

# SEARCHES FOR MONOPOLES, SUPERSYMMETRY, TECHNICOLOR, COMPOSITENESS, EXTRA DIMENSIONS, etc.

## Magnetic Monopole Searches

Isolated supermassive monopole candidate events have not been confirmed. The most sensitive experiments obtain negative results.

Best cosmic-ray supermassive monopole flux limit:

$$< 1.4 \times 10^{-16} \text{ cm}^{-2}\text{sr}^{-1}\text{s}^{-1} \quad \text{for } 1.1 \times 10^{-4} < \beta < 1$$

## Supersymmetric Particle Searches

All supersymmetric mass bounds here are model dependent.

The limits assume:

1)  $\tilde{\chi}_1^0$  is the lightest supersymmetric particle; 2)  $R$ -parity is conserved;

See the Particle Listings for a Note giving details of supersymmetry.

$\tilde{\chi}_i^0$  — neutralinos (mixtures of  $\tilde{\gamma}$ ,  $\tilde{Z}^0$ , and  $\tilde{H}_i^0$ )

Mass  $m_{\tilde{\chi}_1^0} > 0$  GeV, CL = 95%

[general MSSM, non-universal gaugino masses]

Mass  $m_{\tilde{\chi}_1^0} > 46$  GeV, CL = 95%

[all  $\tan\beta$ , all  $m_0$ , all  $m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ ]

Mass  $m_{\tilde{\chi}_2^0} > 620$  GeV, CL = 95%

$[\tilde{\chi}_2^0 \rightarrow \ell^\pm \ell^\mp \tilde{\chi}_1^0]$ , simplified model,  $m_{\tilde{\chi}_1^0} = 0$  GeV

Mass  $m_{\tilde{\chi}_3^0} > 620$  GeV, CL = 95%

$[\tilde{\chi}_3^0 \rightarrow \ell^\pm \ell^\mp \tilde{\chi}_1^0]$ , simplified model,  $m_{\tilde{\chi}_1^0} = 0$  GeV

Mass  $m_{\tilde{\chi}_4^0} > 116$  GeV, CL = 95%

[ $1 < \tan\beta < 40$ , all  $m_0$ , all  $m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$ ]

$\tilde{\chi}_i^\pm$  — charginos (mixtures of  $\widetilde{W}^\pm$  and  $\widetilde{H}_i^\pm$ )

Mass  $m_{\tilde{\chi}_1^\pm} > 94$  GeV, CL = 95%

[ $\tan\beta < 40$ ,  $m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0} > 3$  GeV, all  $m_0$ ]

Mass  $m_{\tilde{\chi}_1^\pm} > 500$  GeV, CL = 95%

[simplified model,  $2\ell^\pm + \cancel{E}_T$ ,  $m_{\tilde{\chi}_1^0} = 0$  GeV]

$\tilde{\chi}^\pm$  — long-lived chargino

Mass  $m_{\tilde{\chi}^\pm} > 620$  GeV, CL = 95% [stable  $\tilde{\chi}^\pm$ ]

$\tilde{\nu}$  — sneutrino

Mass  $m > 41$  GeV, CL = 95% [model independent]

Mass  $m > 94$  GeV, CL = 95%

[CMSSM,  $1 \leq \tan\beta \leq 40$ ,  $m_{\tilde{e}_R} - m_{\tilde{\chi}_1^0} > 10$  GeV]

$\tilde{e}$  — scalar electron (selectron)

Mass  $m(\tilde{e}_L) > 107$  GeV, CL = 95% [all  $m_{\tilde{e}_L} - m_{\tilde{\chi}_1^0}$ ]

$\tilde{\mu}$  — scalar muon (smuon)

Mass  $m > 94$  GeV, CL = 95%

[CMSSM,  $1 \leq \tan\beta \leq 40$ ,  $m_{\tilde{\mu}_R} - m_{\tilde{\chi}_1^0} > 10$  GeV]

$\tilde{\tau}$  — scalar tau (stau)

Mass  $m > 81.9$  GeV, CL = 95%

[ $m_{\tilde{\tau}_R} - m_{\tilde{\chi}_1^0} > 15$  GeV, all  $\theta_\tau$ ,  $B(\tilde{\tau} \rightarrow \tau \tilde{\chi}_1^0) = 100\%$ ]

Mass  $m > 286$  GeV, CL = 95% [long-lived  $\tilde{\tau}$ ]

Where the limits below show a **range** of lower bounds, the bounds depend on different simplified models, different signals, different assumptions, and different luminosities.

$\tilde{q}$  – squarks of the first two quark generations

The first of these limits is within CMSSM with cascade decays, evaluated assuming a fixed value of the parameters  $\mu$  and  $\tan\beta$ , and assuming two-generations of mass degenerate squarks ( $\tilde{q}_L$  and  $\tilde{q}_R$ ) and gaugino mass parameters that are constrained by the unification condition at the grand unification scale.

Mass  $m > 1450$  GeV, CL = 95%

[CMSSM,  $\tan\beta = 30$ ,  $A_0 = -2\max(m_0, m_{1/2})$ ,  $\mu > 0$ ]

Mass  $m > 608$ – $1260$  GeV, CL = 95%

[mass degenerate squarks]

Mass  $m > 490$ – $600$  GeV, CL = 95%

[single light squark bounds]

$\tilde{q}$  — long-lived squark

Mass  $m > 1000$ , CL = 95%

[ $\tilde{t}$ , charge-suppressed interaction model]

Mass  $m > 845$ , CL = 95% [ $\tilde{b}$ , stable, Regge model]

$\tilde{b}$  — scalar bottom (sbottom)

Mass  $m > 323\text{--}880$  GeV, CL = 95%

[There is dependence on mass difference  $\tilde{b}$ -LSP]

$\tilde{t}$  — scalar top (stop)

Mass  $m > 323\text{--}800$  GeV, CL = 95%

[Lower value is a decay via charm, and upper a decay via top]

$\tilde{g}$  — gluino

Mass  $m > 700\text{--}1780$  GeV, CL = 95%

## Technicolor

The limits for technicolor (and top-color) particles are quite varied depending on assumptions. See the Technicolor section of the full *Review* (the data listings).

## Quark and Lepton Compositeness, Searches for

### Scale Limits $\Lambda$ for Contact Interactions (the lowest dimensional interactions with four fermions)

If the Lagrangian has the form

$$\pm \frac{g^2}{2\Lambda^2} \bar{\psi}_L \gamma_\mu \psi_L \bar{\psi}_L \gamma^\mu \psi_L$$

(with  $g^2/4\pi$  set equal to 1), then we define  $\Lambda \equiv \Lambda_{LL}^\pm$ . For the full definitions and for other forms, see the Note in the Listings on Searches for Quark and Lepton Compositeness in the full *Review* and the original literature.

$$\Lambda_{LL}^+(eeee) > 8.3 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^-(eeee) > 10.3 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^+(ee\mu\mu) > 8.5 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^-(ee\mu\mu) > 9.5 \text{ TeV, CL} = 95\%$$

$$\Lambda_{LL}^+(ee\tau\tau) > 7.9 \text{ TeV, CL} = 95\%$$

|                                    |                        |
|------------------------------------|------------------------|
| $\Lambda_{LL}^-(ee\tau\tau)$       | $> 7.2$ TeV, CL = 95%  |
| $\Lambda_{LL}^+(\ell\ell\ell\ell)$ | $> 9.1$ TeV, CL = 95%  |
| $\Lambda_{LL}^-(\ell\ell\ell\ell)$ | $> 10.3$ TeV, CL = 95% |
| $\Lambda_{LL}^+(eeqq)$             | $> 16.4$ TeV, CL = 95% |
| $\Lambda_{LL}^-(eeqq)$             | $> 20.7$ TeV, CL = 95% |
| $\Lambda_{LL}^+(eeuu)$             | $> 23.3$ TeV, CL = 95% |
| $\Lambda_{LL}^-(eeuu)$             | $> 12.5$ TeV, CL = 95% |
| $\Lambda_{LL}^+(eedd)$             | $> 11.1$ TeV, CL = 95% |
| $\Lambda_{LL}^-(eedd)$             | $> 26.4$ TeV, CL = 95% |
| $\Lambda_{LL}^+(eccc)$             | $> 9.4$ TeV, CL = 95%  |
| $\Lambda_{LL}^-(eccc)$             | $> 5.6$ TeV, CL = 95%  |
| $\Lambda_{LL}^+(eebb)$             | $> 9.4$ TeV, CL = 95%  |
| $\Lambda_{LL}^-(eebb)$             | $> 10.2$ TeV, CL = 95% |
| $\Lambda_{LL}^+(\mu\mu qq)$        | $> 15.8$ TeV, CL = 95% |
| $\Lambda_{LL}^-(\mu\mu qq)$        | $> 21.8$ TeV, CL = 95% |
| $\Lambda(\ell\nu\ell\nu)$          | $> 3.10$ TeV, CL = 90% |
| $\Lambda(e\nu qq)$                 | $> 2.81$ TeV, CL = 95% |
| $\Lambda_{LL}^+(qqqq)$             | $> 12.0$ TeV, CL = 95% |
| $\Lambda_{LL}^-(qqqq)$             | $> 17.5$ TeV, CL = 95% |
| $\Lambda_{LL}^+(\nu\nu qq)$        | $> 5.0$ TeV, CL = 95%  |
| $\Lambda_{LL}^-(\nu\nu qq)$        | $> 5.4$ TeV, CL = 95%  |

## Excited Leptons

The limits from  $\ell^{*+}\ell^{*-}$  do not depend on  $\lambda$  (where  $\lambda$  is the  $\ell\ell^*$  transition coupling). The  $\lambda$ -dependent limits assume chiral coupling.

$e^{*\pm}$  — excited electron

Mass  $m > 103.2$  GeV, CL = 95% (from  $e^*e^*$ )

Mass  $m > 3.000 \times 10^3$  GeV, CL = 95% (from  $ee^*$ )

Mass  $m > 356$  GeV, CL = 95% (if  $\lambda_\gamma = 1$ )

$\mu^{*\pm}$  — excited muon

Mass  $m > 103.2$  GeV, CL = 95% (from  $\mu^*\mu^*$ )

Mass  $m > 3.000 \times 10^3$  GeV, CL = 95% (from  $\mu\mu^*$ )

$\tau^{*\pm}$  — excited tauMass  $m > 103.2$  GeV, CL = 95% (from  $\tau^* \tau^*$ )Mass  $m > 2.500 \times 10^3$  GeV, CL = 95% (from  $\tau \tau^*$ ) $\nu^*$  — excited neutrinoMass  $m > 1.600 \times 10^3$  GeV, CL = 95% (from  $\nu^* \nu^*$ )Mass  $m > 213$  GeV, CL = 95% (from  $\nu^* X$ ) $q^*$  — excited quarkMass  $m > 338$  GeV, CL = 95% (from  $q^* q^*$ )Mass  $m > 5.200 \times 10^3$  GeV, CL = 95% (from  $q^* X$ )**Color Sextet and Octet Particles**Color Sextet Quarks ( $q_6$ )Mass  $m > 84$  GeV, CL = 95% (Stable  $q_6$ )Color Octet Charged Leptons ( $\ell_8$ )Mass  $m > 86$  GeV, CL = 95% (Stable  $\ell_8$ )Color Octet Neutrinos ( $\nu_8$ )Mass  $m > 110$  GeV, CL = 90% ( $\nu_8 \rightarrow \nu g$ )**Extra Dimensions**

Please refer to the Extra Dimensions section of the full *Review* for a discussion of the model-dependence of these bounds, and further constraints.

**Constraints on the radius of the extra dimensions, for the case of two-flat dimensions of equal radii** $R < 30 \mu\text{m}$ , CL = 95% (direct tests of Newton's law) $R < 10.9 \mu\text{m}$ , CL = 95% ( $pp \rightarrow jG$ ) $R < 0.16\text{--}916$  nm (astrophysics; limits depend on technique and assumptions)**Constraints on the fundamental gravity scale** $M_{TT} > 6.3$  TeV, CL = 95% ( $pp \rightarrow$  dijet, angular distribution) $M_c > 4.16$  TeV, CL = 95% ( $pp \rightarrow \ell\bar{\ell}$ )**Constraints on the Kaluza-Klein graviton in warped extra dimensions** $M_G > 3.3$  TeV, CL = 95% ( $pp \rightarrow e^+ e^-, \mu^+ \mu^-$ )**Constraints on the Kaluza-Klein gluon in warped extra dimensions** $M_{g_{KK}} > 2.5$  TeV, CL = 95% ( $g_{KK} \rightarrow t\bar{t}$ )