

# Double- $\beta$ Decay

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## Half-life Measurements and Limits for Double- $\beta$ Decay

In all cases of double-beta decay,  $(Z,A) \rightarrow (Z+2,A) + 2e^- + (0 \text{ or } 2)\bar{\nu}_e$ . In the following Listings, only best or comparable limits or lifetimes for each isotope are reported. For  $2\nu$  decay, which is well established, only measured half-lives are reported.

$t_{1/2}(10^{21} \text{ yr})$	$CL\%$	ISOTOPE	TRANSITION	METHOD	DOCUMENT ID	
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>						
>15700	90	$^{76}\text{Ge}$	$0\nu$	Enriched HPGe	<sup>1</sup> AALSETH 02B	
> 58	90	$^{134}\text{Xe}$	$0\nu$	Liquid Xe Scint.	<sup>2</sup> BERNABEI 02D	
> 1200	90	$^{136}\text{Xe}$	$0\nu$	Liquid Xe Scint.	<sup>3</sup> BERNABEI 02D	
$15000 \pm 168000$ $- 7500$		$^{76}\text{Ge}$	$0\nu$	Enriched HPGe	<sup>4</sup> KLAPDOR-K... 02D	
$(7.2 \pm 0.9 \pm 1.8)\text{E-3}$		$^{100}\text{Mo}$	$2\nu$	Liq. Ar ioniz.	<sup>5</sup> ASHITKOV 01	
> 4.9	90	$^{100}\text{Mo}$	$0\nu$	Liq. Ar ioniz.	<sup>6</sup> ASHITKOV 01	
> 1.3	90	$^{160}\text{Gd}$	$0\nu$	$^{160}\text{Gd}_2\text{SiO}_5:\text{Ce}$	<sup>7</sup> DANEVICH 01	
> 1.3	90	$^{160}\text{Gd}$	$0\nu \rightarrow 2^+$	$^{160}\text{Gd}_2\text{SiO}_5:\text{Ce}$	<sup>8</sup> DANEVICH 01	
$0.59 \pm 0.17 \pm 0.06$		$^{100}\text{Mo}$	$0\nu+2\nu \rightarrow 0^+_1$	Ge coinc.	<sup>9</sup> DEBRAECKEL.01	
> 55	90	$^{100}\text{Mo}$	$0\nu, \langle m_\nu \rangle$	ELEGANT V	<sup>10</sup> EJIRI 01	
> 42	90	$^{100}\text{Mo}$	$0\nu, \langle \lambda \rangle$	ELEGANT V	<sup>10</sup> EJIRI 01	
> 49	90	$^{100}\text{Mo}$	$0\nu, \langle \eta \rangle$	ELEGANT V	<sup>10</sup> EJIRI 01	
>19000	90	$^{76}\text{Ge}$	$0\nu$	Enriched HPGe	<sup>11</sup> KLAPDOR-K... 01	
$1.55 \pm 0.001 \pm 0.19$ $- 0.15$	90	$^{76}\text{Ge}$	$2\nu$	Enriched HPGe	<sup>12</sup> KLAPDOR-K... 01	
$(9.4 \pm 3.2)\text{E-3}$	90	$^{96}\text{Zr}$	$0\nu+2\nu$	Geochem	<sup>13</sup> WIESER 01	
> 144	90	$^{130}\text{Te}$	$0\nu$	Cryog. det.	<sup>14</sup> ALESSAND... 00	
> 86	90	$^{128}\text{Te}$	$0\nu$	Cryog. det.	<sup>14</sup> ALESSAND... 00	
$0.042 \pm 0.033$		$^{48}\text{Ca}$	$2\nu$	Ge spectrometer	<sup>15</sup> BRUDANIN 00	
$0.026 \pm 0.001 \pm 0.007$ $- 0.004$		$^{116}\text{Cd}$	$2\nu$	$^{116}\text{CdWO}_4$ scint. <sup>16</sup>	DANEVICH 00	
> 70	90	$^{116}\text{Cd}$	$0\nu$	$^{116}\text{CdWO}_4$ scint. <sup>17</sup>	DANEVICH 00	
$0.021 \pm 0.008 \pm 0.002$ $- 0.004$		$^{96}\text{Zr}$	$2\nu$	NEMO-2	<sup>18</sup> ARNOLD 99	
> 1.0	90	$^{96}\text{Zr}$	$0\nu$	NEMO-2	<sup>18</sup> ARNOLD 99	
$(8.3 \pm 1.0 \pm 0.7)\text{E-2}$		$^{82}\text{Se}$	$2\nu$	NEMO-2	<sup>19</sup> ARNOLD 98	
> 9.5	90	$^{82}\text{Se}$	$0\nu$	NEMO-2	<sup>20</sup> ARNOLD 98	
> 2.8	90	$^{82}\text{Se}$	$0\nu \rightarrow 2^+$	NEMO-2	<sup>21</sup> ARNOLD 98	
$(7.6 \pm 2.2)\text{E-3}$ $- 1.4$		$^{100}\text{Mo}$	$2\nu$	Si(Li)	<sup>22</sup> ALSTON-... 97	
$(6.82 \pm 0.38 \pm 0.68)\text{E-3}$		$^{100}\text{Mo}$	$2\nu$	TPC	<sup>23</sup> DESILVA 97	
$(6.75 \pm 0.37 \pm 0.68)\text{E-3}$		$^{150}\text{Nd}$	$2\nu$	TPC	<sup>24</sup> DESILVA 97	
> 1.2	90	$^{150}\text{Nd}$	$0\nu$	TPC	<sup>25</sup> DESILVA 97	
$1.77 \pm 0.01 \pm 0.13$ $- 0.11$		$^{76}\text{Ge}$	$2\nu$	Enriched HPGe	<sup>26</sup> GUENTHER 97	
$(3.75 \pm 0.35 \pm 0.21)\text{E-2}$		$^{116}\text{Cd}$	$2\nu$	$0^+ \rightarrow 0^+$	NEMO 2	<sup>27</sup> ARNOLD 96

$0.043^{+0.024}_{-0.011} \pm 0.014$	${}^{48}\text{Ca}$	$2\nu$	TPC	28	BALYSH	96	
$0.79 \pm 0.10$	${}^{130}\text{Te}$	$0\nu+2\nu$	Geochem	29	TAKAOKA	96	
$0.61^{+0.18}_{-0.11}$	${}^{100}\text{Mo}$	$0\nu+2\nu$	$0^+ \rightarrow 0^+_1$ $\gamma$ in HPGe	30	BARABASH	95	
$(9.5 \pm 0.4 \pm 0.9)\text{E-}3$	${}^{100}\text{Mo}$	$2\nu$	NEMO 2		DASSIE	95	
$> 0.6$	90	${}^{100}\text{Mo}$	$0\nu$	$0^+ \rightarrow 0^+_1$	NEMO 2	DASSIE	
$0.026^{+0.009}_{-0.005}$	${}^{116}\text{Cd}$	$2\nu$	$0^+ \rightarrow 0^+$	ELEGANT IV	EJIRI	95	
$0.017^{+0.010}_{-0.005} \pm 0.0035$	${}^{150}\text{Nd}$	$2\nu$	$0^+ \rightarrow 0^+$	TPC	ARTEMEV	93	
$0.039 \pm 0.009$	${}^{96}\text{Zr}$	$0\nu+2\nu$	Geochem		KAWASHIMA	93	
$2.7 \pm 0.1$	${}^{130}\text{Te}$	$0\nu+2\nu$	Geochem		BERNATOW...	92	
$7200 \pm 400$	${}^{128}\text{Te}$	$0\nu+2\nu$	Geochem	31	BERNATOW...	92	
$> 27$	68	${}^{82}\text{Se}$	$0\nu$	$0^+ \rightarrow 0^+$	ELLIOTT	92	
$0.108^{+0.026}_{-0.006}$	${}^{82}\text{Se}$	$2\nu$	$0^+ \rightarrow 0^+$	TPC	ELLIOTT	92	
$2.0 \pm 0.6$		${}^{238}\text{U}$	$0\nu+2\nu$	Radiochem	32	TURKEVICH	91
$> 9.5$	76	${}^{48}\text{Ca}$	$0\nu$	$\text{CaF}_2$ scint.	YOU	91	
$2.60 \pm 0.28$		${}^{130}\text{Te}$	$0\nu+2\nu$	Geochem	33	KIRSTEN	83

<sup>1</sup> AALSETH 02B limit is based on 117 mol·yr of data using enriched Ge detectors.

Background reduction by means of pulse shape analysis is applied to part of the data set.

Reported limit is slightly less restrictive than that in Klapdor-Kleingrothaus 01

However, it excludes part of the allowed half-life range reported in Klapdor-Kleingrothaus 01B for the same nuclide.

<sup>2</sup> BERNABEI 02D report a limit for the  $0\nu, 0^+ \rightarrow 0^+$  decay of  ${}^{134}\text{Xe}$ , present in the source at 17%, by considering the maximum number of events for this mode compatible with the fitted smooth background.

<sup>3</sup> BERNABEI 02D report a limit for the  $0\nu, 0^+ \rightarrow 0^+$  decay of  ${}^{136}\text{Xe}$ , by considering the maximum number of events for this mode compatible with the fitted smooth background.

The quoted sensitivity is  $450 \times 10^{21}$  yr. The Feldman and Cousins method is used to obtain the quoted limit.

<sup>4</sup> Klapdor-Kleingrothaus 02D is an expanded version of Klapdor-Kleingrothaus 01B. The authors re-evaluate the data collected by the Heidelberg-Moscow experiment (Klapdor-Kleingrothaus 01) and present a more detailed description of their analysis of an excess of counts at the energy expected for neutrinoless double-beta decay. They interpret this excess, which has a significance of 2.2 to  $3.1\sigma$  depending on the data analysis, as evidence for the observation of Lepton Number violation and violation of Baryon minus Lepton Number. The analysis has been criticized by Aalseth 02. The criticisms have been addressed in Klapdor-Kleingrothaus 02. See also Klapdor-Kleingrothaus 02B.

<sup>5</sup> ASHITKOV 01 result for  $2\nu$  of  ${}^{100}\text{Mo}$  is in agreement with other determinations of that halflife.

<sup>6</sup> ASHITKOV 01 result for  $0\nu$  of  ${}^{100}\text{Mo}$  is less stringent than EJIRI 01.

<sup>7</sup> DANEVICH 01 place limit on  $0\nu$  decay of  ${}^{160}\text{Gd}$  using  $\text{Gd}_2\text{SiO}_5:\text{Ce}$  crystal scintillators. The limit is more stringent than KOBAYASHI 95.

<sup>8</sup> DANEVICH 01 place limits on  $0\nu$  decay of  ${}^{160}\text{Gd}$  into excited  $2^+$  state of daughter nucleus using  $\text{Gd}_2\text{SiO}_5:\text{Ce}$  crystal scintillators.

<sup>9</sup> DEBRAECKEER 01 performed an inclusive measurement of the  $\beta\beta$  decay into the second excited state of the daughter nucleus. A novel coincidence technique counting the de-excitation photons is employed. The result agrees with BARABASH 95.

<sup>10</sup> EJIRI 01 uses tracking calorimeter and isotopically enriched passive source. Efficiencies were calculated assuming  $\langle m_\nu \rangle$ ,  $\langle \lambda \rangle$ , or  $\langle \eta \rangle$  driven decay. This is a continuation of EJIRI 96 which it supersedes.

<sup>11</sup> Klapdor-Kleingrothaus 01 is a continuation of the work published in BAUDIS 99. Isotopically enriched Ge detectors are used in calorimetric measurement. The most stringent bound is derived from the data set in which pulse-shape analysis has been used to

- reduce background. Exposure time is 35.5 kg y. Supersedes BAUDIS 99 as most stringent result.
- 12 KLAUDOR-KLEINGROTHAUS 01 is a measurement of the  $\beta\beta 2\nu$ -decay rate with higher statistics than GUENTHER 97. The reported value has a worse systematic error than their previous result.
- 13 WIESER 01 reports an inclusive geochemical measurement of  $^{96}\text{Zr}$   $\beta\beta$  half life. Their result agrees within  $2\sigma$  with ARNOLD 99 but only marginally, within  $3\sigma$ , with KAWASHIMA 93.
- 14 ALESSANDRELLO 00 limit is based on calorimetric measurement with an array of 20  $\text{TeO}_2$  cryogenic detectors. Uses enriched and natural Te crystals. Replaces ALESSANDRELLO 98.
- 15 BRUDANIN 00 determine the  $2\nu$  halflife of  $^{48}\text{Ca}$ . Their value is less accurate than BALYSH 96.
- 16 DANEVICH 00 provides calorimetric measurement of  $2\nu$  decay of  $^{116}\text{Cd}$  using enriched  $\text{CdWO}_4$  scintillators. Agrees with EJIRI 95 and ARNOLD 96.
- 17 DANEVICH 00 places limits on  $0\nu$  decay of  $^{116}\text{Cd}$  using enriched  $\text{CdWO}_4$  scintillators. Replaces GEORGADZE 95.
- 18 ARNOLD 99 measure directly the  $2\nu$  decay of Zr for the first time, using the NEMO-2 tracking detector and an isotopically enriched source. The lifetime is more accurate than the geochemical result of KAWASHIMA 93.
- 19 ARNOLD 98 measure the  $2\nu$  decay of  $^{82}\text{Se}$  by comparing the spectra in an enriched and natural selenium source using the NEMO-2 tracking detector. The measured half-life is in agreement, perhaps slightly shorter, than ELLIOTT 92.
- 20 ARNOLD 98 determine the limit for  $0\nu$  decay to the ground state of  $^{82}\text{Se}$  using the NEMO-2 tracking detector. The half-life limit is in agreement, but less stringent, than ELLIOTT 92.
- 21 ARNOLD 98 determine the limit for  $0\nu$  decay to the excited  $2^+$  state of  $^{82}\text{Se}$  using the NEMO-2 tracking detector.
- 22 ALSTON-GARNJOST 97 report evidence for  $2\nu$  decay of  $^{100}\text{Mo}$ . This decay has been also observed by EJIRI 91, DASSIE 95, and DESILVA 97.
- 23 DESILVA 97 result for  $2\nu$  decay of  $^{100}\text{Mo}$  is in agreement with ALSTON-GARNJOST 97 and DASSIE 95. This measurement has the smallest errors.
- 24 DESILVA 97 result for  $2\nu$  decay of  $^{150}\text{Nd}$  is in marginal agreement with ARTEMEV 93. It has smaller errors.
- 25 DESILVA 97 do not explain whether their efficiency for  $0\nu$  decay of  $^{150}\text{Nd}$  was calculated under the assumption of a  $\langle m_\nu \rangle$ ,  $\langle \lambda \rangle$ , or  $\langle \eta \rangle$  driven decay.
- 26 GUENTHER 97 half-life for the  $2\nu$  decay of  $^{76}\text{Ge}$  is not in good agreement with the previous measurements of BALYSH 94, AVIGNONE 91, and MILEY 90.
- 27 ARNOLD 96 measure the  $2\nu$  decay of  $^{116}\text{Cd}$ . This result is in agreement with EJIRI 95, but has smaller errors. Supersedes ARNOLD 95.
- 28 BALYSH 96 measure the  $2\nu$  decay of  $^{48}\text{Ca}$ , using a passive source of enriched  $^{48}\text{Ca}$  in a TPC.
- 29 TAKAOKA 96 measure the geochemical half-life of  $^{130}\text{Te}$ . Their value is in disagreement with the quoted values of BERNATOWICZ 92 and KIRSTEN 83; but agrees with several other unquoted determinations, e.g., MANUEL 91.
- 30 BARABASH 95 cannot distinguish  $0\nu$  and  $2\nu$ , but it is inferred indirectly that the  $0\nu$  mode accounts for less than 0.026% of their event sample. They also note that their result disagrees with the previous experiment by the NEMO group (BLUM 92).
- 31 BERNATOWICZ 92 finds  $^{128}\text{Te}/^{130}\text{Te}$  activity ratio from slope of  $^{128}\text{Xe}/^{132}\text{Xe}$  vs  $^{130}\text{Xe}/^{132}\text{Xe}$  ratios during extraction, and normalizes to lead-dated ages for the  $^{130}\text{Te}$  lifetime. The authors state that their results imply that “(a) the double beta decay of  $^{128}\text{Te}$  has been firmly established and its half-life has been determined . . . without any ambiguity due to trapped Xe interferences . . . (b) Theoretical calculations . . . underestimate the [long half-lives of  $^{128}\text{Te}$   $^{130}\text{Te}$ ] by 1 or 2 orders of magnitude, pointing to a real suppression in the  $2\nu$  decay rate of these isotopes. (c) Despite [this], most  $\beta\beta$ -models

predict a *ratio* of  $2\nu$  decay widths . . . in fair agreement with observation." Further details of the experiment are given in BERNATOWICZ 93. Our listed half-life has been revised downward from the published value by the authors, on the basis of reevaluated cosmic-ray  $^{128}\text{Xe}$  production corrections.

- <sup>32</sup>TURKEVICH 91 observes activity in old U sample. The authors compare their results with theoretical calculations. They state "Using the phase-space factors of Boehm and Vogel (BOEHM 87) leads to matrix element values for the  $^{238}\text{U}$  transition in the same range as deduced for  $^{130}\text{Te}$  and  $^{76}\text{Ge}$ . On the other hand, the latest theoretical estimates (STAUDT 90) give an upper limit that is 10 times lower. This large discrepancy implies either a defect in the calculations or the presence of a faster path than the standard two-neutrino mode in this case." See BOEHM 87 and STAUDT 90.  
<sup>33</sup>KIRSTEN 83 reports " $2\sigma$ " error. References are given to earlier determinations of the  $^{130}\text{Te}$  lifetime.

## $\langle m_\nu \rangle$ , The Effective Weighted Sum of Majorana Neutrino Masses Contributing to Neutrinoless Double- $\beta$ Decay

$\langle m_\nu \rangle = |\sum U_{1j}^2 m_{\nu_j}|$ , where the sum goes from 1 to  $n$  and where  $n$  = number of neutrino generations, and  $\nu_j$  is a Majorana neutrino. Note that  $U_{ej}^2$ , not  $|U_{ej}|^2$ , occurs in the sum. The possibility of cancellations has been stressed. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported.

VALUE (eV)	CL%	ISOTOPE	TRANSITION	METHOD	DOCUMENT ID
• • • We do not use the following data for averages, fits, limits, etc.					• • •
< 0.33–1.35	90			Enriched HPGe	<sup>34</sup> AALSETH 02B
<2.9	90	$^{136}\text{Xe}$	$0\nu$	Liquid Xe Scint.	<sup>35</sup> BERNABEI 02D
$0.39^{+0.17}_{-0.28}$		$^{76}\text{Ge}$	$0\nu$	Enriched HPGe	<sup>36</sup> KLAPDOR-K... 02D
< 2.1–4.8	90	$^{100}\text{Mo}$	$0\nu$	ELEGANT V	<sup>37</sup> EJIRI 01
< 0.35	90	$^{76}\text{Ge}$		Enriched HPGe	<sup>38</sup> KLAPDOR-K... 01
< 1.1–2.6	90	$^{130}\text{Te}$	$0\nu$	Cryog. det.	<sup>39</sup> ALESSAND... 00
< 2.4–2.6	90	$^{116}\text{Cd}$	$0\nu$	$^{116}\text{CdWO}_4$ scint	<sup>40</sup> DANEVICH 00
<23	90	$^{96}\text{Zr}$		NEMO-2	<sup>41</sup> ARNOLD 99
< 1.1–1.5		$^{128}\text{Te}$		Geochem	<sup>42</sup> BERNATOW... 92
<5	68	$^{82}\text{Se}$		TPC	<sup>43</sup> ELLIOTT 92
<8.3	76	$^{48}\text{Ca}$	$0\nu$	$\text{CaF}_2$ scint.	YOU 91

<sup>34</sup>AALSETH 02B reported range of limits on  $\langle m_\nu \rangle$  reflects the spread of theoretical nuclear matrix elements. Excludes part of allowed mass range reported in KLAPDOR-KLEINGROTHAUS 01B.

<sup>35</sup>BERNABEI 02D limit is based on the matrix elements of SIMKOVIC 02. The range of neutrino masses based on a variety of matrix elements is 1.1–2.9 eV.

<sup>36</sup>KLAPDOR-KLEINGROTHAUS 02D is a detailed description of the analysis of the data collected by the Heidelberg-Moscow experiment, previously presented in KLAPDOR-KLEINGROTHAUS 01B. Matrix elements in STAUDT 90 have been used. See the footnote in the preceding table for further details. See also KLAPDOR-KLEINGROTHAUS 02B.

<sup>37</sup>The range of the reported  $\langle m_\nu \rangle$  values reflects the spread of the nuclear matrix elements. On axis value assuming  $\langle \lambda \rangle = \langle \eta \rangle = 0$ .

<sup>38</sup>KLAPDOR-KLEINGROTHAUS 01 uses the calculation by STAUDT 90. Using several other models in the literature could worsen the limit up to 1.2 eV. This is the most stringent experimental bound on  $m_\nu$ . It supersedes BAUDIS 99B.

<sup>39</sup>ALESSANDRELLO 00 spread in limit for  $\langle m_\nu \rangle$  reflects the range found for theoretical matrix elements.

- <sup>40</sup> DANEVICH 00 limit for  $\langle m_\nu \rangle$  is based on the nuclear matrix elements of STAUDT 90 (2.6 eV) and ARNOLD 96 (2.4 eV).
- <sup>41</sup> ARNOLD 99 limit based on the nuclear matrix elements of STAUDT 90.
- <sup>42</sup> BERNATOWICZ 92 finds these majorona neutrino mass limits assuming that the measured geochemical decay width is a limit on the  $0\nu$  decay width. The range is the range found using matrix elements from HAXTON 84, TOMODA 87, and SUHONEN 91. Further details of the experiment are given in BERNATOWICZ 93.
- <sup>43</sup> ELLIOTT 92 uses the matrix elements of HAXTON 84.

### Limits on Lepton-Number Violating ( $V+A$ ) Current Admixture

For reasons given in the discussion at the beginning of this section, we list only results from 1989 and later.  $\langle \lambda \rangle = \lambda \sum U_{ej} V_{ej}$  and  $\langle \eta \rangle = \eta \sum U_{ej} V_{ej}$ , where the sum is over the number of neutrino generations. This sum vanishes for massless or unmixed neutrinos. In the following Listings, only best or comparable limits or lifetimes for each isotope are reported.

$\langle \lambda \rangle (10^{-6})$	$CL\%$	$\langle \eta \rangle (10^{-8})$	$CL\%$	ISOTOPE	METHOD	DOCUMENT ID
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>						
< 3.2–4.7	90	< 2.4–2.7	90	<sup>100</sup> Mo	ELEGANT V	<sup>44</sup> EJIRI 01
< 1.9–3.9	90	< 1.2–6.4	90	<sup>130</sup> Te	Cryog. det.	<sup>45</sup> ALESSAND... 00
< 3.4	90	< 3.9	90	<sup>116</sup> Cd	<sup>116</sup> CdWO <sub>4</sub> scint.	<sup>46</sup> DANEVICH 00
< 1.1	90	< 0.64	90	<sup>76</sup> Ge	Enriched HPGe	<sup>47</sup> GUENTHER 97
< 4.4	90	< 2.3	90	<sup>136</sup> Xe	TPC	<sup>48</sup> VUILLEUMIER 93
		< 5.3		<sup>128</sup> Te	Geochem	<sup>49</sup> BERNATOW... 92

- <sup>44</sup> The range of the reported  $\langle \lambda \rangle$  and  $\langle \eta \rangle$  values reflects the spread of the nuclear matrix elements. On axis value assuming  $\langle m_\nu \rangle=0$  and  $\langle \lambda \rangle=\langle \eta \rangle=0$ , respectively.
- <sup>45</sup> ALESSANDRELLO 00 limits for  $\langle \lambda \rangle$  and  $\langle \eta \rangle$  use several nuclear matrix element calculations. Limits reported for  $\langle m_\nu \rangle = \langle \eta \rangle = \langle \lambda \rangle = 0$ .
- <sup>46</sup> DANEVICH 00 limits for  $\langle \lambda \rangle$  and  $\langle \eta \rangle$  are based on nuclear matrix element of STAUDT 90. Replaces DANEVICH 95.
- <sup>47</sup> GUENTHER 97 limits use the matrix elements of STAUDT 90. Supersedes BALYSH 95 and BALYSH 92.
- <sup>48</sup> VUILLEUMIER 93 uses the matrix elements of MUTO 89. Based on a half-life limit  $2.6 \times 10^{23}$  y at 90%CL.
- <sup>49</sup> BERNATOWICZ 92 takes the measured geochemical decay width as a limit on the  $0\nu$  width, and uses the SUHONEN 91 coefficients to obtain the least restrictive limit on  $\eta$ . Further details of the experiment are given in BERNATOWICZ 93.

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### Double- $\beta$ Decay REFERENCES

AALSETH	02	MPL A17 1475	C.E. Aalseth <i>et al.</i>	
AALSETH	02B	PR D65 092007	C.E. Aalseth <i>et al.</i>	(IGEX Collab.)
BERNABEI	02D	PL B546 23	R. Bernabei <i>et al.</i>	(DAMA Collab.)
KLAPDOR-K...	02	hep-ph/0205228	H.V. Klapdor-Kleingrothaus	
KLAPDOR-K...	02B	JINRRC 110 57	H.V. Klapdor-Kleingrothaus, A. Dietz, I.V. Krivosheina	
KLAPDOR-K...	02D	FP 32 1181	H.V. Klapdor-Kleingrothaus, A. Dietz, I.V. Krivosheina	
SIMKOVIC	02	hep-ph/0204278	F. Simkovic, P. Domin, A. Faessler	
ASHITKOV	01	JETPL 74 529	V.D. Ashitkov <i>et al.</i>	
		Translated from ZETFP 74 601.		

DANEVICH	01	NP A694 375	F.A. Danevich <i>et al.</i>
DEBRAECKEL..	01	PRL 86 3510	L. De Braeckeleer <i>et al.</i>
EJIRI	01	PR C63 065501	H. Ejiri <i>et al.</i>
KLAPDOR-K...	01	EPJ A12 147	H.V. Klapdor-Kleingrothaus <i>et al.</i>
KLAPDOR-K...	01B	MPL A16 2409	H.V. Klapdor-Kleingrothaus <i>et al.</i>
WIESER	01	PR C64 024308	M.E. Wieser, J.R. De Laeter
ALESSAND...	00	PL B486 13	A. Alessandrello <i>et al.</i>
BRUDANIN	00	PL B495 63	V.B. Brudanin <i>et al.</i>
DANEVICH	00	PR C62 045501	F.A. Danevich <i>et al.</i>
ARNOLD	99	NP A658 299	R. Arnold <i>et al.</i>
BAUDIS	99	PR D59 022001	L. Baudis <i>et al.</i>
BAUDIS	99B	PRL 83 41	L. Baudis <i>et al.</i>
ALESSAND...	98	PL B433 156	A. Alessandrello <i>et al.</i>
ARNOLD	98	NP A636 209	R. Arnold <i>et al.</i>
ALSTON-...	97	PR C55 474	M. Alston-Garnjost <i>et al.</i>
DESILVA	97	PR C56 2451	A. de Silva <i>et al.</i>
GUENTHER	97	PR D55 54	M. Gunther <i>et al.</i>
ARNOLD	96	ZPHY C72 239	R. Arnold <i>et al.</i>
BALYSH	96	PRL 77 5186	A. Balysh <i>et al.</i>
EJIRI	96	NP A611 85	H. Ejiri <i>et al.</i>
TAKAOKA	96	PR C53 1557	N. Takaoka, Y. Motomura, K. Nagao
ARNOLD	95	JETPL 61 170	R.G. Arnold <i>et al.</i>
		Translated from ZETFP	61 168.
BALYSH	95	PL B356 450	A. Balysh <i>et al.</i>
BARABASH	95	PL B345 408	A.S. Barabash <i>et al.</i>
DANEVICH	95	PL B344 72	F.A. Danevich <i>et al.</i>
DASSIE	95	PR D51 2090	D. Dassie <i>et al.</i>
EJIRI	95	JPSJ 64 339	H. Ejiri <i>et al.</i>
GEORGADZE	95	PAN 58 1093	A.Sh. Georgadze <i>et al.</i>
		Translated from YAF	58 1170.
KOBAYASHI	95	NP A586 457	M. Kobayashi, M. Kobayashi
BALYSH	94	PL B322 176	A. Balysh <i>et al.</i>
ARTEMEV	93	JETPL 58 262	V.A. Artemiev <i>et al.</i>
		Translated from ZETFP	58 256.
BERNATOW...	93	PR C47 806	T. Bernatowicz <i>et al.</i>
KAWASHIMA	93	PR C47 R2452	A. Kawashima, K. Takahashi, A. Masuda
VUILLEUMIER	93	PR D48 1009	J.C. Vuilleumier <i>et al.</i>
BALYSH	92	PL B283 32	A. Balysh <i>et al.</i>
BERNATOW...	92	PRL 69 2341	T. Bernatowicz <i>et al.</i>
BLUM	92	PL B275 506	D. Blum <i>et al.</i>
ELLIOTT	92	PR C46 1535	S.R. Elliott <i>et al.</i>
AVIGNONE	91	PL B256 559	F.T. Avignone <i>et al.</i>
EJIRI	91	PL B258 17	H. Ejiri <i>et al.</i>
MANUEL	91	JPG 17 S221	O.K. Manuel
SUHONEN	91	NP A535 509	J. Suhonen, S.B. Khadkikar, A. Faessler
TURKEVICH	91	PRL 67 3211	A. Turkevich, T.E. Economou, G.A. Cowan
YOU	91	PL B265 53	K. You <i>et al.</i>
MLEY	90	PRL 65 3092	H.S. Miley <i>et al.</i>
STAUDT	90	EPL 13 31	A. Staudt, K. Muto, H.V. Klapdor-Kleingrothaus
MUTO	89	ZPHY A334 187	K. Muto, E. Bender, H.V. Klapdor
BOEHM	87	Massive Neutrinos Cambridge Univ. Press, Cambridge	F. Bohm, P. Vogel
TOMODA	87	PL B199 475	T. Tomoda, A. Faessler
HAXTON	84	PPNP 12 409	W.C. Haxton, G.J. Stevenson
KIRSTEN	83	PRL 50 474	T. Kirsten, H. Richter, E. Jessberger