

***n***

$$I(J^P) = \frac{1}{2}(\frac{1}{2}^+) \text{ Status: } ***$$

We have omitted some results that have been superseded by later experiments. See our earlier editions.

### ***n* MASS (atomic mass units u)**

The mass is known much more precisely in u (atomic mass units) than in MeV. See the next data block.

VALUE (u)	DOCUMENT ID	TECN	COMMENT
<b>1.00866491560±0.00000000055</b>	MOHR 05	RVUE	2002 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.00866491578±0.00000000055	MOHR 99	RVUE	1998 CODATA value
1.008665904 ±0.000000014	COHEN 87	RVUE	1986 CODATA value

### ***n* MASS (MeV)**

The mass is known much more precisely in u (atomic mass units) than in MeV. The conversion from u to MeV,  $1 \text{ u} = 931.494043 \pm 0.000080 \text{ MeV}/c^2$  (MOHR 05, the 2002 CODATA value), involves the relatively poorly known electronic charge.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>939.565360±0.000081</b>	MOHR 05	RVUE	2002 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
939.565331±0.000037	1 KESSLER 99	SPEC	$np \rightarrow d\gamma$
939.565330±0.000038	MOHR 99	RVUE	1998 CODATA value
939.56565 ±0.00028	2,3 DIFILIPPO 94	TRAP	Penning trap
939.56563 ±0.00028	COHEN 87	RVUE	1986 CODATA value
939.56564 ±0.00028	3,4 GREENE 86	SPEC	$np \rightarrow d\gamma$
939.5731 ±0.0027	3 COHEN 73	RVUE	1973 CODATA value

<sup>1</sup> We use the 1998 CODATA u-to-MeV conversion factor (see the heading above) to get this mass in MeV from the much more precisely measured KESSLER 99 value of  $1.00866491637 \pm 0.0000000082 \text{ u}$ .

<sup>2</sup> The mass is known much more precisely in u:  $m = 1.0086649235 \pm 0.0000000023 \text{ u}$ . We use the 1986 CODATA conversion factor to get the mass in MeV.

<sup>3</sup> These determinations are not independent of the  $m_n - m_p$  measurements below.

<sup>4</sup> The mass is known much more precisely in u:  $m = 1.008664919 \pm 0.000000014 \text{ u}$ .

### **$\bar{n}$ MASS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>939.485±0.051</b>	59	5 CRESTI 86	HBC	$\bar{p}p \rightarrow \bar{n}n$

<sup>5</sup> This is a corrected result (see the erratum). The error is statistical. The maximum systematic error is 0.029 MeV.

## $(m_n - m_{\bar{n}})/m_n$

A test of *CPT* invariance. Calculated from the  $n$  and  $\bar{n}$  masses, above.

<u>VALUE</u>	<u>DOCUMENT ID</u>			
<b><math>(9 \pm 5) \times 10^{-5}</math> OUR EVALUATION</b>				
<b><math>m_n - m_p</math></b>				
<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b><math>1.2933317 \pm 0.0000005</math></b>	<sup>6</sup> MOHR	05	RVUE	2002 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.2933318 $\pm 0.0000005$	<sup>7</sup> MOHR	99	RVUE	1998 CODATA value
1.293318 $\pm 0.000009$	<sup>8</sup> COHEN	87	RVUE	1986 CODATA value
1.2933328 $\pm 0.0000072$	GREENE	86	SPEC	$np \rightarrow d\gamma$
1.293429 $\pm 0.000036$	COHEN	73	RVUE	1973 CODATA value
<sup>6</sup> Calculated by us from the MOHR 05 ratio $m_n/m_p = 1.00137841870 \pm 0.00000000058$ . in u, $m_n - m_p = (1.3884487 \pm 0.0000006) \times 10^{-3}$ u.				
<sup>7</sup> Calculated by us from the MOHR 99 ratio $m_n/m_p = 1.00137841887 \pm 0.00000000058$ . In u, $m_n - m_p = (1.3884489 \pm 0.0000006) \times 10^{-3}$ u.				
<sup>8</sup> Calculated by us from the COHEN 87 ratio $m_n/m_p = 1.001378404 \pm 0.000000009$ . In u, $m_n - m_p = 0.001388434 \pm 0.000000009$ u.				

## **$n$ MEAN LIFE**

We now compile only direct measurements of the lifetime, not those inferred from decay correlation measurements. For the average, we only use measurements with an error less than 10 s.

The most recent result, that of SEREBROV 05, is so far from other results that it makes no sense to include it in the average. It is up to workers in this field to resolve this issue. Until this major disagreement is understood our present average of  $885.7 \pm 0.8$  s must be suspect.

For an early review, see EROZOLIMSKII 89 and papers that follow it in an issue of NIM devoted to the "Proceedings of the International Workshop on Fundamental Physics with Slow Neutrons" (Grenoble 1989). For later reviews and/or commentary, see FREEDMAN 90, SCHRECKENBACH 92, and PENDLEBURY 93.

Limits on lifetimes for *bound* neutrons are given in the section "*p* PARTIAL MEAN LIVES."

<u>VALUE (s)</u>	<u>DOCUMENT ID</u>		<u>COMMENT</u>	
<b><math>885.7 \pm 0.8</math> OUR AVERAGE</b>				
886.8 $\pm 1.2 \pm 3.2$	DEWEY	03	CNTR	In-beam $n, p$ trap
885.4 $\pm 0.9 \pm 0.4$	ARZUMANOV	00	CNTR	UCN double bottle
889.2 $\pm 3.0 \pm 3.8$	BYRNE	96	CNTR	Penning trap
882.6 $\pm 2.7$	<sup>9</sup> MAMPE	93	CNTR	Gravitational trap
888.4 $\pm 3.1 \pm 1.1$	NESVIZHEV...	92	CNTR	Gravitational trap
887.6 $\pm 3.0$	MAMPE	89	CNTR	Gravitational trap
891 $\pm 9$	SPIVAK	88	CNTR	Beam

• • • We do not use the following data for averages, fits, limits, etc. • • •

878.5 ± 0.7 ± 0.3	<sup>10</sup> SEREBROV	05	CNTR	Gravitational trap
888.4 ± 2.9	ALFIMENKOV	90	CNTR	See NESVIZHEVSKII 92
893.6 ± 3.8 ± 3.7	BYRNE	90	CNTR	See BYRNE 96
878 ± 27 ± 14	KOSSAKOW...	89	TPC	Pulsed beam
877 ± 10	PAUL	89	CNTR	Storage ring
876 ± 10 ± 19	LAST	88	SPEC	Pulsed beam
903 ± 13	KOSVINTSEV	86	CNTR	Gravitational trap
937 ± 18	<sup>11</sup> BYRNE	80	CNTR	
875 ± 95	KOSVINTSEV	80	CNTR	
881 ± 8	BONDAREN...	78	CNTR	See SPIVAK 88
918 ± 14	CHRISTENSEN72	CNTR		

<sup>9</sup> IGNATOVICH 95 calls into question some of the corrections and averaging procedures used by MAMPE 93. The response, BONDARENKO 96, denies the validity of the criticisms.

<sup>10</sup> This SEREBROV 05 result is 6.5 standard deviations from our average of previous results and 5.6 standard deviations from the previous most precise result (that of ARZUMANOV 00).

<sup>11</sup> This measurement has been withdrawn (J. Byrne, private communication, 1990).

## *n* MAGNETIC MOMENT

See the "Note on Baryon Magnetic Moments" in the *A* Listings.

VALUE ( $\mu_N$ )	DOCUMENT ID	TECN	COMMENT
<b>–1.91304273 ± 0.00000045</b>	MOHR	05	RVUE 2002 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
–1.91304272 ± 0.00000045	MOHR	99	RVUE 1998 CODATA value
–1.91304275 ± 0.00000045	COHEN	87	RVUE 1986 CODATA value
–1.91304277 ± 0.00000048	<sup>12</sup> GREENE	82	MRS

<sup>12</sup> GREENE 82 measures the moment to be  $(1.04187564 \pm 0.00000026) \times 10^{-3}$  Bohr magnetons. The value above is obtained by multiplying this by  $m_p/m_e = 1836.152701 \pm 0.000037$  (the 1986 CODATA value from COHEN 87).

## *n* ELECTRIC DIPOLE MOMENT

A nonzero value is forbidden by both *T* invariance and *P* invariance.

A number of early results have been omitted. See RAMSEY 90 and GOLUB 94 for reviews.

VALUE ( $10^{-25}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
< 0.63	90	<sup>13</sup> HARRIS	99	MRS $d = (-0.1 \pm 0.36) \times 10^{-25}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.97	90	ALTAREV	96	MRS $(+0.26 \pm 0.40 \pm 0.16) \times 10^{-25}$
< 1.1	95	ALTAREV	92	MRS See ALTAREV 96
< 1.2	95	SMITH	90	MRS See HARRIS 99
< 2.6	95	ALTAREV	86	MRS $d = (-1.4 \pm 0.6) \times 10^{-25}$
0.3 ± 4.8		PENDLEBURY	84	MRS Ultracold neutrons
< 6	90	ALTAREV	81	MRS $d = (2.1 \pm 2.4) \times 10^{-25}$
< 16	90	ALTAREV	79	MRS $d = (4.0 \pm 7.5) \times 10^{-25}$

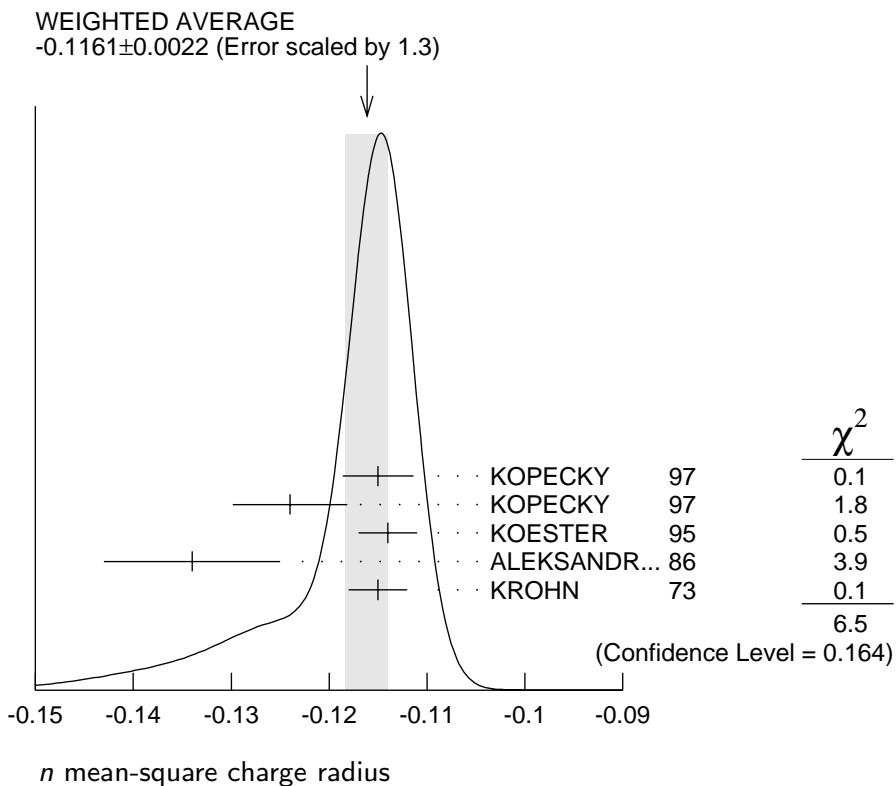
<sup>13</sup>This HARRIS 99 result includes the result of SMITH 90. However, the averaging of the results of these two experiments has been criticized by LAMOREAUX 00.

## *n* MEAN-SQUARE CHARGE RADIUS

The mean-square charge radius of the neutron,  $\langle r_n^2 \rangle$ , is related to the neutron-electron scattering length  $b_{ne}$  by  $\langle r_n^2 \rangle = 3(m_e a_0 / m_n) b_{ne}$ , where  $m_e$  and  $m_n$  are the masses of the electron and neutron, and  $a_0$  is the Bohr radius. Numerically,  $\langle r_n^2 \rangle = 86.34 b_{ne}$ , if we use  $a_0$  for a nucleus with infinite mass.

VALUE (fm <sup>2</sup> )	DOCUMENT ID	COMMENT
<b>-0.1161±0.0022 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.
-0.115 ± 0.002	KOPECKY 97	<i>ne</i> scattering (Pb)
-0.124 ± 0.003	KOPECKY 97	<i>ne</i> scattering (Bi)
-0.114 ± 0.003	KOESTER 95	<i>ne</i> scattering (Pb, Bi)
-0.134 ± 0.009	ALEKSANDR... 86	<i>ne</i> scattering (Bi)
-0.115 ± 0.003	<sup>14</sup> KROHN 73	<i>ne</i> scattering (Ne, Ar, Kr, Xe)
• • • We do not use the following data for averages, fits, limits, etc. • • •		
-0.113 ± 0.003	KOPECKY 95	<i>ne</i> scattering (Pb)
-0.114 ± 0.003	KOESTER 86	<i>ne</i> scattering (Pb, Bi)
-0.118 ± 0.002	KOESTER 76	<i>ne</i> scattering (Pb)
-0.120 ± 0.002	KOESTER 76	<i>ne</i> scattering (Bi)
-0.116 ± 0.003	KROHN 66	<i>ne</i> scattering (Ne, Ar, Kr, Xe)

<sup>14</sup>This value is as corrected by KOESTER 76.



## ***n* ELECTRIC POLARIZABILITY $\alpha_n$**

Following is the electric polarizability  $\alpha_n$  defined in terms of the induced electric dipole moment by  $\mathbf{D} = 4\pi\epsilon_0\alpha_n\mathbf{E}$ . For a review, see SCHMIED-MAYER 89.

VALUE ( $10^{-4}$ fm $^3$ )	DOCUMENT ID	TECN	COMMENT
<b><i>11.6± 1.5 OUR AVERAGE</i></b>			
$12.5 \pm 1.8^{+1.6}_{-1.3}$	15 KOSSERT	03 CNTR	$\gamma d \rightarrow \gamma pn$
$8.8 \pm 2.4 \pm 3.0$	16 LUNDIN	03 CNTR	$\gamma d \rightarrow \gamma d$
$12.0 \pm 1.5 \pm 2.0$	SCHMIEDM...	91 CNTR	<i>n</i> Pb transmission
$10.7^{+3.3}_{-10.7}$	ROSE	90B CNTR	$\gamma d \rightarrow \gamma np$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
13.6	17 KOLB	00 CNTR	$\gamma d \rightarrow \gamma np$
$0.0 \pm 5.0$	18 KOESTER	95 CNTR	<i>n</i> Pb, <i>n</i> Bi transmission
$11.7^{+4.3}_{-11.7}$	ROSE	90 CNTR	See ROSE 90B
$8 \pm 10$	KOESTER	88 CNTR	<i>n</i> Pb, <i>n</i> Bi transmission
12 ± 10	SCHMIEDM...	88 CNTR	<i>n</i> Pb, <i>n</i> C transmission
15 KOSSERT 03 gets $\alpha_n - \beta_n = (9.8 \pm 3.6^{+2.1}_{-1.1} \pm 2.2) \times 10^{-4}$ fm $^3$ , and uses $\alpha_n + \beta_n = (15.2 \pm 0.5) \times 10^{-4}$ fm $^3$ from LEVCHUK 00. Thus the errors on $\alpha_n$ and $\beta_n$ are anti-correlated.			
16 LUNDIN 03 measures $\alpha_N - \beta_N = (6.4 \pm 2.4) \times 10^{-4}$ fm $^3$ and uses accurate values for $\alpha_p$ and $\alpha_p$ and a precise sum-rule result for $\alpha_n + \beta_n$ . The second error is a model uncertainty, and errors on $\alpha_n$ and $\beta_n$ are anticorrelated.			
17 KOLB 00 obtains this value with a lower limit of $7.6 \times 10^{-4}$ fm $^3$ but no upper limit from this experiment alone. Combined with results of ROSE 90, the 1- $\sigma$ range is $(7.6\text{--}14.0) \times 10^{-4}$ fm $^3$ .			
18 KOESTER 95 uses natural Pb and the isotopes 208, 207, and 206. See this paper for a discussion of methods used by various groups to extract $\alpha_n$ from data.			

## ***n* MAGNETIC POLARIZABILITY $\beta_n$**

VALUE ( $10^{-4}$ fm $^3$ )	DOCUMENT ID	TECN	COMMENT
<b><i>3.7±2.0 OUR AVERAGE</i></b>			
$2.7 \pm 1.8^{+1.3}_{-1.6}$	19 KOSSERT	03 CNTR	$\gamma d \rightarrow \gamma pn$
$6.5 \pm 2.4 \pm 3.0$	20 LUNDIN	03 CNTR	$\gamma d \rightarrow \gamma d$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.6	21 KOLB	00 CNTR	$\gamma d \rightarrow \gamma np$
19 KOSSERT 03 gets $\alpha_n - \beta_n = (9.8 \pm 3.6^{+2.1}_{-1.1} \pm 2.2) \times 10^{-4}$ fm $^3$ , and uses $\alpha_n + \beta_n = (15.2 \pm 0.5) \times 10^{-4}$ fm $^3$ from LEVCHUK 00. Thus the errors on $\alpha_n$ and $\beta_n$ are anti-correlated.			
20 LUNDIN 03 measures $\alpha_N - \beta_N = (6.4 \pm 2.4) \times 10^{-4}$ fm $^3$ and uses accurate values for $\alpha_p$ and $\alpha_p$ and a precise sum-rule result for $\alpha_n + \beta_n$ . The second error is a model uncertainty, and errors on $\alpha_n$ and $\beta_n$ are anticorrelated.			
21 KOLB 00 obtains this value with an upper limit of $7.6 \times 10^{-4}$ fm $^3$ but no lower limit from this experiment alone. Combined with results of ROSE 90, the 1- $\sigma$ range is $(1.2\text{--}7.6) \times 10^{-4}$ fm $^3$ .			

## *n* CHARGE

See also “ $|q_p + q_e|/e$ ” in the proton Listings.

VALUE ( $10^{-21} e$ )		DOCUMENT ID	TECN	COMMENT
= <b>0.4 ± 1.1</b>	22	BAUMANN 88		Cold <i>n</i> deflection
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-15 ± 22	23	GAEHLER 82	CNTR	Cold <i>n</i> deflection
22 The BAUMANN 88 error ±1.1 gives the 68% CL limits about the the value -0.4.				
23 The GAEHLER 82 error ±22 gives the 90% CL limits about the the value -15.				

## LIMIT ON *n*–*n* OSCILLATIONS

### Mean Time for *n*–*n* Transition in Vacuum

A test of  $\Delta B=2$  baryon number nonconservation. MOHAPATRA 80 and MOHAPATRA 89 discuss the theoretical motivations for looking for *n*–*n* oscillations. DOVER 83 and DOVER 85 give phenomenological analyses. The best limits come from looking for the decay of neutrons bound in nuclei. However, these analyses require model-dependent corrections for nuclear effects. See KABIR 83, DOVER 89, ALBERICO 91, and GAL 00 for discussions. Direct searches for  $n \rightarrow \bar{n}$  transitions using reactor neutrons are cleaner but give somewhat poorer limits. We include limits for both free and bound neutrons in the Summary Table.

VALUE (s)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt;1.3 × 10<sup>8</sup></b>	90	CHUNG	02B SOU2	<i>n</i> bound in iron
<b>&gt;8.6 × 10<sup>7</sup></b>	90	BALDO-...	94 CNTR	Reactor (free) neutrons
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>1 × 10 <sup>7</sup>	90	BALDO-...	90 CNTR	See BALDO-CEOLIN 94
>1.2 × 10 <sup>8</sup>	90	BERGER	90 FREJ	<i>n</i> bound in iron
>4.9 × 10 <sup>5</sup>	90	BRESSI	90 CNTR	Reactor neutrons
>4.7 × 10 <sup>5</sup>	90	BRESSI	89 CNTR	See BRESSI 90
>1.2 × 10 <sup>8</sup>	90	TAKITA	86 CNTR	<i>n</i> bound in oxygen
>1 × 10 <sup>6</sup>	90	FIDECARO	85 CNTR	Reactor neutrons
>8.8 × 10 <sup>7</sup>	90	PARK	85B CNTR	
>3 × 10 <sup>7</sup>		BATTISTONI	84 NUSX	
>2.7 × 10 <sup>7</sup> –1.1 × 10 <sup>8</sup>		JONES	84 CNTR	
>2 × 10 <sup>7</sup>		CHERRY	83 CNTR	

## *n* DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Confidence level
$\Gamma_1 p e^- \bar{\nu}_e$	100 %	
$\Gamma_2$ hydrogen-atom $\bar{\nu}_e$		
$\Gamma_3 p e^- \bar{\nu}_e \gamma$	[a] < $6.9 \times 10^{-3}$	90%
<b>Charge conservation (<i>Q</i>) violating mode</b>		
$\Gamma_4 p \nu_e \bar{\nu}_e$	<i>Q</i> < $8 \times 10^{-27}$	68%

[a] This limit is for  $\gamma$  energies between 35 and 100 keV.

## *n* BRANCHING RATIOS

### $\Gamma(\text{hydrogen-atom } \bar{\nu}_e)/\Gamma_{\text{total}}$

$\Gamma_2/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$<3 \times 10^{-2}$	95	24 GREEN	90 RVUE	
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<sup>24</sup> GREEN 90 infers that  $\tau(\text{hydrogen-atom } \bar{\nu}_e) > 3 \times 10^4$  s by comparing neutron lifetime measurements made in storage experiments with those made in  $\beta$ -decay experiments. However, the result depends sensitively on the lifetime measurements, and does not of course take into account more recent measurements of same.

### $\Gamma(p e^- \bar{\nu}_e \gamma)/\Gamma_{\text{total}}$

$\Gamma_3/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<6.9 \times 10^{-3}$	90	25 BECK	02 CNTR	Cold <i>n</i>	

<sup>25</sup> This BECK 02 limit is for  $\gamma$  energies between 35 and 100 keV.

### $\Gamma(p \nu_e \bar{\nu}_e)/\Gamma_{\text{total}}$

$\Gamma_4/\Gamma$

Forbidden by charge conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<8 \times 10^{-27}$	68	26 NORMAN	96 RVUE	${}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge}$ neutrals	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<9.7 \times 10^{-18}$	90	ROY	83	CNTR	${}^{113}\text{Cd} \rightarrow {}^{113m}\text{In}$ neut.
$<7.9 \times 10^{-21}$		VAIDYA	83	CNTR	${}^{87}\text{Rb} \rightarrow {}^{87m}\text{Sr}$ neut.
$<9 \times 10^{-24}$	90	BARABANOV	80	CNTR	${}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge}$ X
$<3 \times 10^{-19}$		NORMAN	79	CNTR	${}^{87}\text{Rb} \rightarrow {}^{87m}\text{Sr}$ neut.

<sup>26</sup> NORMAN 96 gets this limit by attributing SAGE and GALLEX counting rates to the charge-nonconserving transition  ${}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge}$ +neutrals rather than to solar-neutrino reactions.

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## *n* $\rightarrow$ *p* $e^- \bar{\nu}_e$ DECAY PARAMETERS

See the above “Note on Baryon Decay Parameters.” For discussions of recent results, see the references cited at the beginning of the section on the neutron mean life. For discussions of the values of the weak coupling constants  $g_A$  and  $g_V$  obtained using the neutron lifetime and asymmetry parameter  $A$ , comparisons with other methods of obtaining these constants, and implications for particle physics and for astrophysics, see DUBBERS 91 and WOOLCOCK 91. For tests of the  $V-A$  theory of neutron decay, see EROZOLIMSKII 91B and MOSTOVOI 96.

### $\lambda \equiv g_A / g_V$

VALUE	DOCUMENT ID	TECN	COMMENT	
<b><math>-1.2695 \pm 0.0029</math> OUR AVERAGE</b>			Error includes scale factor of 2.0. See the ideogram below.	
$-1.2739 \pm 0.0019$	27 ABELE	02 SPEC	Cold <i>n</i> , polarized, A	
$-1.2686 \pm 0.0046 \pm 0.007$	28 MOSTOVOI	01 CNTR	A and <i>B</i> $\times$ polarizations	
$-1.266 \pm 0.004$	LIAUD	97 TPC	Cold <i>n</i> , polarized, A	
$-1.2594 \pm 0.0038$	29 YEROZLIM...	97 CNTR	Cold <i>n</i> , polarized, A	
$-1.262 \pm 0.005$	BOPP	86 SPEC	Cold <i>n</i> , polarized, A	

• • • We do not use the following data for averages, fits, limits, etc. • • •

-1.274 ± 0.003	ABELE	97D	SPEC	Cold $n$ , polarized, A
-1.266 ± 0.004	SCHRECK...	95	TPC	See LIAUD 97
-1.2544 ± 0.0036	EROZOLIM...	91	CNTR	See YEROZOLIM-SKY 97
-1.226 ± 0.042	MOSTOVOY	83	RVUE	
-1.261 ± 0.012	EROZOLIM...	79	CNTR	Cold $n$ , polarized, A
-1.259 ± 0.017	<sup>30</sup> STRATOWA	78	CNTR	$p$ recoil spectrum, $a$
-1.263 ± 0.015	EROZOLIM...	77	CNTR	See EROZOLIMSKII 79
-1.250 ± 0.036	<sup>30</sup> DOBROZE...	75	CNTR	See STRATOWA 78
-1.258 ± 0.015	<sup>31</sup> KROHN	75	CNTR	Cold $n$ , polarized, A
-1.263 ± 0.016	<sup>32</sup> KROPF	74	RVUE	$n$ decay alone
-1.250 ± 0.009	<sup>32</sup> KROPF	74	RVUE	$n$ decay + nuclear ft

<sup>27</sup>This is the combined result of ABELE 02 and ABELE 97D.

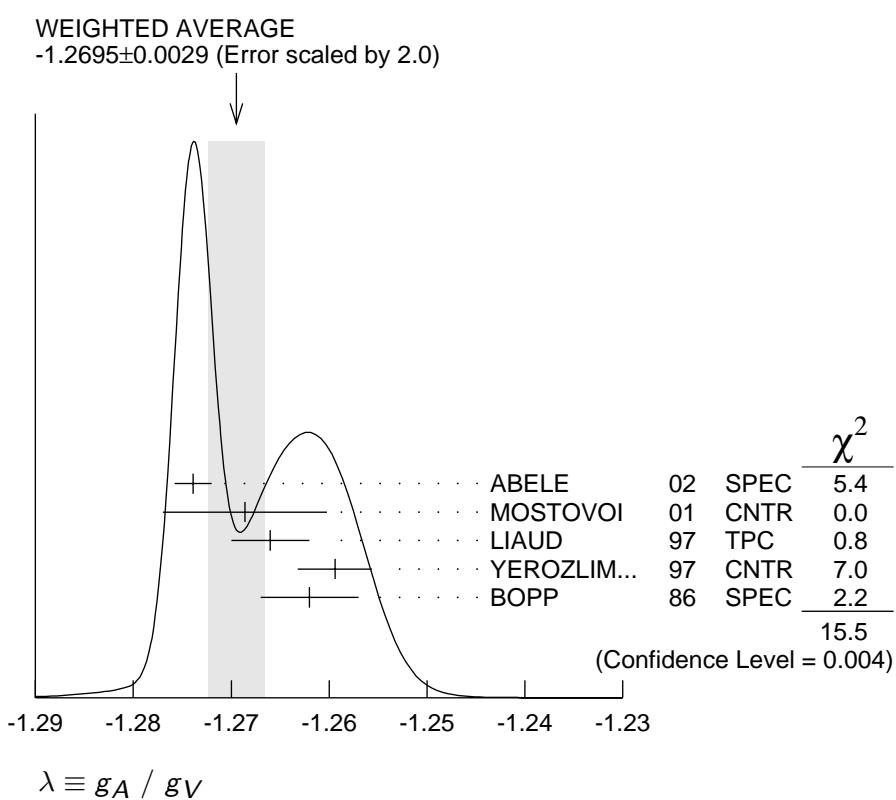
<sup>28</sup>MOSTOVOI 01 measures the two  $P$ -odd correlations  $A$  and  $B$ , or rather  $SA$  and  $SB$ , where  $S$  is the  $n$  polarization, in free neutron decay.

<sup>29</sup>YEROZOLIMSKY 97 makes a correction to the EROZOLIMSKII 91 value.

<sup>30</sup>These experiments measure the absolute value of  $g_A/g_V$  only.

<sup>31</sup>KROHN 75 includes events of CHRISTENSEN 70.

<sup>32</sup>KROPF 74 reviews all data through 1972.



## $\beta$ ASYMMETRY PARAMETER A

This is the neutron-spin electron-momentum correlation coefficient. Unless otherwise noted, the values are corrected for radiative effects and weak magnetism.

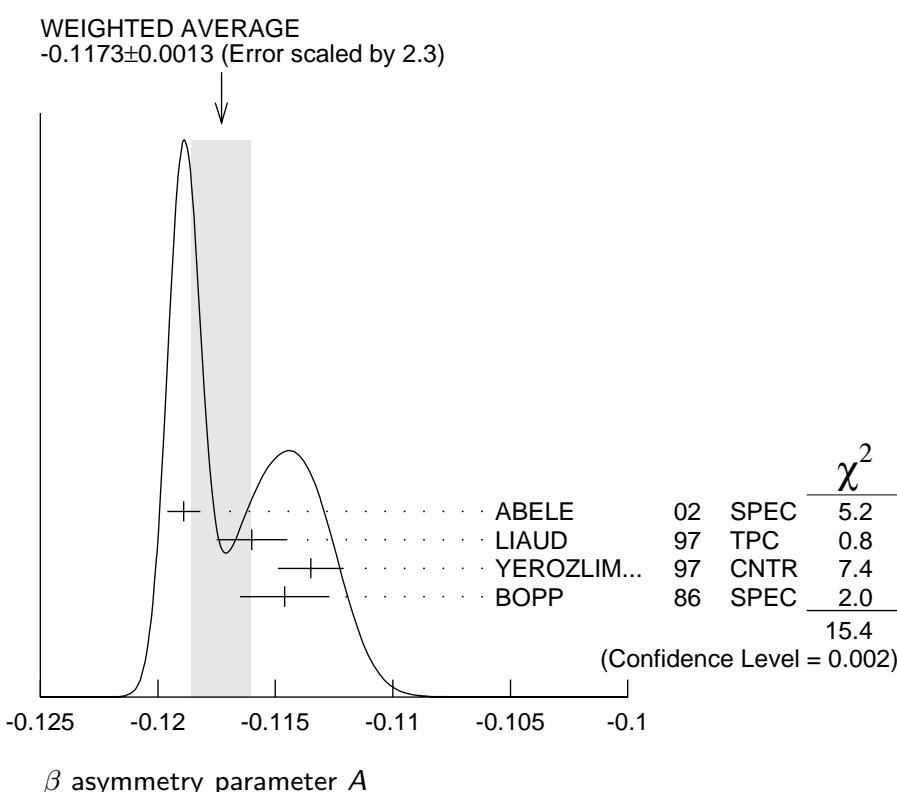
VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.1173±0.0013 OUR AVERAGE</b>	Error includes scale factor of 2.3. See the ideogram below.		
-0.1189±0.0007	<sup>33</sup> ABELE	02 SPEC	Cold $n$ , polarized
-0.1160±0.0009±0.0012	LIAUD	97 TPC	Cold $n$ , polarized
-0.1135±0.0014	<sup>34</sup> YEROZLIM...	97 CNTR	Cold $n$ , polarized
-0.1146±0.0019	BOPP	86 SPEC	Cold $n$ , polarized
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.1168±0.0017	<sup>35</sup> MOSTOVOI	01 CNTR	Inferred
-0.1189±0.0012	ABELE	97D SPEC	Cold $n$ , polarized
-0.1160±0.0009±0.0011	SCHRECK...	95 TPC	See LIAUD 97
-0.1116±0.0014	EROZOLIM...	91 CNTR	See YEROZOLIM-SKY 97
-0.114 ± 0.005	<sup>36</sup> EROZOLIM...	79 CNTR	Cold $n$ , polarized
-0.113 ± 0.006	KROHN	75 CNTR	Cold $n$ , polarized

<sup>33</sup> This is the combined result of ABELE 02 and ABELE 97D.

<sup>34</sup> YEROZOLIMSKY 97 makes a correction to the EROZOLIMSKII 91 value.

<sup>35</sup> MOSTOVOI 01 calculates this from its measurement of  $\lambda=g_A/g_V$  above.

<sup>36</sup> These results are not corrected for radiative effects and weak magnetism, but the corrections are small compared to the errors.



## $\bar{\nu}_e$ ASYMMETRY PARAMETER $B$

This is the neutron-spin antineutrino-momentum correlation coefficient.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.983 ± 0.004 OUR AVERAGE</b>			
0.9801 ± 0.0046	SEREBROV 98	CNTR	Cold $n$ , polarized
0.9894 ± 0.0083	KUZNETSOV 95	CNTR	Cold $n$ , polarized
0.995 ± 0.034	CHRISTENSEN70	CNTR	Cold $n$ , polarized
1.00 ± 0.05	EROZOLIM...	70C	CNTR Cold $n$ , polarized
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.9876 ± 0.0004	<sup>37</sup> MOSTOVOI 01	CNTR	Inferred
<sup>37</sup> MOSTOVOI 01 calculates this from its measurement of $\lambda = g_A/g_V$ above.			

## $e\bar{\nu}_e$ ANGULAR CORRELATION COEFFICIENT $a$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>-0.103 ± 0.004 OUR AVERAGE</b>			
-0.1054 ± 0.0055	BYRNE 02	SPEC	Proton recoil spectrum
-0.1017 ± 0.0051	STRATOWA 78	CNTR	Proton recoil spectrum
-0.091 ± 0.039	GRIGOREV 68	SPEC	Proton recoil spectrum
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-0.1045 ± 0.0014	<sup>38</sup> MOSTOVOI 01	CNTR	Inferred
<sup>38</sup> MOSTOVOI 01 calculates this from its measurement of $\lambda = g_A/g_V$ above.			

## $\phi_{AV}$ , PHASE OF $g_A$ RELATIVE TO $g_V$

Time reversal invariance requires this to be 0 or  $180^\circ$ . This is related to  $D$  given in the next data block and  $\lambda \equiv g_A/g_V$  by  $\sin(\phi_{AV}) = D(1+3\lambda^2)/2\lambda$ .

VALUE ( $^\circ$ )	DOCUMENT ID	TECN	COMMENT
<b>180.06 ± 0.07 OUR AVERAGE</b>			
180.04 ± 0.09	SOLDNER 04	CNTR	Cold $n$ , polarized
180.08 ± 0.13	LISING 00	CNTR	Polarized >93%
179.71 ± 0.39	EROZOLIM...	78	CNTR Cold $n$ , polarized
180.35 ± 0.43	EROZOLIM...	74	CNTR Cold $n$ , polarized
180.14 ± 0.22	STEINBERG 74	CNTR	Cold $n$ , polarized
• • • We do not use the following data for averages, fits, limits, etc. • • •			
181.1 ± 1.3	<sup>39</sup> KROPF 74	RVUE	$n$ decay
<sup>39</sup> KROPF 74 reviews all data through 1972.			

## TRIPLE CORRELATION COEFFICIENT $D$

These are measurements of the component of  $n$  spin perpendicular to the decay plane in  $\beta$  decay. Should be zero if  $T$  invariance is not violated.

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>- 4 ± 6 OUR AVERAGE</b>			
- 2.8 ± 6.4 ± 3.0	SOLDNER 04	CNTR	Cold $n$ , polarized
- 6 ± 12 ± 5	LISING 00	CNTR	Polarized >93%
+ 22 ± 30	EROZOLIM...	78	CNTR Cold $n$ , polarized
- 27 ± 50	<sup>40</sup> EROZOLIM...	74	CNTR Cold $n$ , polarized
- 11 ± 17	STEINBERG 74	CNTR	Cold $n$ , polarized
<sup>40</sup> EROZOLIMSKII 78 says asymmetric proton losses and nonuniform beam polarization may give a systematic error up to $30 \times 10^{-4}$ , thus increasing the EROZOLIMSKII 74 error to $50 \times 10^{-4}$ . STEINBERG 74 and STEINBERG 76 estimate these systematic errors to be insignificant in their experiment.			

## n REFERENCES

We have omitted some papers that have been superseded by later experiments. See our earlier editions.

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SOLDNER	04	PL B581 49	T. Soldner <i>et al.</i>	(ILLG, MUNT)
DEWEY	03	PRL 91 152302	M.S. Dewey <i>et al.</i>	(NIST, TULN, IND+)
KOSSERT	03	EPJ A16 259	K. Kosser <i>et al.</i>	(Mainz MAMI Collab.)
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LUNDIN	03	PRL 90 192501	M. Lundin <i>et al.</i>	
ABELE	02	PRL 88 211801	H. Abele <i>et al.</i>	(PERKEO-II Collab.)
BECK	02	JETPL 76 332	M. Beck <i>et al.</i>	(LEUV, SUSS, KIAE, PNPI)
		Translated from ZETFP 76 392.		
BYRNE	02	JPG 28 1325	J. Byrne <i>et al.</i>	
CHUNG	02B	PR D66 032004	J. Chung <i>et al.</i>	(SOUDAN-2 Collab.)
MOSTOVOI	01	PAN 64 1955	Yu.A. Mostovoi <i>et al.</i>	
		Translated from YAF 64 2040.		
ARZUMANOV	00	PL B483 15	S. Arzumanov <i>et al.</i>	
GAL	00	PR C61 028201	A. Gal	
KOLB	00	PRL 85 1388	N.R. Kolb <i>et al.</i>	
LAMOREAUX	00	PR D61 051301R	S.K. Lamoreaux, R. Golub	
LEVCHUK	00	NP A674 449	M.I. Levchuk, A.I. L'vov	(BELA, LEBD)
LISING	00	PR C62 055501	L.J. Lising <i>et al.</i>	(NIST emiT Collab.)
HARRIS	99	PRL 82 904	P.G. Harris <i>et al.</i>	
KESSLER	99	PL A255 221	E.G. Kessler Jr <i>et al.</i>	
MOHR	99	JPCRD 28 1713	P.J. Mohr, B.N. Taylor	(NIST)
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SERE BROV	98	JETP 86 1074	A.P. Serebrov <i>et al.</i>	
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LIAUD	97	NP A612 53	P. Liaud <i>et al.</i>	(ILLG, LAPP)
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MOSTOVOI	96	PAN 59 968	Y.A. Mostovoy	(KIAE)
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NORMAN	96	PR D53 4086	E.B. Norman, J.N. Bahcall, M. Goldhaber	(LBL+)
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DOBROZE...	75	PR D11 510	R. Dobrozemsky <i>et al.</i>	(SEIB)
KROHN	75	PL 55B 175	V.E. Krohn, G.R. Ringo	(ANL)
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COHEN	73	JPCRD 2 664	E.R. Cohen, B.N. Taylor	(RISC, NBS)
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GRIGOREV	68	SJNP 6 239	V.K. Grigoriev <i>et al.</i>	(ITEP)
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KROHN	66	PR 148 1303	V.E. Krohn, G.R. Ringo	