 1.1]
<b>1.65 ± 0.17 OUR AVERAGE</b>				
1.61 ± 0.10 ± 0.15	856	FRABETTI	94J E687	$\gamma\text{Be } \bar{E}_\gamma = 220 \text{ GeV}$
1.7 ± 0.8	35	AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$
2.8 ± 1.0	116	PICCOLO	77 MRK1	$e^+ e^- 4.03, 4.41 \text{ GeV}$

 $\Gamma(\bar{K}^0 \rho^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{23}/\Gamma_{22}$ 

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.259<sup>+0.014</sup><sub>-0.023</sub> OUR NEW AVERAGE</b>				Error includes scale factor of 1.1. [0.25 ± 0.05 OUR 2002 AVERAGE Scale factor = 1.5]
0.264 ± 0.009 <sup>+0.010</sup> <sub>-0.026</sub>		MURAMATSU 02 CLE2	$e^+ e^- \approx 10 \text{ GeV}$	
0.350 ± 0.028 ± 0.067		FRABETTI 94G E687	$\gamma\text{Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$	
0.227 ± 0.032 ± 0.009		ALBRECHT 93D ARG	$e^+ e^- \approx 10 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.215 ± 0.051 ± 0.037		ANJOS 93 E691	$\gamma\text{Be } 90\text{--}260 \text{ GeV}$	
0.20 ± 0.06 ± 0.03		FRABETTI 92B E687	$\gamma\text{Be } \bar{E}_\gamma = 221 \text{ GeV}$	
0.12 ± 0.01 ± 0.07		ADLER 87 MRK3	$e^+ e^- 3.77 \text{ GeV}$	

 $\Gamma(\bar{K}^0 f_0(980) \times \mathcal{B}(f_0 \rightarrow \pi^+ \pi^-))/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{24}/\Gamma_{22}$ 

This includes only  $\pi^+ \pi^-$  decays of the  $f_0(980)$ , because branching fractions of this resonance are not known.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.047<sup>+0.010</sup><sub>-0.007</sub> OUR NEW AVERAGE</b>				[0.054 ± 0.015 OUR 2002 AVERAGE]
0.043 ± 0.005 <sup>+0.012</sup> <sub>-0.006</sub>		MURAMATSU 02 CLE2	$e^+ e^- \approx 10 \text{ GeV}$	
0.068 ± 0.016 ± 0.018		FRABETTI 94G E687	$\gamma\text{Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$	
0.046 ± 0.018 ± 0.006		ALBRECHT 93D ARG	$e^+ e^- \approx 10 \text{ GeV}$	

$\Gamma(\bar{K}^0 f_2(1270))/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{83}/\Gamma_{22}$ Unseen decay modes of the  $f_2(1270)$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.008</b> $^{+0.007}_{-0.004}$ OUR NEW AVERAGE	[ $0.076 \pm 0.028$ OUR 2002 AVERAGE]		
$0.0048 \pm 0.0027$ $^{+0.0065}_{-0.0029}$	MURAMATSU 02	CLE2	$e^+ e^- \approx 10$ GeV
$0.065 \pm 0.025$ $\pm 0.030$	FRABETTI 94G	E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
$0.088 \pm 0.037$ $\pm 0.014$	ALBRECHT 93D	ARG	$e^+ e^- \approx 10$ GeV

 $\Gamma(\bar{K}^0 f_0(1370) \times B(f_0 \rightarrow \pi^+ \pi^-))/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{26}/\Gamma_{22}$ This includes only  $\pi^+ \pi^-$  decays of the  $f_0(1370)$ , because branching fractions of this resonance are not known.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.085</b> $^{+0.019}_{-0.021}$ OUR NEW AVERAGE	[ $0.080 \pm 0.024$ OUR 2002 AVERAGE]		
$0.099 \pm 0.011$ $^{+0.028}_{-0.044}$	MURAMATSU 02	CLE2	$e^+ e^- \approx 10$ GeV
$0.077 \pm 0.022$ $\pm 0.031$	FRABETTI 94G	E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
$0.082 \pm 0.028$ $\pm 0.013$	ALBRECHT 93D	ARG	$e^+ e^- \approx 10$ GeV

 $\Gamma(K^*(892)^- \pi^+)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{85}/\Gamma_{22}$ Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.997</b> $^{+0.030}_{-0.034}$ OUR NEW UNCHECKED FIT		[ $1.01 \pm 0.06$ OUR 2002 FIT Scale factor = 1.4]		
<b>0.991</b> $^{+0.028}_{-0.040}$ OUR NEW AVERAGE		[ $1.00 \pm 0.07$ OUR 2002 AVERAGE Scale factor = 1.4]		
$0.986 \pm 0.020$ $^{+0.027}_{-0.063}$		MURAMATSU 02	CLE2	$e^+ e^- \approx 10$ GeV
$0.938 \pm 0.054$ $\pm 0.038$		FRABETTI 94G	E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
$1.08 \pm 0.063$ $\pm 0.045$		ALBRECHT 93D	ARG	$e^+ e^- \approx 10$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.720 \pm 0.145$ $\pm 0.185$		ANJOS 93	E691	$\gamma$ Be 90–260 GeV
$0.96 \pm 0.12$ $\pm 0.075$		FRABETTI 92B	E687	$\gamma$ Be $\bar{E}_\gamma = 221$ GeV
$0.84 \pm 0.06$ $\pm 0.08$		ADLER 87	MRK3	$e^+ e^-$ 3.77 GeV
$1.05$ $^{+0.23}_{-0.26}$ $^{+0.07}_{-0.09}$	25	SCHINDLER 81	MRK2	$e^+ e^-$ 3.771 GeV

 $\Gamma(K_0^*(1430)^- \pi^+)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{107}/\Gamma_{22}$ Unseen decay modes of the  $\bar{K}_0^*(1430)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT	
<b>0.165</b> $^{+0.033}_{-0.021}$ OUR NEW UNCHECKED FIT		[ $0.20 \pm 0.04$ OUR 2002 FIT]		
<b>0.154</b> $^{+0.034}_{-0.019}$ OUR NEW AVERAGE		[ $0.19 \pm 0.05$ OUR 2002 AVERAGE]		
$0.118 \pm 0.011$ $^{+0.050}_{-0.018}$		MURAMATSU 02	CLE2	$e^+ e^- \approx 10$ GeV
$0.176 \pm 0.044$ $\pm 0.047$		FRABETTI 94G	E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
$0.208 \pm 0.055$ $\pm 0.034$		ALBRECHT 93D	ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(K_2^*(1430)^-\pi^+)/\Gamma(\bar{K}^0\pi^+\pi^-)$  $\Gamma_{109}/\Gamma_{22}$ Unseen decay modes of the  $\bar{K}_2^*(1430)^-$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**0.033±0.006<sup>+0.020</sup><sub>-0.010</sub>** MURAMATSU 02 CLE2  $e^+e^- \approx 10$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.15 90 ALBRECHT 93D ARG  $e^+e^- \approx 10$  GeV $\Gamma(K^*(1680)^-\pi^+)/\Gamma(\bar{K}^0\pi^+\pi^-)$  $\Gamma_{111}/\Gamma_{22}$ Unseen decay modes of the  $K^*(1680)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.14<sup>+0.07</sup><sub>-0.06</sub> OUR FIT** Error includes scale factor of 1.2.**0.085±0.016<sup>+0.069</sup><sub>-0.059</sub>** MURAMATSU 02 CLE2  $e^+e^- \approx 10$  GeV $\Gamma(\bar{K}^0\pi^+\pi^- \text{ nonresonant})/\Gamma(\bar{K}^0\pi^+\pi^-)$  $\Gamma_{30}/\Gamma_{22}$ 

Neither FRABETTI 94G nor ALBRECHT 93D sees evidence for a nonresonant component.

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.009<sup>+0.020</sup><sub>-0.005</sub> OUR FIT****0.009<sup>+0.020</sup><sub>-0.006</sub> OUR NEW AVERAGE** [0.27 ± 0.04 OUR 2000 AVERAGE]**0.009±0.004<sup>+0.020</sup><sub>-0.004</sub>** MURAMATSU 02 CLE2  $e^+e^- \approx 10$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.263±0.024±0.041 ANJOS 93 E691  $\gamma Be$  90–260 GeV0.26 ± 0.08 ± 0.05 FRABETTI 92B E687  $\gamma Be$   $\bar{E}_\gamma = 221$  GeV0.33 ± 0.05 ± 0.10 ADLER 87 MRK3  $e^+e^-$  3.77 GeV $\Gamma(K^-\pi^+\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{31}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.130±0.008 OUR NEW UNCHECKED FIT** Error includes scale factor of 1.3. [0.131 ± 0.009 OUR 2002 FIT Scale factor = 1.3]**0.131±0.016 OUR AVERAGE**0.133±0.012±0.013 931 ADLER 88C MRK3  $e^+e^-$  3.77 GeV0.117±0.043 37 48 SCHINDLER 81 MRK2  $e^+e^-$  3.771 GeV48 SCHINDLER 81 (MARK-2) measures  $\sigma(e^+e^- \rightarrow \psi(3770)) \times$  branching fraction to be 0.68 ± 0.23 nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb. $\Gamma(K^-\pi^+\pi^0)/\Gamma(K^-\pi^+)$  $\Gamma_{31}/\Gamma_{20}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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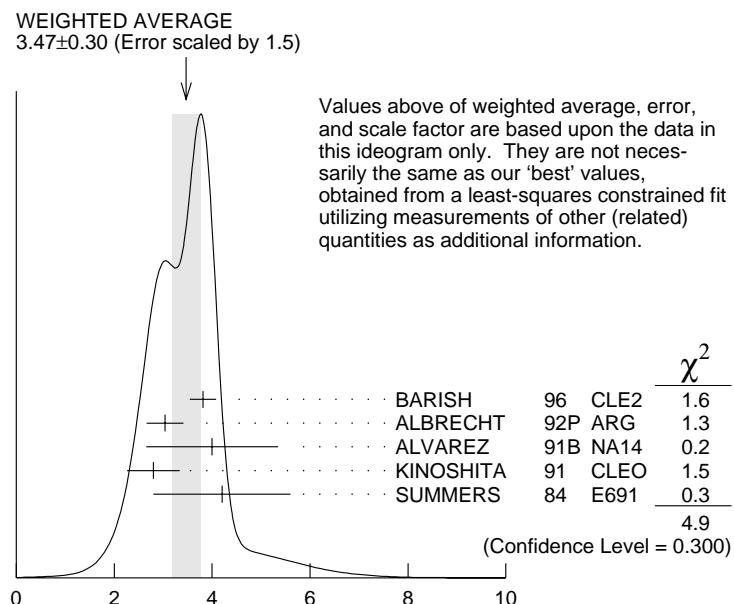
**3.42±0.22 OUR NEW UNCHECKED FIT** Error includes scale factor of 1.3. [3.44 ± 0.22 OUR 2002 FIT Scale factor = 1.3]**3.47±0.30 OUR AVERAGE** Error includes scale factor of 1.5. See the ideogram below.3.81±0.07±0.26 10k BARISH 96 CLE2  $e^+e^- \approx \Upsilon(4S)$ 3.04±0.16±0.34 931 49 ALBRECHT 92P ARG  $e^+e^- \approx 10$  GeV

4.0 ± 0.9 ± 1.0 69 ALVAREZ 91B NA14 Photoproduction

2.8 ± 0.14±0.52 1050 KINOSHITA 91 CLEO  $e^+e^- \sim 10.7$  GeV

4.2 ± 1.4 41 SUMMERS 84 E691 Photoproduction

49 This value is calculated from numbers in Table 1 of ALBRECHT 92P.



$$\Gamma(K^- \pi^+ \pi^0) / \Gamma(K^- \pi^+)$$

$$\Gamma(K^- \rho^+)/\Gamma(K^- \pi^+ \pi^0)$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.78 ±0.04 OUR AVERAGE</b>				
0.788±0.019±0.048		KOPP 01	CLE2	$e^+ e^- \approx 10.6 \text{ GeV}$
0.765±0.041±0.054		FRABETTI 94G	E687	$\gamma \text{Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.647±0.039±0.150		ANJOS 93	E691	$\gamma \text{Be} 90\text{--}260 \text{ GeV}$
0.81 ±0.03 ±0.06		ADLER 87	MRK3	$e^+ e^- 3.77 \text{ GeV}$
0.31 $\begin{array}{l} +0.20 \\ -0.14 \end{array}$	13	SUMMERS 84	E691	Photoproduction
0.85 $\begin{array}{l} +0.11 \\ -0.15 \end{array}$ $\begin{array}{l} +0.09 \\ -0.10 \end{array}$	31	SCHINDLER 81	MRK2	$e^+ e^- 3.771 \text{ GeV}$

$$\Gamma(K^- \rho(1700)^+ \times B(\rho(1700)^+ \rightarrow \pi^+ \pi^0)) / \Gamma(K^- \pi^+ \pi^0)$$

$$\Gamma_{32}/\Gamma_{31}$$

This only includes  $\pi^+ \pi^0$  decays of the  $\rho(1700)^+$ , because branching fractions of this resonance are not known.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.057±0.008±0.009</b>	KOPP 01	CLE2	$e^+ e^- \approx 10.6 \text{ GeV}$

$\Gamma(K^*(892)^-\pi^+)/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{85}/\Gamma_{31}$ Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.457±0.034 OUR NEW UNCHECKED FIT</b>	Error includes scale factor of 1.2. [0.46 ± 0.04 OUR 2002 FIT Scale factor = 1.2]		

**0.48 +0.08 -0.04 OUR AVERAGE**

0.483±0.021 +0.081 -0.032	KOPP	01 CLE2	$e^+e^- \approx 10.6 \text{ GeV}$
0.444±0.084 +0.147	FRABETTI	94G E687	$\gamma\text{Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.252±0.033 +0.035	ANJOS	93 E691	$\gamma\text{Be} 90\text{--}260 \text{ GeV}$
0.36 ± 0.06 ± 0.09	ADLER	87 MRK3	$e^+e^- 3.77 \text{ GeV}$

 $\Gamma(\bar{K}^*(892)^0\pi^0)/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{86}/\Gamma_{31}$ Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.214±0.027 OUR NEW UNCHECKED FIT</b>	Error includes scale factor of 1.1. [0.214 ± 0.026 OUR 2002 FIT Scale factor = 1.1]		

**0.204±0.025 OUR AVERAGE**

0.191±0.014 +0.024	KOPP	01 CLE2	$e^+e^- \approx 10.6 \text{ GeV}$
0.248±0.047 +0.023	FRABETTI	94G E687	$\gamma\text{Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.213±0.027 +0.035	ANJOS	93 E691	$\gamma\text{Be} 90\text{--}260 \text{ GeV}$
0.20 ± 0.03 ± 0.05	ADLER	87 MRK3	$e^+e^- 3.77 \text{ GeV}$

 $\Gamma(K_0^*(1430)^-\pi^+)/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{107}/\Gamma_{31}$ Unseen decay modes of the  $K_0^*(1430)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.075 +0.016 -0.010 OUR NEW UNCHECKED FIT</b>	[0.090 ± 0.020 OUR 2002 FIT]		

**0.107±0.019±0.045**

KOPP	01 CLE2	$e^+e^- \approx 10.6 \text{ GeV}$
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 $\Gamma(\bar{K}_0^*(1430)^0\pi^0)/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{108}/\Gamma_{31}$ Unseen decay modes of the  $\bar{K}_0^*(1430)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.066 +0.051 -0.010</b>	[0.090 ± 0.020 OUR 2002 FIT]		

 $\Gamma(K^*(1680)^-\pi^+)/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{111}/\Gamma_{31}$ Unseen decay modes of the  $K^*(1680)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.063 +0.031 -0.027 OUR FIT</b>	Error includes scale factor of 1.2.		

**0.101±0.023±0.033**

KOPP	01 CLE2	$e^+e^- \approx 10.6 \text{ GeV}$
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$\Gamma(K^-\pi^+\pi^0 \text{ nonresonant})/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{39}/\Gamma_{31}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.080<sup>+0.038</sup><sub>-0.014</sub> OUR AVERAGE</b>				
0.075 $\pm 0.009^{+0.056}_{-0.011}$	KOPP	01 CLE2	$e^+e^- \approx 10.6 \text{ GeV}$	
0.101 $\pm 0.033 \pm 0.040$	FRABETTI	94G E687	$\gamma\text{Be}, \bar{E}_\gamma \approx 220 \text{ GeV}$	
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
0.036 $\pm 0.004 \pm 0.018$	ANJOS	93 E691	$\gamma\text{Be} 90\text{--}260 \text{ GeV}$	
0.09 $\pm 0.02 \pm 0.04$	ADLER	87 MRK3	$e^+e^- 3.77 \text{ GeV}$	
0.51 $\pm 0.22$	21 SUMMERS	84 E691	Photoproduction	

 $\Gamma(\bar{K}^*(892)^0\pi^0)/\Gamma(\bar{K}^0\pi^0)$  $\Gamma_{86}/\Gamma_{21}$ Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.22<math>\pm 0.20</math> OUR NEW UNCHECKED FIT</b>				Error includes scale factor of 1.2. [1.23 $\pm 0.20$ OUR 2002 FIT Scale factor = 1.2]
<b>1.65<sup>+0.39</sup><sub>-0.31</sub> <math>\pm 0.20</math></b>	122	PROCARIO	93B CLE2	$\bar{K}^0\pi^0\pi^0$ Dalitz plot

 $\Gamma(\bar{K}_2^*(1430)^0\pi^0)/\Gamma(\bar{K}^*(892)^0\pi^0)$  $\Gamma_{110}/\Gamma_{86}$ Unseen decay modes of the  $\bar{K}_2^*(1430)^0$  and  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.12	90	PROCARIO	93B CLE2	$\bar{K}^0\pi^0\pi^0$ Dalitz plot

 $\Gamma(\bar{K}^0\pi^0\pi^0 \text{ nonresonant})/\Gamma(\bar{K}^0\pi^0)$  $\Gamma_{42}/\Gamma_{21}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.37<math>\pm 0.08 \pm 0.04</math></b>	76	PROCARIO	93B CLE2	$\bar{K}^0\pi^0\pi^0$ Dalitz plot

 $\Gamma(K^-\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}$  $\Gamma_{43}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0746<math>\pm 0.0031</math> OUR FIT</b>				

**0.075  $\pm 0.006$  OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

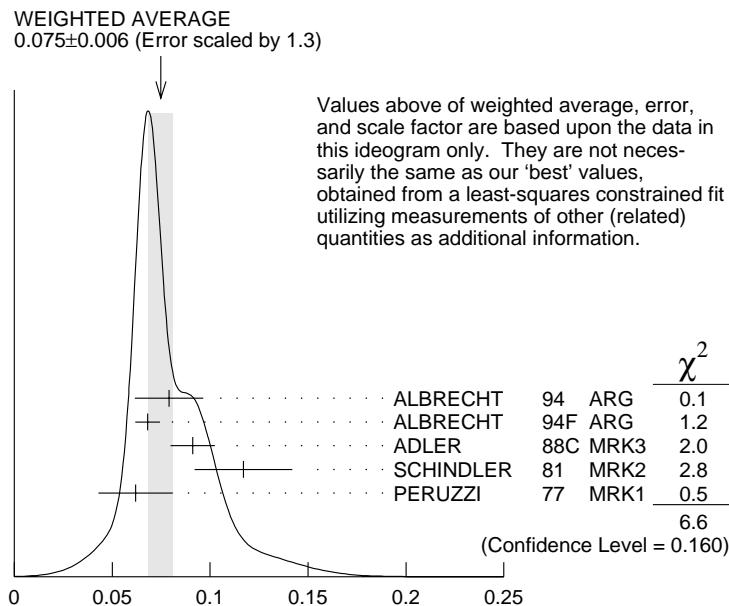
0.079 $\pm 0.015 \pm 0.009$	50 ALBRECHT	94 ARG	$e^+e^- \approx \gamma(4S)$
0.0680 $\pm 0.0027 \pm 0.0057$	1430	51 ALBRECHT	94F ARG $e^+e^- \approx \gamma(4S)$
0.091 $\pm 0.008 \pm 0.008$	992	ADLER	88C MRK3 $e^+e^- 3.77 \text{ GeV}$
0.117 $\pm 0.025$	185	52 SCHINDLER	81 MRK2 $e^+e^- 3.771 \text{ GeV}$
0.062 $\pm 0.019$	44	53 PERUZZI	77 MRK1 $e^+e^- 3.77 \text{ GeV}$

<sup>50</sup> ALBRECHT 94 uses  $D^0$  mesons from  $\bar{B}^0 \rightarrow D^*+\ell^-\bar{\nu}_\ell$  decays. This is a different set of events than used by ALBRECHT 94F.

<sup>51</sup> See the footnote on the ALBRECHT 94F measurement of  $\Gamma(K^-\pi^+)/\Gamma_{\text{total}}$  for the method used.

<sup>52</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+e^- \rightarrow \psi(3770)) \times \text{branching fraction}$  to be  $0.68 \pm 0.11 \text{ nb}$ . We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6 \text{ nb}$ .

<sup>53</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+e^- \rightarrow \psi(3770)) \times \text{branching fraction}$  to be  $0.36 \pm 0.10 \text{ nb}$ . We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6 \text{ nb}$ .



$$\Gamma(K^- \pi^+ \pi^+ \pi^-) / \Gamma_{\text{total}}$$

$$\Gamma(K^- \pi^+ \pi^+ \pi^-) / \Gamma(K^- \pi^+)$$

$$\Gamma_{43}/\Gamma_{20}$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.96±0.08 OUR FIT</b>				
<b>1.97±0.09 OUR AVERAGE</b>				
1.94±0.07 <sup>+0.09</sup> <sub>-0.11</sub>		JUN	00 SELX	$\Sigma^-$ nucleus, 600 GeV
1.7 ± 0.2 ± 0.2	1745	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
1.90±0.25±0.20	337	ALVAREZ	91B NA14	Photoproduction
2.12±0.16±0.09		BORTOLETTI088	CLEO	$e^+ e^-$ 10.55 GeV
2.0 ± 0.9	48	BAILEY	86 ACCM	$\pi^-$ Be fixed target
2.17±0.28±0.23		ALBRECHT	85F ARG	$e^+ e^-$ 10 GeV
2.0 ± 1.0	10	BAILEY	83B SPEC	$\pi^-$ Be → $D^0$
2.2 ± 0.8	214	PICCOLO	77 MRK1	$e^+ e^-$ 4.03, 4.41 GeV

$$\Gamma(K^- \pi^+ \rho^0 \text{total}) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$$

$$\Gamma_{44}/\Gamma_{43}$$

This includes  $K^- a_1(1260)^+$ ,  $\bar{K}^*(892)^0 \rho^0$ , etc. The next entry gives the specifically 3-body fraction. We rely on the MARK III and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.835±0.035 OUR AVERAGE</b>			
0.80 ± 0.03 ± 0.05	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
0.855±0.032±0.030	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.98 ± 0.12 ± 0.10	ALVAREZ	91B NA14	Photoproduction

$\Gamma(K^-\pi^+\rho^0\text{3-body})/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{45}/\Gamma_{43}$ 

We rely on the MARK III and E691 full amplitude analyses of the  $K^-\pi^+\pi^+\pi^-$  channel for values of the resonant substructure.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.063±0.028 OUR AVERAGE</b>				
0.05 ± 0.03 ± 0.02		ANJOS	92C E691	$\gamma$ Be 90–260 GeV
0.084±0.022±0.04		COFFMAN	92B MRK3	$e^+e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.77 ± 0.06 ± 0.06	54	ALVAREZ	91B NA14	Photoproduction
0.85 <sup>+0.11</sup> <sub>-0.22</sub>	180	PICCOLO	77 MRK1	$e^+e^-$ 4.03, 4.41 GeV

<sup>54</sup> This value is for  $\rho^0$  ( $K^-\pi^+$ )-nonresonant. ALVAREZ 91B cannot determine what fraction of this is  $K^-\alpha_1(1260)^+$ .

 $\Gamma(\bar{K}^*(892)^0\rho^0)/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{91}/\Gamma_{43}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included. We rely on the MARK III and E691 full amplitude analyses of the  $K^-\pi^+\pi^+\pi^-$  channel for values of the resonant substructure.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.195±0.03±0.03</b>				
		ANJOS	92C E691	$\gamma$ Be 90–260 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.34 ± 0.09 ± 0.09		ALVAREZ	91B NA14	Photoproduction
0.75 ± 0.3	5	BAILEY	83B SPEC	$\pi$ Be → $D^0$
0.15 <sup>+0.16</sup> <sub>-0.15</sub>	20	PICCOLO	77 MRK1	$e^+e^-$ 4.03, 4.41 GeV

 $\Gamma(\bar{K}^*(892)^0\rho^0\text{transverse})/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{92}/\Gamma_{43}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.21 ± 0.07 OUR FIT</b>				
<b>0.213±0.024±0.075</b>		COFFMAN	92B MRK3	$e^+e^-$ 3.77 GeV

 $\Gamma(\bar{K}^*(892)^0\rho^0S\text{-wave})/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{93}/\Gamma_{43}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.375±0.045±0.06</b>				
		ANJOS	92C E691	$\gamma$ Be 90–260 GeV

 $\Gamma(\bar{K}^*(892)^0\rho^0S\text{-wave long.})/\Gamma_{\text{total}}$   $\Gamma_{94}/\Gamma$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<0.003	90	COFFMAN	92B MRK3	$e^+e^-$ 3.77 GeV

 $\Gamma(\bar{K}^*(892)^0\rho^0P\text{-wave})/\Gamma_{\text{total}}$   $\Gamma_{95}/\Gamma$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<0.003				
	90	COFFMAN	92B MRK3	$e^+e^-$ 3.77 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.009	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
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$\Gamma(\bar{K}^*(892)^0 \rho^0 D\text{-wave})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{96}/\Gamma_{43}$ Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.255 \pm 0.045 \pm 0.06</math></b>	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

 $\Gamma(K^- \pi^+ f_0(980))/\Gamma_{\text{total}}$   $\Gamma_{101}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.011	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

 $\Gamma(\bar{K}^*(892)^0 f_0(980))/\Gamma_{\text{total}}$   $\Gamma_{102}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.007	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

 $\Gamma(K^- a_1(1260)^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{81}/\Gamma_{43}$ Unseen decay modes of the  $a_1(1260)^+$  are included, assuming that the  $a_1(1260)^+$  decays entirely to  $\rho\pi$  [or at least to  $(\pi\pi)_{I=1}\pi$ ].

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.97 \pm 0.14</math> OUR AVERAGE</b>			
0.94 $\pm 0.13 \pm 0.20$	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
$0.984 \pm 0.048 \pm 0.16$	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(K^- a_2(1320)^+)/\Gamma_{\text{total}}$   $\Gamma_{84}/\Gamma$ Unseen decay modes of the  $a_2(1320)^+$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.002</b>	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.006	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(K_1(1270)^- \pi^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{103}/\Gamma_{43}$ Unseen decay modes of the  $K_1(1270)^-$  are included. The MARK3 and E691 experiments disagree considerably here.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>0.15 \pm 0.04</math> OUR FIT</b>				
<b><math>0.194 \pm 0.056 \pm 0.088</math></b>		COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.013	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

 $\Gamma(K_1(1400)^- \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{104}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.012</b>	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(K^*(1410)^- \pi^+)/\Gamma_{\text{total}}$   $\Gamma_{106}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.012	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \text{total})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{87}/\Gamma_{43}$ 

This includes  $\bar{K}^*(892)^0 \rho^0$ , etc. The next entry gives the specifically 3-body fraction.  
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.30±0.06±0.03</b>	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

 $\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \text{3-body})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{88}/\Gamma_{43}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.19 ±0.04 OUR FIT</b>			
<b>0.18 ±0.04 OUR AVERAGE</b>			

0.165±0.03 ±0.045 ANJOS 92C E691  $\gamma$ Be 90–260 GeV  
0.210±0.027±0.06 COFFMAN 92B MRK3  $e^+ e^-$  3.77 GeV

 $\Gamma(K^- \pi^+ \pi^+ \pi^- \text{nonresonant})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{51}/\Gamma_{43}$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.233±0.032 OUR AVERAGE</b>			

0.23 ±0.02 ±0.03 ANJOS 92C E691  $\gamma$ Be 90–260 GeV  
0.242±0.025±0.06 COFFMAN 92B MRK3  $e^+ e^-$  3.77 GeV

 $\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{52}/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.109±0.013 OUR NEW UNCHECKED FIT</b>		[0.108 ± 0.013 OUR 2002 FIT]		
<b>0.103±0.022±0.025</b>	140	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.134^{+0.032}_{-0.033}$  55 BARLAG 92C ACCM  $\pi^-$  Cu 230 GeV

55 BARLAG 92C computes the branching fraction using topological normalization.

 $\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{52}/\Gamma_{22}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.82±0.20 OUR FIT</b>				
<b>1.86±0.23 OUR AVERAGE</b>				

1.80±0.20±0.21 190 56 ALBRECHT 92P ARG  $e^+ e^- \approx 10$  GeV  
2.8 ±0.8 ±0.8 46 ANJOS 92C E691  $\gamma$ Be 90–260 GeV  
1.85±0.26±0.30 158 KINOSHITA 91 CLEO  $e^+ e^- \sim 10.7$  GeV

56 This value is calculated from numbers in Table 1 of ALBRECHT 92P.

 $\Gamma(\bar{K}^0 \eta)/\Gamma(K^- \pi^+)$   $\Gamma_{75}/\Gamma_{20}$ 

Unseen decay modes of the  $\eta$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •				

<0.64 90 ALBRECHT 89D ARG  $e^+ e^-$  10 GeV

 $\Gamma(\bar{K}^0 \eta)/\Gamma(\bar{K}^0 \pi^0)$   $\Gamma_{75}/\Gamma_{21}$ 

Unseen decay modes of the  $\eta$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.33±0.04 OUR FIT</b>				
<b>0.32±0.04±0.03</b>	225	PROCARIO	93B CLE2	$\eta \rightarrow \gamma\gamma$

$\Gamma(\bar{K}^0\eta)/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{75}/\Gamma_{22}$ 
Unseen decay modes of the  $\eta$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.128±0.017 OUR NEW UNCHECKED FIT</b>		[0.129 ± 0.017 OUR 2002 FIT]		
<b>0.14 ±0.02 ±0.02</b>	80	PROCARIO	93B CLE2	$\eta \rightarrow \pi^+\pi^-\pi^0$

 $\Gamma(\bar{K}^0\omega)/\Gamma(\bar{K}^-\pi^+)$   $\Gamma_{78}/\Gamma_{20}$ 
Unseen decay modes of the  $\omega$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.61±0.09 OUR NEW UNCHECKED FIT</b>		[0.59 ± 0.10 OUR 2002 FIT]		
<b>1.00±0.36±0.20</b>		ALBRECHT	89D ARG	$e^+e^- 10 \text{ GeV}$

 $\Gamma(\bar{K}^0\omega)/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{78}/\Gamma_{22}$ 
Unseen decay modes of the  $\omega$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.39±0.06 OUR NEW UNCHECKED FIT</b>		[0.38 ± 0.06 OUR 2002 FIT]		
<b>0.36±0.07 OUR NEW AVERAGE</b>		[0.33 ± 0.09 OUR 2002 AVERAGE Scale factor = 1.1]		
0.42±0.11 <sup>+0.06</sup> <sub>-0.05</sub>		MURAMATSU 02	CLE2	$e^+e^- \approx 10 \text{ GeV}$
0.29±0.08±0.05	16	57 ALBRECHT	92P ARG	$e^+e^- \approx 10 \text{ GeV}$
0.54±0.14±0.16	40	KINOSHITA 91	CLEO	$e^+e^- \sim 10.7 \text{ GeV}$

<sup>57</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.
 $\Gamma(\bar{K}^0\omega)/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$   $\Gamma_{78}/\Gamma_{52}$ 
Unseen decay modes of the  $\omega$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.212±0.034 OUR NEW UNCHECKED FIT</b>		[0.21 ± 0.04 OUR 2002 FIT]		
<b>0.220±0.048±0.0116</b>		COFFMAN 92B	MRK3	$e^+e^- 3.77 \text{ GeV}$

 $\Gamma(\bar{K}^0\eta'(958))/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{79}/\Gamma_{22}$ 
Unseen decay modes of the  $\eta'(958)$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.32±0.04 OUR AVERAGE</b>				
0.31±0.02±0.04	594	PROCARIO 93B	CLE2	$\eta' \rightarrow \eta\pi^+\pi^-, \rho^0\gamma$
0.37±0.13±0.06	18	58 ALBRECHT 92P	ARG	$e^+e^- \approx 10 \text{ GeV}$

<sup>58</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.
 $\Gamma(K^*(892)^-\rho^+)/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$   $\Gamma_{97}/\Gamma_{52}$ 
Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.606±0.188±0.126</b>		COFFMAN 92B	MRK3	$e^+e^- 3.77 \text{ GeV}$

 $\Gamma(K^*(892)^-\rho^+\text{longitudinal})/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$   $\Gamma_{98}/\Gamma_{52}$ 
Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.290±0.111</b>		COFFMAN 92B	MRK3	$e^+e^- 3.77 \text{ GeV}$

 $\Gamma(K^*(892)^-\rho^+\text{transverse})/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$   $\Gamma_{99}/\Gamma_{52}$ 
Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.317±0.180</b>		COFFMAN 92B	MRK3	$e^+e^- 3.77 \text{ GeV}$

$\Gamma(K^*(892)^-\rho^+ P\text{-wave})/\Gamma_{\text{total}}$  $\Gamma_{100}/\Gamma$ Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.015	90	59 COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

59 Obtained using other  $\bar{K}^*(892)\rho$   $P$ -wave limits and isospin relations. $\Gamma(\bar{K}^*(892)^0\rho^0 \text{transverse})/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$  $\Gamma_{92}/\Gamma_{52}$ Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.14 ± 0.05 OUR FIT</b>			
<b>0.126 ± 0.111</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(\bar{K}^0 a_1(1260)^0)/\Gamma_{\text{total}}$  $\Gamma_{82}/\Gamma$ Unseen decay modes of the  $a_1(1260)^+$  are included, assuming that the  $a_1(1260)^+$  decays entirely to  $\rho\pi$  [or at least to  $(\pi\pi)_{I=1}\pi$ ].

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.019	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(K_1(1270)^-\pi^+)/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$  $\Gamma_{103}/\Gamma_{52}$ Unseen decay modes of the  $K_1(1270)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.105 ± 0.028 OUR FIT</b>			
<b>0.10 ± 0.03</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(\bar{K}_1(1400)^0\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{105}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.037	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(\bar{K}^*(892)^0\pi^+\pi^- 3\text{-body})/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$  $\Gamma_{88}/\Gamma_{52}$ Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.130 ± 0.034 OUR NEW UNCHECKED FIT</b>	Error includes scale factor of 1.1. [0.131 ± 0.035 OUR 2002 FIT Scale factor = 1.1]		
<b>0.191 ± 0.105</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(\bar{K}^0\pi^+\pi^-\pi^0 \text{nonresonant})/\Gamma(\bar{K}^0\pi^+\pi^-\pi^0)$  $\Gamma_{59}/\Gamma_{52}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.210 ± 0.147 ± 0.150</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

 $\Gamma(K^-\pi^+\pi^0\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{60}/\Gamma$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.177 ± 0.029	60 BARLAG	92C ACCM	$\pi^-$ Cu	230 GeV
0.149 ± 0.037 ± 0.030	24 ADLER	88C MRK3	$e^+ e^-$	3.77 GeV
0.209 ± 0.074 ± 0.012	9 AGUILAR-...	87F HYBR	$\pi p, pp$	360, 400 GeV

60 AGUILAR-BENITEZ 87F and BARLAG 92C compute the branching fraction using topological normalization. They do not distinguish the presence of a third  $\pi^0$ , and thus are not included in the average.

61 ADLER 88C uses an absolute normalization method finding this decay channel opposite a detected  $\bar{D}^0 \rightarrow K^+\pi^-$  in pure  $D\bar{D}$  events.

$\Gamma(K^-\pi^+\pi^+\pi^-\pi^0)/\Gamma(K^-\pi^+)$   $\Gamma_{61}/\Gamma_{20}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.06±0.10 OUR NEW UNCHECKED FIT</b>	[1.05 ± 0.10 OUR 2002 FIT]			

**0.98±0.11±0.11** 225 62 ALBRECHT 92P ARG  $e^+e^- \approx 10$  GeV

62 This value is calculated from numbers in Table 1 of ALBRECHT 92P.

 $\Gamma(K^-\pi^+\pi^+\pi^-\pi^0)/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{61}/\Gamma_{43}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.54±0.05 OUR FIT</b>				

**0.56±0.07 OUR AVERAGE**

$0.55 \pm 0.07^{+0.12}_{-0.09}$	167	KINOSHITA	91 CLEO	$e^+e^- \sim 10.7$ GeV
$0.57 \pm 0.06 \pm 0.05$	180	ANJOS	90D E691	Photoproduction

 $\Gamma(\bar{K}^*(892)^0\pi^+\pi^-\pi^0)/\Gamma(K^-\pi^+\pi^+\pi^-\pi^0)$   $\Gamma_{112}/\Gamma_{61}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.45±0.15±0.15</b>				

 $\Gamma(\bar{K}^*(892)^0\eta)/\Gamma(K^-\pi^+)$   $\Gamma_{113}/\Gamma_{20}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\eta$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.46±0.12 OUR FIT</b>				
<b>0.58±0.19^{+0.24}_{-0.28}</b>	46	KINOSHITA	91 CLEO	$e^+e^- \sim 10.7$ GeV

 $\Gamma(\bar{K}^*(892)^0\eta)/\Gamma(K^-\pi^+\pi^0)$   $\Gamma_{113}/\Gamma_{31}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\eta$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.135±0.034 OUR FIT</b>				
<b>0.13 ±0.02 ±0.03</b>	214	PROCARIO	93B CLE2	$\bar{K}^*{}^0\eta \rightarrow K^-\pi^+/\gamma\gamma$

 $\Gamma(K^-\pi^+\omega)/\Gamma(K^-\pi^+)$   $\Gamma_{114}/\Gamma_{20}$ 

Unseen decay modes of the  $\omega$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.78±0.12±0.10</b>	99	63 ALBRECHT	92P ARG	$e^+e^- \approx 10$ GeV

63 This value is calculated from numbers in Table 1 of ALBRECHT 92P.

 $\Gamma(\bar{K}^*(892)^0\omega)/\Gamma(K^-\pi^+)$   $\Gamma_{115}/\Gamma_{20}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\omega$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.28±0.11±0.04</b>	17	64 ALBRECHT	92P ARG	$e^+e^- \approx 10$ GeV

64 This value is calculated from numbers in Table 1 of ALBRECHT 92P.

 $\Gamma(\bar{K}^*(892)^0\omega)/\Gamma(K^-\pi^+\pi^+\pi^-\pi^0)$   $\Gamma_{115}/\Gamma_{61}$ 

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\omega$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

<0.44 90 65 ANJOS 90D E691 Photoproduction

65 Recovered from the published limit,  $\Gamma(\bar{K}^*(892)^0\omega)/\Gamma_{\text{total}}$ , in order to make our normalization consistent.

### $\Gamma(K^-\pi^+\eta'(958))/\Gamma(K^-\pi^+\pi^+\pi^-)$

$\Gamma_{116}/\Gamma_{43}$

Unseen decay modes of the  $\eta'(958)$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.093 \pm 0.014 \pm 0.019</math></b>	286	PROCARIO	93B CLE2	$\eta' \rightarrow \eta\pi^+\pi^-, \rho^0\gamma$

### $\Gamma(\bar{K}^*(892)^0\eta'(958))/\Gamma(K^-\pi^+\eta'(958))$

$\Gamma_{117}/\Gamma_{116}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

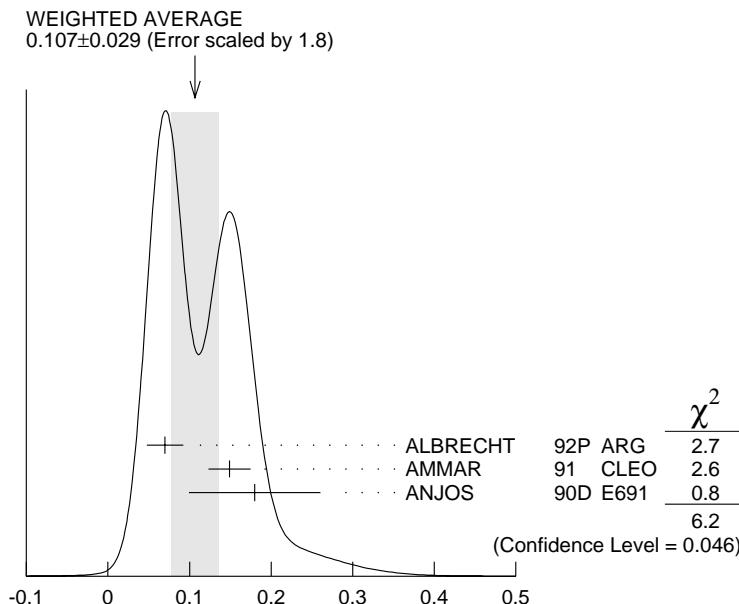
VALUE	CL%	DOCUMENT ID	TECN
<b>&lt;0.15</b>	90	PROCARIO	93B CLE2

### $\Gamma(\bar{K}^0\pi^+\pi^+\pi^-\pi^-)/\Gamma(\bar{K}^0\pi^+\pi^-)$

$\Gamma_{66}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.107 \pm 0.029</math> OUR AVERAGE</b>				Error includes scale factor of 1.8. See the ideogram below.
0.07 $\pm 0.02$ $\pm 0.01$	11	66 ALBRECHT	92P ARG	$e^+e^- \approx 10$ GeV
$0.149 \pm 0.026$	56	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
$0.18 \pm 0.07 \pm 0.04$	6	ANJOS	90D E691	Photoproduction

<sup>66</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.



### $\Gamma(\bar{K}^0\pi^+\pi^+\pi^-\pi^-)/\Gamma(\bar{K}^0\pi^+\pi^-)$

### $\Gamma(\bar{K}^0\pi^+\pi^-\pi^0\pi^0(\pi^0))/\Gamma_{\text{total}}$

$\Gamma_{67}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$0.106^{+0.073}_{-0.029} \pm 0.006$  4 <sup>67</sup> AGUILAR-BENITEZ 87F HYBR  $\pi p, pp$  360, 400 GeV

<sup>67</sup> AGUILAR-BENITEZ 87F computes the branching fraction using topological normalization, and does not distinguish the presence of a third  $\pi^0$ .

$\Gamma(\bar{K}^0 K^+ K^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.172±0.014 OUR FIT</b>				
<b>0.178±0.019 OUR AVERAGE</b>				
0.20 ± 0.05 ± 0.04	47	FRABETTI	92B E687	$\gamma$ Be $\bar{E}_\gamma = 221$ GeV
0.170±0.022	136	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
0.24 ± 0.08		BEBEK	86 CLEO	$e^+ e^-$ near $\Upsilon(4S)$
0.185±0.055	52	ALBRECHT	85B ARG	$e^+ e^-$ 10 GeV

$\Gamma_{68}/\Gamma_{22} = (\Gamma_{70} + \frac{1}{2}\Gamma_{80})/\Gamma_{22}$

$\Gamma(\bar{K}^0 \phi)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

Unseen decay modes of the  $\phi$  are included.

$\Gamma_{80}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.158±0.016 OUR FIT</b>				
<b>0.156±0.017 OUR AVERAGE</b>				
0.13 ± 0.06 ± 0.02	13	FRABETTI	92B E687	$\gamma$ Be $\bar{E}_\gamma = 221$ GeV
0.163±0.023	63	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
0.155±0.033	56	ALBRECHT	87E ARG	$e^+ e^-$ 10 GeV
0.14 ± 0.05	29	BEBEK	86 CLEO	$e^+ e^-$ near $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.186±0.052	26	ALBRECHT	85B ARG	See ALBRECHT 87E

$\Gamma(\bar{K}^0 K^+ K^- \text{non-}\phi)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

$\Gamma_{70}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.093±0.014 OUR FIT</b>				
<b>0.088±0.019 OUR AVERAGE</b>				
0.11 ± 0.04 ± 0.03	20	FRABETTI	92B E687	$\gamma$ Be $\bar{E}_\gamma = 221$ GeV
0.084±0.020		ALBRECHT	87E ARG	$e^+ e^-$ 10 GeV

$\Gamma(K_S^0 K_S^0 K_S^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

$\Gamma_{71}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0154±0.0025 OUR AVERAGE</b>				
0.0139±0.0019±0.0024	61	ASNER	96B CLE2	$e^+ e^-$ ≈ $\Upsilon(4S)$
0.035 ± 0.012 ± 0.006	10	FRABETTI	94J E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
0.016 ± 0.005	22	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
0.017 ± 0.007 ± 0.005	5	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(K^+ K^- K^- \pi^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$

$\Gamma_{73}/\Gamma_{43}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0032±0.0009 OUR AVERAGE</b>				Error includes scale factor of 1.4.
0.0054±0.0016±0.0008	18	AITALA	01D E791	$\pi^-$ nucleus, 500 GeV
0.0028±0.0007±0.0001	20	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$\Gamma(K^- \pi^+ \phi)/\Gamma(K^+ K^- K^- \pi^+)$

$\Gamma_{118}/\Gamma_{73}$

Unseen decay modes of the  $\phi$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.4±0.6</b>	13	68 AITALA	01D E791	$\pi^-$ nucleus, 500 GeV

<sup>68</sup> This AITALA 01D result is from a projection fit, not a full amplitude analysis.

$\Gamma(K^+ K^- \bar{K}^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{74}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
$0.0072^{+0.0048}_{-0.0035}$	69 BARLAG	92C ACCM $\pi^-$ Cu 230 GeV	
69 BARLAG 92C computes the branching fraction using topological normalization.			

**Pionic modes**

$\Gamma(\pi^+ \pi^-)/\Gamma(K^- \pi^+)$   $\Gamma_{119}/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0376 ± 0.0017 OUR AVERAGE</b>				Error includes scale factor of 1.1.
$0.0351 \pm 0.0016 \pm 0.0017$	710	CSORNA	02 CLE2	$e^+ e^- \approx \gamma(4S)$
$0.040 \pm 0.002 \pm 0.003$	2043	AITALA	98C E791	$\pi^-$ nucleus, 500 GeV
$0.043 \pm 0.007 \pm 0.003$	177	FRABETTI	94C E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
$0.0348 \pm 0.0030 \pm 0.0023$	227	SELEN	93 CLE2	$e^+ e^- \approx \gamma(4S)$
$0.048 \pm 0.013 \pm 0.008$	51	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
$0.055 \pm 0.008 \pm 0.005$	120	ANJOS	91D E691	Photoproduction
$0.040 \pm 0.007 \pm 0.006$	57	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV
$0.050 \pm 0.007 \pm 0.005$	110	ALEXANDER	90 CLEO	$e^+ e^-$ 10.5–11 GeV
$0.033 \pm 0.010 \pm 0.006$	39	BALTRUSAIT..85E	MRK3	$e^+ e^-$ 3.77 GeV
$0.033 \pm 0.015$		ABRAMS	79D MRK2	$e^+ e^-$ 3.77 GeV

$\Gamma(\pi^0 \pi^0)/\Gamma(K^- \pi^+)$   $\Gamma_{120}/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.022 \pm 0.004 \pm 0.004$	40	SELEN	93 CLE2	$e^+ e^- \approx \gamma(4S)$

$\Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{121}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$0.011 \pm 0.004 \pm 0.002$	10	70 BALTRUSAIT..85E	MRK3	$e^+ e^-$ 3.77 GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$0.0390^{+0.0100}_{-0.0095}$	71 BARLAG	92C ACCM $\pi^-$ Cu 230 GeV		

70 All the BALTRUSAITIS 85E events are consistent with  $\rho^0 \pi^0$ .

71 BARLAG 92C computes the branching fraction using topological normalization. Possible contamination by extra  $\pi^0$ 's may partly explain the unexpectedly large value.

$\Gamma(\pi^+ \pi^+ \pi^- \pi^-)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{122}/\Gamma_{43}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.098 ± 0.006 OUR AVERAGE</b>				
$0.095 \pm 0.007 \pm 0.002$	814	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
$0.115 \pm 0.023 \pm 0.016$	64	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
$0.108 \pm 0.024 \pm 0.008$	79	FRABETTI	92 E687	$\gamma$ Be
$0.102 \pm 0.013$	345	72 AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
$0.096 \pm 0.018 \pm 0.007$	66	ANJOS	91 E691	$\gamma$ Be 80–240 GeV

72 AMMAR 91 finds  $1.25 \pm 0.25 \pm 0.25$   $\rho^0$ 's per  $\pi^+ \pi^+ \pi^- \pi^-$  decay, but can't untangle the resonant substructure ( $\rho^0 \rho^0$ ,  $a_1^\pm \pi^\mp$ ,  $\rho^0 \pi^+ \pi^-$ ).

$\Gamma(\pi^+\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{123}/\Gamma$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
$0.0192^{+0.0041}_{-0.0038}$	73 BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

73 BARLAG 92C computes the branching fraction using topological normalization.

 $\Gamma(\pi^+\pi^+\pi^+\pi^-\pi^-\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{124}/\Gamma$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
$0.0004 \pm 0.0003$	74 BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

74 BARLAG 92C computes the branching fraction using topological normalization.

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**Hadronic modes with a  $K\bar{K}$  pair**


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 $\Gamma(K^+K^-)/\Gamma(K^-\pi^+)$   $\Gamma_{125}/\Gamma_{20}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.1083 <math>\pm</math> 0.0027 OUR FIT</b>				
<b>0.1084 <math>\pm</math> 0.0026 OUR AVERAGE</b>				
$0.1040 \pm 0.0033 \pm 0.0027$	1900	CSORNA	02 CLE2	$e^+e^- \approx \gamma(4S)$
$0.109 \pm 0.003 \pm 0.003$	3317	AITALA	98C E791	$\pi^-$ nucleus, 500 GeV
$0.116 \pm 0.007 \pm 0.007$	1102	ASNER	96B CLE2	$e^+e^- \approx \gamma(4S)$
$0.109 \pm 0.007 \pm 0.009$	581	FRABETTI	94C E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
$0.107 \pm 0.029 \pm 0.015$	103	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
$0.138 \pm 0.027 \pm 0.010$	155	FRABETTI	92 E687	$\gamma$ Be
$0.16 \pm 0.05$	34	ALVAREZ	91B NA14	Photoproduction
$0.107 \pm 0.010 \pm 0.009$	193	ANJOS	91D E691	Photoproduction
$0.10 \pm 0.02 \pm 0.01$	131	ALBRECHT	90C ARG	$e^+e^- \approx 10$ GeV
$0.117 \pm 0.010 \pm 0.007$	249	ALEXANDER	90 CLEO	$e^+e^-$ 10.5–11 GeV
$0.122 \pm 0.018 \pm 0.012$	118	BALTRUSAIT..85E	MRK3	$e^+e^-$ 3.77 GeV
$0.113 \pm 0.030$		ABRAMS	79D MRK2	$e^+e^-$ 3.77 GeV

 $\Gamma(K^+K^-)/\Gamma(\pi^+\pi^-)$   $\Gamma_{125}/\Gamma_{119}$ 

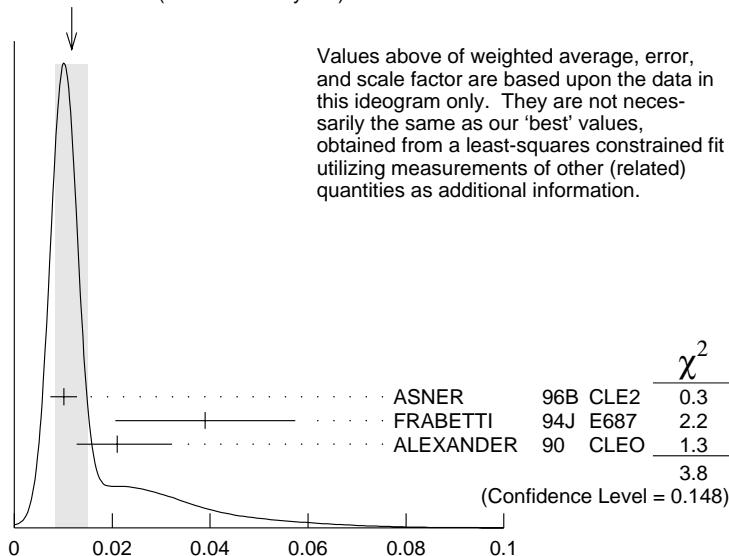
The unused results here are redundant with  $\Gamma(K^+K^-)/\Gamma(K^-\pi^+)$  and  $\Gamma(\pi^+\pi^-)/\Gamma(K^-\pi^+)$  measurements by the same experiments.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$2.96 \pm 0.16 \pm 0.15$	710	CSORNA	02 CLE2	$e^+e^- \approx \gamma(4S)$
$2.75 \pm 0.15 \pm 0.16$		AITALA	98C E791	$\pi^-$ nucleus, 500 GeV
$2.53 \pm 0.46 \pm 0.19$		FRABETTI	94C E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
$2.23 \pm 0.81 \pm 0.46$		ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
$1.95 \pm 0.34 \pm 0.22$		ANJOS	91D E691	Photoproduction
$2.5 \pm 0.7$		ALBRECHT	90C ARG	$e^+e^- \approx 10$ GeV
$2.35 \pm 0.37 \pm 0.28$		ALEXANDER	90 CLEO	$e^+e^-$ 10.5–11 GeV

### $\Gamma(K^0\bar{K}^0)/\Gamma(\bar{K}^0\pi^+\pi^-)$ $\Gamma_{126}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0120±0.0033 OUR FIT</b>	Error includes scale factor of 1.3.			
<b>0.0117±0.0033 OUR AVERAGE</b>	Error includes scale factor of 1.3. See the ideogram below.			
0.0101±0.0022±0.0016	26	ASNER	96B CLE2	$e^+e^- \approx \gamma(4S)$
0.039 ± 0.013 ± 0.013	20	FRABETTI	94J E687	$\gamma Be \bar{E}_\gamma = 220$ GeV
0.021 +0.011 -0.008	5	ALEXANDER	90 CLEO	$e^+e^-$ 10.5–11 GeV

WEIGHTED AVERAGE  
0.0117±0.0033 (Error scaled by 1.3)



### $\Gamma(K^0\bar{K}^0)/\Gamma(\bar{K}^0\pi^+\pi^-)$

### $\Gamma(K^0\bar{K}^0)/\Gamma(K^+\bar{K}^-)$ $\Gamma_{126}/\Gamma_{125}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.17±0.05 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.24±0.16</b>	4	75 CUMALAT	88 SPEC	$nN$ 0–800 GeV

<sup>75</sup> Includes a correction communicated to us by the authors of CUMALAT 88.

### $\Gamma(K^0\bar{K}^0)/\Gamma(K^+\bar{K}^-)$ $\Gamma_{127}/\Gamma_{20}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.183±0.027 OUR NEW UNCHECKED FIT</b>	[0.182 ± 0.027 OUR 2002 FIT Scale factor = 1.1]		
<b>0.16 ± 0.06</b>	76 ANJOS	91 E691	$\gamma Be$ 80–240 GeV

<sup>76</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K^0 K^- \pi^+)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{127}/\Gamma_{22}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.116±0.017 OUR NEW UNCHECKED FIT</b>	Error includes scale factor of 1.1. [0.117 ± 0.017 OUR 2002 FIT Scale factor = 1.1]			
<b>0.119±0.021 OUR AVERAGE</b>				Error includes scale factor of 1.3.
0.108±0.019	61	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
0.16 ± 0.03 ± 0.02	39	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV

 $\Gamma(\bar{K}^*(892)^0 K^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{147}/\Gamma_{22}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.029	90	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.03	90	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV

 $\Gamma(K^*(892)^+ K^-)/\Gamma(K^- \pi^+)$   $\Gamma_{148}/\Gamma_{20}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.100±0.021 OUR NEW UNCHECKED FIT</b> [0.099 ± 0.021 OUR 2002 FIT]			
<b>0.16 +0.08 -0.06</b>	77 ANJOS	91 E691	$\gamma$ Be 80–240 GeV

77 The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

 $\Gamma(K^*(892)^+ K^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{148}/\Gamma_{22}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.064±0.013 OUR FIT</b>				
<b>0.058±0.014 OUR AVERAGE</b>				
0.064±0.018	23	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
0.05 ± 0.02 ± 0.01	15	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV

 $\Gamma(K^0 K^- \pi^+ \text{nonresonant})/\Gamma(K^- \pi^+)$   $\Gamma_{130}/\Gamma_{20}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.06±0.06</b>	78 ANJOS	91 E691	$\gamma$ Be 80–240 GeV

78 The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

 $\Gamma(\bar{K}^0 K^+ \pi^-)/\Gamma(K^- \pi^+)$   $\Gamma_{131}/\Gamma_{20}$ 

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.139±0.027 OUR NEW UNCHECKED FIT</b> [0.138 ± 0.026 OUR 2002 FIT]			
<b>0.10 ± 0.05</b>			79 ANJOS 91 E691 $\gamma$ Be 80–240 GeV

79 The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

 $\Gamma(\bar{K}^0 K^+ \pi^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{131}/\Gamma_{22}$ 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.088±0.017 OUR NEW UNCHECKED FIT</b> [0.089 ± 0.017 OUR 2002 FIT]				
<b>0.098±0.020</b>	55	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV

$\Gamma(K^*(892)^0 \bar{K}^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{149}/\Gamma_{22}$ Unseen decay modes of the  $K^*(892)^0$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.015</b>	90	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV

 $\Gamma(K^*(892)^- K^+)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$  $\Gamma_{150}/\Gamma_{22}$ Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.034 ± 0.019</b>	12	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV

 $\Gamma(\bar{K}^0 K^+ \pi^- \text{nonresonant})/\Gamma(K^- \pi^+)$  $\Gamma_{134}/\Gamma_{20}$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.10<sup>+0.06</sup><sub>-0.05</sub></b>	80	ANJOS	91 $\gamma$ Be 80–240 GeV

80 The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

 $\Gamma(K^+ K^- \pi^0)/\Gamma(K^- \pi^+ \pi^0)$  $\Gamma_{135}/\Gamma_{31}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0095 ± 0.0026</b>	151	ASNER	96B	CLE2 $e^+ e^- \approx \gamma(4S)$

 $\Gamma(K_S^0 K_S^0 \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{136}/\Gamma$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.00059</b>	ASNER	96B	CLE2 $e^+ e^- \approx \gamma(4S)$

 $\Gamma(\phi \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{151}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.0014</b>	90	ALBRECHT	94I	ARG $e^+ e^- \approx 10$ GeV

 $\Gamma(\phi \eta)/\Gamma_{\text{total}}$  $\Gamma_{152}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.0028</b>	90	ALBRECHT	94I	ARG $e^+ e^- \approx 10$ GeV

 $\Gamma(\phi \omega)/\Gamma_{\text{total}}$  $\Gamma_{153}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.0021</b>	90	ALBRECHT	94I	ARG $e^+ e^- \approx 10$ GeV

 $\Gamma(K^+ K^- \pi^+ \pi^-)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$  $\Gamma_{137}/\Gamma_{43}$ 

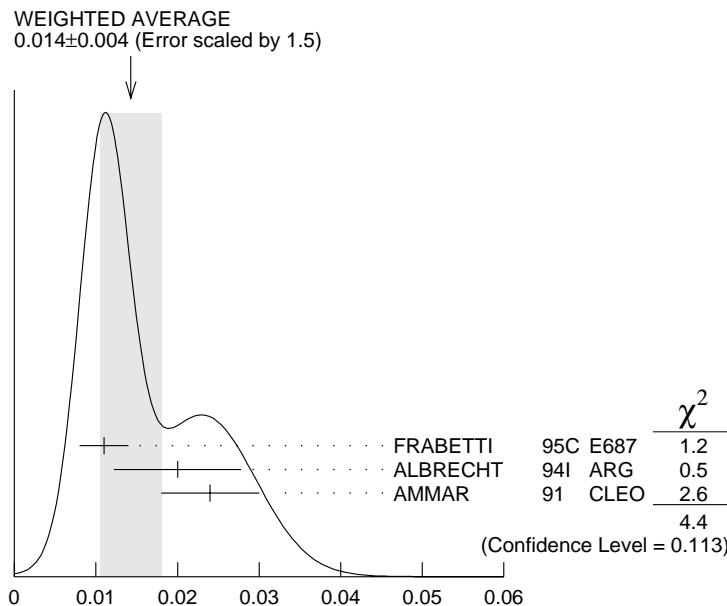
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0334 ± 0.0028 OUR AVERAGE</b>				
0.0313 ± 0.0037 ± 0.0036	136	AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
0.035 ± 0.004 ± 0.002	244	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
0.041 ± 0.007 ± 0.005	114	ALBRECHT	94I ARG	$e^+ e^- \approx 10$ GeV
0.0314 ± 0.010	89	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
0.028 <sup>+0.008</sup> <sub>-0.007</sub>		ANJOS	91 E691	$\gamma$ Be 80–240 GeV

## $\Gamma(\phi\pi^+\pi^-)/\Gamma(K^-\pi^+\pi^+\pi^-)$

Unseen decay modes of the  $\phi$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.014 ±0.004 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
0.011 ±0.003		FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
0.020 ±0.006 ±0.005	28	ALBRECHT	94I ARG	$e^+e^- \approx 10$ GeV
0.024 ±0.006	34	<sup>81</sup> AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
0.0076 <sup>+0.0066</sup> <sub>-0.0049</sub>	3	ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>81</sup>AMMAR 91 measures  $\phi\rho^0$ , but notes that  $\phi\rho^0$  dominates  $\phi\pi^+\pi^-$ . We put the measurement here to keep from having more  $\phi\rho^0$  than  $\phi\pi^+\pi^-$ .

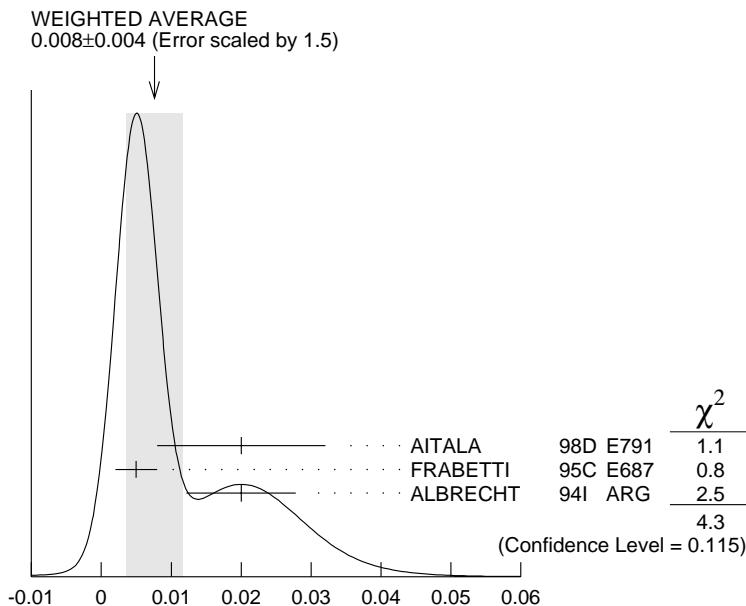


## $\Gamma(\phi\pi^+\pi^-)/\Gamma(K^-\pi^+\pi^+\pi^-)$

## $\Gamma(\phi\rho^0)/\Gamma(K^-\pi^+\pi^+\pi^-)$

Unseen decay modes of the  $\phi$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.008±0.004 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
0.02 ±0.009±0.008		AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
0.005±0.003		FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
0.020±0.006±0.005	28	ALBRECHT	94I ARG	$e^+e^- \approx 10$ GeV



$$\Gamma(\phi\rho^0)/\Gamma(K^-\pi^+\pi^+\pi^-)$$

$$\Gamma(\phi\pi^+\pi^- \text{3-body})/\Gamma(K^-\pi^+\pi^+\pi^-)$$

Unseen decay modes of the  $\phi$  are included.

$$\Gamma_{156}/\Gamma_{43}$$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.009±0.004±0.005</b>		AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
<0.006	90	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$$\Gamma(K^+K^-\rho^0 \text{3-body})/\Gamma(K^-\pi^+\pi^+\pi^-)$$

$$\Gamma_{140}/\Gamma_{43}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.012 ±0.003</b>	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$$\Gamma(K^*(892)^0 K^-\pi^+ + \text{c.c.})/\Gamma(K^-\pi^+\pi^+\pi^-)$$

$$\Gamma_{157}/\Gamma_{43}$$

Unseen decay modes of the  $K^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.01</b>	90	82 AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>				
<0.017	90	82 FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
$0.010^{+0.016}_{-0.010}$		ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>82</sup> These upper limits are in conflict with values in the next two data blocks.

$\Gamma(K^*(892)^0 K^- \pi^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{158}/\Gamma_{43}$ 

The  $K^{*0} K^- \pi^+$  and  $\bar{K}^{*0} K^+ \pi^-$  modes are distinguished by the charge of the pion in  $D^*(2010)^\pm \rightarrow D^0 \pi^\pm$  decays. Unseen decay modes of the  $K^*(892)^0$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.043 \pm 0.014 \pm 0.009$  55 83 ALBRECHT 94I ARG  $e^+ e^- \approx 10$  GeV

83 This ALBRECHT 94I value is in conflict with upper limits given above.

 $\Gamma(\bar{K}^*(892)^0 K^+ \pi^-)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{159}/\Gamma_{43}$ 

The  $K^{*0} K^- \pi^+$  and  $\bar{K}^{*0} K^+ \pi^-$  modes are distinguished by the charge of the pion in  $D^*(2010)^\pm \rightarrow D^0 \pi^\pm$  decays. Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.023 \pm 0.013 \pm 0.009$  30 84 ALBRECHT 94I ARG  $e^+ e^- \approx 10$  GeV

84 This ALBRECHT 94I value is in conflict with upper limits given above.

 $\Gamma(K^*(892)^0 \bar{K}^*(892)^0)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{160}/\Gamma_{43}$ 

Unseen decay modes of the  $K^*(892)^0$  and  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.018 ± 0.007 OUR AVERAGE** Error includes scale factor of 1.2.

$0.016 \pm 0.006$  FRABETTI 95C E687  $\gamma$ Be,  $\bar{E}_\gamma \approx 200$  GeV

$0.036^{+0.020}_{-0.016}$  11 ANJOS 91 E691  $\gamma$ Be 80–240 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<0.02$  90 AITALA 98D E791  $\pi^-$  nucleus, 500 GeV

$<0.033$  90 85 AMMAR 91 CLEO  $e^+ e^- \approx 10.5$  GeV

85 A corrected value (G. Moneti, private communication).

 $\Gamma(K^+ K^- \pi^+ \pi^- \text{non-}\phi)/\Gamma_{\text{total}}$   $\Gamma_{143}/\Gamma$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.0017 \pm 0.0005$  86 BARLAG 92C ACCM  $\pi^-$  Cu 230 GeV

86 BARLAG 92C computes the branching fraction using topological normalization.

 $\Gamma(K^+ K^- \pi^+ \pi^- \text{nonresonant})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{144}/\Gamma_{43}$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**<0.011** 90 FRABETTI 95C E687  $\gamma$ Be,  $\bar{E}_\gamma \approx 200$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.001^{+0.011}_{-0.001}$  ANJOS 91 E691  $\gamma$ Be 80–240 GeV

 $\Gamma(K^0 \bar{K}^0 \pi^+ \pi^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{145}/\Gamma_{22}$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.126 ± 0.038 ± 0.030** 25 ALBRECHT 94I ARG  $e^+ e^- \approx 10$  GeV

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$ VALUE **$0.0031 \pm 0.0020$** DOCUMENT ID

87 BARLAG

TECN92C ACCM  $\pi^-$  Cu 230 GeV $\Gamma_{146}/\Gamma$ 

87 BARLAG 92C computes the branching fraction using topological normalization.

**Radiative modes** $\Gamma(\rho^0 \gamma)/\Gamma_{\text{total}}$ VALUE **$<2.4 \times 10^{-4}$** DOCUMENT ID

ASNER

TECN

CLE2

 $\Gamma_{161}/\Gamma$  $\Gamma(\omega \gamma)/\Gamma_{\text{total}}$ VALUE **$<2.4 \times 10^{-4}$** DOCUMENT ID

ASNER

TECN $\Gamma_{162}/\Gamma$  $\Gamma(\phi \gamma)/\Gamma_{\text{total}}$ VALUE **$<1.9 \times 10^{-4}$** DOCUMENT ID

ASNER

TECN $\Gamma_{163}/\Gamma$  $\Gamma(\bar{K}^*(892)^0 \gamma)/\Gamma_{\text{total}}$ VALUE **$<7.6 \times 10^{-4}$** DOCUMENT ID

ASNER

TECN $\Gamma_{164}/\Gamma$

**Rare or forbidden modes** **$\Gamma(K^+\ell^-\bar{\nu}_\ell(\text{via } \bar{D}^0))/\Gamma(K^-\ell^+\nu_\ell)$**  **$\Gamma_{165}/\Gamma_8$** 

This is a limit on  $R_M$  without the complications of possible doubly-Cabibbo-suppressed decays that occur when using hadronic modes.

For the limits on  $|m_1 - m_2|$  and  $(\Gamma_1 - \Gamma_2)/\Gamma$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.005	90	88 AITALA	96C E791	$\pi^-$ nucleus, 500 GeV

88 AITALA 96C uses  $D^{*+} \rightarrow D^0\pi^+$  (and charge conjugate) decays to identify the charm at production and  $D^0 \rightarrow K^-\ell^+\nu_\ell$  (and charge conjugate) decays to identify the charm at decay.

 **$\Gamma(K^+\pi^-)/\Gamma(K^-\pi^+)$**  **$\Gamma_{166}/\Gamma_{20}$** 

This is  $R_b$  in the note on “ $D^0$ - $\bar{D}^0$  Mixing,” near the start of the  $D^0$  Listings. The experiments here use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0)\pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. The  $D^0 \rightarrow K^+\pi^-$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+\pi^-$  decay. Some of the experiments can use the decay-time information to disentangle the two mechanisms. Here, we list the experimental branching ratio, which if there is no mixing is the DCS ratio; in the next data block we give the limits on the mixing ratio.

Some early limits have been omitted from this Listing; see our 1998 edition (EPJ **C3** 1).

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0039 ± 0.0006 OUR AVERAGE</b>					
0.00404 ± 0.00085 ± 0.00025	149	89 LINK	01 FOCS	$\gamma$ nucleus	
0.00332 $^{+0.00063}_{-0.00065}$ ± 0.00040	45	90 GODANG	00 CLE2	$e^+e^-$	
0.0068 $^{+0.0034}_{-0.0033}$ ± 0.0007	34	91 AITALA	98 E791	$\pi^-$ nucleus, 500 GeV	
0.0184 ± 0.0059 ± 0.0034	19	92 BARATE	98W ALEP	$e^+e^-$ at $Z^0$	
0.0077 ± 0.0025 ± 0.0025	19	93 CINABRO	94 CLE2	$e^+e^- \approx \gamma(4S)$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.011	90	93 AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
<0.015	90	1 ± 6	94 ANJOS	88C E691 Photoproduction
<0.014	90	93 ALBRECHT	87K ARG	$e^+e^-$ 10 GeV

89 This LINK 01 result assumes no  $D^0$ - $\bar{D}^0$  mixing; see Fig. 4 of the paper for the DCS value as a function of the (unknown) mixing parameters  $x'$  and  $y'$ .

90 This GODANG 00 result assumes no  $D^0$ - $\bar{D}^0$  mixing ( $R_M=0$  in the note on “ $D^0$ - $\bar{D}^0$  Mixing” near the start of the  $D^0$  Listings). The DCS ratio becomes  $0.0048 \pm 0.0012 \pm 0.0004$  when mixing is allowed.

91 This AITALA 98 result assumes no  $D^0$ - $\bar{D}^0$  mixing ( $R_M=0$  in the note on “ $D^0$ - $\bar{D}^0$  Mixing” near the start of the  $D^0$  Listings). The DCS ratio becomes  $0.0090  $^{+0.0120}_{-0.0109}$  ± 0.0044$  when mixing is allowed.

- <sup>92</sup> BARATE 98W gets  $0.0177^{+0.0060}_{-0.0056} \pm 0.0031$  for the DCS ratio when mixing is allowed, assuming no interference between the DCS and mixing amplitudes ( $y' = 0$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings).
- <sup>93</sup> CINABRO 94, AMMAR 91, and ALBRECHT 87K cannot distinguish between doubly Cabibbo-suppressed decay and  $D^0$ - $\bar{D}^0$  mixing.
- <sup>94</sup> ANJOS 88C allows mixing but assumes no interference between the DCS and mixing amplitudes ( $y' = 0$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.049.

 $\Gamma(K^+\pi^-(\text{via } \bar{D}^0))/\Gamma(K^-\pi^+)$  $\Gamma_{167}/\Gamma_{20}$ 

This is  $R_M$  in the note on " $D^0$ - $\bar{D}^0$  Mixing" near the start of the  $D^0$  Listings. The experiments here (1) use the charge of the pion in  $D^*(2010)^{\pm} \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^{\pm}$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_1 - m_2|$  and  $(\Gamma_1 - \Gamma_2)/\Gamma$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.00041</b>	95		95 GODANG	00 CLE2	$e^+ e^-$
<b>• • •</b> We do not use the following data for averages, fits, limits, etc. <b>• • •</b>					
<0.0092	95		96 BARATE	98W ALEP	$e^+ e^-$ at $Z^0$
<0.005	90	$1 \pm 4$	97 ANJOS	88C E691	Photoproduction
<sup>95</sup> This GODANG 00 result assumes that the strong phase between $D^0 \rightarrow K^+\pi^-$ and $\bar{D}^0 \rightarrow K^+\pi^-$ is small, and limits only $D^0 \rightarrow \bar{D}^0$ transitions via off-shell intermediate states. The limit on transitions via on-shell intermediate states is 0.0017.					
<sup>96</sup> This BARATE 98W result assumes no interference between the DCS and mixing amplitudes ( $y' = 0$ in the note on " $D^0$ - $\bar{D}^0$ Mixing" near the start of the $D^0$ Listings). When interference is allowed, the limit degrades to 0.036 (95%CL).					
<sup>97</sup> This ANJOS 88C result assumes no interference between the DCS and mixing amplitudes ( $y' = 0$ in the note on " $D^0$ - $\bar{D}^0$ Mixing" near the start of the $D^0$ Listings). When interference is allowed, the limit degrades to 0.019. Combined with results on $K^{\pm}\pi^{\mp}\pi^+\pi^-$ , the limit is, assuming no interference, 0.0037.					

 $\Gamma(K^*(892)^+\pi^-)/\Gamma(\bar{K}^0\pi^+\pi^-)$  $\Gamma_{168}/\Gamma_{22}$ 

Unseen decay modes of the  $K^*(892)^+$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.005 <math>\pm 0.002</math> <math>^{+0.006}_{-0.001}</math></b>	MURAMATSU 02	CLE2	$e^+ e^- \approx 10 \text{ GeV}$

 $\Gamma(K^+\pi^-\pi^0)/\Gamma(K^-\pi^+\pi^0)$  $\Gamma_{169}/\Gamma_{31}$ 

The experiments here use the charge of the pion in  $D^*(2010)^{\pm} \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^{\pm}$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. The  $D^0 \rightarrow K^+\pi^-\pi^0$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+\pi^-\pi^0$  decay.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0043 <math>\pm 0.0011</math> <math>\pm 0.0007</math></b>	38	98 BRANDENB... 01	CLE2	$e^+ e^- \approx \gamma(4S)$

- <sup>98</sup> BRANDENBURG 01 cannot distinguish between doubly Cabibbo-suppressed decay and  $D^0$ - $\bar{D}^0$  mixing.

$\Gamma(K^+\pi^-\pi^+\pi^-)/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{170}/\Gamma_{43}$ 

The experiments here use the charge of the pion in  $D^*(2010)^{\pm} \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^{\pm}$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. The  $D^0 \rightarrow K^+\pi^-\pi^+\pi^-$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+\pi^-\pi^+\pi^-$  decay. Some of the experiments can use the decay-time information to disentangle the two mechanisms. Here, we list the experimental branching ratio, which if there is no mixing is the DCS ratio; in the next data block we give the limits on the mixing ratio.

Some early limits have been omitted from this Listing; see our 1998 edition (EPJ **C3** 1).

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0042±0.0013 OUR AVERAGE</b>					
0.0044 <sup>+0.0013</sup> <sub>-0.0012</sub> ± 0.0006	54	99	DYTMAN	01 CLE2	$e^+e^- \approx \gamma(4S)$
0.0025 <sup>+0.0036</sup> <sub>-0.0034</sub> ± 0.0003		100	AITALA	98 E791	$\pi^-$ nucleus, 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.018	90	99	AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
<0.018	90	5 ± 12	101 ANJOS	88C E691	Photoproduction

<sup>99</sup> AMMAR 91 and DYTMAN 01 cannot distinguish between doubly Cabibbo-suppressed decay and  $D^0$ - $\bar{D}^0$  mixing.

<sup>100</sup> This AITALA 98 result assumes no  $D^0$ - $\bar{D}^0$  mixing ( $R_M$  in the note on “ $D^0$ - $\bar{D}^0$  Mixing”). It becomes  $-0.0020^{+0.0117}_{-0.0106} \pm 0.0035$  when mixing is allowed and decay-time information is used to distinguish doubly Cabibbo-suppressed decays from mixing.

<sup>101</sup> ANJOS 88C uses decay-time information to distinguish doubly Cabibbo-suppressed (DCS) decays from  $D^0$ - $\bar{D}^0$  mixing. However, the result assumes no interference between the DCS and mixing amplitudes ( $y' = 0$  in the note on “ $D^0$ - $\bar{D}^0$  Mixing” near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.033.

 $\Gamma(K^+\pi^-\pi^+\pi^-(\text{via } \bar{D}^0))/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{171}/\Gamma_{43}$ 

This is a  $D^0$ - $\bar{D}^0$  mixing limit. The experiments here (1) use the charge of the pion in  $D^*(2010)^{\pm} \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^{\pm}$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma_{D^0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<0.005	90	0 ± 4	102 ANJOS	88C E691	Photoproduction

<sup>102</sup> ANJOS 88C uses decay-time information to distinguish doubly Cabibbo-suppressed (DCS) decays from  $D^0$ - $\bar{D}^0$  mixing. However, the result assumes no interference between the DCS and mixing amplitudes ( $y' = 0$  in the note on “ $D^0$ - $\bar{D}^0$  Mixing” near the start of the  $D^0$  Listings). When interference is allowed, the limit degrades to 0.007. Combined with results on  $K^{\pm}\pi^{\mp}$ , the limit is, assuming no interference, 0.0037.

$\Gamma(K^+\pi^- \text{ or } K^+\pi^-\pi^+\pi^- \text{ (via } \bar{D}^0\text{)})/\Gamma(K^-\pi^+ \text{ or } K^-\pi^+\pi^+\pi^-)$   $\Gamma_{172}/\Gamma_0$ 

This is a  $D^0$ - $\bar{D}^0$  mixing limit. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma_{D^0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.0085	90	103 AITALA	98 E791	$\pi^-$ nucleus, 500 GeV

<sup>103</sup> AITALA 98 uses decay-time information to distinguish doubly Cabibbo-suppressed decays from  $D^0$ - $\bar{D}^0$  mixing. The fit allows interference between the two amplitudes, and also allows  $CP$  violation in this term. The central value obtained is  $0.0039^{+0.0036}_{-0.0032} \pm 0.0016$ . When interference is disallowed, the result becomes  $0.0021 \pm 0.0009 \pm 0.0002$ .

 $\Gamma(\mu^- \text{ anything (via } \bar{D}^0\text{)})/\Gamma(\mu^+ \text{ anything})$   $\Gamma_{173}/\Gamma_2$ 

This is a  $D^0$ - $\bar{D}^0$  mixing limit. See the somewhat better limits above.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.0056	90	LOUIS	86 SPEC	$\pi^- W$ 225 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<0.012	90	BENVENUTI	85 CNTR	$\mu C$ , 200 GeV
<0.044	90	BODEK	82 SPEC	$\pi^-$ , $pFe \rightarrow D^0$

 $\Gamma(e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{174}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< $6.2 \times 10^{-6}$	90		AITALA	99G E791	$\pi^- N$ 500 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
< $8.19 \times 10^{-6}$	90		PRIPSTEIN	00 E789	$p$ nucleus, 800 GeV
< $1.3 \times 10^{-5}$	90	0	FREYBERGER	96 CLE2	$e^+ e^- \approx \Upsilon(4S)$
< $1.3 \times 10^{-4}$	90		ADLER	88 MRK3	$e^+ e^-$ 3.77 GeV
< $1.7 \times 10^{-4}$	90	7	ALBRECHT	88G ARG	$e^+ e^-$ 10 GeV
< $2.2 \times 10^{-4}$	90	8	HAAS	88 CLEO	$e^+ e^-$ 10 GeV

 $\Gamma(\mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{175}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< $4.1 \times 10^{-6}$	90		ADAMOVICH	97 BEAT	$\pi^- Cu, W$ 350 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
< $1.56 \times 10^{-5}$	90		PRIPSTEIN	00 E789	$p$ nucleus, 800 GeV
< $5.2 \times 10^{-6}$	90		AITALA	99G E791	$\pi^- N$ 500 GeV
< $4.2 \times 10^{-6}$	90		ALEXOPOU...	96 E771	$p Si$ , 800 GeV
< $3.4 \times 10^{-5}$	90	1	FREYBERGER	96 CLE2	$e^+ e^- \approx \Upsilon(4S)$
< $7.6 \times 10^{-6}$	90	0	ADAMOVICH	95 BEAT	See ADAMOVICH 97
< $4.4 \times 10^{-5}$	90	0	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
< $3.1 \times 10^{-5}$	90		MISHRA	94 E789	$-4.1 \pm 4.8$ events
< $7.0 \times 10^{-5}$	90	3		88G ARG	$e^+ e^-$ 10 GeV
< $1.1 \times 10^{-5}$	90		LOUIS	86 SPEC	$\pi^- W$ 225 GeV
< $3.4 \times 10^{-4}$	90		AUBERT	85 EMC	Deep inelast. $\mu^- N$

<sup>104</sup> Here MISHRA 94 uses "the statistical approach advocated by the PDG." For an alternate approach, giving a limit of  $9 \times 10^{-6}$  at 90% confidence level, see the paper.

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{176}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<4.5 \times 10^{-5}$	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(\pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{177}/\Gamma$ 

A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-4}$	90	2	KODAMA	95	E653 $\pi^-$ emulsion 600 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$<5.4 \times 10^{-4}$	90	3	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{178}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-4}$	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(\eta \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{179}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<5.3 \times 10^{-4}$	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(\pi^+ \pi^- e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{180}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.73 \times 10^{-4}$	90	9	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\rho^0 e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{181}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.0 \times 10^{-4}$	90	2	105 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

$\bullet \bullet \bullet$  We do not use the following data for averages, fits, limits, etc.  $\bullet \bullet \bullet$

$<1.24 \times 10^{-4}$  90 1 AITALA 01c E791  $\pi^-$  nucleus, 500 GeV

$<4.5 \times 10^{-4}$  90 2 HAAS 88 CLEO  $e^+ e^-$  10 GeV

105 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 1.8 \times 10^{-4}$  using a photon pole amplitude model.

 $\Gamma(\pi^+ \pi^- \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{182}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.0 \times 10^{-5}$	90	2	AITALA	01c E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\rho^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{183}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.2 \times 10^{-5}$	90	0	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$<4.9 \times 10^{-4}$	90	1	106 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$
$<2.3 \times 10^{-4}$	90	0	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
$<8.1 \times 10^{-4}$	90	5	HAAS	88 CLEO	$e^+ e^-$ 10 GeV

106 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 4.5 \times 10^{-4}$  using a photon pole amplitude model.

 $\Gamma(\omega e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{184}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-4}$	90	1	107 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

107 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 2.7 \times 10^{-4}$  using a photon pole amplitude model.

 $\Gamma(\omega \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{185}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<8.3 \times 10^{-4}$	90	0	108 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

108 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 6.5 \times 10^{-4}$  using a photon pole amplitude model.

 $\Gamma(K^- K^+ e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{186}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.15 \times 10^{-4}$	90	9	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\phi e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{187}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<5.2 \times 10^{-5}$	90	2	109 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

$\bullet \bullet \bullet$  We do not use the following data for averages, fits, limits, etc.  $\bullet \bullet \bullet$

$<5.9 \times 10^{-5}$  90 0 AITALA 01C E791  $\pi^-$  nucleus, 500 GeV

109 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 7.6 \times 10^{-5}$  using a photon pole amplitude model.

 $\Gamma(K^- K^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{188}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.3 \times 10^{-5}$	90	0	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\phi\mu^+\mu^-)/\Gamma_{\text{total}}$  $\Gamma_{189}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-5}$	90	0	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<4.1 \times 10^{-4}$	90	0	110 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$
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110 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 2.4 \times 10^{-4}$  using a photon pole amplitude model.

 $\Gamma(\bar{K}^0 e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{190}/\Gamma$ 

Not a useful test for  $\Delta C = 1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-4}$	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.7 \times 10^{-3}$	90		ADLER		$e^+ e^-$ 3.77 GeV
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 $\Gamma(\bar{K}^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{191}/\Gamma$ 

Not a useful test for  $\Delta C = 1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.6 \times 10^{-4}$	90	2	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.7 \times 10^{-4}$	90	1	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$
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 $\Gamma(K^-\pi^+e^+e^-)/\Gamma_{\text{total}}$  $\Gamma_{192}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.85 \times 10^{-4}$	90	6	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\bar{K}^*(892)^0 e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{193}/\Gamma$ 

Not a useful test for  $\Delta C = 1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<4.7 \times 10^{-5}$	90	2	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.4 \times 10^{-4}$	90	1	111 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$
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111 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 2.0 \times 10^{-4}$  using a photon pole amplitude model.

 $\Gamma(K^-\pi^+\mu^+\mu^-)/\Gamma_{\text{total}}$  $\Gamma_{194}/\Gamma$ 

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.59 \times 10^{-4}$	90	12	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\bar{K}^*(892)^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{195}/\Gamma$ 

Not a useful test for  $\Delta C=1$  weak neutral current because both quarks must change flavor.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.4 \times 10^{-5}$	90	3	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.18 \times 10^{-3}$	90	1	112 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$
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112 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 1.0 \times 10^{-3}$  using a photon pole amplitude model.

 $\Gamma(\pi^+ \pi^- \pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{196}/\Gamma$ 

A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<8.1 \times 10^{-4}$	90	1	KODAMA	95	$\pi^-$ emulsion 600 GeV

 $\Gamma(\mu^\pm e^\mp)/\Gamma_{\text{total}}$   $\Gamma_{197}/\Gamma$ 

A test of lepton family number conservation.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 8.1 \times 10^{-6}$	90		AITALA	99G E791	$\pi^- N$ 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.72 \times 10^{-5}$	90		PRIPSTEIN	00	E789 $p$ nucleus, 800 GeV
$< 1.9 \times 10^{-5}$	90	2	113 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$
$< 1.0 \times 10^{-4}$	90	4	ALBRECHT	88G ARG	$e^+ e^-$ 10 GeV
$< 2.7 \times 10^{-4}$	90	9	HAAS	88	CLEO $e^+ e^-$ 10 GeV
$< 1.2 \times 10^{-4}$	90		BECKER	87C	MRK3 $e^+ e^-$ 3.77 GeV
$< 9 \times 10^{-4}$	90		PALKA	87	SILI 200 GeV $\pi p$
$< 21 \times 10^{-4}$	90	0	114 RILES	87	MRK2 $e^+ e^-$ 29 GeV

113 This is the corrected result given in the erratum to FREYBERGER 96.

114 RILES 87 assumes  $B(D \rightarrow K\pi) = 3.0\%$  and has production model dependency.

 $\Gamma(\pi^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{198}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<8.6 \times 10^{-5}$	90	2	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(\eta e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{199}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.0 \times 10^{-4}$	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(\pi^+ \pi^- e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{200}/\Gamma$ 

A test of lepton family-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.5 \times 10^{-5}$	90	1	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\rho^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{201}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<4.9 \times 10^{-5}$	90	0	115 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.6 \times 10^{-5}$	90	1	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV
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115 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 5.0 \times 10^{-5}$  using a photon pole amplitude model.

 $\Gamma(\omega e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{202}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.2 \times 10^{-4}$	90	0	116 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

116 This FREYBERGER 96 limit is obtained using a phase-space model. The same limit is obtained using a photon pole amplitude model.

 $\Gamma(K^- K^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{203}/\Gamma$ 

A test of lepton family-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-4}$	90	5	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\phi e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{204}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-5}$	90	0	117 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<4.7 \times 10^{-5}$	90	0	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV
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117 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 3.3 \times 10^{-5}$  using a photon pole amplitude model.

 $\Gamma(\bar{K}^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{205}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.0 \times 10^{-4}$	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

 $\Gamma(K^- \pi^+ e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{206}/\Gamma$ 

A test of lepton family-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<5.53 \times 10^{-4}$	90	15	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(\overline{K}^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$  $\Gamma_{207}/\Gamma$ 

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<8.3 \times 10^{-5}$	90	9	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$<1.0 \times 10^{-4}$	90	0	<sup>118</sup> FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

<sup>118</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The same limit is obtained using a photon pole amplitude model.

 $\Gamma(\pi^- \pi^- e^+ e^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{208}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.12 \times 10^{-4}$	90	1	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\pi^- \pi^- \mu^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{209}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.9 \times 10^{-5}$	90	1	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(K^- \pi^- e^+ e^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{210}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.06 \times 10^{-4}$	90	2	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(K^- \pi^- \mu^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{211}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<3.9 \times 10^{-4}$	90	14	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(K^- K^- e^+ e^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{212}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.52 \times 10^{-4}$	90	2	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(K^- K^- \mu^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{213}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<9.4 \times 10^{-5}$	90	1	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(\pi^- \pi^- \mu^+ \mu^+ + \text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{214}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<7.9 \times 10^{-5}$	90	4	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

$\Gamma(K^-\pi^-e^+\mu^++\text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{215}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.18 \times 10^{-4}$	90	7	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 $\Gamma(K^+K^-e^+\mu^++\text{c.c.})/\Gamma_{\text{total}}$  $\Gamma_{216}/\Gamma$ 

A test of lepton-number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<5.7 \times 10^{-5}$	90	0	AITALA	01C E791	$\pi^-$ nucleus, 500 GeV

 **$D^0$  CP-VIOLATING DECAY-RATE ASYMMETRIES** $A_{CP}(K^+K^-)$  in  $D^0, \bar{D}^0 \rightarrow K^+K^-$ 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.005±0.016 OUR AVERAGE</b>				
0.000±0.022±0.008	3023	119 CSORNA	02 CLE2	$e^+e^- \approx \gamma(4S)$
-0.001±0.022±0.015	3330	119 LINK	00B FOCS	
-0.010±0.049±0.012	609	119 AITALA	98C E791	$-0.093 < A_{CP} < +0.073$ (90% CL)
+0.080±0.061		BARTEL	95 CLE2	$-0.022 < A_{CP} < +0.18$ (90% CL)
+0.024±0.084		119 FRABETTI	94I E687	$-0.11 < A_{CP} < +0.16$ (90% CL)

119 FRABETTI 94I, AITALA 98C, LINK 00B, and CSORNA 02 measure  $N(D^0 \rightarrow K^+K^-)/N(D^0 \rightarrow K^-\pi^+)$ , the ratio of numbers of events observed, and similarly for the  $\bar{D}^0$ .

 $A_{CP}(K_S^0K_S^0)$  in  $D^0, \bar{D}^0 \rightarrow K_S^0K_S^0$ 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
-0.23±0.19	65	BONVICINI	01 CLE2	$e^+e^- \approx 10.6$ GeV

 $A_{CP}(\pi^+\pi^-)$  in  $D^0, \bar{D}^0 \rightarrow \pi^+\pi^-$ 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.021±0.026 OUR AVERAGE</b>				
0.019±0.032±0.008	1136	120 CSORNA	02 CLE2	$e^+e^- \approx \gamma(4S)$
+0.048±0.039±0.025	1177	120 LINK	00B FOCS	
-0.049±0.078±0.030	343	120 AITALA	98C E791	$-0.186 < A_{CP} < +0.088$ (90% CL)

120 AITALA 98C, LINK 00B, and CSORNA 02 measure  $N(D^0 \rightarrow \pi^+\pi^-)/N(D^0 \rightarrow K^-\pi^+)$ , the ratio of numbers of events observed, and similarly for the  $\bar{D}^0$ .

**$A_{CP}(\pi^0\pi^0)$  in  $D^0, \bar{D}^0 \rightarrow \pi^0\pi^0$** 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*: D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.001±0.048</b>	810	BONVICINI	01	CLE2 $e^+e^- \approx 10.6$ GeV

 **$A_{CP}(K_S^0\phi)$  in  $D^0, \bar{D}^0 \rightarrow K_S^0\phi$** 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*: D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.028±0.094</b>	BARTEL	95	CLE2 $-0.182 < A_{CP} < +0.126$ (90%CL)

 **$A_{CP}(K_S^0\pi^0)$  in  $D^0, \bar{D}^0 \rightarrow K_S^0\pi^0$** 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*: D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.001±0.013</b>	9099	BONVICINI	01	CLE2 $e^+e^- \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

-0.018±0.030	BARTEL	95	CLE2	See BONVICINI 01
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 **$A_{CP}(K^\pm\pi^\mp)$  in  $D^0 \rightarrow K^+\pi^-, \bar{D}^0 \rightarrow K^-\pi^+$** 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*: D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.02<sup>+0.19</sup><sub>-0.20</sub>±0.01</b>	45	121 GODANG	00	CLE2 $-0.43 < A_{CP} < +0.34$ (95%CL)

121 This GODANG 00 result assumes no  $D^0\bar{D}^0$  mixing; it becomes  $-0.01^{+0.16}_{-0.17} \pm 0.01$  when mixing is allowed.

 **$A_{CP}(K^\mp\pi^\pm\pi^0)$  in  $D^0 \rightarrow K^-\pi^+\pi^0, \bar{D}^0 \rightarrow K^+\pi^-\pi^0$** 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*: D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.031±0.086</b>	122 KOPP	01	CLE2 $e^+e^- \approx 10.6$ GeV

122 KOPP 01 fits separately the  $D^0$  and  $\bar{D}^0$  Dalitz plots and then calculates the integrated difference of normalized densities divided by the integrated sum.

 **$A_{CP}(K^\pm\pi^\mp\pi^0)$  in  $D^0 \rightarrow K^+\pi^-\pi^0, \bar{D}^0 \rightarrow K^-\pi^+\pi^0$** 

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*: D^{*+} \rightarrow D^0\pi^+$  and  $D^{*-} \rightarrow \bar{D}^0\pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>+0.09<sup>+0.25</sup><sub>-0.22</sub></b>	38	BRANDENB... 01	CLE2	$e^+e^- \approx \gamma(4S)$

## **D<sup>0</sup> PRODUCTION CROSS SECTION AT $\psi(3770)$**

A compilation of the cross sections for the direct production of  $D^0$  mesons at or near the  $\psi(3770)$  peak in  $e^+ e^-$  production.

<u>VALUE (nanobarns)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>			
5.8 $\pm$ 0.5 $\pm$ 0.6	123 ADLER	88C MRK3	$e^+ e^-$ 3.768 GeV
7.3 $\pm$ 1.3	124 PARTRIDGE	84 CBAL	$e^+ e^-$ 3.771 GeV
8.00 $\pm$ 0.95 $\pm$ 1.21	125 SCHINDLER	80 MRK2	$e^+ e^-$ 3.771 GeV
11.5 $\pm$ 2.5	126 PERUZZI	77 MRK1	$e^+ e^-$ 3.774 GeV
123 This measurement compares events with one detected $D$ to those with two detected $D$ mesons, to determine the absolute cross section. ADLER 88C find the ratio of cross sections (neutral to charged) to be $1.36 \pm 0.23 \pm 0.14$ .			
124 This measurement comes from a scan of the $\psi(3770)$ resonance and a fit to the cross section. PARTRIDGE 84 measures $6.4 \pm 1.15$ nb for the cross section. We take the phase space division of neutral and charged $D$ mesons in $\psi(3770)$ decay to be 1.33, and we assume that the $\psi(3770)$ is an isosinglet to evaluate the cross sections. The noncharm decays (e.g. radiative) of the $\psi(3770)$ are included in this measurement and may amount to a few percent correction.			
125 This measurement comes from a scan of the $\psi(3770)$ resonance and a fit to the cross section. SCHINDLER 80 assume the phase space division of neutral and charged $D$ mesons in $\psi(3770)$ decay to be 1.33, and that the $\psi(3770)$ is an isosinglet. The noncharm decays (e.g. radiative) of the $\psi(3770)$ are included in this measurement and may amount to a few percent correction.			
126 This measurement comes from a scan of the $\psi(3770)$ resonance and a fit to the cross section. The phase space division of neutral and charged $D$ mesons in $\psi(3770)$ decay is taken to be 1.33, and $\psi(3770)$ is assumed to be an isosinglet. The noncharm decays (e.g. radiative) of the $\psi(3770)$ are included in this measurement and may amount to a few percent correction. We exclude this measurement from the average because of uncertainties in the contamination from $\tau$ lepton pairs. Also see RAPIDIS 77.			

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Also	96B	PRL 77 2147 (errata)	A. Freyberger <i>et al.</i>	(CLEO Collab.)
KUBOTA	96B	PR D54 2994	Y. Kubota <i>et al.</i>	(CLEO Collab.)
ADAMOVICH	95	PL B353 563	M.I. Adamovich <i>et al.</i>	(CERN BEATRICE Collab.)
BARTELT	95	PR D52 4860	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BUTLER	95	PR D52 2656	F. Butler <i>et al.</i>	(CLEO Collab.)
FRAZETTI	95C	PL B354 486	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRAZETTI	95G	PL B364 127	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	95	PL B345 85	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ALBRECHT	94	PL B324 249	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	94F	PL B340 125	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	94I	ZPHY C64 375	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
CINABRO	94	PRL 72 1406	D. Cinabro <i>et al.</i>	(CLEO Collab.)
FRAZETTI	94C	PL B321 295	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRAZETTI	94D	PL B323 459	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRAZETTI	94G	PL B331 217	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRAZETTI	94I	PR D50 R2953	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRAZETTI	94J	PL B340 254	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	94	PL B336 605	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
MISHRA	94	PR D50 R9	C.S. Mishra <i>et al.</i>	(FNAL E789 Collab.)
AKERIB	93	PRL 71 3070	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
ALBRECHT	93D	PL B308 435	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	93	PR D48 56	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BEAN	93C	PL B317 647	A. Bean <i>et al.</i>	(CLEO Collab.)
FRAZETTI	93I	PL B315 203	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	93B	PL B313 260	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
PROCARIO	93B	PR D48 4007	M. Procario <i>et al.</i>	(CLEO Collab.)
SELEN	93	PRL 71 1973	M.A. Selen <i>et al.</i>	(CLEO Collab.)
ADAMOVICH	92	PL B280 163	M.I. Adamovich <i>et al.</i>	(CERN WA82 Collab.)
ALBRECHT	92P	ZPHY C56 7	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	92B	PR D46 R1	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ANJOS	92C	PR D46 1941	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BARLAG	92C	ZPHY C55 383	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
Also	90D	ZPHY C48 29	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
COFFMAN	92B	PR D45 2196	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
Also	90	PRL 64 2615	J. Adler <i>et al.</i>	(Mark III Collab.)
FRAZETTI	92	PL B281 167	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRAZETTI	92B	PL B286 195	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALVAREZ	91B	ZPHY C50 11	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
AMMAR	91	PR D44 3383	R. Ammar <i>et al.</i>	(CLEO Collab.)
ANJOS	91	PR D43 R635	J.C. Anjos <i>et al.</i>	(FNAL-TPS Collab.)
ANJOS	91D	PR D44 R3371	J.C. Anjos <i>et al.</i>	(FNAL-TPS Collab.)
BAI	91	PRL 66 1011	Z. Bai <i>et al.</i>	(Mark III Collab.)
COFFMAN	91	PL B263 135	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
CRAWFORD	91B	PR D44 3394	G. Crawford <i>et al.</i>	(CLEO Collab.)
DECAMP	91J	PL B266 218	D. Decamp <i>et al.</i>	(ALEPH Collab.)
FRAZETTI	91	PL B263 584	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KINOSHITA	91	PR D43 2836	K. Kinoshita <i>et al.</i>	(CLEO Collab.)
KODAMA	91	PRL 66 1819	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ALBRECHT	90C	ZPHY C46 9	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	90	PRL 65 1184	J. Alexander <i>et al.</i>	(CLEO Collab.)
ALEXANDER	90B	PR D65 1531	J. Alexander <i>et al.</i>	(CLEO Collab.)
ALVAREZ	90	ZPHY C47 539	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
ANJOS	90D	PR D42 2414	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BARLAG	90C	ZPHY C46 563	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
ADLER	89	PRL 62 1821	J. Adler <i>et al.</i>	(Mark III Collab.)
ADLER	89C	PR D40 906	J. Adler <i>et al.</i>	(Mark III Collab.)
ALBRECHT	89D	ZPHY C43 181	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	89F	PRL 62 1587	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)

ABACHI	88	PL B205 411	S. Abachi <i>et al.</i>	(HRS Collab.)
ADLER	88	PR D37 2023	J. Adler <i>et al.</i>	(Mark III Collab.)
ADLER	88C	PRL 60 89	J. Adler <i>et al.</i>	(Mark III Collab.)
ALBRECHT	88G	PL B209 380	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88I	PL B210 267	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	88C	PRL 60 1239	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BORTOLETTO	88	PR D37 1719	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
Also	89D	PR D39 1471 erratum	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
CUMALAT	88	PL B210 253	J.P. Cumalat <i>et al.</i>	(E-400 Collab.)
HAAS	88	PRL 60 1614	P. Haas <i>et al.</i>	(CLEO Collab.)
RAAB	88	PR D37 2391	J.R. Raab <i>et al.</i>	(FNAL E691 Collab.)
ADAMOVICH	87	EPL 4 887	M.I. Adamovich <i>et al.</i>	(Photon Emulsion Collab.)
ADLER	87	PL B196 107	J. Adler <i>et al.</i>	(Mark III Collab.)
AGUILAR-...	87E	ZPHY C36 551	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also	88B	ZPHY C40 321	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
AGUILAR-...	87F	ZPHY C36 559	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also	88	ZPHY C38 520 erratum	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
ALBRECHT	87E	ZPHY C33 359	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87K	PL B199 447	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARLAG	87B	ZPHY C37 17	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
BECKER	87C	PL B193 147	J.J. Becker <i>et al.</i>	(Mark III Collab.)
Also	87D	PL B198 590 erratum	J.J. Becker <i>et al.</i>	(Mark III Collab.)
PALKA	87	PL B189 238	H. Palka <i>et al.</i>	(ACCMOR Collab.)
RILES	87	PR D35 2914	K. Riles <i>et al.</i>	(Mark II Collab.)
BAILEY	86	ZPHY C30 51	R. Bailey <i>et al.</i>	(ACCMOR Collab.)
BEBEK	86	PRL 56 1893	C. Bebek <i>et al.</i>	(CLEO Collab.)
LOUIS	86	PRL 56 1027	W.C. Louis <i>et al.</i>	(PRIN, CHIC, ISU)
ALBRECHT	85B	PL 158B 525	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	85F	PL 150B 235	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AUBERT	85	PL 155B 461	J.J. Aubert <i>et al.</i>	(EMC Collab.)
BALTRUSAIT...	85B	PRL 54 1976	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BALTRUSAIT...	85E	PRL 55 150	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BENVENUTI	85	PL 158B 531	A.C. Benvenuti <i>et al.</i>	(BCDMS Collab.)
ADAMOVICH	84B	PL 140B 123	M.I. Adamovich <i>et al.</i>	(CERN WA58 Collab.)
DERRICK	84	PRL 53 1971	M. Derrick <i>et al.</i>	(HRS Collab.)
PARTRIDGE	84	Thesis CALT-68-1150	R.A. Partridge	(Crystal Ball Collab.)
SUMMERS	84	PRL 52 410	D.J. Summers <i>et al.</i>	(UCSB, CARL, COLO+)
BAILEY	83B	PL 132B 237	R. Bailey <i>et al.</i>	(ACCMOR Collab.)
BODEK	82	PL 113B 82	A. Bodek <i>et al.</i>	(ROCH, CIT, CHIC, FNAL+)
FIORINO	81	LNC 30 166	A. Fiorino <i>et al.</i>	
SCHINDLER	81	PR D24 78	R.H. Schindler <i>et al.</i>	(Mark II Collab.)
TRILLING	81	PRPL 75 57	G.H. Trilling	(LBL, UCB) J
ASTON	80E	PL 94B 113	D. Aston <i>et al.</i>	(BONN, CERN, EPOL, GLAS+)
AVERY	80	PRL 44 1309	P. Avery <i>et al.</i>	(ILL, FNAL, COLU)
SCHINDLER	80	PR D21 2716	R.H. Schindler <i>et al.</i>	(Mark II Collab.)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
Also	81	SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)
		Translated from YAF 34	1471.	
ABRAMS	79D	PRL 43 481	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
ATIYA	79	PRL 43 414	M.S. Atiya <i>et al.</i>	(COLU, ILL, FNAL)
BALTAY	78C	PRL 41 73	C. Baltay <i>et al.</i>	(COLU, BNL)
VUILLEMIN	78	PRL 41 1149	V. Vuillemin <i>et al.</i>	(Mark I Collab.)
GOLDHABER	77	PL 69B 503	G. Goldhaber <i>et al.</i>	(Mark I Collab.)
PERUZZI	77	PRL 39 1301	I. Peruzzi <i>et al.</i>	(Mark I Collab.)
PICCOLO	77	PL 70B 260	M. Piccolo <i>et al.</i>	(Mark I Collab.)
RAPIDIS	77	PRL 39 526	P.A. Rapidis <i>et al.</i>	(Mark I Collab.)
GOLDHABER	76	PRL 37 255	G. Goldhaber <i>et al.</i>	(Mark I Collab.)

## OTHER RELATED PAPERS

RICHMAN	95	RMP 67 893	J.D. Richman, P.R. Burchat	(UCSB, STAN)
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