

**T**

$$J = \frac{1}{2}$$

$\tau$  discovery paper was PERL 75.  $e^+ e^- \rightarrow \tau^+ \tau^-$  cross-section threshold behavior and magnitude are consistent with pointlike spin-1/2 Dirac particle. BRANDELIK 78 ruled out pointlike spin-0 or spin-1 particle. FELDMAN 78 ruled out  $J = 3/2$ . KIRKBY 79 also ruled out  $J=\text{integer}$ ,  $J = 3/2$ .

NODE=S035

NODE=S035

<b><math>\tau</math> MASS</b>					
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1776.93±0.09 OUR AVERAGE</b>					
1777.09±0.08±0.11	175M	<sup>1</sup> ADACHI	23C BEL2	$190 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035M
$1776.69^{+0.17}_{-0.19} \pm 0.15$		<sup>2</sup> ANASHIN	23A KEDR	$(6.7 \pm 8.5) \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.78 \text{ GeV}$	NODE=S035M
$1776.91^{+0.12}_{-0.13} \pm 0.10$	1171	<sup>3</sup> ABLIKIM	14D BES3	$23.3 \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.60 \text{ GeV}$	
1776.68±0.12±0.41	682k	<sup>1</sup> AUBERT	09AK BABR	$423 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
1776.61±0.13±0.35		<sup>1</sup> BELOUS	07 BELL	$414 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$1775.1 \pm 1.6 \pm 1.0$	13.3k	<sup>4</sup> ABBIENDI	00A OPAL	1990–1995 LEP runs	
$1778.2 \pm 0.8 \pm 1.2$		ANASTASSOV 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$1776.96^{+0.18}_{-0.21} \pm 0.25$	65	<sup>5</sup> BAI	96 BES	$E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.57 \text{ GeV}$	
$1776.3 \pm 2.4 \pm 1.4$	11k	<sup>6</sup> ALBRECHT	92M ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
$1783^{+3}_{-4}$	692	<sup>7</sup> BACINO	78B DLCO	$E_{\text{cm}}^{\text{ee}} = 3.1\text{--}7.4 \text{ GeV}$	
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
$1776.81^{+0.25}_{-0.23} \pm 0.15$	81	ANASHIN	07 KEDR	$6.7 \text{ pb}^{-1}, E_{\text{cm}}^{\text{ee}} = 3.54\text{--}3.78 \text{ GeV}$	
$1777.8 \pm 0.7 \pm 1.7$	35k	<sup>8</sup> BALEST	93 CLEO	Repl. by ANASTASSOV 97	
$1776.9^{+0.4}_{-0.5} \pm 0.2$	14	<sup>9</sup> BAI	92 BES	Repl. by BAI 96	
1 ADACHI 23c, AUBERT 09AK and BELOUS 07 fit $\tau$ pseudomass spectrum in $\tau \rightarrow \pi\pi^+\pi^-\nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$ .					
2 Previously also reported LEVICHEV 14. Superseedes ANASHIN 07.					
3 ABLIKIM 14D fit $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$ at different energies near threshold.					
4 ABBIENDI 00A fit $\tau$ pseudomass spectrum in $\tau \rightarrow \pi^\pm \leq 2\pi^0\nu_\tau$ and $\tau \rightarrow \pi^\pm\pi^+\pi^- \leq 1\pi^0\nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$ .					
5 BAI 96 fit $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$ at different energies near threshold.					
6 ALBRECHT 92M fit $\tau$ pseudomass spectrum in $\tau^- \rightarrow 2\pi^-\pi^+\nu_\tau$ decays. Result assumes $m_{\nu_\tau} = 0$ .					
7 BACINO 78B value comes from $e^\pm X^\mp$ threshold. Published mass 1782 MeV increased by 1 MeV using the high precision $\psi(2S)$ mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.					
8 BALEST 93 fit spectra of minimum kinematically allowed $\tau$ mass in events of the type $e^+ e^- \rightarrow \tau^+ \tau^- \rightarrow (\pi^+ n\pi^0\nu_\tau)(\pi^- m\pi^0\nu_\tau)$ $n \leq 2, m \leq 2, 1 \leq n+m \leq 3$ . If $m_{\nu_\tau} \neq 0$ , result increases by $(m_{\nu_\tau}^2/1100 \text{ MeV})$ .					
9 BAI 92 fit $\sigma(e^+ e^- \rightarrow \tau^+ \tau^-)$ near threshold using $e\mu$ events.					

$$(m_{\tau^+} - m_{\tau^-})/m_{\text{average}}$$

A test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<2.8 \times 10^{-4}$	90	BELOUS	07 BELL	$414 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035MDF
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
$<5.5 \times 10^{-4}$	90	<sup>1</sup> AUBERT	09AK BABR	$423 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035MDF
$<3.0 \times 10^{-3}$	90	ABBIENDI	00A OPAL	1990–1995 LEP runs	NODE=S035MDF

<sup>1</sup>AUBERT 09AK quote both the listed upper limit and  $(m_{\tau^+} - m_{\tau^-})/m_{\text{average}} = (-3.4 \pm 1.3 \pm 0.3) \times 10^{-4}$ .

<b><math>\tau</math> MEAN LIFE</b>					
<u>VALUE</u> ( $10^{-15}$ s)	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b><math>290.3 \pm 0.5</math> OUR AVERAGE</b>					
290.17 $\pm$ 0.53 $\pm$ 0.33	1.1M	BELOUS	14	BELL	$711 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
290.9 $\pm$ 1.4 $\pm$ 1.0		ABDALLAH	04T	DLPH	1991–1995 LEP runs
293.2 $\pm$ 2.0 $\pm$ 1.5		ACCIARRI	00B	L3	1991–1995 LEP runs
290.1 $\pm$ 1.5 $\pm$ 1.1		BARATE	97R	ALEP	1989–1994 LEP runs
289.2 $\pm$ 1.7 $\pm$ 1.2		ALEXANDER	96E	OPAL	1990–1994 LEP runs
289.0 $\pm$ 2.8 $\pm$ 4.0	57.4k	BALEST	96	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
291.2 $\pm$ 2.0 $\pm$ 1.2		BARATE	97I	ALEP	Repl. by BARATE 97R
291.4 $\pm$ 3.0		ABREU	96B	DLPH	Repl. by ABDALLAH 04T
290.1 $\pm$ 4.0	34k	ACCIARRI	96K	L3	Repl. by ACCIARRI 00B
297 $\pm$ 9 $\pm$ 5	1671	ABE	95Y	SLD	1992–1993 SLC runs
304 $\pm$ 14 $\pm$ 7	4100	BATTLE	92	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
301 $\pm$ 29	3780	KLEINWORT	89	JADE	$E_{\text{cm}}^{\text{ee}} = 35\text{--}46 \text{ GeV}$
288 $\pm$ 16 $\pm$ 17	807	AMIDEI	88	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
306 $\pm$ 20 $\pm$ 14	695	BRAUNSCH...	88C	TASS	$E_{\text{cm}}^{\text{ee}} = 36 \text{ GeV}$
299 $\pm$ 15 $\pm$ 10	1311	ABACHI	87C	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
295 $\pm$ 14 $\pm$ 11	5696	ALBRECHT	87P	ARG	$E_{\text{cm}}^{\text{ee}} = 9.3\text{--}10.6 \text{ GeV}$
309 $\pm$ 17 $\pm$ 7	3788	BAND	87B	MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
325 $\pm$ 14 $\pm$ 18	8470	BEBEK	87C	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
460 $\pm$ 190	102	FELDMAN	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

### $(\tau_{\tau^+} - \tau_{\tau^-}) / \tau_{\text{average}}$

Test of *CPT* invariance.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 7.0 \times 10^{-3}$	90	1 BELOUS	14	BELL $711 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup>BELOUS 14 quote limit on the absolute value of the relative lifetime difference.

### $\tau$ MAGNETIC MOMENT ANOMALY

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

<u><math>\mu_{\tau}/(e\hbar/2m_{\tau}) - 1 = (g_{\tau} - 2)/2</math></u>				
For a theoretical calculation $[(g_{\tau} - 2)/2 = 117.721(5) \times 10^{-8}]$ , see EIDELMAN 07.				
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>-0.057 to 0.024</b>	95	1 AAD	23BMATLS	$\gamma\gamma \rightarrow \tau^+\tau^-$ , Pb–Pb
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
-0.041 $^{+0.012}_{-0.009}$		1,2 AAD	23BMATLS	$\gamma\gamma \rightarrow \tau^+\tau^-$ , Pb–Pb
0.001 $^{+0.055}_{-0.089}$		2,3 TUMASYAN	23AS CMS	$\gamma\gamma \rightarrow \tau^+\tau^-$ , Pb–Pb
-0.018 $\pm$ 0.017		2,4 ABDALLAH	04K DLPH	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$
< 0.107	95	5 ACHARD	04G L3	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$
-0.007 to 0.005	95	6 GONZALEZ-S...	00 RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ and $W \rightarrow \tau\nu_{\tau}$
-0.052 to 0.058	95	7 ACCIARRI	98E L3	1991–1995 LEP runs
-0.068 to 0.065	95	8 ACKERSTAFF	98N OPAL	1990–1995 LEP runs
-0.004 to 0.006	95	9 ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.01	95	10 ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.12	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau\tau\gamma$ at LEP
< 0.023	95	11 SILVERMAN	83 RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ at PETRA

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- <sup>1</sup> AAD 23BM measurement is derived from  $\gamma\gamma \rightarrow \tau^+\tau^-$  total cross-section from 1.44 nb<sup>-1</sup> LHC Pb-Pb collisions at  $\sqrt{S_{NN}} = 5.02$  TeV. Authors report both the measured value and the corresponding 95% CL limit.
- <sup>2</sup> Measurement ill-suited for a standard average because its likelihood appears to be remarkably non-Gaussian and asymmetric according to the model-dependent extraction procedure and the reported 95% CL limits.
- <sup>3</sup> TUMASYAN 23AS measurement is derived from  $\gamma\gamma \rightarrow \tau^+\tau^-$  total cross-section from 404  $\mu b^{-1}$  LHC Pb-Pb collisions at  $\sqrt{S_{NN}} = 5.02$  TeV.
- <sup>4</sup> ABDALLAH 04K measurement is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV. In addition to the measurement, the authors also quote 95% CL limits of  $> -0.052$  and  $< 0.013$ .
- <sup>5</sup> ACHARD 04G limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of the magnetic moment anomaly.
- <sup>6</sup> GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.
- <sup>7</sup> ACCIARRI 98E use  $Z \rightarrow \tau^+\tau^-\gamma$  events. In addition to the limits, the authors also quote a value of  $0.004 \pm 0.027 \pm 0.023$ .
- <sup>8</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z-m_\tau)^2$ .
- <sup>9</sup> ESCRIBANO 97 use preliminary experimental results.
- <sup>10</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+\tau^-)$ , and is on the absolute value of the magnetic moment anomaly.
- <sup>11</sup> SILVERMAN 83 limit is derived from  $e^+e^- \rightarrow \tau^+\tau^-$  total cross-section measurements for  $q^2$  up to  $(37 \text{ GeV})^2$ .

## $\tau$ ELECTRIC DIPOLE MOMENT ( $d_\tau$ )

A nonzero value is forbidden by both  $T$  invariance and  $P$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

### $\text{Re}(d_\tau)$

VALUE (10 <sup>-16</sup> e cm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>-0.185 to 0.061</b>	95	<sup>1</sup> INAMI	22 BELL	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
< 2.3	90	<sup>2</sup> GROZIN	09A RVUE	From $e$ EDM limit
< 3.7	95	<sup>3</sup> ABDALLAH	04K DLPH	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
< 11.4	95	<sup>4</sup> ACHARD	04G L3	$e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ at LEP2
- 0.22 to 0.45	95	<sup>5</sup> INAMI	03 BELL	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
< 4.6	95	<sup>6</sup> ALBRECHT	00 ARG	$E_{\text{cm}}^{\text{ee}} = 10.4 \text{ GeV}$
> -3.1 and < 3.1	95	ACCIARRI 98E	L3	1991–1995 LEP runs
> -3.8 and < 3.6	95	<sup>7</sup> ACKERSTAFF 98N	OPAL	1990–1995 LEP runs
< 0.11	95	<sup>8,9</sup> ESCRIBANO 97	RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.5	95	<sup>10</sup> ESCRIBANO 93	RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 7	90	GRIFOLS 91	RVUE	$Z \rightarrow \tau\tau\gamma$ at LEP
< 1.6	90	DELAGUILA 90	RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$

<sup>1</sup> INAMI 22 use  $e^+e^- \rightarrow \tau^+\tau^-$  events from 833 fb<sup>-1</sup> of data. Also report a measurement of  $\text{Re}(d_\tau) = (-0.62 \pm 0.63) \times 10^{-17}$  e cm.

<sup>2</sup> GROZIN 09A calculate the contribution to the electron electric dipole moment from the  $\tau$  electric dipole moment appearing in loops, which is  $\Delta d_e = 6.9 \times 10^{-12} d_\tau$ . Dividing the REGAN 02 upper limit  $|d_e| \leq 1.6 \times 10^{-27}$  e cm at CL=90% by  $6.9 \times 10^{-12}$  gives this limit.

<sup>3</sup> ABDALLAH 04K limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 183 and 208 GeV and is on the absolute value of  $d_\tau$ .

<sup>4</sup> ACHARD 04G limit is derived from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at  $\sqrt{s}$  between 189 and 206 GeV, and is on the absolute value of  $d_\tau$ .

<sup>5</sup> INAMI 03 use  $e^+e^- \rightarrow \tau^+\tau^-$  events.

<sup>6</sup> ALBRECHT 00 use  $e^+e^- \rightarrow \tau^+\tau^-$  events. Limit is on the absolute value of  $\text{Re}(d_\tau)$ .

<sup>7</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z-m_\tau)^2$ .

<sup>8</sup> ESCRIBANO 97 derive the relationship  $|d_\tau| = \cot \theta_W |d_\tau^W|$  using effective Lagrangian methods, and use a conference result  $|d_\tau^W| < 5.8 \times 10^{-18}$  e cm at 95% CL (L. Silvestris, ICHEP96) to obtain this result.

<sup>9</sup> ESCRIBANO 97 use preliminary experimental results.

<sup>10</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+\tau^-)$ , and is on the absolute value of the electric dipole moment.

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NODE=S035EDM;LINKAGE=C

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**Im( $d_\tau$ )**

<i>VALUE</i> ( $10^{-16}$ ecm)	<i>CL%</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
<b>-0.103 to 0.023 (CL = 95%)</b>		$[-0.103 \text{ to } 0.0230 \times 10^{-16} \text{ ecm (CL = 95%) OUR}$		
2025 BEST LIMIT]				
<b>-0.103 to 0.023</b>	95	<sup>1</sup> INAMI	22	BELL $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.25 to 0.008	95	<sup>2</sup> INAMI	03	BELL $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
< 1.8	95	<sup>3</sup> ALBRECHT	00	ARG $E_{\text{cm}}^{\text{ee}} = 10.4 \text{ GeV}$
1 INAMI 22 use $e^+ e^- \rightarrow \tau^+ \tau^-$ events from $833 \text{ fb}^{-1}$ of data. Also report a measurement of $\text{Im}(d_\tau) = (-0.40 \pm 0.32) \times 10^{-17} \text{ ecm}.$				
2 INAMI 03 use $e^+ e^- \rightarrow \tau^+ \tau^-$ events.				
3 ALBRECHT 00 use $e^+ e^- \rightarrow \tau^+ \tau^-$ events. Limit is on the absolute value of $\text{Im}(d_\tau).$				

 **$\tau$  WEAK DIPOLE MOMENT ( $d_\tau^w$ )**

A nonzero value is forbidden by *CP* invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

**Re( $d_\tau^w$ )**

<i>VALUE</i> ( $10^{-17}$ ecm)	<i>CL%</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
<b>&lt;0.50</b>	95	<sup>1</sup> HEISTER	03F	ALEP 1990–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<3.0	90	<sup>1</sup> ACCIARRI	98C	L3 1991–1995 LEP runs
<0.56	95	ACKERSTAFF	97L	OPAL 1991–1995 LEP runs
<0.78	95	<sup>2</sup> AKERS	95F	OPAL Repl. by ACKERSTAFF 97L
<1.5	95	<sup>2</sup> BUSKULIC	95C	ALEP Repl. by HEISTER 03F
<7.0	95	<sup>2</sup> ACTON	92F	OPAL $Z \rightarrow \tau^+ \tau^-$ at LEP
<3.7	95	<sup>2</sup> BUSKULIC	92J	ALEP Repl. by BUSKULIC 95C

1 Limit is on the absolute value of the real part of the weak dipole moment.

2 Limit is on the absolute value of the real part of the weak dipole moment, and applies for  $q^2 = m_Z^2.$

**Im( $d_\tau^w$ )**

<i>VALUE</i> ( $10^{-17}$ ecm)	<i>CL%</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
<b>&lt;1.1</b>	95	<sup>1</sup> HEISTER	03F	ALEP 1990–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.5	95	ACKERSTAFF	97L	OPAL 1991–1995 LEP runs
<4.5	95	<sup>2</sup> AKERS	95F	OPAL Repl. by ACKERSTAFF 97L

1 HEISTER 03F limit is on the absolute value of the imaginary part of the weak dipole moment.

2 Limit is on the absolute value of the imaginary part of the weak dipole moment, and applies for  $q^2 = m_Z^2.$

 **$\tau$  WEAK ANOMALOUS MAGNETIC DIPOLE MOMENT ( $\alpha_\tau^w$ )**

Electroweak radiative corrections are expected to contribute at the  $10^{-6}$  level. See BERNABEU 95.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

**Re( $\alpha_\tau^w$ )**

<i>VALUE</i>	<i>CL%</i>	<i>DOCUMENT ID</i>	<i>TECN</i>	<i>COMMENT</i>
<b>&lt;1.1 × 10<sup>-3</sup></b>	95	<sup>1</sup> HEISTER	03F	ALEP 1990–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
> -0.0024 and < 0.0025	95	<sup>2</sup> GONZALEZ-S...00	RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_\tau$

<4.5 × 10<sup>-3</sup> 90 <sup>1</sup> ACCIARRI 98C L3 1991–1995 LEP runs

1 Limit is on the absolute value of the real part of the weak anomalous magnetic dipole moment.

2 GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

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NODE=S035243

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NODE=S035WMM;LINKAGE=GS

**Im( $\alpha_\tau^w$ )**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.7 \times 10^{-3}$	95	1 HEISTER	03F ALEP	1990–1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<9.9 \times 10^{-3}$	90	1 ACCIARRI	98C L3	1991–1995 LEP runs
1 Limit is on the absolute value of the imaginary part of the weak anomalous magnetic dipole moment.				

NODE=S035WMI  
NODE=S035WMI **$\tau^-$  DECAY MODES**

$\tau^+$  modes are charge conjugates of the modes below. “ $h^\pm$ ” stands for  $\pi^\pm$  or  $K^\pm$ . “ $\ell$ ” stands for e or  $\mu$ . “Neutrals” stands for  $\gamma$ 's and/or  $\pi^0$ 's.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Modes with one charged particle</b>		
$\Gamma_1$ particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ ("1-prong")	(85.24 $\pm$ 0.06 ) %	NODE=S035;CLUMP=B DESIG=219
$\Gamma_2$ particle $^- \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	(84.58 $\pm$ 0.06 ) %	DESIG=191
$\Gamma_3$ $\mu^- \bar{\nu}_\mu \nu_\tau$	[a] (17.39 $\pm$ 0.04 ) %	DESIG=1
$\Gamma_4$ $\mu^- \bar{\nu}_\mu \nu_\tau \gamma$	[b] ( 3.67 $\pm$ 0.08 ) $\times 10^{-3}$	DESIG=76
$\Gamma_5$ $e^- \bar{\nu}_e \nu_\tau$	[a] (17.82 $\pm$ 0.04 ) %	DESIG=2
$\Gamma_6$ $e^- \bar{\nu}_e \nu_\tau \gamma$	[b] ( 1.83 $\pm$ 0.05 ) %	DESIG=274
$\Gamma_7$ $h^- \geq 0 K_L^0 \nu_\tau$	(12.03 $\pm$ 0.05 ) %	DESIG=181
$\Gamma_8$ $h^- \nu_\tau$	(11.51 $\pm$ 0.05 ) %	DESIG=217
$\Gamma_9$ $\pi^- \nu_\tau$	[a] (10.82 $\pm$ 0.05 ) %	DESIG=12
$\Gamma_{10}$ $K^- \nu_\tau$	[a] ( 6.96 $\pm$ 0.10 ) $\times 10^{-3}$	DESIG=7
$\Gamma_{11}$ $h^- \geq 1$ neutrals $\nu_\tau$	(37.00 $\pm$ 0.09 ) %	DESIG=192
$\Gamma_{12}$ $h^- \geq 1 \pi^0 \nu_\tau$ (ex. $K^0$ )	(36.50 $\pm$ 0.09 ) %	DESIG=299
$\Gamma_{13}$ $h^- \pi^0 \nu_\tau$	(25.93 $\pm$ 0.09 ) %	DESIG=171
$\Gamma_{14}$ $\pi^- \pi^0 \nu_\tau$	[a] (25.49 $\pm$ 0.09 ) %	DESIG=16
$\Gamma_{15}$ $\pi^- \pi^0$ non- $\rho(770) \nu_\tau$	( 3.0 $\pm$ 3.2 ) $\times 10^{-3}$	DESIG=24
$\Gamma_{16}$ $K^- \pi^0 \nu_\tau$	[a] ( 4.33 $\pm$ 0.15 ) $\times 10^{-3}$	DESIG=182
$\Gamma_{17}$ $h^- \geq 2 \pi^0 \nu_\tau$	(10.81 $\pm$ 0.09 ) %	DESIG=19
$\Gamma_{18}$ $h^- 2 \pi^0 \nu_\tau$	( 9.48 $\pm$ 0.10 ) %	DESIG=218
$\Gamma_{19}$ $h^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 9.32 $\pm$ 0.10 ) %	DESIG=27
$\Gamma_{20}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 9.26 $\pm$ 0.10 ) %	DESIG=201
$\Gamma_{21}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), scalar	< 9 $\times 10^{-3}$ CL=95%	DESIG=282
$\Gamma_{22}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), vector	< 7 $\times 10^{-3}$ CL=95%	DESIG=283
$\Gamma_{23}$ $K^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 6.5 $\pm$ 2.2 ) $\times 10^{-4}$	DESIG=115
$\Gamma_{24}$ $h^- \geq 3 \pi^0 \nu_\tau$	( 1.34 $\pm$ 0.07 ) %	DESIG=79
$\Gamma_{25}$ $h^- \geq 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.25 $\pm$ 0.07 ) %	DESIG=300
$\Gamma_{26}$ $h^- 3 \pi^0 \nu_\tau$	( 1.18 $\pm$ 0.07 ) %	DESIG=26
$\Gamma_{27}$ $\pi^- 3 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 1.04 $\pm$ 0.07 ) %	DESIG=203
$\Gamma_{28}$ $K^- 3 \pi^0 \nu_\tau$ (ex. $K^0$ , $\eta$ )	[a] ( 4.8 $\pm$ 2.1 ) $\times 10^{-4}$	DESIG=116
$\Gamma_{29}$ $h^- 4 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 1.6 $\pm$ 0.4 ) $\times 10^{-3}$	DESIG=220
$\Gamma_{30}$ $h^- 4 \pi^0 \nu_\tau$ (ex. $K^0$ , $\eta$ )	[a] ( 1.1 $\pm$ 0.4 ) $\times 10^{-3}$	DESIG=110
$\Gamma_{31}$ $a_1(1260) \nu_\tau \rightarrow \pi^- \gamma \nu_\tau$	( 4.0 $\pm$ 1.5 ) $\times 10^{-4}$	DESIG=344
$\Gamma_{32}$ $K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau$	( 1.552 $\pm$ 0.029 ) %	DESIG=198
$\Gamma_{33}$ $K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau$	( 8.59 $\pm$ 0.28 ) $\times 10^{-3}$	DESIG=11

<b>Modes with <math>K^0</math>'s</b>			NODE=S035;CLUMP=C
$\Gamma_{34}$	$K_S^0$ (particles) $\bar{\nu}_\tau$	$( 9.43 \pm 0.28 ) \times 10^{-3}$	DESIG=127
$\Gamma_{35}$	$h^- \bar{K}^0 \bar{\nu}_\tau$	$( 9.87 \pm 0.14 ) \times 10^{-3}$	DESIG=212
$\Gamma_{36}$	$\pi^- \bar{K}^0 \bar{\nu}_\tau$	[a] $( 8.38 \pm 0.14 ) \times 10^{-3}$	DESIG=117
$\Gamma_{37}$	$\pi^- \bar{K}^0 (\text{non-}K^*(892)^-) \bar{\nu}_\tau$	$( 5.4 \pm 2.1 ) \times 10^{-4}$	DESIG=142
$\Gamma_{38}$	$K^- K^0 \bar{\nu}_\tau$	[a] $( 1.486 \pm 0.034 ) \times 10^{-3}$	DESIG=62
$\Gamma_{39}$	$K^- K^0 \geq 0 \pi^0 \bar{\nu}_\tau$	$( 2.99 \pm 0.07 ) \times 10^{-3}$	DESIG=273
$\Gamma_{40}$	$h^- \bar{K}^0 \pi^0 \bar{\nu}_\tau$	$( 5.32 \pm 0.13 ) \times 10^{-3}$	DESIG=213
$\Gamma_{41}$	$\pi^- \bar{K}^0 \pi^0 \bar{\nu}_\tau$	[a] $( 3.82 \pm 0.13 ) \times 10^{-3}$	DESIG=118
$\Gamma_{42}$	$\bar{K}^0 \rho^- \bar{\nu}_\tau$	$( 2.2 \pm 0.5 ) \times 10^{-3}$	DESIG=249
$\Gamma_{43}$	$K^- K^0 \pi^0 \bar{\nu}_\tau$	[a] $( 1.50 \pm 0.07 ) \times 10^{-3}$	DESIG=119
$\Gamma_{44}$	$\pi^- \bar{K}^0 \geq 1 \pi^0 \bar{\nu}_\tau$	$( 4.08 \pm 0.25 ) \times 10^{-3}$	DESIG=272
$\Gamma_{45}$	$\pi^- \bar{K}^0 \pi^0 \pi^0 \bar{\nu}_\tau (\text{ex. } K^0)$	[a] $( 2.6 \pm 2.3 ) \times 10^{-4}$	DESIG=238
$\Gamma_{46}$	$K^- K^0 \pi^0 \pi^0 \bar{\nu}_\tau$	$< 1.6 \times 10^{-4}$ CL=95%	DESIG=239
$\Gamma_{47}$	$\pi^- K^0 \bar{K}^0 \bar{\nu}_\tau$	$( 1.55 \pm 0.24 ) \times 10^{-3}$	DESIG=141
$\Gamma_{48}$	$\pi^- K_S^0 K_S^0 \bar{\nu}_\tau$	[a] $( 2.35 \pm 0.06 ) \times 10^{-4}$	DESIG=214
$\Gamma_{49}$	$\pi^- K_S^0 K_L^0 \bar{\nu}_\tau$	[a] $( 1.08 \pm 0.24 ) \times 10^{-3}$	DESIG=240
$\Gamma_{50}$	$\pi^- K_L^0 K_L^0 \bar{\nu}_\tau$	$( 2.35 \pm 0.06 ) \times 10^{-4}$	DESIG=342
$\Gamma_{51}$	$\pi^- K^0 \bar{K}^0 \pi^0 \bar{\nu}_\tau$	$( 3.6 \pm 1.2 ) \times 10^{-4}$	DESIG=278
$\Gamma_{52}$	$\pi^- K_S^0 K_S^0 \pi^0 \bar{\nu}_\tau$	[a] $( 1.82 \pm 0.21 ) \times 10^{-5}$	DESIG=241
$\Gamma_{53}$	$K^* - K^0 \pi^0 \bar{\nu}_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \bar{\nu}_\tau$	$( 1.08 \pm 0.21 ) \times 10^{-5}$	DESIG=336
$\Gamma_{54}$	$f_1(1285) \pi^- \bar{\nu}_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \bar{\nu}_\tau$	$( 6.8 \pm 1.5 ) \times 10^{-6}$	DESIG=337
$\Gamma_{55}$	$f_1(1420) \pi^- \bar{\nu}_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \bar{\nu}_\tau$	$( 2.4 \pm 0.8 ) \times 10^{-6}$	DESIG=338
$\Gamma_{56}$	$\pi^- K_S^0 K_L^0 \pi^0 \bar{\nu}_\tau$	[a] $( 3.2 \pm 1.2 ) \times 10^{-4}$	DESIG=242
$\Gamma_{57}$	$\pi^- K_L^0 K_L^0 \pi^0 \bar{\nu}_\tau$	$( 1.82 \pm 0.21 ) \times 10^{-5}$	DESIG=343
$\Gamma_{58}$	$K^- K_S^0 K_S^0 \bar{\nu}_\tau$	$< 6.3 \times 10^{-7}$ CL=90%	DESIG=332
$\Gamma_{59}$	$K^- K_S^0 K_S^0 \pi^0 \bar{\nu}_\tau$	$< 4.0 \times 10^{-7}$ CL=90%	DESIG=333
$\Gamma_{60}$	$K^0 h^+ h^- \geq 0 \text{ neutrals } \bar{\nu}_\tau$	$< 1.7 \times 10^{-3}$ CL=95%	DESIG=29
$\Gamma_{61}$	$K^0 h^+ h^- \bar{\nu}_\tau$	[a] $( 2.5 \pm 2.0 ) \times 10^{-4}$	DESIG=244
<b>Modes with three charged particles</b>			NODE=S035;CLUMP=C
$\Gamma_{62}$	$h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \bar{\nu}_\tau$	$( 15.20 \pm 0.06 ) \%$	DESIG=194
$\Gamma_{63}$	$h^- h^- h^+ \geq 0 \text{ neutrals } \bar{\nu}_\tau$ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$ ) ("3-prong")	$( 14.55 \pm 0.06 ) \%$	DESIG=209
$\Gamma_{64}$	$h^- h^- h^+ \bar{\nu}_\tau$	$( 9.80 \pm 0.05 ) \%$	DESIG=17
$\Gamma_{65}$	$h^- h^- h^+ \bar{\nu}_\tau (\text{ex. } K^0)$	$( 9.46 \pm 0.05 ) \%$	DESIG=208
$\Gamma_{66}$	$h^- h^- h^+ \bar{\nu}_\tau (\text{ex. } K^0, \omega)$	$( 9.43 \pm 0.05 ) \%$	DESIG=215
$\Gamma_{67}$	$\pi^- \pi^+ \pi^- \bar{\nu}_\tau$	$( 9.31 \pm 0.05 ) \%$	DESIG=257
$\Gamma_{68}$	$\pi^- \pi^+ \pi^- \bar{\nu}_\tau (\text{ex. } K^0)$	$( 9.02 \pm 0.05 ) \%$	DESIG=258
$\Gamma_{69}$	$\pi^- \pi^+ \pi^- \bar{\nu}_\tau (\text{ex. } K^0)$ non-axial vector	$< 2.4 \%$ CL=95%	DESIG=284
$\Gamma_{70}$	$\pi^- \pi^+ \pi^- \bar{\nu}_\tau (\text{ex. } K^0, \omega)$	[a] $( 8.99 \pm 0.05 ) \%$	DESIG=259
$\Gamma_{71}$	$h^- h^- h^+ \geq 1 \text{ neutrals } \bar{\nu}_\tau$	$( 5.29 \pm 0.05 ) \%$	DESIG=18
$\Gamma_{72}$	$h^- h^- h^+ \geq 1 \pi^0 \bar{\nu}_\tau (\text{ex. } K^0)$ $h^- h^- h^+ \pi^0 \bar{\nu}_\tau$	$( 5.09 \pm 0.05 ) \%$ $( 4.76 \pm 0.05 ) \%$	DESIG=222 DESIG=25
$\Gamma_{73}$	$h^- h^- h^+ \pi^0 \bar{\nu}_\tau$	$( 4.57 \pm 0.05 ) \%$	DESIG=143
$\Gamma_{74}$	$h^- h^- h^+ \pi^0 \bar{\nu}_\tau (\text{ex. } K^0)$	$( 2.79 \pm 0.07 ) \%$	DESIG=202
$\Gamma_{75}$	$h^- h^- h^+ \pi^0 \bar{\nu}_\tau (\text{ex. } K^0, \omega)$	$( 4.62 \pm 0.05 ) \%$	DESIG=261
$\Gamma_{76}$	$\pi^- \pi^+ \pi^- \pi^0 \bar{\nu}_\tau$	$( 4.49 \pm 0.05 ) \%$	DESIG=262
$\Gamma_{77}$	$\pi^- \pi^+ \pi^- \pi^0 \bar{\nu}_\tau (\text{ex. } K^0)$	$( 2.74 \pm 0.07 ) \%$	DESIG=263
$\Gamma_{78}$	$\pi^- \pi^+ \pi^- \pi^0 \bar{\nu}_\tau (\text{ex. } K^0, \omega)$	[a] $( 5.17 \pm 0.31 ) \times 10^{-3}$	DESIG=91
$\Gamma_{79}$	$h^- \rho \pi^0 \bar{\nu}_\tau$	$( 4.95 \pm 0.31 ) \times 10^{-3}$	DESIG=92
$\Gamma_{80}$	$h^- \rho^+ h^- \bar{\nu}_\tau$	$( 4.95 \pm 0.31 ) \times 10^{-3}$	DESIG=93
$\Gamma_{81}$	$h^- \rho^- h^+ \bar{\nu}_\tau$	$( 4.95 \pm 0.31 ) \times 10^{-3}$	DESIG=301
$\Gamma_{82}$	$h^- h^- h^+ \geq 2 \pi^0 \bar{\nu}_\tau (\text{ex. } K^0)$	$( 5.05 \pm 0.31 ) \times 10^{-3}$	DESIG=221
$\Gamma_{83}$	$h^- h^- h^+ 2 \pi^0 \bar{\nu}_\tau$	$( 4.95 \pm 0.31 ) \times 10^{-3}$	DESIG=112

$\Gamma_{85}$	$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[a]	$(10 \pm 4) \times 10^{-4}$	DESIG=216	
$\Gamma_{86}$	$h^- h^- h^+ 3\pi^0 \nu_\tau$		$(2.13 \pm 0.30) \times 10^{-4}$	DESIG=204	
$\Gamma_{87}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0$ )		$(1.94 \pm 0.30) \times 10^{-4}$	DESIG=318	
$\Gamma_{88}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta, f_1(1285)$ )		$(1.7 \pm 0.4) \times 10^{-4}$	DESIG=319	
$\Gamma_{89}$	$2\pi^- \pi^+ 3\pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1(1285)$ )	[a]	$(1.4 \pm 2.7) \times 10^{-5}$	DESIG=320	
$\Gamma_{90}$	$K^- h^+ h^- \geq 0$ neutrals $\nu_\tau$		$(6.29 \pm 0.14) \times 10^{-3}$	DESIG=28	
$\Gamma_{91}$	$K^- h^+ \pi^- \nu_\tau$ (ex. $K^0$ )		$(4.37 \pm 0.07) \times 10^{-3}$	DESIG=270	
$\Gamma_{92}$	$K^- h^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		$(8.6 \pm 1.2) \times 10^{-4}$	DESIG=271	
$\Gamma_{93}$	$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$		$(4.77 \pm 0.14) \times 10^{-3}$	DESIG=6	
$\Gamma_{94}$	$K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau$ (ex. $K^0$ )		$(3.73 \pm 0.13) \times 10^{-3}$	DESIG=275	
$\Gamma_{95}$	$K^- \pi^+ \pi^- \nu_\tau$		$(3.45 \pm 0.07) \times 10^{-3}$	DESIG=245	
$\Gamma_{96}$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )		$(2.93 \pm 0.07) \times 10^{-3}$	DESIG=260	
$\Gamma_{97}$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[a]	$(2.93 \pm 0.07) \times 10^{-3}$	DESIG=340	
$\Gamma_{98}$	$K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$		$(1.4 \pm 0.5) \times 10^{-3}$	DESIG=286	
$\Gamma_{99}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$		$(1.31 \pm 0.12) \times 10^{-3}$	DESIG=246	
$\Gamma_{100}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		$(7.9 \pm 1.2) \times 10^{-4}$	DESIG=264	
$\Gamma_{101}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )		$(7.6 \pm 1.2) \times 10^{-4}$	DESIG=285	
$\Gamma_{102}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )		$(3.7 \pm 0.9) \times 10^{-4}$	DESIG=294	
$\Gamma_{103}$	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[a]	$(3.9 \pm 1.4) \times 10^{-4}$	DESIG=341	
$\Gamma_{104}$	$K^- \pi^+ K^- \geq 0$ neut. $\nu_\tau$	< 9	$\times 10^{-4}$	CL=95%	DESIG=123
$\Gamma_{105}$	$K^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$		$(1.496 \pm 0.033) \times 10^{-3}$	DESIG=122	
$\Gamma_{106}$	$K^- K^+ \pi^- \nu_\tau$	[a]	$(1.435 \pm 0.027) \times 10^{-3}$	DESIG=5	
$\Gamma_{107}$	$K^- K^+ \pi^- \pi^0 \nu_\tau$	[a]	$(6.1 \pm 1.8) \times 10^{-5}$	DESIG=247	
$\Gamma_{108}$	$K^- K^+ K^- \nu_\tau$		$(2.2 \pm 0.8) \times 10^{-5}$	S=5.4	DESIG=248
$\Gamma_{109}$	$K^- K^+ K^- \nu_\tau$ (ex. $\phi$ )	< 2.5	$\times 10^{-6}$	CL=90%	DESIG=303
$\Gamma_{110}$	$K^- K^+ K^- \pi^0 \nu_\tau$	< 4.8	$\times 10^{-6}$	CL=90%	DESIG=296
$\Gamma_{111}$	$\pi^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$	< 2.5	$\times 10^{-3}$	CL=95%	DESIG=121
$\Gamma_{112}$	$e^- e^- e^+ \bar{\nu}_e \nu_\tau$		$(2.8 \pm 1.5) \times 10^{-5}$	DESIG=210	
$\Gamma_{113}$	$\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau$	< 3.2	$\times 10^{-5}$	CL=90%	DESIG=211
$\Gamma_{114}$	$\pi^- e^- e^+ \nu_\tau$	seen			DESIG=350
$\Gamma_{115}$	$\pi^- \mu^- \mu^+ \nu_\tau$	< 1.14	$\times 10^{-5}$	CL=90%	DESIG=349

**Modes with five charged particles**

$\Gamma_{116}$	$3h^- 2h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^- \pi^+$ ) ("5-prong")		$(9.9 \pm 0.4) \times 10^{-4}$	NODE=S035;CLUMP=D DESIG=20
$\Gamma_{117}$	$3h^- 2h^+ \nu_\tau$ (ex. $K^0$ )		$(8.29 \pm 0.31) \times 10^{-4}$	DESIG=3
$\Gamma_{118}$	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega$ )		$(8.27 \pm 0.31) \times 10^{-4}$	DESIG=321
$\Gamma_{119}$	$3\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0, \omega, f_1(1285)$ )	[a]	$(7.75 \pm 0.30) \times 10^{-4}$	DESIG=322
$\Gamma_{120}$	$K^- 2\pi^- 2\pi^+ \nu_\tau$ (ex. $K^0$ )	[a]	$(6 \pm 12) \times 10^{-7}$	DESIG=327
$\Gamma_{121}$	$K^+ 3\pi^- \pi^+ \nu_\tau$	< 5.0	$\times 10^{-6}$	CL=90% DESIG=328
$\Gamma_{122}$	$K^+ K^- 2\pi^- \pi^+ \nu_\tau$	< 4.5	$\times 10^{-7}$	CL=90% DESIG=329
$\Gamma_{123}$	$3h^- 2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )		$(1.65 \pm 0.11) \times 10^{-4}$	DESIG=4
$\Gamma_{124}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )		$(1.64 \pm 0.11) \times 10^{-4}$	DESIG=323
$\Gamma_{125}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, f_1(1285)$ )		$(1.11 \pm 0.10) \times 10^{-4}$	DESIG=324
$\Gamma_{126}$	$3\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0, \eta, \omega, f_1(1285)$ )	[a]	$(3.8 \pm 0.9) \times 10^{-5}$	DESIG=325
$\Gamma_{127}$	$K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	[a]	$(1.1 \pm 0.6) \times 10^{-6}$	DESIG=330
$\Gamma_{128}$	$K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau$	< 8	$\times 10^{-7}$	CL=90% DESIG=331
$\Gamma_{129}$	$3h^- 2h^+ 2\pi^0 \nu_\tau$	< 3.4	$\times 10^{-6}$	CL=90% DESIG=128

**Miscellaneous other allowed modes**

				NODE=S035;CLUMP=E
$\Gamma_{130}$	$(5\pi)^-\nu_\tau$	$( 7.8 \pm 0.5 ) \times 10^{-3}$		DESIG=129
$\Gamma_{131}$	$4h^-3h^+ \geq 0$ neutrals $\nu_\tau$ ("7-prong")	$< 3.0 \times 10^{-7}$	CL=90%	DESIG=61
$\Gamma_{132}$	$4h^-3h^+\nu_\tau$	$< 4.3 \times 10^{-7}$	CL=90%	DESIG=290
$\Gamma_{133}$	$4h^-3h^+\pi^0\nu_\tau$	$< 2.5 \times 10^{-7}$	CL=90%	DESIG=291
$\Gamma_{134}$	$X^-(S=-1)\nu_\tau$	$( 2.92 \pm 0.04 ) \%$		DESIG=281
$\Gamma_{135}$	$K^*(892)^- \geq 0$ neutrals $\geq 0K_L^0\nu_\tau$	$( 1.42 \pm 0.18 ) \%$	S=1.4	DESIG=23
$\Gamma_{136}$	$K^*(892)^-\nu_\tau$	$( 1.20 \pm 0.07 ) \%$	S=1.8	DESIG=21
$\Gamma_{137}$	$K^*(892)^-\nu_\tau \rightarrow \pi^-\bar{K}^0\nu_\tau$	$( 7.82 \pm 0.26 ) \times 10^{-3}$		DESIG=304
$\Gamma_{138}$	$K^*(892)^0K^- \geq 0$ neutrals $\nu_\tau$	$( 3.2 \pm 1.4 ) \times 10^{-3}$		DESIG=97
$\Gamma_{139}$	$K^*(892)^0K^-\nu_\tau$	$( 2.1 \pm 0.4 ) \times 10^{-3}$		DESIG=206
$\Gamma_{140}$	$\bar{K}^*(892)^0\pi^- \geq 0$ neutrals $\nu_\tau$	$( 3.8 \pm 1.7 ) \times 10^{-3}$		DESIG=98
$\Gamma_{141}$	$\bar{K}^*(892)^0\pi^-\nu_\tau$	$( 2.2 \pm 0.5 ) \times 10^{-3}$		DESIG=205
$\Gamma_{142}$	$(\bar{K}^*(892)\pi)^-\nu_\tau \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$	$( 1.0 \pm 0.4 ) \times 10^{-3}$		DESIG=250
$\Gamma_{143}$	$K_1(1270)^-\nu_\tau$	$( 4.7 \pm 1.1 ) \times 10^{-3}$		DESIG=125
$\Gamma_{144}$	$K_1(1400)^-\nu_\tau$	$( 1.7 \pm 2.6 ) \times 10^{-3}$	S=1.7	DESIG=126
$\Gamma_{145}$	$K^*(1410)^-\nu_\tau$	$( 1.5 \pm 1.4 ) \times 10^{-3}$		DESIG=279
$\Gamma_{146}$	$K_0^*(1430)^-\nu_\tau$	$< 5 \times 10^{-4}$	CL=95%	DESIG=280
$\Gamma_{147}$	$K_2^*(1430)^-\nu_\tau$	$< 3 \times 10^{-3}$	CL=95%	DESIG=22
$\Gamma_{148}$	$a_0(980)^- \geq 0$ neutrals $\nu_\tau$			DESIG=99
$\Gamma_{149}$	$\eta\pi^-\nu_\tau$	$< 9.9 \times 10^{-5}$	CL=95%	DESIG=14
$\Gamma_{150}$	$\eta\pi^-\pi^0\nu_\tau$	$[a] ( 1.39 \pm 0.07 ) \times 10^{-3}$		DESIG=58
$\Gamma_{151}$	$\eta\pi^-\pi^0\pi^0\nu_\tau$	$[a] ( 1.9 \pm 0.4 ) \times 10^{-4}$		DESIG=68
$\Gamma_{152}$	$\eta K^-\nu_\tau$	$[a] ( 1.55 \pm 0.08 ) \times 10^{-4}$		DESIG=109
$\Gamma_{153}$	$\eta K^*(892)^-\nu_\tau$	$( 1.38 \pm 0.15 ) \times 10^{-4}$		DESIG=265
$\Gamma_{154}$	$\eta K^-\pi^0\nu_\tau$	$[a] ( 4.8 \pm 1.2 ) \times 10^{-5}$		DESIG=266
$\Gamma_{155}$	$\eta K^-\pi^0(\text{non-}K^*(892))\nu_\tau$	$< 3.5 \times 10^{-5}$	CL=90%	DESIG=309
$\Gamma_{156}$	$\eta\bar{K}^0\pi^-\nu_\tau$	$[a] ( 9.4 \pm 1.5 ) \times 10^{-5}$		DESIG=267
$\Gamma_{157}$	$\eta\bar{K}^0\pi^-\pi^0\nu_\tau$	$< 5.0 \times 10^{-5}$	CL=90%	DESIG=310
$\Gamma_{158}$	$\eta K^-K^0\nu_\tau$	$< 9.0 \times 10^{-6}$	CL=90%	DESIG=311
$\Gamma_{159}$	$\eta\pi^+\pi^-\pi^- \geq 0$ neutrals $\nu_\tau$	$< 3 \times 10^{-3}$	CL=90%	DESIG=66
$\Gamma_{160}$	$\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0)$	$[a] ( 2.20 \pm 0.13 ) \times 10^{-4}$		DESIG=230
$\Gamma_{161}$	$\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0, f_1(1285))$	$( 9.9 \pm 1.6 ) \times 10^{-5}$		DESIG=314
$\Gamma_{162}$	$\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau$	$< 3.9 \times 10^{-4}$	CL=90%	DESIG=231
$\Gamma_{163}$	$\eta\eta\pi^-\nu_\tau$	$< 7.4 \times 10^{-6}$	CL=90%	DESIG=69
$\Gamma_{164}$	$\eta\eta\pi^-\pi^0\nu_\tau$	$< 2.0 \times 10^{-4}$	CL=95%	DESIG=70
$\Gamma_{165}$	$\eta\eta K^-\nu_\tau$	$< 3.0 \times 10^{-6}$	CL=90%	DESIG=312
$\Gamma_{166}$	$\eta'(958)\pi^-\nu_\tau$	$< 4.0 \times 10^{-6}$	CL=90%	DESIG=232
$\Gamma_{167}$	$\eta'(958)\pi^-\pi^0\nu_\tau$	$< 1.2 \times 10^{-5}$	CL=90%	DESIG=233
$\Gamma_{168}$	$\eta'(958)K^-\nu_\tau$	$< 2.4 \times 10^{-6}$	CL=90%	DESIG=326
$\Gamma_{169}$	$\phi\pi^-\nu_\tau$	$( 3.4 \pm 0.6 ) \times 10^{-5}$		DESIG=207
$\Gamma_{170}$	$\phi K^-\nu_\tau$	$[a] ( 4.4 \pm 1.6 ) \times 10^{-5}$		DESIG=223
$\Gamma_{171}$	$f_1(1285)\pi^-\nu_\tau$	$( 3.9 \pm 0.5 ) \times 10^{-4}$	S=1.9	DESIG=234
$\Gamma_{172}$	$f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau$	$( 1.18 \pm 0.07 ) \times 10^{-4}$	S=1.3	DESIG=235
$\Gamma_{173}$	$f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^-2\pi^+\nu_\tau$	$[a] ( 5.2 \pm 0.4 ) \times 10^{-5}$		DESIG=315
$\Gamma_{174}$	$\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$	$< 1.0 \times 10^{-4}$	CL=90%	DESIG=276
$\Gamma_{175}$	$\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_S\text{-wave }\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau$	$< 1.9 \times 10^{-4}$	CL=90%	DESIG=277
$\Gamma_{176}$	$h^-\omega \geq 0$ neutrals $\nu_\tau$	$( 2.40 \pm 0.08 ) \%$		DESIG=71
$\Gamma_{177}$	$h^-\omega\nu_\tau$	$( 1.99 \pm 0.06 ) \%$		DESIG=8
$\Gamma_{178}$	$\pi^-\omega\nu_\tau$	$[a] ( 1.95 \pm 0.06 ) \%$		DESIG=339
$\Gamma_{179}$	$K^-\omega\nu_\tau$	$[a] ( 4.1 \pm 0.9 ) \times 10^{-4}$		DESIG=295
$\Gamma_{180}$	$h^-\omega\pi^0\nu_\tau$	$[a] ( 4.1 \pm 0.4 ) \times 10^{-3}$		DESIG=113
$\Gamma_{181}$	$h^-\omega 2\pi^0\nu_\tau$	$( 1.4 \pm 0.5 ) \times 10^{-4}$		DESIG=237
$\Gamma_{182}$	$\pi^-\omega 2\pi^0\nu_\tau$	$[a] ( 7.2 \pm 1.6 ) \times 10^{-5}$		DESIG=317
$\Gamma_{183}$	$h^-\omega\nu_\tau$	$< 5.4 \times 10^{-7}$	CL=90%	DESIG=302
$\Gamma_{184}$	$2h^-h^+\omega\nu_\tau$	$( 1.20 \pm 0.22 ) \times 10^{-4}$		DESIG=287
$\Gamma_{185}$	$2\pi^-\pi^+\omega\nu_\tau (\text{ex. } K^0)$	$[a] ( 8.4 \pm 0.6 ) \times 10^{-5}$		DESIG=316

**Lepton Family number (*LF*), Lepton number (*L*),  
or Baryon number (*B*) violating modes**

NODE=S035;CLUMP=A

*L* means lepton number violation (e.g.  $\tau^- \rightarrow e^+ \pi^- \pi^-$ ). Following common usage, *LF* means lepton family violation *and not* lepton number violation (e.g.  $\tau^- \rightarrow e^- \pi^+ \pi^-$ ). *B* means baryon number violation.

NODE=S035

$\Gamma_{186}$	$e^- \gamma$	<i>LF</i>	$< 3.3$	$\times 10^{-8}$	CL=90%	DESIG=32
$\Gamma_{187}$	$e^- \gamma \gamma$	<i>LF</i>	$< 2.5$	$\times 10^{-4}$	CL=90%	DESIG=355
$\Gamma_{188}$	$\mu^- \gamma$	<i>LF</i>	$< 4.2$	$\times 10^{-8}$	CL=90%	DESIG=31
$\Gamma_{189}$	$\mu^- \gamma \gamma$	<i>LF</i>	$< 5.8$	$\times 10^{-4}$	CL=90%	DESIG=356
$\Gamma_{190}$	$e^- \pi^0$	<i>LF</i>	$< 8.0$	$\times 10^{-8}$	CL=90%	DESIG=40
$\Gamma_{191}$	$\mu^- \pi^0$	<i>LF</i>	$< 1.1$	$\times 10^{-7}$	CL=90%	DESIG=39
$\Gamma_{192}$	$e^- K_S^0$	<i>LF</i>	$< 2.6$	$\times 10^{-8}$	CL=90%	DESIG=42
$\Gamma_{193}$	$\mu^- K_S^0$	<i>LF</i>	$< 2.3$	$\times 10^{-8}$	CL=90%	DESIG=41
$\Gamma_{194}$	$e^- \eta$	<i>LF</i>	$< 9.2$	$\times 10^{-8}$	CL=90%	DESIG=67
$\Gamma_{195}$	$\mu^- \eta$	<i>LF</i>	$< 6.5$	$\times 10^{-8}$	CL=90%	DESIG=114
$\Gamma_{196}$	$e^- \rho^0$	<i>LF</i>	$< 2.2$	$\times 10^{-8}$	CL=90%	DESIG=44
$\Gamma_{197}$	$\mu^- \rho^0$	<i>LF</i>	$< 1.7$	$\times 10^{-8}$	CL=90%	DESIG=43
$\Gamma_{198}$	$e^- \omega$	<i>LF</i>	$< 2.4$	$\times 10^{-8}$	CL=90%	DESIG=305
$\Gamma_{199}$	$\mu^- \omega$	<i>LF</i>	$< 3.9$	$\times 10^{-8}$	CL=90%	DESIG=306
$\Gamma_{200}$	$e^- K^*(892)^0$	<i>LF</i>	$< 1.9$	$\times 10^{-8}$	CL=90%	DESIG=53
$\Gamma_{201}$	$\mu^- K^*(892)^0$	<i>LF</i>	$< 2.9$	$\times 10^{-8}$	CL=90%	DESIG=54
$\Gamma_{202}$	$e^- \bar{K}^*(892)^0$	<i>LF</i>	$< 1.7$	$\times 10^{-8}$	CL=90%	DESIG=131
$\Gamma_{203}$	$\mu^- \bar{K}^*(892)^0$	<i>LF</i>	$< 4.3$	$\times 10^{-8}$	CL=90%	DESIG=132
$\Gamma_{204}$	$e^- \eta'(958)$	<i>LF</i>	$< 1.6$	$\times 10^{-7}$	CL=90%	DESIG=292
$\Gamma_{205}$	$\mu^- \eta'(958)$	<i>LF</i>	$< 1.3$	$\times 10^{-7}$	CL=90%	DESIG=293
$\Gamma_{206}$	$e^- f_0(980) \rightarrow e^- \pi^+ \pi^-$	<i>LF</i>	$< 3.2$	$\times 10^{-8}$	CL=90%	DESIG=307
$\Gamma_{207}$	$\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-$	<i>LF</i>	$< 3.4$	$\times 10^{-8}$	CL=90%	DESIG=308
$\Gamma_{208}$	$e^- \phi$	<i>LF</i>	$< 2.0$	$\times 10^{-8}$	CL=90%	DESIG=255
$\Gamma_{209}$	$\mu^- \phi$	<i>LF</i>	$< 2.3$	$\times 10^{-8}$	CL=90%	DESIG=256
$\Gamma_{210}$	$e^- e^+ e^-$	<i>LF</i>	$< 2.7$	$\times 10^{-8}$	CL=90%	DESIG=38
$\Gamma_{211}$	$e^- \mu^+ \mu^-$	<i>LF</i>	$< 2.7$	$\times 10^{-8}$	CL=90%	DESIG=36
$\Gamma_{212}$	$e^+ \mu^- \mu^-$	<i>LF</i>	$< 1.7$	$\times 10^{-8}$	CL=90%	DESIG=55
$\Gamma_{213}$	$\mu^- e^+ e^-$	<i>LF</i>	$< 1.8$	$\times 10^{-8}$	CL=90%	DESIG=37
$\Gamma_{214}$	$\mu^+ e^- e^-$	<i>LF</i>	$< 1.5$	$\times 10^{-8}$	CL=90%	DESIG=56
$\Gamma_{215}$	$\mu^- \mu^+ \mu^-$	<i>LF</i>	$< 1.9$	$\times 10^{-8}$	CL=90%	DESIG=35
$\Gamma_{216}$	$e^- \pi^+ \pi^-$	<i>LF</i>	$< 2.3$	$\times 10^{-8}$	CL=90%	DESIG=45
$\Gamma_{217}$	$e^+ \pi^- \pi^-$	<i>L</i>	$< 2.0$	$\times 10^{-8}$	CL=90%	DESIG=46
$\Gamma_{218}$	$\mu^- \pi^+ \pi^-$	<i>LF</i>	$< 2.1$	$\times 10^{-8}$	CL=90%	DESIG=47
$\Gamma_{219}$	$\mu^+ \pi^- \pi^-$	<i>L</i>	$< 3.9$	$\times 10^{-8}$	CL=90%	DESIG=48
$\Gamma_{220}$	$e^- \pi^+ K^-$	<i>LF</i>	$< 3.7$	$\times 10^{-8}$	CL=90%	DESIG=49
$\Gamma_{221}$	$e^- \pi^- K^+$	<i>LF</i>	$< 3.1$	$\times 10^{-8}$	CL=90%	DESIG=77
$\Gamma_{222}$	$e^+ \pi^- K^-$	<i>L</i>	$< 3.2$	$\times 10^{-8}$	CL=90%	DESIG=50
$\Gamma_{223}$	$e^- K_S^0 K_S^0$	<i>LF</i>	$< 7.1$	$\times 10^{-8}$	CL=90%	DESIG=288
$\Gamma_{224}$	$e^- K^+ K^-$	<i>LF</i>	$< 3.4$	$\times 10^{-8}$	CL=90%	DESIG=251
$\Gamma_{225}$	$e^+ K^- K^-$	<i>L</i>	$< 3.3$	$\times 10^{-8}$	CL=90%	DESIG=252
$\Gamma_{226}$	$\mu^- \pi^+ K^-$	<i>LF</i>	$< 8.6$	$\times 10^{-8}$	CL=90%	DESIG=51
$\Gamma_{227}$	$\mu^- \pi^- K^+$	<i>LF</i>	$< 4.5$	$\times 10^{-8}$	CL=90%	DESIG=78
$\Gamma_{228}$	$\mu^+ \pi^- K^-$	<i>L</i>	$< 4.8$	$\times 10^{-8}$	CL=90%	DESIG=52
$\Gamma_{229}$	$\mu^- K_S^0 K_S^0$	<i>LF</i>	$< 8.0$	$\times 10^{-8}$	CL=90%	DESIG=289
$\Gamma_{230}$	$\mu^- K^+ K^-$	<i>LF</i>	$< 4.4$	$\times 10^{-8}$	CL=90%	DESIG=253
$\Gamma_{231}$	$\mu^+ K^- K^-$	<i>L</i>	$< 4.7$	$\times 10^{-8}$	CL=90%	DESIG=254
$\Gamma_{232}$	$e^- \pi^0 \pi^0$	<i>LF</i>	$< 6.5$	$\times 10^{-6}$	CL=90%	DESIG=224
$\Gamma_{233}$	$\mu^- \pi^0 \pi^0$	<i>LF</i>	$< 1.4$	$\times 10^{-5}$	CL=90%	DESIG=225
$\Gamma_{234}$	$e^- \eta \eta$	<i>LF</i>	$< 3.5$	$\times 10^{-5}$	CL=90%	DESIG=226
$\Gamma_{235}$	$\mu^- \eta \eta$	<i>LF</i>	$< 6.0$	$\times 10^{-5}$	CL=90%	DESIG=227
$\Gamma_{236}$	$e^- \pi^0 \eta$	<i>LF</i>	$< 2.4$	$\times 10^{-5}$	CL=90%	DESIG=228
$\Gamma_{237}$	$\mu^- \pi^0 \eta$	<i>LF</i>	$< 2.2$	$\times 10^{-5}$	CL=90%	DESIG=229
$\Gamma_{238}$	$p e^- e^-$	<i>L,B</i>	$< 3.0$	$\times 10^{-8}$	CL=90%	DESIG=351

$\Gamma_{239}$	$\bar{p} e^+ e^-$	$L,B$	$< 3.0$	$\times 10^{-8}$	CL=90%	DESIG=352
$\Gamma_{240}$	$\bar{p} e^+ \mu^-$	$L,B$	$< 2.0$	$\times 10^{-8}$	CL=90%	DESIG=353
$\Gamma_{241}$	$\bar{p} e^- \mu^+$	$L,B$	$< 1.8$	$\times 10^{-8}$	CL=90%	DESIG=354
$\Gamma_{242}$	$p \mu^- \mu^-$	$L,B$	$< 4.0$	$\times 10^{-8}$	CL=90%	DESIG=334
$\Gamma_{243}$	$\bar{p} \mu^+ \mu^-$	$L,B$	$< 1.8$	$\times 10^{-8}$	CL=90%	DESIG=335
$\Gamma_{244}$	$\bar{p} \gamma$	$L,B$	$< 3.5$	$\times 10^{-6}$	CL=90%	DESIG=104
$\Gamma_{245}$	$\bar{p} \pi^0$	$L,B$	$< 1.5$	$\times 10^{-5}$	CL=90%	DESIG=105
$\Gamma_{246}$	$\bar{p} 2\pi^0$	$L,B$	$< 3.3$	$\times 10^{-5}$	CL=90%	DESIG=268
$\Gamma_{247}$	$\bar{p} \eta$	$L,B$	$< 8.9$	$\times 10^{-6}$	CL=90%	DESIG=106
$\Gamma_{248}$	$\bar{p} \pi^0 \eta$	$L,B$	$< 2.7$	$\times 10^{-5}$	CL=90%	DESIG=269
$\Gamma_{249}$	$\Lambda \pi^-$	$L,B$	$< 4.7$	$\times 10^{-8}$	CL=90%	DESIG=297
$\Gamma_{250}$	$\bar{\Lambda} \pi^-$	$L,B$	$< 4.3$	$\times 10^{-8}$	CL=90%	DESIG=298
$\Gamma_{251}$	$e^-$ light boson	$LF$	$< 9$	$\times 10^{-4}$	CL=95%	DESIG=102
$\Gamma_{252}$	$\mu^-$ light boson	$LF$	$< 6$	$\times 10^{-4}$	CL=95%	DESIG=103

[a] Basis mode for the  $\tau$ .

[b] See the Particle Listings below for the energy limits used in this measurement.

LINKAGE=BB

LINKAGE=KDM

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## CONSTRAINED FIT INFORMATION

An overall fit to 85 branching ratios uses 170 measurements and one constraint to determine 46 parameters. The overall fit has a  $\chi^2 = 135.0$  for 125 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

	$x_5$	18								
	$x_9$	2	-1							
	$x_{10}$	3	4	5						
$x_{14}$	-18	-19	-17	-5						
$x_{16}$	-1	-1	0	-2	-9					
$x_{20}$	-11	-11	-14	-4	-46	-1				
$x_{23}$	-1	0	-2	-3	-1	-14	-10			
$x_{27}$	-6	-5	-10	-1	0	1	-39	1		
$x_{28}$	0	-1	-1	-2	0	-13	-3	-23	-11	
$x_{30}$	-4	-3	-11	-1	-9	0	7	-2	-44	2
$x_{36}$	-2	-2	-3	-1	-1	-1	-2	0	-1	0
$x_{38}$	-1	-1	1	0	0	0	0	-2	0	-2
$x_{41}$	-2	-2	-2	-1	-1	0	-2	0	-1	0
$x_{43}$	-1	-1	-1	-1	0	-3	0	-5	0	-5
$x_{45}$	-5	-5	-5	-2	-3	-1	-4	-2	-1	-2
$x_{48}$	-1	-1	2	0	-1	2	-1	-1	0	-1
$x_{49}$	-5	-5	-5	-2	-3	-1	-5	-2	-1	-2
$x_{52}$	0	0	0	0	0	0	0	-1	0	-1
$x_{56}$	-2	-2	-2	-1	-1	-1	-2	-1	-1	-1
$x_{61}$	-5	-5	-5	-2	-3	-1	-4	-2	-1	-2
$x_{70}$	-7	-9	4	-2	-6	3	-12	-2	-7	-1
$x_{78}$	-4	-4	-5	0	-9	0	1	1	-1	1
$x_{85}$	0	0	-2	0	-2	0	0	0	2	0
$x_{89}$	0	0	0	0	0	0	0	0	0	0
$x_{97}$	-2	-2	-1	-1	-1	-1	-3	-1	-2	-1
$x_{103}$	1	1	0	-1	1	-1	-1	-1	0	-1
$x_{106}$	-1	-2	2	-1	-1	1	-2	-1	-1	-1
$x_{107}$	0	0	0	0	0	0	0	0	0	0
$x_{119}$	-1	-1	1	0	-1	1	-2	-1	-1	0
$x_{120}$	0	0	0	0	0	0	0	0	0	0
$x_{126}$	0	0	0	0	0	0	0	0	0	0
$x_{127}$	0	0	0	0	0	0	0	0	0	0
$x_{150}$	-1	-1	-1	0	-1	0	-2	-1	0	-1
$x_{151}$	-1	-1	0	0	-1	0	-1	0	0	0
$x_{152}$	0	0	0	0	0	0	0	-1	0	-1
$x_{154}$	0	0	0	0	0	0	0	0	0	0
$x_{156}$	0	0	0	0	0	0	0	0	0	0
$x_{160}$	-1	-1	1	0	-1	1	-1	0	-1	0
$x_{170}$	0	0	0	0	0	0	0	0	0	0
$x_{173}$	0	-1	1	0	0	1	-1	0	0	0
$x_{178}$	-3	-3	-3	-1	-4	0	-1	0	-1	0
$x_{179}$	0	0	0	0	0	0	0	0	0	0
$x_{180}$	-2	-2	-5	-1	-3	0	-2	-1	2	-1
$x_{182}$	0	0	0	0	0	0	0	0	0	0
$x_{185}$	-1	-1	0	0	-1	1	-1	0	0	0

$x_3$     $x_5$     $x_9$     $x_{10}$     $x_{14}$     $x_{16}$     $x_{20}$     $x_{23}$     $x_{27}$     $x_{28}$

	$x_{36}$	$x_{38}$	$x_{41}$	$x_{43}$	$x_{45}$	$x_{48}$	$x_{49}$	$x_{52}$	$x_{56}$	$x_{61}$	$x_{70}$	$x_{78}$	$x_{85}$	$x_{89}$	$x_{97}$	$x_{103}$	$x_{106}$	$x_{107}$	$x_{119}$	$x_{120}$	$x_{126}$	$x_{127}$	$x_{150}$	$x_{151}$	$x_{152}$	$x_{154}$	$x_{156}$	$x_{160}$	$x_{170}$	$x_{173}$	$x_{178}$	$x_{179}$	$x_{180}$	$x_{182}$	$x_{185}$			
	0	0	0	0	0	0	0	0	0	0	-5	3	2	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
$x_{36}$	0																																					
$x_{38}$	0	-15																																				
$x_{41}$	0	-13	2																																			
$x_{43}$	0	-1	-14	-20																																		
$x_{45}$	0	-3	0	-6	0																																	
$x_{48}$	0	-2	3	-4	1	0																																
$x_{49}$	0	-5	0	-4	-1	-10	-1																															
$x_{52}$	0	1	5	-1	6	0	-7	0																														
$x_{56}$	0	-2	0	-2	-1	-4	0	-8	0																													
$x_{61}$	0	-2	0	-2	0	-4	0	-4	0																													
$x_{70}$	-5	-2	3	-2	-1	-4	5	-4	0																													
$x_{78}$	3	1	-1	1	0	2	-1	2	0																													
$x_{85}$	2	0	0	0	0	0	0	0	0																													
$x_{89}$	0	0	0	0	0	0	0	0	-1	0																												
$x_{97}$	-1	-1	0	-1	0	-2	0	-2	0																													
$x_{103}$	-1	-1	0	-1	0	-1	0	-1	0																													
$x_{106}$	-1	-1	1	0	0	-1	2	-1	0																													
$x_{107}$	0	0	0	0	0	0	0	0	0																													
$x_{119}$	-1	-1	1	0	0	-1	2	-1	0																													
$x_{120}$	0	0	0	0	0	0	0	0	0																													
$x_{126}$	0	0	0	0	0	0	0	0	0																													
$x_{127}$	0	0	0	0	0	0	0	0	0																													
$x_{150}$	-2	-1	0	0	0	-1	0	-1	0																													
$x_{151}$	0	0	0	0	0	-1	0	-1	0																													
$x_{152}$	0	0	1	0	0	0	1	0	0																													
$x_{154}$	0	0	0	0	0	0	0	0	0																													
$x_{156}$	0	0	0	0	0	0	0	-1	0																													
$x_{160}$	-1	0	1	0	0	-1	1	-1	0																													
$x_{170}$	0	0	0	0	0	0	0	0	0																													
$x_{173}$	-1	0	1	0	0	0	1	0	0																													
$x_{178}$	1	0	0	0	0	-1	0	-1	0																													
$x_{179}$	0	0	0	0	0	0	0	0	0																													
$x_{180}$	2	-1	0	0	0	-1	0	-1	0																													
$x_{182}$	0	0	0	0	0	0	0	0	0																													
$x_{185}$	-1	0	1	0	0	0	1	-1	0																													

$x_{30} \quad x_{36} \quad x_{38} \quad x_{41} \quad x_{43} \quad x_{45} \quad x_{48} \quad x_{49} \quad x_{52} \quad x_{56}$

x <sub>70</sub>	-4									
x <sub>78</sub>	2	-19								
x <sub>85</sub>	0	-1	-8							
x <sub>89</sub>	0	-1	-1	0						
x <sub>97</sub>	-2	19	-6	0	0					
x <sub>103</sub>	-1	-4	-14	-1	0	-1				
x <sub>106</sub>	-1	15	-4	0	0	0	-1			
x <sub>107</sub>	0	-1	-1	0	0	0	-3	0		
x <sub>119</sub>	-1	3	-1	0	-4	-1	0	1	0	
x <sub>120</sub>	0	0	0	0	0	0	0	0	-1	
x <sub>126</sub>	0	0	0	0	0	0	0	0	3	
x <sub>127</sub>	0	0	0	0	0	0	0	0	-1	
x <sub>150</sub>	-1	0	0	-5	0	0	0	0	0	0
x <sub>151</sub>	-1	0	0	0	-11	0	0	0	0	9
x <sub>152</sub>	0	2	0	0	0	0	-1	1	0	1
x <sub>154</sub>	0	0	0	-1	0	0	0	0	0	0
x <sub>156</sub>	0	0	0	0	-2	0	0	0	0	0
x <sub>160</sub>	-1	1	-1	0	-8	-1	0	1	0	46
x <sub>170</sub>	0	-1	0	0	0	1	0	1	0	0
x <sub>173</sub>	0	1	0	0	-2	0	0	1	0	34
x <sub>178</sub>	-1	-9	-67	-3	0	-2	10	-2	0	-1
x <sub>179</sub>	0	0	12	0	0	-2	-58	0	0	0
x <sub>180</sub>	-1	-2	-11	-64	-1	-1	-1	-1	0	0
x <sub>182</sub>	0	0	0	0	-16	0	0	0	0	7
x <sub>185</sub>	0	1	-1	0	-4	0	0	1	0	39

x<sub>61</sub>   x<sub>70</sub>   x<sub>78</sub>   x<sub>85</sub>   x<sub>89</sub>   x<sub>97</sub>   x<sub>103</sub>   x<sub>106</sub>   x<sub>107</sub>   x<sub>119</sub>

x <sub>126</sub>	0									
x <sub>127</sub>	0	-1								
x <sub>150</sub>	0	0	0							
x <sub>151</sub>	0	2	0	0						
x <sub>152</sub>	0	0	0	4	0					
x <sub>154</sub>	0	0	0	1	0	1				
x <sub>156</sub>	0	0	0	2	-1	1	0			
x <sub>160</sub>	-1	3	-1	0	25	0	0	0		
x <sub>170</sub>	0	0	0	0	0	0	0	0	0	
x <sub>173</sub>	-1	1	0	0	4	0	0	0	20	0
x <sub>178</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>179</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>180</sub>	0	0	0	0	0	0	0	0	0	0
x <sub>182</sub>	0	2	0	0	10	0	0	-1	20	0
x <sub>185</sub>	-1	-2	-1	0	17	0	0	0	38	0

x<sub>120</sub>   x<sub>126</sub>   x<sub>127</sub>   x<sub>150</sub>   x<sub>151</sub>   x<sub>152</sub>   x<sub>154</sub>   x<sub>156</sub>   x<sub>160</sub>   x<sub>170</sub>

x <sub>178</sub>	0									
x <sub>179</sub>	0	-14								
x <sub>180</sub>	0	-4	0							
x <sub>182</sub>	3	0	0	0						
x <sub>185</sub>	17	0	0	0	14					

x<sub>173</sub>   x<sub>178</sub>   x<sub>179</sub>   x<sub>180</sub>   x<sub>182</sub>

See the related review(s):

τ Branching Fractions

$$(\Gamma(\tau^+) - \Gamma(\tau^-)) / (\Gamma(\tau^+) + \Gamma(\tau^-))$$

$\tau^\pm \rightarrow \pi^\pm K_S^0 \nu_\tau$  (RATE DIFFERENCE) / (RATE SUM)

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>-0.36 ± 0.23 ± 0.11</b>	LEES	12M BABR 476 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\tau^-$ BRANCHING RATIOS

$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")}) / \Gamma_{\text{total}} \quad \Gamma_1 / \Gamma$$

$$\Gamma_1 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + \Gamma_{36} + \Gamma_{38} + \Gamma_{41} + \Gamma_{43} + \Gamma_{45} + \Gamma_{48} + \Gamma_{49} + \Gamma_{50} + \Gamma_{52} + \Gamma_{56} + \Gamma_{57} + 0.7195 \Gamma_{150} + 0.7195 \Gamma_{152} + 0.7195 \Gamma_{154} + 0.7195 \Gamma_{156} + 0.336 \Gamma_{170} + 0.0833 \Gamma_{178} + 0.0833 \Gamma_{179} + 0.0833 \Gamma_{180}) / \Gamma$$

The charged particle here can be  $e$ ,  $\mu$ , or hadron. In many analyses, the sum of the topological branching fractions (1, 3, and 5 prongs) is constrained to be unity. Since the 5-prong fraction is very small, the measured 1-prong and 3-prong fractions are highly correlated and cannot be treated as independent quantities in our overall fit. We arbitrarily choose to use the 3-prong fraction in our fit, and leave the 1-prong fraction out. We do, however, use these 1-prong measurements in our average below.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**85.24 ± 0.06 OUR FIT** (Produced by HFLAV)

**85.26 ± 0.13 OUR AVERAGE** Error includes scale factor of 1.6. See the ideogram below.

• • • We use the following data for averages but not for fits. • • •

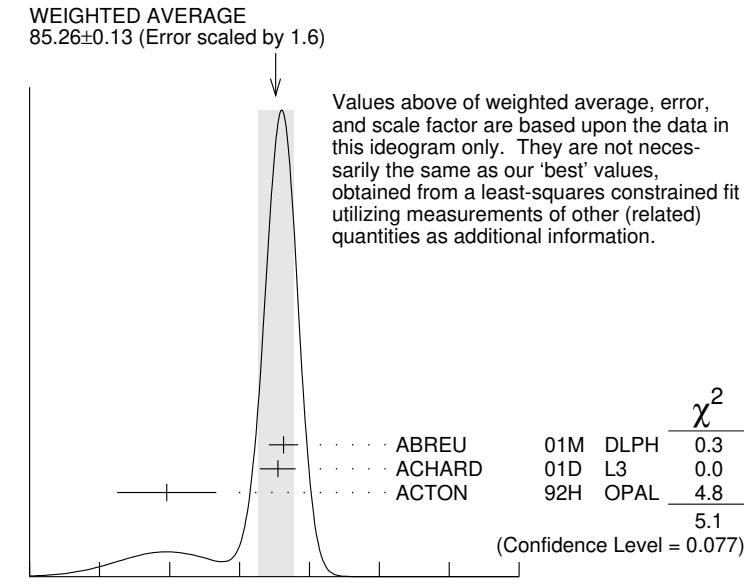
85.316 ± 0.093 ± 0.049	78k	1 ABREU	01M DLPH	1992–1995 LEP runs
85.274 ± 0.105 ± 0.073		2 ACHARD	01D L3	1992–1995 LEP runs
84.48 ± 0.27 ± 0.23		ACTON	92H OPAL	1990–1991 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

85.45	$+0.69$	$\pm 0.65$	DECAMP	92C ALEP	Repl. by SCHABEL 05C
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1 The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{3-prong})$  and  $B(\tau \rightarrow \text{5-prong})$  are  $-0.98$  and  $-0.08$  respectively.

2 The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"3-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.082$  respectively.



$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau \text{ ("1-prong")}) / \Gamma_{\text{total}} (\%)$$

NODE=S035250

NODE=S035DR1

NODE=S035DR1

NODE=S035220

NODE=S035220

NODE=S035B75

NODE=S035B75

NODE=S035B75

NOTFITTED

NOTFITTED

NOTFITTED

NODE=S035B75;LINKAGE=M1

NODE=S035B75;LINKAGE=CH

$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$	$\Gamma_2 / \Gamma$	NODE=S035R24
$\Gamma_2 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{14} + \Gamma_{16} + \Gamma_{20} + \Gamma_{23} + \Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.6534\Gamma_{36} + 0.6534\Gamma_{38} + 0.6534\Gamma_{41} + 0.6534\Gamma_{43} + 0.6534\Gamma_{45} + 0.0942\Gamma_{48} + 0.3069\Gamma_{49} + \Gamma_{50} + 0.0942\Gamma_{52} + 0.3069\Gamma_{56} + \Gamma_{57} + 0.7195\Gamma_{150} + 0.7195\Gamma_{152} + 0.7195\Gamma_{154} + 0.4701\Gamma_{156} + 0.1049\Gamma_{170} + 0.0833\Gamma_{178} + 0.0833\Gamma_{179} + 0.0833\Gamma_{180}) / \Gamma$		NODE=S035R24
		NODE=S035R24

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	NODE=S035R24
<b>84.58 ± 0.06 OUR FIT</b>	(Produced by HFLAV)				

### 85.1 ± 0.4 OUR AVERAGE

• • • We use the following data for averages but not for fits. • • •

85.6 ± 0.6 ± 0.3	3300	1 ADEVA	91F L3	$E_{\text{cm}}^{\text{ee}} = 88.3\text{--}94.3 \text{ GeV}$	NOTFITTED
84.9 ± 0.4 ± 0.3		BEHREND	89B CELL	$E_{\text{cm}}^{\text{ee}} = 14\text{--}47 \text{ GeV}$	NOTFITTED
84.7 ± 0.8 ± 0.6		2 AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	NOTFITTED

• • • We do not use the following data for averages, fits, limits, etc. • • •

86.4 ± 0.3 ± 0.3		ABACHI	89B HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
87.1 ± 1.0 ± 0.7		3 BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
87.2 ± 0.5 ± 0.8		SCHMIDKE	86 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
84.7 ± 1.1 ± 1.6	169	4 ALTHOFF	85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$	
86.1 ± 0.5 ± 0.9		BARTEL	85F JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
87.8 ± 1.3 ± 3.9		5 BERGER	85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
86.7 ± 0.3 ± 0.6		FERNANDEZ	85 MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	

1 Not independent of ADEVA 91F  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  value.

2 Not independent of AIHARA 87B  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  values.

3 Not independent of SCHMIDKE 86 value (also not independent of BURCHAT 87 value for  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ ).

4 Not independent of ALTHOFF 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  values.

5 Not independent of (1-prong + 0 $\pi^0$ ) and (1-prong + ≥ 1 $\pi^0$ ) values.

### $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$

### $\Gamma_3 / \Gamma$

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	NODE=S035R1
<b>17.39 ± 0.04 OUR FIT</b>	(Produced by HFLAV)				
<b>17.33 ± 0.05 OUR AVERAGE</b>					
17.319 ± 0.070 ± 0.032	54k	1 SCHAEL	05C ALEP	1991-1995 LEP runs	
17.34 ± 0.09 ± 0.06	31.4k	ABBIENDI	03 OPAL	1990-1995 LEP runs	
17.342 ± 0.110 ± 0.067	21.5k	2 ACCIARRI	01F L3	1991-1995 LEP runs	
17.325 ± 0.095 ± 0.077	27.7k	ABREU	99X DLPH	1991-1995 LEP runs	
• • • We use the following data for averages but not for fits. • • •					
17.37 ± 0.08 ± 0.18		3 ANASTASSOV 97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NOTFITTED
• • • We do not use the following data for averages, fits, limits, etc. • • •					
17.31 ± 0.11 ± 0.05	20.7k	BUSKULIC	96C ALEP	Repl. by SCHAEL 05C	
17.02 ± 0.19 ± 0.24	6586	ABREU	95T DLPH	Repl. by ABREU 99X	
17.36 ± 0.27	7941	AKERS	95I OPAL	Repl. by ABBIENDI 03	
17.6 ± 0.4 ± 0.4	2148	ADRIANI	93M L3	Repl. by ACCIARRI 01F	
17.4 ± 0.3 ± 0.5		4 ALBRECHT	93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
17.35 ± 0.41 ± 0.37		DECAMP	92C ALEP	1989-1990 LEP runs	
17.7 ± 0.8 ± 0.4	568	BEHREND	90 CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$	
17.4 ± 1.0	2197	ADEVA	88 MRKJ	$E_{\text{cm}}^{\text{ee}} = 14\text{--}16 \text{ GeV}$	
17.7 ± 1.2 ± 0.7		AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
18.3 ± 0.9 ± 0.8		BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
18.6 ± 0.8 ± 0.7	558	5 BARTEL	86D JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	NOTFITTED
12.9 ± 1.7 ± 0.7		ALTHOFF	85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$	
18.0 ± 0.9 ± 0.5	473	5 ASH	85B MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	NOTFITTED
18.0 ± 1.0 ± 0.6		6 BALTRUSAIT..85	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$	
19.4 ± 1.6 ± 1.7	153	BERGER	85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	
17.6 ± 2.6 ± 2.1	47	BEHREND	83C CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$	
17.8 ± 2.0 ± 1.8		BERGER	81B PLUT	$E_{\text{cm}}^{\text{ee}} = 9\text{--}32 \text{ GeV}$	

- <sup>1</sup> See footnote to SCHABEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.
- <sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  is 0.08.
- <sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(e\bar{\nu}_e \nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu \nu_\tau)/B(e\bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e\bar{\nu}_e \nu_\tau)$  are 0.50, 0.58, 0.50, and 0.08 respectively.
- <sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}^2$  values.
- <sup>5</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .
- <sup>6</sup> Error correlated with BALTRUSAITIS 85  $e\nu\bar{\nu}$  value.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}}$					$\Gamma_4/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.367 ± 0.008 OUR AVERAGE</b>					
0.363 ± 0.002 ± 0.015	22k	<sup>1</sup> SHIMIZU	18A BELL	$711 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
0.369 ± 0.003 ± 0.010	16k	<sup>2</sup> LEES	15G BABR	$431 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
0.361 ± 0.016 ± 0.035		<sup>3</sup> BERGFELD	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.30 ± 0.04 ± 0.05	116	<sup>4</sup> ALEXANDER	96S OPAL	1991–1994 LEP runs	
0.23 ± 0.10	10	<sup>5</sup> WU	90 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	

- <sup>1</sup> SHIMIZU 18A impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .
- <sup>2</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .
- <sup>3</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ . For  $E_\gamma^* > 20 \text{ MeV}$ , they quote  $(3.04 \pm 0.14 \pm 0.30) \times 10^{-3}$ .
- <sup>4</sup> ALEXANDER 96S impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma > 20 \text{ MeV}$ .
- <sup>5</sup> WU 90 reports  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) = 0.013 \pm 0.006$ , which is converted to  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}}$  using  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}} = 17.35\%$ . Requirements on detected  $\gamma$ 's correspond to a  $\tau$  rest frame energy cutoff  $E_\gamma > 37 \text{ MeV}$ .

$\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_5/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>17.82 ± 0.04 OUR FIT</b> (Produced by HFLAV)					
<b>17.82 ± 0.05 OUR AVERAGE</b>					
17.837 ± 0.072 ± 0.036	56k	<sup>1</sup> SCHABEL	05C ALEP	1991–1995 LEP runs	
17.806 ± 0.104 ± 0.076	24.7k	<sup>2</sup> ACCIARRI	01F L3	1991–1995 LEP runs	
17.81 ± 0.09 ± 0.06	33.1k	ABBIENDI	99H OPAL	1991–1995 LEP runs	
17.877 ± 0.109 ± 0.110	23.3k	ABREU	99X DLPH	1991–1995 LEP runs	
17.76 ± 0.06 ± 0.17		<sup>3</sup> ANASTASSOV	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
17.78 ± 0.10 ± 0.09	25.3k	ALEXANDER	96D OPAL	Repl. by ABBIENDI 99H	
17.79 ± 0.12 ± 0.06	20.6k	BUSKULIC	96C ALEP	Repl. by SCHABEL 05C	
17.51 ± 0.23 ± 0.31	5059	ABREU	95T DLPH	Repl.. by ABREU 99X	
17.9 ± 0.4 ± 0.4	2892	ADRIANI	93M L3	Repl. by ACCIARRI 01F	
17.5 ± 0.3 ± 0.5		<sup>4</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
17.97 ± 0.14 ± 0.23	3970	AKERIB	92 CLEO	Repl. by ANASTASSOV 97	
19.1 ± 0.4 ± 0.6	2960	<sup>5</sup> AMMAR	92 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5\text{--}10.9 \text{ GeV}$	NOTFITTED
18.09 ± 0.45 ± 0.45		DECAMP	92C ALEP	Repl. by SCHABEL 05C	
17.0 ± 0.5 ± 0.6	1.7k	ABACHI	90 HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
18.4 ± 0.8 ± 0.4	644	BEHREND	90 CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$	
16.3 ± 0.3 ± 3.2		JANSSEN	89 CBAL	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$	
18.4 ± 1.2 ± 1.0		AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
19.1 ± 0.8 ± 1.1		BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
16.8 ± 0.7 ± 0.9	515	<sup>5</sup> BARTEL	86D JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$	NOTFITTED
20.4 ± 3.0 ± 1.4	–0.9	ALTHOFF	85 TASS	$E_{\text{cm}}^{\text{ee}} = 34.5 \text{ GeV}$	

NODE=S035R1;LINKAGE=SC

NODE=S035R1;LINKAGE=FA

NODE=S035R1;LINKAGE=N1

NODE=S035R1;LINKAGE=BB

NODE=S035R1;LINKAGE=Q

NODE=S035R1;LINKAGE=H

NODE=S035B81

NODE=S035B81

NODE=S035B81;LINKAGE=E

NODE=S035B81;LINKAGE=D

NODE=S035B81;LINKAGE=C

NODE=S035B81;LINKAGE=A

NODE=S035B81;LINKAGE=B

NODE=S035R2

NODE=S035R2

NODE=S035R2

NOTFITTED

NOTFITTED

17.8	$\pm 0.9$	$\pm 0.6$	390	5 ASH	85B MAC	$E_{cm}^{ee} = 29$ GeV
18.2	$\pm 0.7$	$\pm 0.5$		6 BALTRUSAIT..85	MRK3	$E_{cm}^{ee} = 3.77$ GeV
13.0	$\pm 1.9$	$\pm 2.9$		BERGER	85 PLUT	$E_{cm}^{ee} = 34.6$ GeV
18.3	$\pm 2.4$	$\pm 1.9$	60	BEHREND	83C CELL	$E_{cm}^{ee} = 34$ GeV
16.0	$\pm 1.3$		459	7 BACINO	78B DLCO	$E_{cm}^{ee} = 3.1\text{--}7.4$ GeV

<sup>1</sup> Correlation matrix for SCHael 05C branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{total}$
- (2)  $\Gamma(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma_{total}$
- (3)  $\Gamma(\tau^- \rightarrow \pi^-\nu_\tau)/\Gamma_{total}$
- (4)  $\Gamma(\tau^- \rightarrow \pi^-\pi^0\nu_\tau)/\Gamma_{total}$
- (5)  $\Gamma(\tau^- \rightarrow \pi^-2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{total}$
- (6)  $\Gamma(\tau^- \rightarrow \pi^-3\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{total}$
- (7)  $\Gamma(\tau^- \rightarrow h^-4\pi^0\nu_\tau(\text{ex. } K^0, \eta))/\Gamma_{total}$
- (8)  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0, \omega))/\Gamma_{total}$
- (9)  $\Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{total}$
- (10)  $\Gamma(\tau^- \rightarrow h^-h^-h^+2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{total}$
- (11)  $\Gamma(\tau^- \rightarrow h^-h^-h^+3\pi^0\nu_\tau)/\Gamma_{total}$
- (12)  $\Gamma(\tau^- \rightarrow 3h^-2h^+\nu_\tau(\text{ex. } K^0))/\Gamma_{total}$
- (13)  $\Gamma(\tau^- \rightarrow 3h^-2h^+\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{total}$

	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)	(12)
(2)	-20											
(3)	-9	-6										
(4)	-16	-12	2									
(5)	-5	-5	-17	-37								
(6)	0	-4	-15	2	-27							
(7)	-2	-4	-24	-15	20	-47						
(8)	-14	-9	15	-5	-17	-14	-8					
(9)	-13	-12	-25	-30	4	-2	16	-15				
(10)	0	-2	-23	-14	4	10	13	-6	-17			
(11)	1	0	-5	1	4	6	0	-9	-2	-11		
(12)	0	1	9	4	-8	-4	-6	9	-5	-4	-2	
(13)	1	-4	-3	-5	3	2	-4	-3	-1	4	1	-24

<sup>2</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau)$  is 0.08.

NODE=S035R2;LINKAGE=FA

<sup>3</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$ ,  $B(h^-\nu_\tau)$ , and  $B(h^-\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$  are 0.50, -0.42, 0.48, and -0.39 respectively.

NODE=S035R2;LINKAGE=N1

<sup>4</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau) \times \Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{total}^2$  values.

NODE=S035R2;LINKAGE=BB

<sup>5</sup> Modified using  $B(e^-\bar{\nu}_e\nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

NODE=S035R2;LINKAGE=Q

<sup>6</sup> Error correlated with BALTRUSAITIS 85  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma_{total}$ .

NODE=S035R2;LINKAGE=H

<sup>7</sup> BACINO 78B value comes from fit to events with  $e^\pm$  and one other nonelectron charged prong.

NODE=S035R2;LINKAGE=B

### $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma(e^-\bar{\nu}_e\nu_\tau)$

### $\Gamma_3/\Gamma_5$

Standard Model prediction including mass effects is 0.9726.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**97.62 ± 0.28 OUR FIT** (Produced by HFLAV)

NODE=S035R5

NODE=S035R5

NODE=S035R5

**97.3 ± 0.4 OUR AVERAGE** Error includes scale factor of 1.6. See the ideogram below.

96.75 $\pm 0.07 \pm 0.36$	333E6 ADACHI	24K BELL
97.96 $\pm 0.16 \pm 0.36$	731k <sup>1</sup> AUBERT	10F BABR 467 $fb^{-1}$ $E_{cm}^{ee} = 10.6$ GeV
97.77 $\pm 0.63 \pm 0.87$	<sup>2</sup> ANASTASSOV 97 CLEO	$E_{cm}^{ee} = 10.6$ GeV
99.7 $\pm 3.5 \pm 4.0$	ALBRECHT 92D ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV

<sup>1</sup> Correlation matrix for AUBERT 10F branching fractions:

NODE=S035R5;LINKAGE=AU

- (1)  $\Gamma(\tau^- \rightarrow \mu^-\bar{\nu}_\mu\nu_\tau) / \Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$

- (2)  $\Gamma(\tau^- \rightarrow \pi^-\nu_\tau) / \Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$

- (3)  $\Gamma(\tau^- \rightarrow K^-\nu_\tau) / \Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)$

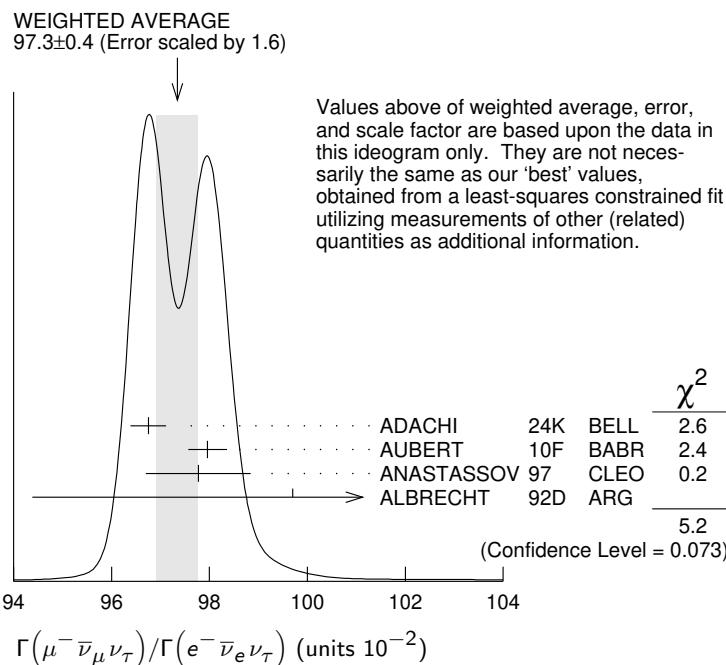
(1) (2)

(2) 0.25

(3) 0.12 0.33

<sup>2</sup>The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(h^-\nu_\tau)$ , and  $B(h^-\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$  are 0.58, -0.42, 0.07, and 0.45 respectively.

NODE=S035R5;LINKAGE=N1



### Γ(e⁻̄νeντγ) / Γ<sub>total</sub>

VALUE (%)	EVTS	DOCUMENT ID	TECN.	COMMENT	Γ <sub>6</sub> /Γ
<b>1.83 ± 0.05 OUR AVERAGE</b>					
1.79 ± 0.02 ± 0.10	12k	<sup>1</sup> SHIMIZU	18A BELL	$711 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
1.847 ± 0.015 ± 0.052	18k	<sup>2</sup> LEES	15G BABR	$431 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
1.75 ± 0.06 ± 0.17		<sup>3</sup> BERGFELD	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

<sup>1</sup> SHIMIZU 18A impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .

NODE=S035C39  
NODE=S035C39

<sup>2</sup> LEES 15G impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .

NODE=S035C39;LINKAGE=B

<sup>3</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10 \text{ MeV}$ .

NODE=S035C39;LINKAGE=A

• • • We use the following data for averages but not for fits. • • •

NODE=S035C39;LINKAGE=C

### Γ(h⁻ ≥ 0 K<sub>L</sub><sup>0</sup> ν<sub>τ</sub>) / Γ<sub>total</sub>

$$\Gamma_7/\Gamma = (\Gamma_9 + \Gamma_{10} + \frac{1}{2}\Gamma_{36} + \frac{1}{2}\Gamma_{38} + \Gamma_{50})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN.	COMMENT	Γ <sub>7</sub> /Γ
<b>12.03 ± 0.05 OUR FIT</b> (Produced by HFLAV)					
<b>12.2 ± 0.4 OUR AVERAGE</b>					

12.47 ± 0.26 ± 0.43	2967	<sup>1</sup> ACCIARRI	95 L3	1992 LEP run
12.4 ± 0.7 ± 0.7	283	<sup>2</sup> ABREU	92N DLPH	1990 LEP run
12.1 ± 0.7 ± 0.5	309	ALEXANDER	91D OPAL	1990 LEP run

• • • We use the following data for averages but not for fits. • • •

NODE=S035R43

11.3 ± 0.5 ± 0.8      798      <sup>3</sup> FORD      87 MAC  $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

NODE=S035R43

• • • We do not use the following data for averages, fits, limits, etc. • • •

NOTFITTED

12.44 ± 0.11 ± 0.11	15k	<sup>4</sup> BUSKULIC	96 ALEP	Repl. by SCHAEEL 05C
11.7 ± 0.6 ± 0.8		<sup>5</sup> ALBRECHT	92D ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
12.98 ± 0.44 ± 0.33		<sup>6</sup> DECOMP	92C ALEP	Repl. by SCHAEEL 05C
12.3 ± 0.9 ± 0.5	1338	BEHREND	90 CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
11.1 ± 1.1 ± 1.4		<sup>7</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
12.3 ± 0.6 ± 1.1	328	<sup>8</sup> BARTEL	86D JADE	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
13.0 ± 2.0 ± 4.0		BERGER	85 PLUT	$E_{\text{cm}}^{\text{ee}} = 34.6 \text{ GeV}$
11.2 ± 1.7 ± 1.2	34	<sup>9</sup> BEHREND	83C CELL	$E_{\text{cm}}^{\text{ee}} = 34 \text{ GeV}$

- 1 ACCIARRI 95 with 0.65% added to remove their correction for  $\pi^- K_L^0$  backgrounds.  
 2 ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.  
 3 FORD 87 result for  $B(\pi^- \nu_\tau)$  with 0.67% added to remove their  $K^-$  correction and adjusted for 1992 B("1 prong").  
 4 BUSKULIC 96 quote  $11.78 \pm 0.11 \pm 0.13$  We add 0.66 to undo their correction for unseen  $K_L^0$  and modify the systematic error accordingly.  
 5 Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$ ,  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  values.  
 6 DECAMP 92C quote  $B(h^- \geq 0 K_L^0 \geq 0 (K_S^0 \rightarrow \pi^+ \pi^-) \nu_\tau) = 13.32 \pm 0.44 \pm 0.33$ . We subtract 0.35 to correct for their inclusion of the  $K_S^0$  decays.  
 7 BURCHAT 87 with 1.1% added to remove their correction for  $K^-$  and  $K^*(892)^-$  backgrounds.  
 8 BARTEL 86D result for  $B(\pi^- \nu_\tau)$  with 0.59% added to remove their  $K^-$  correction and adjusted for 1992 B("1 prong").  
 9 BEHREND 83C quote  $B(\pi^- \nu_\tau) = 9.9 \pm 1.7 \pm 1.3$  after subtracting  $1.3 \pm 0.5$  to correct for  $B(K^- \nu_\tau)$ .

$\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_8/\Gamma = (\Gamma_9 + \Gamma_{10})/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT

**11.51 ± 0.05 OUR FIT** (Produced by HFLAV)

**11.63 ± 0.12 OUR AVERAGE** Error includes scale factor of 1.4. See the ideogram below.

$11.571 \pm 0.120 \pm 0.114$	19k	<sup>1</sup> ABDALLAH 06A DLPH 1992–1995 LEP runs
$11.98 \pm 0.13 \pm 0.16$		ACKERSTAFF 98M OPAL 1991–1995 LEP runs
$11.52 \pm 0.05 \pm 0.12$		<sup>2</sup> ANASTASSOV 97 CLEO $E_{cm}^{ee} = 10.6$ GeV

<sup>1</sup> Correlation matrix for ABDALLAH 06A branching fractions, in percent:

- (1)  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$
- (2)  $\Gamma(\tau^- \rightarrow h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$
- (3)  $\Gamma(\tau^- \rightarrow h^- \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (4)  $\Gamma(\tau^- \rightarrow h^- 2 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (5)  $\Gamma(\tau^- \rightarrow h^- \geq 3 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (6)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (7)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (8)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 1 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (9)  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \geq 2 \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (10)  $\Gamma(\tau^- \rightarrow 3 h^- 2 h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$
- (11)  $\Gamma(\tau^- \rightarrow 3 h^- 2 h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$

	(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
(2)	-34									
(3)	-47	56								
(4)	6	-66	15							
(5)	-6	38	11	-86						
(6)	-7	-8	15	0	-2					
(7)	-2	-1	-5	-3	3	-53				
(8)	-4	-4	-13	-4	-2	-56	75			
(9)	-1	-1	-4	3	-6	26	-78	-16		
(10)	-1	-1	1	0	0	-2	-3	-1	3	
(11)	0	0	0	0	0	1	0	-5	5	-57

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(e \bar{\nu}_e \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.50, 0.48, 0.07, and 0.63 respectively.

NODE=S035R43;LINKAGE=C

NODE=S035R43;LINKAGE=E

NODE=S035R43;LINKAGE=B

NODE=S035R43;LINKAGE=F

NODE=S035R43;LINKAGE=BB

NODE=S035R43;LINKAGE=Q

NODE=S035R43;LINKAGE=A

NODE=S035R43;LINKAGE=D

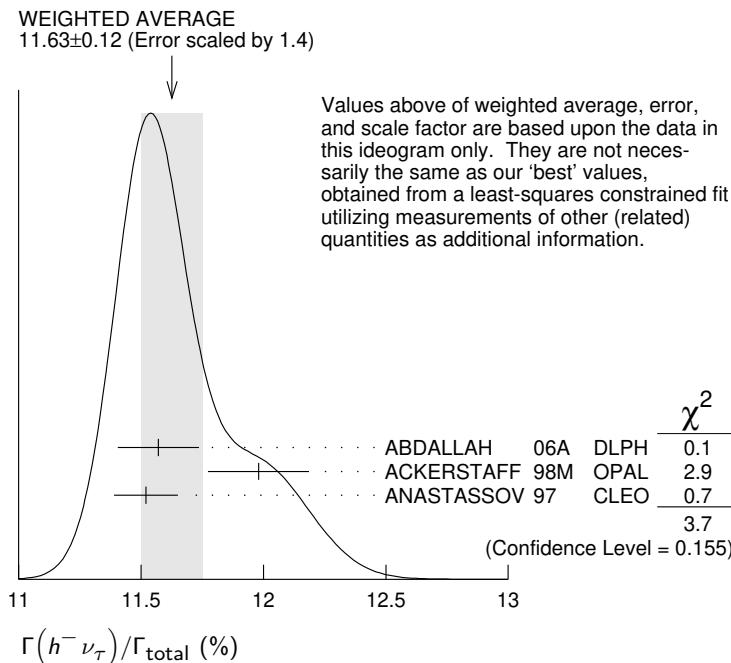
NODE=S035R43;LINKAGE=Q1

NODE=S035B73

NODE=S035B73

NODE=S035B73;LINKAGE=SC

NODE=S035B73;LINKAGE=N1



### $\Gamma(h^- \nu_\tau)/\Gamma_{\text{total}}$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**64.61±0.33 OUR FIT** (Produced by HFLAV)

**64.0 ±0.7 OUR AVERAGE** Error includes scale factor of 1.6.

• • • We use the following data for averages but not for fits. • • •

63.33±0.14±0.61	394k	<sup>1</sup> AUBERT	10F BABR	$467 \text{ fb}^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$
64.84±0.41±0.60		<sup>2</sup> ANASTASSOV 97	CLEO	$E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$

<sup>1</sup> Not independent of AUBERT 10F  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ .

<sup>2</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(e \bar{\nu}_e \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ , and  $B(h^- \nu_\tau)$  are 0.08, -0.39, 0.45, and 0.63 respectively.

### $\Gamma(\pi^- \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**10.82 ±0.05 OUR FIT** (Produced by HFLAV)

**10.828±0.070±0.078** 38k <sup>1</sup> SCHael 05C ALEP 1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

11.06 ±0.11 ±0.14		<sup>2</sup> BUSKULIC 96	ALEP	Repl. by SCHael 05C
11.7 ±0.4 ±1.8	1138	BLOCKER	82D MRK2	$E_{\text{cm}}^{\text{ee}}=3.5-6.7 \text{ GeV}$

<sup>1</sup> See footnote to SCHael 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of BUSKULIC 96  $B(h^- \nu_\tau)$  and  $B(K^- \nu_\tau)$  values.

### $\Gamma(\pi^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**60.71±0.32 OUR FIT** (Produced by HFLAV)

**59.45±0.14±0.61** 369k <sup>1</sup> AUBERT 10F BABR  $467 \text{ fb}^{-1}$   $E_{\text{cm}}^{\text{ee}}=10.6 \text{ GeV}$

<sup>1</sup> See footnote to AUBERT 10F  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  for correlations with other measurements.

### $\Gamma(K^- \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.696±0.010 OUR FIT** (Produced by HFLAV)

**0.685±0.023 OUR AVERAGE**

0.658±0.027±0.029		<sup>1</sup> ABBIENDI 01J	OPAL	1990-1995 LEP runs
0.696±0.025±0.014	2032	BARATE 99K	ALEP	1991-1995 LEP runs
0.85 ±0.18	27	ABREU 94K	DLPH	LEP 1992 Z data
0.66 ±0.07 ±0.09	99	BATTLE 94	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

### $\Gamma_8/\Gamma_5 = (\Gamma_9 + \Gamma_{10})/\Gamma_5$

NODE=S035B97  
NODE=S035B97

NOTFITTED  
NOTFITTED

NODE=S035B97;LINKAGE=AU  
NODE=S035B97;LINKAGE=N1

NODE=S035R6  
NODE=S035R6

NOTFITTED

NODE=S035R6;LINKAGE=SC

NODE=S035R6;LINKAGE=D

NODE=S035C78  
NODE=S035C78

NODE=S035C78;LINKAGE=AU

NODE=S035R7  
NODE=S035R7

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.72 \pm 0.04 \pm 0.04$	728	BUSKULIC	96	ALEP	Repl. by BARATE 99K
$0.59 \pm 0.18$	16	MILLS	84	DLCO	$E_{cm}^{ee} = 29$ GeV
$1.3 \pm 0.5$	15	BLOCKER	82B	MRK2	$E_{cm}^{ee} = 3.9\text{--}6.7$ GeV

<sup>1</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$  is 0.60.

NODE=S035R7;LINKAGE=AJ

$\Gamma(K^- \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$	$\Gamma_{10}/\Gamma_5$			
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.91 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			

<b>3.882 ± 0.032 ± 0.057</b>	25k	<sup>1</sup> AUBERT	10F	BABR	$467 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
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<sup>1</sup> See footnote to AUBERT 10F  $\Gamma(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  for correlations with other measurements.

NODE=S035C79  
NODE=S035C79

$\Gamma(K^- \nu_\tau)/\Gamma(\pi^- \nu_\tau)$	$\Gamma_{10}/\Gamma_9$		
VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>6.44 ± 0.09 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

<b>6.531 ± 0.056 ± 0.093</b>	<sup>1</sup> AUBERT	10F	BABR	$467 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
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<sup>1</sup> Not independent of AUBERT 10F  $\Gamma(\tau^- \rightarrow \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $\Gamma(\tau^- \rightarrow K^- \nu_\tau)/\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$ .

NODE=S035C79;LINKAGE=AU

$\Gamma(h^- \geq 1 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{11}/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>37.00 ± 0.09 OUR FIT</b>	(Produced by HFLAV)		

• • • We do not use the following data for averages, fits, limits, etc. • • •

$36.14 \pm 0.33 \pm 0.58$	<sup>1</sup> AKERS	94E	OPAL	1991–1992 LEP runs
$38.4 \pm 1.2 \pm 1.0$	<sup>2</sup> BURCHAT	87	MRK2	$E_{cm}^{ee} = 29$ GeV
$42.7 \pm 2.0 \pm 2.9$	BERGER	85	PLUT	$E_{cm}^{ee} = 34.6$ GeV

<sup>1</sup> Not independent of ACKERSTAFF 98M  $B(h^- \pi^0 \nu_\tau)$  and  $B(h^- \geq 2\pi^0 \nu_\tau)$  values.

<sup>2</sup> BURCHAT 87 quote for  $B(\pi^\pm \geq 1 \text{ neutral} \nu_\tau) = 0.378 \pm 0.012 \pm 0.010$ . We add 0.006 to account for contribution from  $(K^{*-} \nu_\tau)$  which they fixed at BR = 0.013.

NOTFITTED

NODE=S035C80;LINKAGE=AU

$\Gamma(h^- \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$	$\Gamma_{12}/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>36.50 ± 0.09 OUR FIT</b>	(Produced by HFLAV)		

• • • We use the following data for averages but not for fits. • • •

<b>36.641 ± 0.155 ± 0.127</b>	45k	<sup>1</sup> ABDALLAH	06A	DLPH	1992–1995 LEP runs
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<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

NODE=S035R42;LINKAGE=AE  
NODE=S035R42;LINKAGE=B

$\Gamma(h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma = (\Gamma_{14} + \Gamma_{16})/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>25.93 ± 0.09 OUR FIT</b>	(Produced by HFLAV)			
<b>25.73 ± 0.16 OUR AVERAGE</b>				

$25.67 \pm 0.01 \pm 0.39$	5.4M	FUJIKAWA	08	BELL	$72 \text{ fb}^{-1} E_{cm}^{ee} = 10.6$ GeV
$25.740 \pm 0.201 \pm 0.138$	35k	<sup>1</sup> ABDALLAH	06A	DLPH	1992–1995 LEP runs
$25.89 \pm 0.17 \pm 0.29$		ACKERSTAFF	98M	OPAL	1991–1995 LEP runs
$25.05 \pm 0.35 \pm 0.50$	6613	ACCIARRI	95	L3	1992 LEP run
$25.87 \pm 0.12 \pm 0.42$	51k	<sup>2</sup> ARTUSO	94	CLEO	$E_{cm}^{ee} = 10.6$ GeV

NODE=S035C01

NODE=S035C01

• • • We do not use the following data for averages, fits, limits, etc. • • •

$25.76 \pm 0.15 \pm 0.13$	31k	BUSKULIC	96	ALEP	Repl. by SCHAEFEL 05C
$25.98 \pm 0.36 \pm 0.52$		<sup>3</sup> AKERS	94E	OPAL	Repl. by ACKERSTAFF 98M
$22.9 \pm 0.8 \pm 1.3$	283	<sup>4</sup> ABREU	92N	DLPH	$E_{cm}^{ee} = 88.2\text{--}94.2$ GeV
$23.1 \pm 0.4 \pm 0.9$	1249	<sup>5</sup> ALBRECHT	92Q	ARG	$E_{cm}^{ee} = 10$ GeV
$25.02 \pm 0.64 \pm 0.88$	1849	DECAMP	92C	ALEP	1989–1990 LEP runs
$22.0 \pm 0.8 \pm 1.9$	779	ANTREASYAN	91	CBAL	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
$22.6 \pm 1.5 \pm 0.7$	1101	BEHREND	90	CELL	$E_{cm}^{ee} = 35$ GeV
$23.1 \pm 1.9 \pm 1.6$		BEHREND	84	CELL	$E_{cm}^{ee} = 14,22$ GeV

NOTFITTED

NODE=S035C01;LINKAGE=AH

NODE=S035R84

NODE=S035R84

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> ARTUSO 94 reports the combined result from three independent methods, one of which ( $23\%$  of the  $\tau^- \rightarrow h^- \pi^0 \nu_\tau$ ) is normalized to the inclusive one-prong branching fraction, taken as  $0.854 \pm 0.004$ . Renormalization to the present value causes negligible change.

<sup>3</sup> AKERS 94E quote  $(26.25 \pm 0.36 \pm 0.52) \times 10^{-2}$ ; we subtract  $0.27\%$  from their number to correct for  $\tau^- \rightarrow h^- K_L^0 \nu_\tau$ .

<sup>4</sup> ABREU 92N with  $0.5\%$  added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>5</sup> ALBRECHT 92Q with  $0.5\%$  added to remove their correction for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  background.

### $\Gamma(\pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{14}/\Gamma$
<b>25.49 ± 0.09 OUR FIT</b>	(Produced by HFLAV)				
<b>25.46 ± 0.12 OUR AVERAGE</b>					

$25.471 \pm 0.097 \pm 0.085$     81k    <sup>1</sup> SCHAEL    05C ALEP    1991-1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

$25.36 \pm 0.44$     <sup>2</sup> ARTUSO    94 CLEO     $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$25.30 \pm 0.15 \pm 0.13$     <sup>3</sup> BUSKULIC    96 ALEP    Repl. by SCHAEL 05C

$21.5 \pm 0.4 \pm 1.9$     4,5 ALBRECHT    88L ARG     $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

$23.0 \pm 1.3 \pm 1.7$     582 ADLER    87B MRK3     $E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$

$25.8 \pm 1.7 \pm 2.5$     <sup>6</sup> BURCHAT    87 MRK2     $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

$22.3 \pm 0.6 \pm 1.4$     629 YELTON    86 MRK2     $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of ARTUSO 94 B( $h^- \pi^0 \nu_\tau$ ) and BATTLE 94 B( $K^- \pi^0 \nu_\tau$ ) values.

<sup>3</sup> Not independent of BUSKULIC 96 B( $h^- \pi^0 \nu_\tau$ ) and B( $K^- \pi^0 \nu_\tau$ ) values.

<sup>4</sup> The authors divide by  $(\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10})/\Gamma = 0.467$  to obtain this result.

<sup>5</sup> Experiment had no hadron identification. Kaon corrections were made, but insufficient information is given to permit their removal.

<sup>6</sup> BURCHAT 87 value is not independent of YELTON 86 value. Nonresonant decays included.

### $\Gamma(\pi^- \pi^0 \text{non-}\rho(770)\nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{15}/\Gamma$
<b>0.3±0.1±0.3</b>		<sup>1</sup> BEHREND	84	CELL $E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$	

<sup>1</sup> BEHREND 84 assume a flat nonresonant mass distribution down to the  $\rho(770)$  mass, using events with mass above 1300 to set the level.

### $\Gamma(K^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{16}/\Gamma$
<b>0.433±0.015 OUR FIT</b>	(Produced by HFLAV)				
<b>0.426±0.016 OUR AVERAGE</b>					

$0.416 \pm 0.003 \pm 0.018$     78k    AUBERT    07AP BABR     $230 \text{ fb}^{-1}$      $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$0.471 \pm 0.059 \pm 0.023$     360    ABBIENDI    04J OPAL    1991-1995 LEP runs

$0.444 \pm 0.026 \pm 0.024$     923    BARATE    99K ALEP    1991-1995 LEP runs

$0.51 \pm 0.10 \pm 0.07$     37    BATTLE    94 CLEO     $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.52 \pm 0.04 \pm 0.05$     395    BUSKULIC    96 ALEP    Repl. by BARATE 99K

### $\Gamma(h^- \geq 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{17}/\Gamma$
<b>10.81±0.09 OUR FIT</b>	(Produced by HFLAV)				

**9.91±0.31±0.27**    ACKERSTAFF 98M OPAL    1991-1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

$9.89 \pm 0.34 \pm 0.55$     <sup>1</sup> AKERS    94E OPAL    Repl. by ACKER-STAFF 98M

$14.0 \pm 1.2 \pm 0.6$     938    <sup>2</sup> BEHREND    90 CELL     $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$

$12.0 \pm 1.4 \pm 2.5$     <sup>3</sup> BURCHAT    87 MRK2     $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

$13.9 \pm 2.0 \begin{array}{l} +1.9 \\ -2.2 \end{array}$     <sup>4</sup> AIHARA    86E TPC     $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> AKERS 94E not independent of AKERS 94E B( $h^- \geq 1\pi^0 \nu_\tau$ ) and B( $h^- \pi^0 \nu_\tau$ ) measurements.

<sup>2</sup> No independent of BEHREND 90  $\Gamma(h^- 2\pi^0 \nu_\tau \text{ (exp. } K^0\text{)})$  and  $\Gamma(h^- \geq 3\pi^0 \nu_\tau)$ .

<sup>3</sup> Error correlated with BURCHAT 87  $\Gamma(\rho^- \nu_e)/\Gamma(\text{total})$  value.

<sup>4</sup> AIHARA 86E (TPC) quote  $B(2\pi^0 \pi^- \nu_\tau) + 1.6B(3\pi^0 \pi^- \nu_\tau) + 1.1B(\pi^0 \eta \pi^- \nu_\tau)$ .

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$\Gamma(h^- 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$ 

$$\Gamma_{18}/\Gamma = (\Gamma_{20} + \Gamma_{23} + 0.15344\Gamma_{36} + 0.15344\Gamma_{38})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.48 ± 0.10 OUR FIT</b>	(Produced by HFLAV)			

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.48 ± 0.13 ± 0.10	12k	<sup>1</sup> BUSKULIC	96	ALEP Repl. by SCHAEEL 05C 1 BUSKULIC 96 quote $9.29 \pm 0.13 \pm 0.10$ . We add 0.19 to undo their correction for $\tau^- \rightarrow h^- K^0 \nu_\tau$ .
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 $\Gamma_{18}/\Gamma$ 

NODE=S035B74

NODE=S035B74

NODE=S035B74

 $\Gamma(h^- 2\pi^0 \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ 

$$\Gamma_{19}/\Gamma = (\Gamma_{20} + \Gamma_{23})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.32 ± 0.10 OUR FIT</b>	(Produced by HFLAV)			

**9.17 ± 0.27 OUR AVERAGE**

9.498 ± 0.320 ± 0.275	9.5k	<sup>1</sup> ABDALLAH	06A	DLPH 1992–1995 LEP runs
8.88 ± 0.37 ± 0.42	1060	ACCIARRI	95	L3 1992 LEP run

• • • We use the following data for averages but not for fits. • • •

8.96 ± 0.16 ± 0.44	<sup>2</sup> PROCARIO	93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

10.38 ± 0.66 ± 0.82	809	<sup>3</sup> DECAMP	92C	ALEP Repl. by SCHAEEL 05C
5.7 ± 0.5 ± 1.7	133	<sup>4</sup> ANTREASYAN	91	CBAL $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
10.0 ± 1.5 ± 1.1	333	<sup>5</sup> BEHREND	90	CELL $E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$
8.7 ± 0.4 ± 1.1	815	<sup>6</sup> BAND	87	MAC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
6.2 ± 0.6 ± 1.2		<sup>7</sup> GAN	87	MRK2 $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
6.0 ± 3.0 ± 1.8		BEHREND	84	CELL $E_{\text{cm}}^{\text{ee}} = 14.22 \text{ GeV}$

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> PROCARIO 93 entry is obtained from  $B(h^- 2\pi^0 \nu_\tau)/B(h^- \pi^0 \nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$ .

<sup>3</sup> We subtract 0.0015 to account for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>4</sup> ANTREASYAN 91 subtract 0.001 to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>5</sup> BEHREND 90 subtract 0.002 to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>6</sup> BAND 87 assume  $B(\pi^- 3\pi^0 \nu_\tau) = 0.01$  and  $B(\pi^- \pi^0 \eta \nu_\tau) = 0.005$ .

<sup>7</sup> GAN 87 analysis use photon multiplicity distribution.

 $\Gamma(h^- 2\pi^0 \nu_\tau(\text{ex. } K^0))/\Gamma(h^- \pi^0 \nu_\tau)$  $\Gamma_{19}/\Gamma_{13}$ 

$$\Gamma_{19}/\Gamma_{13} = (\Gamma_{20} + \Gamma_{23})/(\Gamma_{14} + \Gamma_{16})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>36.0 ± 0.4 OUR FIT</b>	(Produced by HFLAV)		

<b>34.2 ± 0.6 ± 1.6</b>	<sup>1</sup> PROCARIO	93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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<sup>1</sup> PROCARIO 93 quote  $0.345 \pm 0.006 \pm 0.016$  after correction for 2 kaon backgrounds assuming  $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We multiply by  $0.990 \pm 0.010$  to remove these corrections to  $B(h^- \pi^0 \nu_\tau)$ .

 $\Gamma(\pi^- 2\pi^0 \nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$  $\Gamma_{20}/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.26 ± 0.10 OUR FIT</b>	(Produced by HFLAV)			

<b>9.239 ± 0.086 ± 0.090</b>	31k	<sup>1</sup> SCHAEEL	05C	ALEP 1991–1995 LEP runs
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• • • We do not use the following data for averages, fits, limits, etc. • • •

9.21 ± 0.13 ± 0.11	<sup>2</sup> BUSKULIC	96	ALEP	Repl. by SCHAEEL 05C
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<sup>1</sup> See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of BUSKULIC 96  $B(h^- 2\pi^0 \nu_\tau(\text{ex. } K^0))$  and  $B(K^- 2\pi^0 \nu_\tau(\text{ex. } K^0))$  values.

 $\Gamma(\pi^- 2\pi^0 \nu_\tau(\text{ex. } K^0), \text{scalar})/\Gamma(\pi^- 2\pi^0 \nu_\tau(\text{ex. } K^0))$  $\Gamma_{21}/\Gamma_{20}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.094</b>	95	<sup>1</sup> BROWDER	00	CLEO $4.7 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau(\text{ex. } K^0))$  from scalars.

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$\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{vector}) / \Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{22}/\Gamma_{20}$
<b>&lt;0.073</b>	95	1 BROWDER	00	CLEO 4.7 fb <sup>-1</sup> $E_{cm}^{ee} = 10.6$ GeV	

<sup>1</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  from vectors.

 $\Gamma(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$ 

VALUE (units 10 <sup>-4</sup> )	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{23}/\Gamma$
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**6.5± 2.2 OUR FIT** (Produced by HFLAV)

**5.8± 2.4 OUR AVERAGE**

$5.6 \pm 2.0 \pm 1.5$  131 BARATE 99K ALEP 1991–1995 LEP runs

$9 \pm 10 \pm 3$  3 1 BATTLE 94 CLEO  $E_{cm}^{ee} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$8 \pm 2 \pm 2$  59 BUSKULIC 96 ALEP Repl. by BARATE 99K

<sup>1</sup> BATTLE 94 quote  $(14 \pm 10 \pm 3) \times 10^{-4}$  or  $< 30 \times 10^{-4}$  at 90% CL. We subtract  $(5 \pm 2) \times 10^{-4}$  to account for  $\tau^- \rightarrow K^- (K^0 \rightarrow \pi^0 \pi^0) \nu_\tau$  background.

 $\Gamma(h^- \geq 3\pi^0 \nu_\tau) / \Gamma_{\text{total}}$ 

$$\Gamma_{24}/\Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.0942\Gamma_{48} + 0.0942\Gamma_{52} + 0.3256\Gamma_{150} + 0.3256\Gamma_{152} + 0.3256\Gamma_{154} + 0.0501\Gamma_{156})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{24}/\Gamma$
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**1.34±0.07 OUR FIT** (Produced by HFLAV)

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.53 \pm 0.40 \pm 0.46$  186 DECAMP 92C ALEP Repl. by SCHAEEL 05C

$3.2 \pm 1.0 \pm 1.0$  BEHREND 90 CELL  $E_{cm}^{ee} = 35$  GeV

 $\Gamma(h^- \geq 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$ 

$$\Gamma_{25}/\Gamma = (\Gamma_{27} + \Gamma_{28} + \Gamma_{30} + 0.3256\Gamma_{150} + 0.3256\Gamma_{152} + 0.3256\Gamma_{154})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{25}/\Gamma$
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**1.25 ± 0.07 OUR FIT** (Produced by HFLAV)

**1.403±0.214±0.224** 1.1k 1 ABDALLAH 06A DLPH 1992–1995 LEP runs

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

 $\Gamma(h^- 3\pi^0 \nu_\tau) / \Gamma_{\text{total}}$ 

$$\Gamma_{26}/\Gamma = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3256\Gamma_{152})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{26}/\Gamma$
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**1.18±0.07 OUR FIT** (Produced by HFLAV)

**1.21±0.17 OUR AVERAGE** Error includes scale factor of 1.2.

$1.70 \pm 0.24 \pm 0.38$  293 ACCIARRI 95 L3 1992 LEP run

• • • We use the following data for averages but not for fits. • • •

$1.15 \pm 0.08 \pm 0.13$  1 PROCARIO 93 CLEO  $E_{cm}^{ee} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.24 \pm 0.09 \pm 0.11$  2.3k 2 BUSKULIC 96 ALEP Repl. by SCHAEEL 05C

$0.0 \begin{matrix} +1.4 \\ -0.1 \end{matrix} \begin{matrix} +1.1 \\ -0.1 \end{matrix}$  3 GAN 87 MRK2  $E_{cm}^{ee} = 29$  GeV

<sup>1</sup> PROCARIO 93 entry is obtained from  $B(h^- 3\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$ .

<sup>2</sup> BUSKULIC 96 quote  $B(h^- 3\pi^0 \nu_\tau (\text{ex. } K^0)) = 1.17 \pm 0.09 \pm 0.11$ . We add 0.07 to remove their correction for  $K^0$  backgrounds.

<sup>3</sup> Highly correlated with GAN 87  $\Gamma(\eta \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  value. Authors quote  $B(\pi^\pm 3\pi^0 \nu_\tau) + 0.67B(\pi^\pm \eta \pi^0 \nu_\tau) = 0.047 \pm 0.010 \pm 0.011$ .

 $\Gamma(h^- 3\pi^0 \nu_\tau) / \Gamma(h^- \pi^0 \nu_\tau)$ 

$$\Gamma_{26}/\Gamma_{13} = (\Gamma_{27} + \Gamma_{28} + 0.15344\Gamma_{41} + 0.15344\Gamma_{43} + 0.3256\Gamma_{152}) / (\Gamma_{14} + \Gamma_{16})$$

VALUE (units 10 <sup>-2</sup> )	DOCUMENT ID	TECN	COMMENT	$\Gamma_{26}/\Gamma_{13}$
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**4.54±0.28 OUR FIT** (Produced by HFLAV)

**4.4 ± 0.3 ± 0.5** 1 PROCARIO 93 CLEO  $E_{cm}^{ee} \approx 10.6$  GeV

<sup>1</sup> PROCARIO 93 quote  $0.041 \pm 0.003 \pm 0.005$  after correction for 2 kaon backgrounds assuming  $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We add  $0.003 \pm 0.003$  and multiply the sum by  $0.990 \pm 0.010$  to remove these corrections.

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$\Gamma(\pi^- 3\pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$					$\Gamma_{27}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.04 ± 0.07 OUR FIT</b>	(Produced by HFLAV)				NODE=S035B56 NODE=S035B56
<b>0.977±0.069±0.058</b>	6.1k	<sup>1</sup> SCHAEEL	05C ALEP	1991-1995 LEP runs	

<sup>1</sup> See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(K^- 3\pi^0 \nu_\tau (\text{ex.} K^0, \eta)) / \Gamma_{\text{total}}$					$\Gamma_{28}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.8± 2.1 OUR FIT</b>	(Produced by HFLAV)				NODE=S035B31 NODE=S035B31
<b>3.7± 2.1±1.1</b>	22	BARATE	99K ALEP	1991-1995 LEP runs	

• • • We do not use the following data for averages, fits, limits, etc. • • •

5 ± 13	<sup>1</sup> BUSKULIC	94E ALEP	Repl. by BARATE 99K
1 BUSKULIC 94E quote $B(K^- \geq 0\pi^0 \geq 0K^0 \nu_\tau) = [B(K^- \nu_\tau) + B(K^- \pi^0 \nu_\tau) + B(K^- K^0 \nu_\tau) + B(K^- \pi^0 \pi^0 \nu_\tau) + B(K^- \pi^0 K^0 \nu_\tau)] = (5 \pm 13) \times 10^{-4}$ accounting for common systematic errors in BUSKULIC 94E and BUSKULIC 94F measurements of these modes. We assume $B(K^- \geq 2K^0 \nu_\tau)$ and $B(K^- \geq 4\pi^0 \nu_\tau)$ are negligible.			

$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$					$\Gamma_{29}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.16±0.04 OUR FIT</b>	(Produced by HFLAV)				NODE=S035B79 NODE=S035B79
<b>0.16±0.05±0.05</b>		<sup>1</sup> PROCARIO	93 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$	NODE=S035B79

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.16±0.04±0.09	232	<sup>2</sup> BUSKULIC	96 ALEP	Repl. by SCHAEEL 05C
1 PROCARIO 93 quotes $B(h^- 4\pi^0 \nu_\tau) / B(h^- \pi^0 \nu_\tau) = 0.006 \pm 0.002 \pm 0.002$ . We multiply by the ARTUSO 94 result for $B(h^- \pi^0 \nu_\tau)$ to obtain $B(h^- 4\pi^0 \nu_\tau)$ . PROCARIO 93 assume $B(h^- \geq 5\pi^0 \nu_\tau)$ is small and do not correct for it.				NODE=S035B79;LINKAGE=A
2 BUSKULIC 96 quote result for $\tau^- \rightarrow h^- \geq 4\pi^0 \nu_\tau$ . We assume $B(h^- \geq 5\pi^0 \nu_\tau)$ is negligible.				NODE=S035B79;LINKAGE=D

$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex.} K^0, \eta)) / \Gamma_{\text{total}}$					$\Gamma_{30}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.11 ± 0.04 OUR FIT</b>	(Produced by HFLAV)				NODE=S035B23 NODE=S035B23
<b>0.112±0.037±0.035</b>	957	<sup>1</sup> SCHAEEL	05C ALEP	1991-1995 LEP runs	

<sup>1</sup> See footnote to SCHAEEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

$\Gamma(a_1(1260)\nu_\tau \rightarrow \pi^- \gamma \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{31}/\Gamma = (0.0022\Gamma_{20} + 0.0022\Gamma_{70}) / \Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.0±1.5 OUR FIT</b>	(Produced by HFLAV)				NODE=S035R08 NODE=S035R08

$\Gamma(K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{32}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.552±0.029 OUR FIT</b>	(Produced by HFLAV)				NODE=S035R26 NODE=S035R26 NODE=S035R26

**1.53 ± 0.04 OUR AVERAGE**

1.528±0.039±0.040	<sup>1</sup> ABBIENDI	01J OPAL	1990-1995 LEP runs
1.54 ± 0.24	ABREU	94K DLPH	LEP 1992 Z data
1.70 ± 0.12 ± 0.19	202	<sup>2</sup> BATTLE	94 CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •			
1.520±0.040±0.041	4006	<sup>3</sup> BARATE	99K ALEP 1991-1995 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.70 ± 0.05 ± 0.06	1610	<sup>4</sup> BUSKULIC	96 ALEP Repl. by BARATE 99K
1.6 ± 0.4 ± 0.2	35	AIHARA	87B TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
1.71 ± 0.29	53	MILLS	84 DLCO $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

NODE=S035B56  
NODE=S035B56

NODE=S035B56;LINKAGE=SC

NODE=S035B31  
NODE=S035B31

NODE=S035B31;LINKAGE=A

NODE=S035B79  
NODE=S035B79

NODE=S035B79;LINKAGE=A

NODE=S035B79;LINKAGE=D

NODE=S035B23  
NODE=S035B23

NODE=S035B23;LINKAGE=SC

NODE=S035R08  
NODE=S035R08

NODE=S035R08

NODE=S035R26  
NODE=S035R26  
NODE=S035R26

NOTFITTED

NOTFITTED

<sup>1</sup>The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  is 0.60.

<sup>2</sup>BATTLE 94 quote  $1.60 \pm 0.12 \pm 0.19$ . We add  $0.10 \pm 0.02$  to correct for their rejection of  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>3</sup>Not independent of BARATE 99K  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$  (ex.  $K^0$ ),  $B(K^- 3\pi^0 \nu_\tau)$  (ex.  $K^0$ )),  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

<sup>4</sup>Not independent of BUSKULIC 96  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

### $\Gamma(K^- \geq 1(\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{33}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.859 ± 0.028 OUR FIT</b> (Produced by HFLAV)				

#### 0.86 ± 0.05 OUR AVERAGE

• • • We use the following data for averages but not for fits. • • •

$0.869 \pm 0.031 \pm 0.034$	<sup>1</sup> ABBIENDI	01J OPAL	1990–1995 LEP runs	NOTFITTED
$0.69 \pm 0.25$	<sup>2</sup> ABREU	94K DLPH	LEP 1992 $Z$ data	NOTFITTED

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.2 $\pm 0.5$ $^{+0.2}_{-0.4}$	9	AIHARA	87B TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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<sup>1</sup>Not independent of ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  and  $B(\tau^- \rightarrow K^- \geq 0 \pi^0 \geq 0 K^0 \geq 0 \gamma \nu_\tau)$  values.

<sup>2</sup>Not independent of ABREU 94K  $B(K^- \nu_\tau)$  and  $B(K^- \geq 0 \text{ neutrals} \nu_\tau)$  measurements.

### $\Gamma(K_S^0(\text{particles})^- \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{34}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.943 ± 0.028 OUR FIT</b> (Produced by HFLAV)				

#### 0.918 ± 0.015 OUR AVERAGE

$0.970 \pm 0.058 \pm 0.062$	929	BARATE	98E ALEP	1991–1995 LEP runs
$0.97 \pm 0.09 \pm 0.06$	141	AKERS	94G OPAL	$E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

$0.915 \pm 0.001 \pm 0.015$	398k	<sup>1</sup> RYU	14 BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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<sup>1</sup>Not independent of RYU 14 measurements of  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$ ,  $B(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \nu_\tau)$ , and  $B(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)$ .

### $\Gamma(h^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{35}/\Gamma = (\Gamma_{36} + \Gamma_{38})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.987 ± 0.014 OUR FIT</b> (Produced by HFLAV)				

#### 0.90 ± 0.07 OUR AVERAGE

$0.855 \pm 0.036 \pm 0.073$	1242	COAN	96 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
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• • • We use the following data for averages but not for fits. • • •

1.01 $\pm 0.11 \pm 0.07$	555	<sup>1</sup> BARATE	98E ALEP	1991–1995 LEP runs
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<sup>1</sup>Not independent of BARATE 98E  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$  and  $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$  values.

### $\Gamma(\pi^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{36}/\Gamma$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.38 ± 0.14 OUR FIT</b> (Produced by HFLAV)				

#### 8.39 ± 0.22 OUR AVERAGE Error includes scale factor of 1.5. See the ideogram below.

$8.32 \pm 0.02 \pm 0.16$	158k	<sup>1</sup> RYU	14 BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$9.33 \pm 0.68 \pm 0.49$	377	ABBIENDI	00C OPAL	1991–1995 LEP runs
$9.28 \pm 0.45 \pm 0.34$	937	<sup>2</sup> BARATE	99K ALEP	1991–1995 LEP runs
$9.5 \pm 1.5 \pm 0.6$		<sup>3</sup> ACCIARRI	95F L3	1991–1993 LEP runs

• • • We use the following data for averages but not for fits. • • •

8.55 $\pm 1.17 \pm 0.66$	509	<sup>4</sup> BARATE	98E ALEP	1991–1995 LEP runs
7.04 $\pm 0.41 \pm 0.72$		<sup>5</sup> COAN	96 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

8.08 $\pm 0.04 \pm 0.26$	53k	EPIFANOV	07 BELL	Repl. by RYU 14
7.9 $\pm 1.0 \pm 0.9$	98	<sup>6</sup> BUSKULIC	96 ALEP	Repl. by BARATE 99k

NODE=S035R26;LINKAGE=AJ

NODE=S035R26;LINKAGE=C

NODE=S035R26;LINKAGE=9K

NODE=S035R26;LINKAGE=AA

NODE=S035R27

NODE=S035R27

NODE=S035R27

NODE=S035R27;LINKAGE=AB

NODE=S035R27;LINKAGE=A

NODE=S035B43

NODE=S035B43

NODE=S035B43

NOTFITTED

NODE=S035B43;LINKAGE=RY

NOTFITTED

NODE=S035B67;LINKAGE=B9

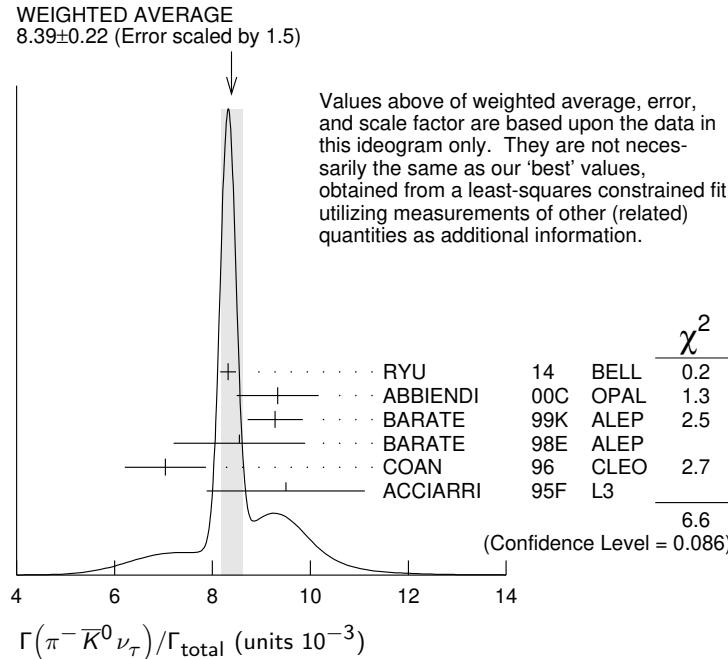
NODE=S035B32

NODE=S035B32

NOTFITTED

NOTFITTED

- 1 RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.
- 2 BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.
- 3 ACCIARRI 95F do not identify  $\pi^- / K^-$  and assume  $B(K^- K^0 \nu_\tau) = (0.29 \pm 0.12)\%$ .
- 4 BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. Not independent of BARATE 98E  $B(K^0 \nu_\tau)$  value.
- 5 Not independent of COAN 96  $B(h^- K^0 \nu_\tau)$  and  $B(K^- K^0 \nu_\tau)$  measurements.
- 6 BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.



### $\Gamma(\pi^- \bar{K}^0 (\text{non-}K^*(892)^-) \nu_\tau) / \Gamma_{\text{total}}$ $\Gamma_{37}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>5.4±2.1</b>		1 EPIFANOV	07 BELL	$351 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<17 95 ACCIARRI 95F L3 1991–1993 LEP runs

1 EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) / B(\tau^- \rightarrow K_S^0 \pi^- \nu_\tau) = 0.933 \pm 0.027$ . We multiply their  $B(\tau^- \rightarrow \bar{K}^0 \pi^- \nu_\tau)$  by  $[1 - (0.933 \pm 0.027)]$  to obtain this result.

NODE=S035B52  
NODE=S035B52

### $\Gamma(K^- K^0 \nu_\tau) / \Gamma_{\text{total}}$ $\Gamma_{38}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.86±0.34 OUR FIT</b>	(Produced by HFLAV)			

#### 14.83±0.35 OUR AVERAGE

14.78±0.22±0.40	29k	1 LEES	18B BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
14.80±0.14±0.54	33k	2 RYU	14 BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
16.2 ± 2.1 ± 1.1	150	3 BARATE	99K ALEP	1991–1995 LEP runs
15.8 ± 4.2 ± 1.7	46	4 BARATE	98E ALEP	1991–1995 LEP runs
15.1 ± 2.1 ± 2.2	111	COAN	96 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

26 ± 9 ± 2 13 5 BUSKULIC 96 ALEP Repl. by BARATE 99K

1 LEES 18B reconstructs  $K_S^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

2 RYU 14 reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

3 BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

4 BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

5 BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

NODE=S035R46  
NODE=S035R46

OCCUR=2

### $\Gamma(K^- K^0 \geq 0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$ $\Gamma_{39}/\Gamma = (\Gamma_{38} + \Gamma_{43})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.299±0.007 OUR FIT</b>	(Produced by HFLAV)			
<b>0.330±0.055±0.039</b>	124	ABBIENDI	00C OPAL	1991–1995 LEP runs

NODE=S035R46;LINKAGE=C

NODE=S035R46;LINKAGE=A

NODE=S035R46;LINKAGE=9K

NODE=S035R46;LINKAGE=B9

NODE=S035R46;LINKAGE=B6

NODE=S035C38

NODE=S035C38

$\Gamma(h^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{40}/\Gamma = (\Gamma_{41} + \Gamma_{43})/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.532 ± 0.013 OUR FIT</b>	(Produced by HFLAV)			
<b>0.50 ± 0.06 OUR AVERAGE</b>	Error includes scale factor of 1.2.			
0.562 ± 0.050 ± 0.048	264	COAN	96	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •				
0.446 ± 0.052 ± 0.046	157	<sup>1</sup> BARATE	98E	ALEP 1991–1995 LEP runs
1 Not independent of BARATE 98E $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$ and $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$ values.				

NODE=S035B68  
NODE=S035B68

$\Gamma(\pi^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{41}/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.382 ± 0.013 OUR FIT</b>	(Produced by HFLAV)			
<b>0.383 ± 0.014 OUR AVERAGE</b>				
0.386 ± 0.004 ± 0.014	27k	<sup>1</sup> RYU	14	BELL $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
0.347 ± 0.053 ± 0.037	299	<sup>2</sup> BARATE	99K	ALEP 1991–1995 LEP runs
0.294 ± 0.073 ± 0.037	142	<sup>3</sup> BARATE	98E	ALEP 1991–1995 LEP runs
0.41 ± 0.12 ± 0.03		<sup>4</sup> ACCIARRI	95F	L3 1991–1993 LEP runs
• • • We use the following data for averages but not for fits. • • •				
0.417 ± 0.058 ± 0.044		<sup>5</sup> COAN	96	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.32 ± 0.11 ± 0.05	23	<sup>6</sup> BUSKULIC	96	ALEP Repl. by BARATE 99K
1 RYU 14 reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
2 BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.				
3 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
4 ACCIARRI 95F do not identify $\pi^- / K^-$ and assume $B(K^- K^0 \pi^0 \nu_\tau) = (0.05 \pm 0.05)\%$ .				
5 Not independent of COAN 96 $B(h^- \bar{K}^0 \pi^0 \nu_\tau)$ and $B(K^- K^0 \pi^0 \nu_\tau)$ measurements.				
6 BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.				

NOTFITTED  
NODE=S035B68;LINKAGE=B9

$\Gamma(\bar{K}^0 \rho^- \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{42}/\Gamma$			
VALUE (%)	DOCUMENT ID	TECN	COMMENT	
<b>0.22 ± 0.05 OUR AVERAGE</b>				
0.250 ± 0.057 ± 0.044		<sup>1</sup> BARATE	99K	ALEP 1991–1995 LEP runs
0.188 ± 0.054 ± 0.038		<sup>2</sup> BARATE	98E	ALEP 1991–1995 LEP runs
1 BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in hadron calorimeter. They determine the $\bar{K}^0 \rho^-$ fraction in $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ decays to be $(0.72 \pm 0.12 \pm 0.10)$ and multiply their $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$ measurement by this fraction to obtain the quoted result.				
2 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays. They determine the $\bar{K}^0 \rho^-$ fraction in $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$ decays to be $(0.64 \pm 0.09 \pm 0.10)$ and multiply their $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$ measurement by this fraction to obtain the quoted result.				

NODE=S035B33;LINKAGE=C  
NODE=S035B33;LINKAGE=9K  
NODE=S035B33;LINKAGE=B9  
NODE=S035B33;LINKAGE=A  
NODE=S035B33;LINKAGE=B  
NODE=S035B33;LINKAGE=B6

$\Gamma(K^- K^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{43}/\Gamma$			
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>15.0 ± 0.7 OUR FIT</b>	(Produced by HFLAV)			
<b>14.9 ± 0.7 OUR AVERAGE</b>				
14.96 ± 0.20 ± 0.74	8.3k	<sup>1</sup> RYU	14	BELL $669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
14.3 ± 2.5 ± 1.5	78	<sup>2</sup> BARATE	99K	ALEP 1991–1995 LEP runs
15.2 ± 7.6 ± 2.1	15	<sup>3</sup> BARATE	98E	ALEP 1991–1995 LEP runs
14.5 ± 3.6 ± 2.0	32	COAN	96	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
10 ± 5 ± 3	5	<sup>4</sup> BUSKULIC	96	ALEP Repl. by BARATE 99K
1 RYU 14 reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
2 BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.				
3 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.				
4 BUSKULIC 96 measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.				

NODE=S035C10  
NODE=S035C10

NODE=S035C10;LINKAGE=9K

NODE=S035C10;LINKAGE=B9

NODE=S035B34  
NODE=S035B34

NODE=S035B34;LINKAGE=A  
NODE=S035B34;LINKAGE=9K  
NODE=S035B34;LINKAGE=B9  
NODE=S035B34;LINKAGE=B6

$\Gamma(\pi^- \bar{K}^0 \geq 1 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{44}/\Gamma = (\Gamma_{41} + \Gamma_{45})/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.408 ± 0.025 OUR FIT</b>	(Produced by HFLAV)			
<b>0.324 ± 0.074 ± 0.066</b>	148	ABBIENDI	00C	OPAL 1991–1995 LEP runs

NODE=S035C37  
NODE=S035C37

$\Gamma(\pi^- \bar{K}^0 \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$						$\Gamma_{45}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.26±0.23 OUR FIT</b>		(Produced by HFLAV)				
<b>0.26±0.24</b>		1 BARATE	99R ALEP	1991–1995 LEP runs		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
<0.66	95	17	2 BARATE	99K ALEP	1991–1995 LEP runs	
$0.58 \pm 0.33 \pm 0.14$	5	3 BARATE	98E ALEP	1991–1995 LEP runs		
1 BARATE 99R combine the BARATE 98E and BARATE 99K measurements to obtain this value.						
2 BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter.						
3 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.						

NODE=S035B98  
NODE=S035B98

$\Gamma(K^- K^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$						$\Gamma_{46}/\Gamma$
VALUE	CL%		DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.16 × 10<sup>-3</sup></b>	95	1	BARATE	99R ALEP	1991–1995 LEP runs	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
< $0.18 \times 10^{-3}$	95	2	BARATE	99K ALEP	1991–1995 LEP runs	
< $0.39 \times 10^{-3}$	95	3	BARATE	98E ALEP	1991–1995 LEP runs	
1 BARATE 99R combine the BARATE 98E and BARATE 99K bounds to obtain this value.						
2 BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in hadron calorimeter.						
3 BARATE 98E reconstruct $K^0$ 's by using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.						

NODE=S035B99  
NODE=S035B99

$\Gamma(\pi^- K^0 \bar{K}^0 \nu_\tau) / \Gamma_{\text{total}}$						$\Gamma_{47}/\Gamma = (\Gamma_{48} + \Gamma_{49} + \Gamma_{50}) / \Gamma$
VALUE (%)	EVTS		DOCUMENT ID	TECN	COMMENT	
<b>0.155±0.024 OUR FIT</b>		(Produced by HFLAV)				
• • • We use the following data for averages but not for fits. • • •						
<b>0.153±0.030±0.016</b>	74	1	BARATE	98E ALEP	1991–1995 LEP runs	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
$0.31 \pm 0.12 \pm 0.04$		2	ACCIARRI	95F L3	1991–1993 LEP runs	
1 BARATE 98E obtain this value by adding twice their $B(\pi^- K_S^0 K_S^0 \nu_\tau)$ value to their $B(\pi^- K_S^0 K_L^0 \nu_\tau)$ value.						
2 ACCIARRI 95F assume $B(\pi^- K_S^0 K_S^0 \nu) = B(\pi^- K_S^0 K_L^0 \nu) = 1/2 B(\pi^- K_S^0 K_L^0 \nu)$ .						

NODE=S035B51  
NODE=S035B51

$\Gamma(\pi^- K_S^0 K_S^0 \nu_\tau) / \Gamma_{\text{total}}$						$\Gamma_{48}/\Gamma$
Bose-Einstein correlations might make the mixing fraction different than 1/4.						
<b>2.35±0.06 OUR FIT</b>		(Produced by HFLAV)				
<b>2.32±0.06 OUR AVERAGE</b>						
2.33±0.03±0.09	6.7k	RYU	14	BELL	$669 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
2.31±0.04±0.08	5.0k	1 LEES	12Y	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$2.6 \pm 1.0 \pm 0.5$	6	BARATE	98E ALEP	1991–1995 LEP runs		
$2.3 \pm 0.5 \pm 0.3$	42	COAN	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$	

NOTFITTED  
NODE=S035B69  
NODE=S035B69  
NODE=S035B69

$\Gamma(\pi^- K_S^0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$						$\Gamma_{49}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS		DOCUMENT ID	TECN	COMMENT	
<b>10.8±2.4 OUR FIT</b>		(Produced by HFLAV)				
<b>10.1±2.3±1.3</b>	68	BARATE	98E ALEP	1991–1995 LEP runs		

NODE=S035C1  
NODE=S035C1

$\Gamma(\pi^- K_L^0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}$						$\Gamma_{50}/\Gamma = \Gamma_{48}/\Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID				
<b>2.35±0.06 OUR FIT</b>		(Produced by HFLAV)				

NODE=S035R06  
NODE=S035R06

$\Gamma(\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$						$\Gamma_{51}/\Gamma = (\Gamma_{52} + \Gamma_{56} + \Gamma_{57}) / \Gamma$
VALUE (units $10^{-4}$ )		DOCUMENT ID	TECN	COMMENT		
<b>3.6±1.2 OUR FIT</b>		(Produced by HFLAV)				
• • • We use the following data for averages but not for fits. • • •						
<b>3.1±2.3</b>		1 BARATE	99R ALEP	1991–1995 LEP runs		
1 BARATE 99R combine BARATE 98E $\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$ and $\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$ measurements to obtain this value.						

NODE=S035C44  
NODE=S035C44

NOTFITTED  
NODE=S035C44;LINKAGE=A

$\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{52}/\Gamma$
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL% EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>1.82±0.21 OUR FIT</b> (Produced by HFLAV)					
<b>1.80±0.21 OUR AVERAGE</b>					
2.00±0.22±0.20	303	RYU	14	BELL 669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
1.60±0.20±0.22	409	1 LEES	12Y	BABR 468 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<20	95	BARATE	98E	ALEP 1991–1995 LEP runs	
1 The correlation coefficient between this measurement and the LEES 12Y $\Gamma(\tau^- \rightarrow \pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$ one is 0.0828.					
$\Gamma(K^* - K^0 \pi^0 \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{53}/\Gamma$
<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>		
<b>10.8±1.4±1.5</b>	RYU	14	BELL	669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$\Gamma(f_1(1285) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{54}/\Gamma$
<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>		
<b>6.8±1.3±0.7</b>	RYU	14	BELL	669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$\Gamma(f_1(1420) \pi^- \nu_\tau \rightarrow \pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{55}/\Gamma$
<u>VALUE (units <math>10^{-6}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>		
<b>2.4±0.5±0.6</b>	RYU	14	BELL	669 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{56}/\Gamma$
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>3.2±1.2 OUR FIT</b> (Produced by HFLAV)					
<b>3.1±1.1±0.5</b>	11	BARATE	98E	ALEP 1991–1995 LEP runs	
$\Gamma(\pi^- K_L^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{57}/\Gamma = \Gamma_{52}/\Gamma$
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>DOCUMENT ID</u>				
<b>1.82±0.21 OUR FIT</b> (Produced by HFLAV)					
$\Gamma(K^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{58}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;6.3 × 10<math>^{-7}</math></b>	90	LEES	12Y	BABR 468 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$\Gamma(K^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{59}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;4.0 × 10<math>^{-7}</math></b>	90	LEES	12Y	BABR 468 fb $^{-1}$ $E_{\text{cm}}^{\text{ee}}=10.6$ GeV	
$\Gamma(K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{60}/\Gamma$
<u>VALUE (%)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;0.17</b>	95	TSCHIRHART	88	HRS $E_{\text{cm}}^{\text{ee}}=29$ GeV	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.27	90	BELTRAMI	85	HRS $E_{\text{cm}}^{\text{ee}}=29$ GeV	
$\Gamma(K^0 h^+ h^- h^- \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{61}/\Gamma$
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>2.5±2.0 OUR FIT</b> (Produced by HFLAV)					
<b>2.3±1.9±0.7</b>	6	1 BARATE	98E	ALEP 1991–1995 LEP runs	
1 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.					
$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{62}/\Gamma$
$\Gamma_{62}/\Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.4247\Gamma_{52} + 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2805\Gamma_{150} + 0.2805\Gamma_{152} + 0.2805\Gamma_{154} + 0.2628\Gamma_{156} + 0.7259\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma$					
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>15.20±0.06 OUR FIT</b> (Produced by HFLAV)					
<b>14.8 ± 0.4 OUR AVERAGE</b>					
14.4 ± 0.6 ± 0.3		ADEVA	91F	L3 $E_{\text{cm}}^{\text{ee}}=88.3\text{--}94.3$ GeV	
15.0 ± 0.4 ± 0.3		BEHREND	89B	CELL $E_{\text{cm}}^{\text{ee}}=14\text{--}47$ GeV	
15.1 ± 0.8 ± 0.6		AIHARA	87B	TPC $E_{\text{cm}}^{\text{ee}}=29$ GeV	

• • • We do not use the following data for averages, fits, limits, etc. • • •

13.5 ± 0.3 ± 0.3		ABACHI	89B	HRS	$E_{cm}^{ee} = 29$ GeV
12.8 ± 1.0 ± 0.7		<sup>1</sup> BURCHAT	87	MRK2	$E_{cm}^{ee} = 29$ GeV
12.1 ± 0.5 ± 1.2		RUCKSTUHL	86	DLCO	$E_{cm}^{ee} = 29$ GeV
12.8 ± 0.5 ± 0.8	1420	SCHMIDKE	86	MRK2	$E_{cm}^{ee} = 29$ GeV
15.3 ± 1.1 <sup>+1.3</sup> <sub>-1.6</sub>	367	ALTHOFF	85	TASS	$E_{cm}^{ee} = 34.5$ GeV
13.6 ± 0.5 ± 0.8		BARTEL	85F	JADE	$E_{cm}^{ee} = 34.6$ GeV
12.2 ± 1.3 ± 3.9		<sup>2</sup> BERGER	85	PLUT	$E_{cm}^{ee} = 34.6$ GeV
13.3 ± 0.3 ± 0.6		FERNANDEZ	85	MAC	$E_{cm}^{ee} = 29$ GeV
24 ± 6	35	BRANDELIK	80	TASS	$E_{cm}^{ee} = 30$ GeV
32 ± 5	692	<sup>3</sup> BACINO	78B	DLCO	$E_{cm}^{ee} = 3.1\text{--}7.4$ GeV
35 ± 11		<sup>3</sup> BRANDELIK	78	DASP	Assumes $V\text{--}A$ decay
18 ± 6.5	33	<sup>3</sup> JAROS	78	LGW	$E_{cm}^{ee} > 6$ GeV

<sup>1</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>2</sup> Not independent of BERGER 85  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)/\Gamma_{total}$ ,  $\Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{total}$ ,  $\Gamma(h^- \geq 1$  neutrals  $\nu_\tau)/\Gamma_{total}$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{total}$ , and therefore not used in the fit.

<sup>3</sup> Low energy experiments are not in average or fit because the systematic errors in background subtraction are judged to be large.

### $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^- \text{)}/\Gamma_{total} \quad \Gamma_{63}/\Gamma$

$$\Gamma_{63}/\Gamma = (\Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2805\Gamma_{150} + 0.2805\Gamma_{152} + 0.2805\Gamma_{154} + 0.499\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>14.55 ± 0.06 OUR FIT</b>	(Produced by HFLAV)			
<b>14.61 ± 0.06 OUR AVERAGE</b>				

14.556 ± 0.105 ± 0.076		<sup>1</sup> ACHARD	01D	L3	1992–1995 LEP runs
14.96 ± 0.09 ± 0.22	10.4k	AKERS	95Y	OPAL	1991–1994 LEP runs

• • • We use the following data for averages but not for fits. • • •

14.652 ± 0.067 ± 0.086		SCHAEL	05C	ALEP	1991–1995 LEP runs
14.569 ± 0.093 ± 0.048	23k	<sup>2</sup> ABREU	01M	DLPH	1992–1995 LEP runs
14.22 ± 0.10 ± 0.37		<sup>3</sup> BAlest	95C	CLEO	$E_{cm}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

15.26 ± 0.26 ± 0.22		ACTON	92H	OPAL	Repl. by AKERS 95Y
13.3 ± 0.3 ± 0.8		<sup>4</sup> ALBRECHT	92D	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
14.35 <sup>+0.40</sup> <sub>-0.45</sub> ± 0.24		DECAMP	92C	ALEP	1989–1990 LEP runs

<sup>1</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.19$  respectively.

<sup>2</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow \text{1-prong})$  and  $B(\tau \rightarrow \text{5-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>3</sup> Not independent of BAlest 95C  $B(h^- h^- h^+ \nu_\tau)$  and  $B(h^- h^- h^+ \pi^0 \nu_\tau)$  values, and BORTOLETTO 93  $B(h^- h^- h^+ 2\pi^0 \nu_\tau)/B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau)$  value.

<sup>4</sup> This ALBRECHT 92D value is not independent of their  $\Gamma(\mu^-\bar{\nu}_\mu\nu_\tau)\Gamma(e^-\bar{\nu}_e\nu_\tau)/\Gamma_{total}^2$  value.

### $\Gamma(h^- h^- h^+ \nu_\tau)/\Gamma_{total} \quad \Gamma_{64}/\Gamma$

$$\Gamma_{64}/\Gamma = (0.34598\Gamma_{36} + 0.34598\Gamma_{38} + \Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.499\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.80 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			

• • • We use the following data for averages but not for fits. • • •

<b>7.6 ± 0.1 ± 0.5</b>	7.5k	<sup>1</sup> ALBRECHT	96E	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
------------------------	------	-----------------------	-----	-----	--------------------------------------

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.92 ± 0.10 ± 0.09	11.2k	<sup>2</sup> BUSKULIC	96	ALEP	Repl. by SCHAEL 05C
9.49 ± 0.36 ± 0.63		DECAMP	92C	ALEP	Repl. by SCHAEL 05C
8.7 ± 0.7 ± 0.3	694	<sup>3</sup> BEHREND	90	CELL	$E_{cm}^{ee} = 35$ GeV
7.0 ± 0.3 ± 0.7	1566	<sup>4</sup> BAND	87	MAC	$E_{cm}^{ee} = 29$ GeV
6.7 ± 0.8 ± 0.9		<sup>5</sup> BURCHAT	87	MRK2	$E_{cm}^{ee} = 29$ GeV
6.4 ± 0.4 ± 0.9		<sup>6</sup> RUCKSTUHL	86	DLCO	$E_{cm}^{ee} = 29$ GeV
7.8 ± 0.5 ± 0.8	890	SCHMIDKE	86	MRK2	$E_{cm}^{ee} = 29$ GeV
8.4 ± 0.4 ± 0.7	1255	<sup>6</sup> FERNANDEZ	85	MAC	$E_{cm}^{ee} = 29$ GeV
9.7 ± 2.0 ± 1.3		BEHREND	84	CELL	$E_{cm}^{ee} = 14,22$ GeV

NODE=S035R31;LINKAGE=K

NODE=S035R31;LINKAGE=G

NODE=S035R31;LINKAGE=E

NODE=S035B63

NODE=S035B63

NODE=S035B63

NOTFITTED

NOTFITTED

NOTFITTED

NODE=S035B63;LINKAGE=CH

NODE=S035B63;LINKAGE=M1

NODE=S035B63;LINKAGE=B

NODE=S035B63;LINKAGE=A

NODE=S035R28

NODE=S035R28

NODE=S035R28

NOTFITTED

- <sup>1</sup> ALBRECHT 96E not independent of ALBRECHT 93C  $\Gamma(h^- h^- h^+ \nu_\tau)$  (ex.  $K^0$ )  $\times$   $\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}^2$  value.
- <sup>2</sup> BUSKULIC 96 quote  $B(h^- h^- h^+ \nu_\tau)$  (ex.  $K^0$ ) =  $9.50 \pm 0.10 \pm 0.11$ . We add 0.42 to remove their  $K^0$  correction and reduce the systematic error accordingly.
- <sup>3</sup> BEHREND 90 subtract 0.3% to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution to measured events.
- <sup>4</sup> BAND 87 subtract for charged kaon modes; not independent of FERNANDEZ 85 value.
- <sup>5</sup> BURCHAT 87 value is not independent of SCHMIDKE 86 value.
- <sup>6</sup> Value obtained by multiplying paper's  $R = B(h^- h^- h^+ \nu_\tau) / B(3\text{-prong})$  by  $B(3\text{-prong}) = 0.143$  and subtracting 0.3% for  $K^*(892)$  background.

### $\Gamma(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0)) / \Gamma_{\text{total}}$

$\Gamma_{65}/\Gamma$

VALUE (%)	$\Gamma_{65}/\Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.499\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179}) / \Gamma$			
	EVTS	DOCUMENT ID	TECN	COMMENT

**9.46 ± 0.05 OUR FIT** (Produced by HFLAV)

**9.44 ± 0.14 OUR AVERAGE** Error includes scale factor of 1.4. See the ideogram below.

$9.317 \pm 0.090 \pm 0.082$     12.2k    <sup>1</sup> ABDALLAH    06A DLPH 1992–1995 LEP runs  
 $9.51 \pm 0.07 \pm 0.20$     37.7k    BALEST    95C CLEO  $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

$9.87 \pm 0.10 \pm 0.24$     <sup>2</sup> AKERS    95Y OPAL 1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

$9.50 \pm 0.10 \pm 0.11$     11.2k    <sup>3</sup> BUSKULIC    96 ALEP Repl. by SCHAEEL 05C

<sup>1</sup> See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup> Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0) / B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

<sup>3</sup> Not independent of BUSKULIC 96  $B(h^- h^- h^+ \nu_\tau)$  value.

NODE=S035R28;LINKAGE=BB

NODE=S035R28;LINKAGE=EE

NODE=S035R28;LINKAGE=E

NODE=S035R28;LINKAGE=B

NODE=S035R28;LINKAGE=D

NODE=S035R28;LINKAGE=F

NODE=S035B62

NODE=S035B62

NODE=S035B62

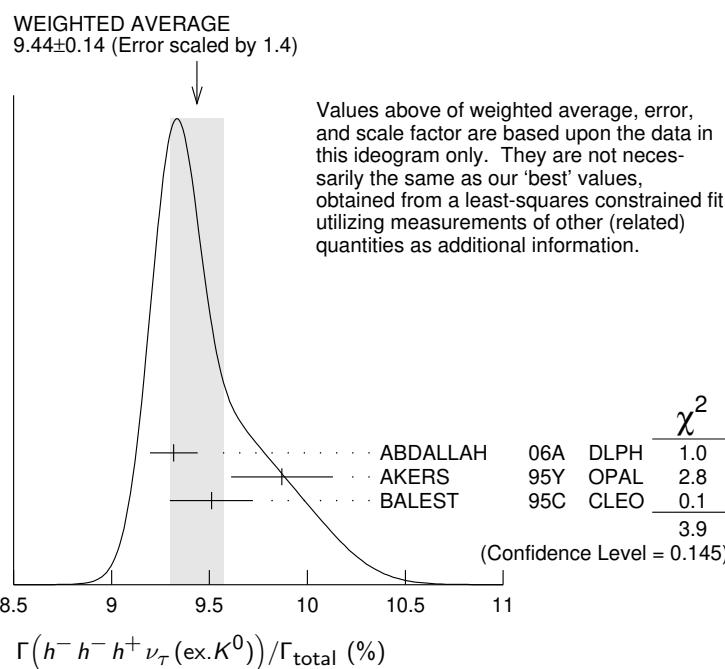
NOTFITTED

NOTFITTED

NODE=S035B62;LINKAGE=AH

NODE=S035B62;LINKAGE=A

NODE=S035B62;LINKAGE=B



### $\Gamma(h^- h^- h^+ \nu_\tau \text{ (ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$

$\Gamma_{65}/\Gamma_{63}$

$$\Gamma_{65}/\Gamma_{63} = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.499\Gamma_{170} + 0.0153\Gamma_{178} + 0.0153\Gamma_{179}) / (0.4247\Gamma_{52} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2805\Gamma_{150} + 0.2302\Gamma_{151} + 0.2805\Gamma_{152} + 0.2805\Gamma_{154} + 0.1131\Gamma_{156} + 0.3256\Gamma_{160} + 0.499\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182})$$

VALUE (units $10^{-2}$ )	DOCUMENT ID	TECN	COMMENT
<b>64.98 ± 0.31 OUR FIT</b> (Produced by HFLAV)			

**66.0 ± 0.4 ± 1.4**    AKERS    95Y OPAL 1991–1994 LEP runs

NODE=S035B64

NODE=S035B64

NODE=S035B64

$\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$	$\Gamma_{66} / \Gamma$	NODE=S035B71			
$\Gamma_{66} / \Gamma = (\Gamma_{70} + \Gamma_{97} + \Gamma_{106} + 0.499\Gamma_{170}) / \Gamma$		NODE=S035B71			
<u>VALUE (%)</u>	<u>DOCUMENT ID</u>	NODE=S035B71			
<b>9.43±0.05 OUR FIT</b>	(Produced by HFLAV)				
$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$	$\Gamma_{67} / \Gamma = (0.34598\Gamma_{36} + \Gamma_{70} + 0.0153\Gamma_{178}) / \Gamma$	NODE=S035C18			
<u>VALUE (%)</u>	<u>DOCUMENT ID</u>	NODE=S035C18			
<b>9.31±0.05 OUR FIT</b>	(Produced by HFLAV)				
$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$	$\Gamma_{68} / \Gamma = (\Gamma_{70} + 0.0153\Gamma_{178}) / \Gamma$	NODE=S035C19			
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	NODE=S035C19
<b>9.02±0.05 OUR FIT</b>	(Produced by HFLAV)				
<b>8.77±0.13 OUR AVERAGE</b>	Error includes scale factor of 1.1.				
8.42±0.00 <sup>+0.26</sup> <sub>-0.25</sub>	8.9M	<sup>1</sup> LEE	10	BELL	$666 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
8.83±0.01±0.13	1.6M	<sup>2</sup> AUBERT	08	BABR	$342 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
9.13±0.05±0.46	43k	<sup>3</sup> BRIERE	03	CLE3	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
1 Quoted statistical error is 0.003%. Correlation matrix for LEE 10 branching fractions:					
(1) $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$					
(2) $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$					
(3) $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$					
(4) $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$					
(1) (2) (3)					
(2) 0.175					
(3) 0.049 0.080					
(4) -0.053 0.035 -0.008					
2 Correlation matrix for AUBERT 08 branching fractions:					
(1) $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$					
(2) $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}}$					
(3) $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$					
(4) $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$					
(1) (2) (3)					
(2) 0.544					
(3) 0.390 0.177					
(4) 0.031 0.093 0.087					
3 47% correlated with BRIERE 03 $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$ and 71% correlated with $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$ because of a common 5% normalization error.					
$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0), \text{non-axial vector}) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$					
$\Gamma_{69} / \Gamma_{68} = \Gamma_{69} / (\Gamma_{70} + 0.0153\Gamma_{177})$					
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;0.261</b>	95	<sup>1</sup> ACKERSTAFF	97R	OPAL	1992–1994 LEP runs
1 Model-independent limit from structure function analysis on contribution to $B(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0))$ from non-axial vectors.					
$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau(\text{ex. } K^0, \omega)) / \Gamma_{\text{total}}$	$\Gamma_{70} / \Gamma$	NODE=S035C20			
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	NODE=S035C20
<b>8.99 ±0.05 OUR FIT</b>	(Produced by HFLAV)				
<b>9.041±0.060±0.076</b>	29k	<sup>1</sup> SCHAEL	05C	ALEP	1991–1995 LEP runs
1 See footnote to SCHAEL 05C $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ measurement for correlations with other measurements.					
$\Gamma(h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}}$	$\Gamma_{71} / \Gamma$	NODE=S035R30			
$\Gamma_{71} / \Gamma = (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.4247\Gamma_{52} + \Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{103} + \Gamma_{107} + 0.2805\Gamma_{150} + 0.2805\Gamma_{152} + 0.2805\Gamma_{154} + 0.2926\Gamma_{156} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.9078\Gamma_{180}) / \Gamma$		NODE=S035R30			
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	NODE=S035R30
<b>5.29±0.05 OUR FIT</b>	(Produced by HFLAV)				
• • • We do not use the following data for averages, fits, limits, etc. • • •					

5.6 ± 0.7 ± 0.3	352	1 BEHREND	90	CELL	$E_{\text{cm}}^{\text{ee}} = 35 \text{ GeV}$	NOTFITTED
4.2 ± 0.5 ± 0.9	203	2 ALBRECHT	87L	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	
6.1 ± 0.8 ± 0.9		3 BURCHAT	87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
7.6 ± 0.4 ± 0.9		4,5 RUCKSTUHL	86	DLCO	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	NOTFITTED
4.7 ± 0.5 ± 0.8	530	6 SCHMIDKE	86	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
5.6 ± 0.4 ± 0.7		5 FERNANDEZ	85	MAC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$	
6.2 ± 2.3 ± 1.7		BEHREND	84	CELL	$E_{\text{cm}}^{\text{ee}} = 14,22 \text{ GeV}$	
1 BEHREND 90 value is not independent of BEHREND 90 $B(3h\nu_\tau \geq 1 \text{ neutrals}) + B(5\text{-prong})$ .						
2 ALBRECHT 87L measure the product of branching ratios $B(3\pi^\pm \pi^0 \nu_\tau) B((e\bar{\nu} \text{ or } \mu\bar{\nu} \text{ or } \pi \text{ or } K \text{ or } \rho) \nu_\tau) = 0.029$ and use the PDG 86 values for the second branching ratio which sum to $0.69 \pm 0.03$ to get the quoted value.						
3 BURCHAT 87 value is not independent of SCHMIDKE 86 value.						
4 Contributions from kaons and from $>1\pi^0$ are subtracted. Not independent of (3-prong + $0\pi^0$ ) and (3-prong + $\geq 0\pi^0$ ) values.						
5 Value obtained using paper's $R = B(h^- h^- h^+ \nu_\tau)/B(3\text{-prong})$ and current $B(3\text{-prong}) = 0.143$ .						
6 Not independent of SCHMIDKE 86 $h^- h^- h^+ \nu_\tau$ and $h^- h^- h^+(\geq 0\pi^0)\nu_\tau$ values.						

### $\Gamma(h^- h^- h^+ \geq 1\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ $\Gamma_{72}/\Gamma$

$$\Gamma_{72}/\Gamma = (\Gamma_{78} + \Gamma_{85} + \Gamma_{86} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{150} + 0.2302\Gamma_{152} + 0.2302\Gamma_{154} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.9078\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.09 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			

#### 5.10 ± 0.12 OUR AVERAGE

• • • We use the following data for averages but not for fits. • • •

5.106 ± 0.083 ± 0.103	10.1k	1 ABDALLAH	06A	DLPH	1992–1995 LEP runs
5.09 ± 0.10 ± 0.23		2 AKERS	95Y	OPAL	1991–1994 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.95 ± 0.29 ± 0.65	570	DECAMP	92C	ALEP	Repl. by SCHAEEL 05C
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1 See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

2 Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau (\text{ex. } K^0))/B(h^- h^- h^+ \geq 0 \text{ neutrals} \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

### $\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{73}/\Gamma$

$$\Gamma_{73}/\Gamma = (0.34598\Gamma_{41} + 0.34598\Gamma_{43} + \Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.76 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.45 ± 0.09 ± 0.07	6.1k	1 BUSKULIC	96	ALEP	Repl. by SCHAEEL 05C
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1 BUSKULIC 96 quote  $B(h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) = 4.30 \pm 0.09 \pm 0.09$ . We add 0.15 to remove their  $K^0$  correction and reduce the systematic error accordingly.

### $\Gamma(h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ $\Gamma_{74}/\Gamma$

$$\Gamma_{74}/\Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
-----------	------	-------------	------	---------

<b>4.57 ± 0.05 OUR FIT</b>	(Produced by HFLAV)			
<b>4.45 ± 0.14 OUR AVERAGE</b>	Error includes scale factor of 1.2.			

4.545 ± 0.106 ± 0.103	8.9k	1 ABDALLAH	06A	DLPH	1992–1995 LEP runs
4.23 ± 0.06 ± 0.22	7.2k	BALEST	95C	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

1 See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

### $\Gamma(h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$ $\Gamma_{75}/\Gamma = (\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152})/\Gamma$

VALUE (%)	DOCUMENT ID
<b>2.79 ± 0.07 OUR FIT</b>	(Produced by HFLAV)

NOTFITTED

NOTFITTED

NODE=S035R30;LINKAGE=E

NODE=S035R30;LINKAGE=C

NODE=S035R30;LINKAGE=D

NODE=S035R30;LINKAGE=A

NODE=S035R30;LINKAGE=Q

NODE=S035R30;LINKAGE=S

NODE=S035B78

NODE=S035B78

NODE=S035B78

NODE=S035B78;LINKAGE=AH

NODE=S035B78;LINKAGE=A

NODE=S035B76

NODE=S035B76

NODE=S035B76;LINKAGE=A

NODE=S035B53

NODE=S035B53

NODE=S035B53

NODE=S035B53;LINKAGE=AH

NODE=S035B54

NODE=S035B54

$\Gamma(\pi^-\pi^+\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ 

$$\Gamma_{76}/\Gamma = (0.34598\Gamma_{41} + \Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)

DOCUMENT ID

**4.62±0.05 OUR FIT** (Produced by HFLAV) $\Gamma_{76}/\Gamma$ 

NODE=S035C22

NODE=S035C22

NODE=S035C22

 $\Gamma(\pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ 

$$\Gamma_{77}/\Gamma = (\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})/\Gamma$$

VALUE (%)

EVTS

DOCUMENT ID

TECN COMMENT

**4.49 ±0.05 OUR FIT** (Produced by HFLAV)**4.55 ±0.13 OUR AVERAGE** Error includes scale factor of 1.6.4.598±0.057±0.064 16k <sup>1</sup>SCHAEL 05C ALEP 1991-1995 LEP runs  
4.19 ±0.10 ±0.21 <sup>2</sup>EDWARDS 00A CLEO 4.7 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6$  GeV

<sup>1</sup>SCHAEL 05C quote (4.590±0.057±0.064)%. We add 0.008% to remove their correction for  $\tau^- \rightarrow \pi^-\pi^0\nu_\tau \rightarrow \pi^-\pi^0\pi^+\pi^-\nu_\tau$  decays. See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>2</sup>EDWARDS 00A quote  $(4.19 \pm 0.10) \times 10^{-2}$  with a 5% systematic error.

 $\Gamma_{77}/\Gamma$ 

NODE=S035C23

NODE=S035C23

NODE=S035C23

 $\Gamma(\pi^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex. } K^0,\omega))/\Gamma_{\text{total}}$  $\Gamma_{78}/\Gamma$ 

VALUE (%)

DOCUMENT ID

**2.74±0.07 OUR FIT** (Produced by HFLAV) $\Gamma(h^-\rho\pi^0\nu_\tau)/\Gamma(h^-\bar{h}^+h^+\pi^0\nu_\tau)$  $\Gamma_{79}/\Gamma_{73}$ 

VALUE

EVTS

DOCUMENT ID

TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.30±0.04±0.02 393 ALBRECHT 91D ARG  $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$  GeV $\Gamma(h^-\rho^+h^-\nu_\tau)/\Gamma(h^-\bar{h}^+h^+\pi^0\nu_\tau)$  $\Gamma_{80}/\Gamma_{73}$ 

VALUE

EVTS

DOCUMENT ID

TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.10±0.03±0.04 142 ALBRECHT 91D ARG  $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$  GeV $\Gamma(h^-\rho^=h^+\nu_\tau)/\Gamma(h^-\bar{h}^+h^+\pi^0\nu_\tau)$  $\Gamma_{81}/\Gamma_{73}$ 

VALUE

EVTS

DOCUMENT ID

TECN COMMENT

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.26±0.05±0.01 370 ALBRECHT 91D ARG  $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$  GeV $\Gamma(h^-\bar{h}^+h^+\geq 2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$  $\Gamma_{82}/\Gamma$ 

$$\Gamma_{82}/\Gamma = (\Gamma_{85} + \Gamma_{86} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$$

VALUE (%)

EVTS

DOCUMENT ID

TECN COMMENT

**0.517±0.031 OUR FIT** (Produced by HFLAV)**0.561±0.068±0.095** 1.3k <sup>1</sup>ABDALLAH 06A DLPH 1992–1995 LEP runs

<sup>1</sup>See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^-\nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

 $\Gamma(h^-\bar{h}^+h^+2\pi^0\nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{83}/\Gamma$ 

$$\Gamma_{83}/\Gamma = (0.4247\Gamma_{48} + \Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$$

VALUE (%)

DOCUMENT ID

**0.505±0.031 OUR FIT** (Produced by HFLAV) $\Gamma(h^-\bar{h}^+h^+2\pi^0\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$  $\Gamma_{84}/\Gamma$ 

$$\Gamma_{84}/\Gamma = (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180})/\Gamma$$

VALUE (%)

EVTS

DOCUMENT ID

TECN COMMENT

**0.495±0.031 OUR FIT** (Produced by HFLAV)**0.435±0.030±0.035** 2.6k <sup>1</sup>SCHAEL 05C ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.50 ±0.07 ±0.07 1.8k BUSKULIC 96 ALEP Repl. by SCHAEL 05C

<sup>1</sup>SCHAEL 05C quote  $(0.392 \pm 0.030 \pm 0.035)\%$ . We add 0.043% to remove their correction for  $\tau^- \rightarrow \pi^-\eta\pi^0\nu_\tau \rightarrow \pi^-\pi^+\pi^-\pi^0\nu_\tau$  and  $\tau^- \rightarrow K^*(892)^-\eta\nu_\tau \rightarrow K^-\pi^+\pi^-2\pi^0\nu_\tau$  decays. See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^-\bar{\nu}_e\nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

NODE=S035B59

NODE=S035B59

NODE=S035B59

NODE=S035B77

NODE=S035B77

NODE=S035B77

NODE=S035C03

NODE=S035C03

NODE=S035C03

NODE=S035B0

NODE=S035B0

NODE=S035C03

NODE=S035C03

NODE=S035C03

NODE=S035B77

NODE=S035B77

NODE=S035B77

NODE=S035B59

NODE=S035B59

NODE=S035B59

NODE=S035B59;LINKAGE=SC

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) \quad \Gamma_{84}/\Gamma_{62}$$

$$\begin{aligned} \Gamma_{84}/\Gamma_{62} = & (\Gamma_{85} + 0.2302\Gamma_{150} + 0.2302\Gamma_{154} + 0.892\Gamma_{180}) / (0.34598\Gamma_{36} + \\ & 0.34598\Gamma_{38} + 0.34598\Gamma_{41} + 0.34598\Gamma_{43} + 0.4247\Gamma_{48} + 0.6920\Gamma_{49} + 0.8494\Gamma_{52} + \\ & 0.6920\Gamma_{56} + 0.6534\Gamma_{61} + \Gamma_{70} + \Gamma_{78} + \Gamma_{85} + \Gamma_{89} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + \\ & 0.2805\Gamma_{150} + 0.2302\Gamma_{151} + 0.2805\Gamma_{152} + 0.2805\Gamma_{154} + 0.3759\Gamma_{156} + 0.3256\Gamma_{160} + \\ & 0.7259\Gamma_{170} + 0.9078\Gamma_{178} + 0.9078\Gamma_{179} + 0.9078\Gamma_{180} + 0.892\Gamma_{182}) \end{aligned}$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.26 ± 0.20 OUR FIT</b>	(Produced by HFLAV)			
<b>3.4 ± 0.2 ± 0.3</b>	668	BORTOLETTO93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

NODE=S035B25

NODE=S035B25

NODE=S035B25

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta)) / \Gamma_{\text{total}} \quad \Gamma_{85}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>10 ± 4 OUR FIT</b>	(Produced by HFLAV)

NODE=S035B72

NODE=S035B72

$$\Gamma(h^- h^- h^+ 3\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{86}/\Gamma = (0.4247\Gamma_{52} + \Gamma_{87} + 0.1131\Gamma_{156}) / \Gamma$$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.13 ± 0.30 OUR FIT</b>		(Produced by HFLAV)			

**2.2 ± 0.3 ± 0.4**      139      ANASTASSOV 01      CLEO       $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ 

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 4.9	95	SCHAEL	05C	ALEP	1991-1995 LEP runs
$2.85 \pm 0.56 \pm 0.51$	57	ANDERSON	97	CLEO	Repl. by ANAS-TASSOV 01
11 ± 4 ± 5	440	<sup>1</sup> BUSKULIC	96	ALEP	Repl. by SCHAEL 05C
<sup>1</sup> BUSKULIC 96 state their measurement is for $B(h^- h^- h^+ \geq 3\pi^0 \nu_\tau)$ . We assume that $B(h^- h^- h^+ \geq 4\pi^0 \nu_\tau)$ is very small.					

NODE=S035B57

NODE=S035B57

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{87}/\Gamma$$

$$\Gamma_{87}/\Gamma = (\Gamma_{89} + 0.2302\Gamma_{151} + 0.3256\Gamma_{160} + 0.892\Gamma_{182}) / \Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.94 ± 0.30 OUR FIT</b>	(Produced by HFLAV)		

NODE=S035B57;LINKAGE=B

• • • We use the following data for averages but not for fits. • • •	
<b>2.07 ± 0.18 ± 0.37</b>	<sup>1</sup> LEES      12X BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C85

NODE=S035C85

NODE=S035C85

1 Not independent of LEES 12X $\Gamma(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0)) / \Gamma$ , $\Gamma(\tau^- \rightarrow \eta\pi^-\pi^0\pi^0\nu_\tau) / \Gamma$ , $\Gamma(\tau^- \rightarrow \pi^-\omega\pi^0\nu_\tau) / \Gamma$ , and $\Gamma(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau) / \Gamma$ values.	
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NOTFITTED

NODE=S035C85;LINKAGE=LE

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{88}/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.69 ± 0.08 ± 0.43</b>	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C86

NODE=S035C86

$$\Gamma(2\pi^- \pi^+ 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}} \quad \Gamma_{89}/\Gamma$$

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	COMMENT
<b>1.4 ± 2.7 OUR FIT</b>	(Produced by HFLAV)		

NODE=S035C87

NODE=S035C87

<b>1.0 ± 0.8 ± 3.0</b>	<sup>1</sup> LEES      12X BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
------------------------	--

1 LEES 12X measurement corresponds to the lower limit of  $< 5.8 \times 10^{-5}$  at 90% CL.

NODE=S035C87;LINKAGE=LE

$$\Gamma(K^- h^+ h^- \geq 0 \text{ neutrals} \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{90}/\Gamma$$

$$\Gamma_{90}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + \Gamma_{106} + \Gamma_{107} + 0.2805\Gamma_{152} + \\ 0.499\Gamma_{170} + 0.9078\Gamma_{179}) / \Gamma$$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.629 ± 0.014 OUR FIT</b>		(Produced by HFLAV)		

NODE=S035R34

NODE=S035R34

NODE=S035R34

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{91}/\Gamma = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{179}) / \Gamma$$

VALUE (%)	DOCUMENT ID
<b>0.437 ± 0.007 OUR FIT</b>	(Produced by HFLAV)

NODE=S035C31

NODE=S035C31

$$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{91}/\Gamma_{68}$$

$$\Gamma_{91}/\Gamma_{68} = (\Gamma_{97} + \Gamma_{106} + 0.0153\Gamma_{179}) / (\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.84 ± 0.08 OUR FIT</b>	(Produced by HFLAV)			

NODE=S035C32

NODE=S035C32

NODE=S035C32

$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau(\text{ex.} K^0))/\Gamma_{\text{total}}$ 

$$\Gamma_{92}/\Gamma = (\Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>8.6±1.2 OUR FIT</b>	(Produced by HFLAV)

 $\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau(\text{ex.} K^0))/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau(\text{ex.} K^0))$ 

$$\Gamma_{92}/\Gamma_{77}$$

$$\Gamma_{92}/\Gamma_{77} = (\Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/(\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.91±0.26 OUR FIT</b>	(Produced by HFLAV)			
<b>2.61±0.45±0.42</b>	719	RICHICHI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ 

$$\Gamma_{93}/\Gamma$$

$$\Gamma_{93}/\Gamma = (0.34598\Gamma_{38} + 0.34598\Gamma_{43} + \Gamma_{97} + \Gamma_{103} + 0.2805\Gamma_{152} + 0.9078\Gamma_{179})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.477±0.014 OUR FIT</b>	(Produced by HFLAV)			

**0.58 ±0.15 ±0.12**      20      <sup>1</sup> BAUER      94      TPC       $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.22 <sup>+0.16</sup><sub>-0.13</sub> ±0.05      9      <sup>2</sup> MILLS      85      DLCO       $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>1</sup> We multiply 0.58% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Error correlated with MILLS 85 ( $K \pi \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.

 $\Gamma(K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau(\text{ex.} K^0))/\Gamma_{\text{total}}$ 

$$\Gamma_{94}/\Gamma$$

$$\Gamma_{94}/\Gamma = (\Gamma_{97} + \Gamma_{103} + 0.2302\Gamma_{152} + 0.9078\Gamma_{179})/\Gamma$$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.373±0.013 OUR FIT</b>	(Produced by HFLAV)		
<b>0.30 ±0.05 OUR AVERAGE</b>			

• • • We use the following data for averages but not for fits. • • •

0.343±0.073±0.031      ABBIENDI      00D OPAL      1990–1995 LEP runs  
 0.275±0.064      <sup>1</sup> BARATE      98 ALEP      1991–1995 LEP runs

<sup>1</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  values.

 $\Gamma(K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ 

$$\Gamma_{95}/\Gamma = (0.34598\Gamma_{38} + \Gamma_{97} + 0.0153\Gamma_{179})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>0.345±0.007 OUR FIT</b>	(Produced by HFLAV)

 $\Gamma(K^- \pi^+ \pi^- \nu_\tau(\text{ex.} K^0))/\Gamma_{\text{total}}$ 

$$\Gamma_{96}/\Gamma = (\Gamma_{97} + 0.0153\Gamma_{179})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.293±0.007 OUR FIT</b>	(Produced by HFLAV)			
<b>0.290±0.018 OUR AVERAGE</b>				Error includes scale factor of 2.4. See the ideogram below.

0.330±0.001<sup>+0.016</sup><sub>-0.017</sub> 794k      <sup>1</sup> LEE      10 BELL       $666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$   
 0.273±0.002±0.009 70k      <sup>2</sup> AUBERT      08 BABR       $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$   
 0.415±0.053±0.040 269      ABBIENDI      04J OPAL      1991–1995 LEP runs  
 0.384±0.014±0.038 3.5k      <sup>3</sup> BRIERE      03 CLE3       $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$   
 0.214±0.037±0.029      BARATE      98 ALEP      1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

0.346±0.023±0.056 158      <sup>4</sup> RICHICHI      99 CLEO       $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.360±0.082±0.048      ABBIENDI      00D OPAL      1990–1995 LEP runs

<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.} K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau(\text{ex.} K^0))/\Gamma_{\text{total}}$  value.

<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.} K^0))/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

<sup>3</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

<sup>4</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau(\text{ex.} K^0))/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau(\text{ex.} K^0)), \Gamma(\tau^- \rightarrow$

$$\Gamma_{92}/\Gamma$$

NODE=S035C33

NODE=S035C33

NODE=S035C33

NODE=S035C34

NODE=S035C34

NODE=S035C34

NODE=S035R41

NODE=S035R41

NODE=S035R41

NODE=S035R41;LINKAGE=A

NODE=S035R41;LINKAGE=M

NODE=S035C40

NODE=S035C40

NODE=S035C40

NOTFITTED

NOTFITTED

NODE=S035C40;LINKAGE=B8

NODE=S035C6

NODE=S035C6

NODE=S035C21

NODE=S035C21

NOTFITTED

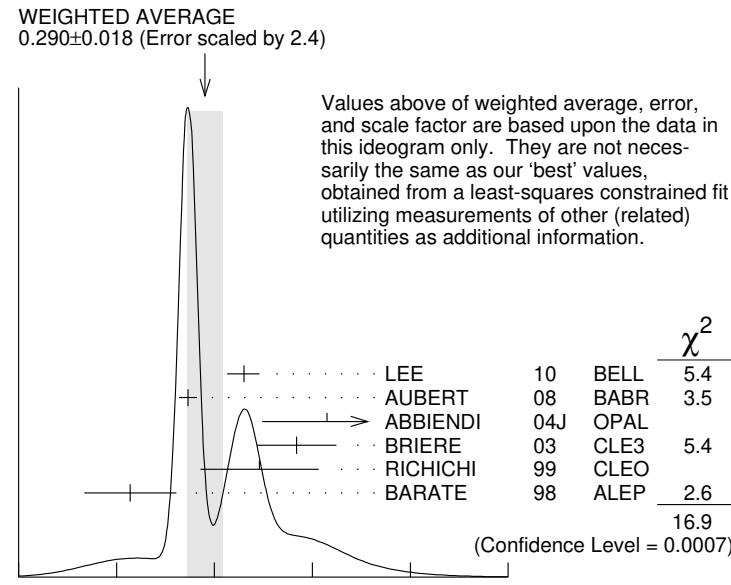
NODE=S035C21;LINKAGE=LE

NODE=S035C21;LINKAGE=AU

NODE=S035C21;LINKAGE=A

NODE=S035C21;LINKAGE=R1

$K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BAEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.



$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}} (\%)$$

$$\Gamma_{96}/\Gamma_{68} = (\Gamma_{97} + 0.0153\Gamma_{179})/(\Gamma_{70} + 0.0153\Gamma_{178})$$

$$\Gamma_{96}/\Gamma_{68}$$

NODE=S035C76

NODE=S035C76

NODE=S035C76

VALUE (units  $10^{-2}$ )

EVTS

DOCUMENT ID

TECN

COMMENT

**3.25 +/- 0.07 OUR FIT**

(Produced by HFLAV)

• • • We use the following data for averages but not for fits. • • •

$$3.92 \pm 0.02^{+0.15}_{-0.16}$$

794k

<sup>1</sup> LEE

10

BELL

$666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.

NOTFITTED

NODE=S035C76;LINKAGE=LE

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega))/\Gamma_{\text{total}}$$

$$\Gamma_{97}/\Gamma$$

NODE=S035R04

NODE=S035R04

VALUE (units  $10^{-3}$ )

DOCUMENT ID

**2.93 +/- 0.07 OUR FIT**

(Produced by HFLAV)

$$\Gamma(K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$$

$$\Gamma_{98}/\Gamma_{96}$$

NODE=S035C52

NODE=S035C52

VALUE

DOCUMENT ID

TECN

COMMENT

**0.48 +/- 0.14 +/- 0.10**

<sup>1</sup> ASNER

00B

CLEO

$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$$0.39 \pm 0.14$$

<sup>2</sup> BARATE

99R

ALEP 1991–1995 LEP runs

<sup>1</sup> ASNER 00B assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays proceed only through  $K\rho$  and  $K^*\pi$  intermediate states. They assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances, and assume  $B(K_1(1270) \rightarrow K^*(892)\pi) = (16 \pm 5)\%$ ,  $B(K_1(1270) \rightarrow K\rho) = (42 \pm 6)\%$ , and  $B(K_1(1400) \rightarrow K\rho) = 0$ .

NODE=S035C52;LINKAGE=BA

<sup>2</sup> BARATE 99R assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)$  decays proceed only through  $K\rho$  and  $K^*\pi$  intermediate states. The quoted error is statistical only.

NODE=S035C52;LINKAGE=AB

$$\Gamma(K^- \pi^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$$

$$\Gamma_{99}/\Gamma$$

NODE=S035C7

NODE=S035C7

NODE=S035C7

$$\Gamma_{99}/\Gamma = (0.34598\Gamma_{43} + \Gamma_{103} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/\Gamma$$

VALUE (units  $10^{-4}$ )

DOCUMENT ID

**13.1 +/- 1.2 OUR FIT**

(Produced by HFLAV)

$\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0))/\Gamma_{\text{total}}$ 

$$\Gamma_{100}/\Gamma = (\Gamma_{103} + 0.2302\Gamma_{152} + 0.892\Gamma_{179})/\Gamma$$

 $\Gamma_{100}/\Gamma$ 

NODE=S035C25

NODE=S035C25

NODE=S035C25

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>7.9±1.2 OUR FIT</b>	(Produced by HFLAV)			
<b>7.3±1.2 OUR AVERAGE</b>				

7.4±0.8±1.1	<sup>1</sup> ARMS	05	CLE3	$7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
6.1±3.9±1.8	BARATE	98	ALEP	1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

7.5±2.6±1.8	<sup>2</sup> RICHICHI	99	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<17	95	ABBIENDI	00D	OPAL	1990–1995 LEP runs
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<sup>1</sup> Not independent of ARMS 05  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega)) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-\omega\nu_\tau) / \Gamma_{\text{total}}$  values.

<sup>2</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^-h^+\pi^-\nu_\tau(\text{ex.}K^0)) / \Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$ ,  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau) / \Gamma(\tau^- \rightarrow \pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  and BAEST 95C  $\Gamma(\tau^- \rightarrow h^-h^-h^+\nu_\tau(\text{ex.}K^0)) / \Gamma_{\text{total}}$  values.

 $\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\eta))/\Gamma_{\text{total}}$ 

$$\Gamma_{101}/\Gamma = (\Gamma_{103} + 0.892\Gamma_{179})/\Gamma$$

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>7.6±1.2 OUR FIT</b>	(Produced by HFLAV)

 $\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega))/\Gamma_{\text{total}}$  $\Gamma_{102}/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.7±0.5±0.8</b>	833	ARMS	05	CLE3 $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(K^-\pi^+\pi^-\pi^0\nu_\tau(\text{ex.}K^0,\omega,\eta))/\Gamma_{\text{total}}$  $\Gamma_{103}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b>3.9±1.4 OUR FIT</b>	(Produced by HFLAV)

 $\Gamma(K^-\pi^+K^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{104}/\Gamma$ 

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.09</b>	95	BAUER	94	TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

 $\Gamma(K^-\pi^+K^-\geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{105}/\Gamma = (\Gamma_{106} + \Gamma_{107})/\Gamma$ 

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.1496±0.0033 OUR FIT</b>	(Produced by HFLAV)			

**0.203 ± 0.031 OUR AVERAGE**

0.159 ± 0.053 ± 0.020	ABBIENDI	00D	OPAL	1990–1995 LEP runs
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0.15 <sup>+0.09</sup> <sub>-0.07</sub>	± 0.03	4	<sup>1</sup> BAUER	94 TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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• • • We use the following data for averages but not for fits. • • •

0.238 ± 0.042	<sup>2</sup> BARATE	98	ALEP	1991–1995 LEP runs
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<sup>1</sup> We multiply 0.15% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>2</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^-K^+\pi^-\pi^0\nu_\tau) / \Gamma_{\text{total}}$  values.

 $\Gamma(K^-K^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$  $\Gamma_{106}/\Gamma$ 

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.435±0.027 OUR FIT</b>	(Produced by HFLAV)			

**1.43 ± 0.07 OUR AVERAGE** Error includes scale factor of 2.4. See the ideogram below.

1.55 ± 0.01 <sup>+0.06</sup> <sub>-0.05</sub>	108k	<sup>1</sup> LEE	10	BELL $666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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1.346 ± 0.010 ± 0.036	18k	<sup>2</sup> AUBERT	08	BABR $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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1.55 ± 0.06 ± 0.09	932	<sup>3</sup> BRIERE	03	CLE3 $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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1.63 ± 0.21 ± 0.17	BARATE	98	ALEP	1991–1995 LEP runs
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• • • We use the following data for averages but not for fits. • • •

0.87 ± 0.56 ± 0.40	ABBIENDI	00D	OPAL	1990–1995 LEP runs
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1.45 ± 0.13 ± 0.28	2.3k	<sup>4</sup> RICHICHI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

2.2 <sup>+1.7</sup> <sub>-1.1</sub>	± 0.5	9	<sup>5</sup> MILLS	85 DLCO $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
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NOTFITTED

NOTFITTED

NOTFITTED

NODE=S035C25;LINKAGE=AR

NODE=S035C25;LINKAGE=R1

NODE=S035C54

NODE=S035C54

NODE=S035R05

NODE=S035R05

NODE=S035B37

NODE=S035B37

NOTFITTED

NODE=S035B37;LINKAGE=A

NODE=S035B37;LINKAGE=B8

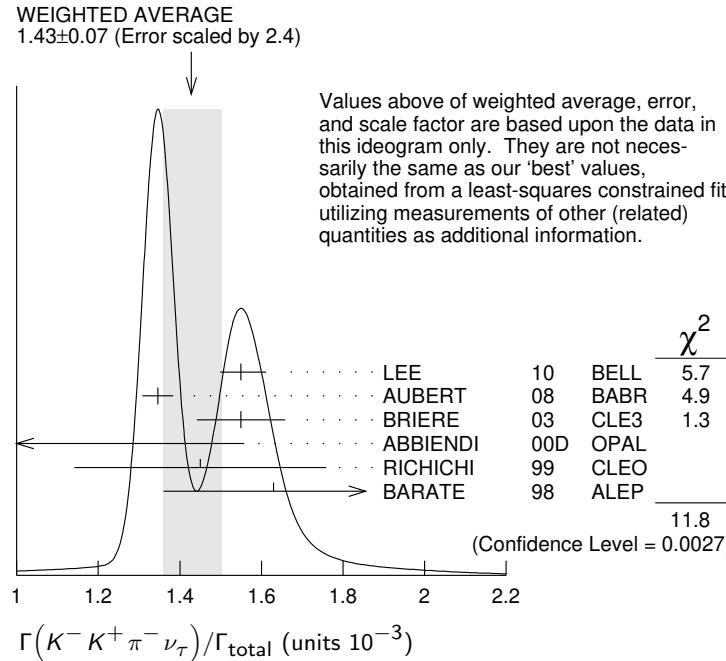
NODE=S035R40

NODE=S035R40

NOTFITTED

NOTFITTED

- <sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))$  value.
- <sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.
- <sup>3</sup> 71% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau$  because of a common 5% normalization error.
- <sup>4</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))$  and BAILEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  values.
- <sup>5</sup> Error correlated with MILLS 85 ( $K \pi \pi \pi^0 \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.



$$\Gamma(K^- K^+ \pi^- \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))$$

$$\Gamma_{106}/\Gamma_{68} = \Gamma_{106}/(\Gamma_{70} + 0.0153\Gamma_{178})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.592±0.030 OUR FIT</b> (Produced by HFLAV)				

**1.83 ± 0.05 OUR AVERAGE**

1.60 ± 0.15 ± 0.30 2.3k RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

1.84 ± 0.01 ± 0.05 108k <sup>1</sup> LEE 10 BELL 666 fb<sup>-1</sup>  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  values.

$$\Gamma_{106}/\Gamma_{68}$$

NODE=S035C35

NODE=S035C35

NODE=S035C35

$$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$$

$$\Gamma_{107}/\Gamma$$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.61±0.18 OUR FIT** (Produced by HFLAV)

**0.60±0.18 OUR AVERAGE**

0.55±0.14±0.12 48 ARMS 05 CLE3 7.6 fb<sup>-1</sup>,  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

7.5 ± 2.9 ± 1.5 BARATE 98 ALEP 1991–1995 LEP runs

• • • We use the following data for averages but not for fits. • • •

3.3 ± 1.8 ± 0.7 158 <sup>1</sup> RICHICHI 99 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<27 95 ABBIENDI 00D OPAL 1990–1995 LEP runs

NOTFITTED

NODE=S035C35;LINKAGE=LE

NODE=S035C8

NODE=S035C8

NOTFITTED

NODE=S035C8;LINKAGE=R1

<sup>1</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex.} K^0))$  and BAILEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex.} K^0)) / \Gamma_{\text{total}}$  values.

$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))$ 

$$\Gamma_{107}/\Gamma_{77} = \Gamma_{107}/(\Gamma_{78} + 0.892\Gamma_{178} + 0.0153\Gamma_{180})$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.14±0.04 OUR FIT</b>	(Produced by HFLAV)			
<b>0.79±0.44±0.16</b>	158	<sup>1</sup> RICHICHI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> RICHICHI 99 also quote a 95%CL upper limit of 0.0157 for this measurement.

 $\Gamma_{107}/\Gamma_{77}$ 

NODE=S035C36

NODE=S035C36

NODE=S035C36

 $\Gamma(K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{108}/\Gamma = 0.499 \Gamma_{170}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.2 ± 0.8 OUR FIT</b>	Error includes scale factor of 5.4. (Produced by HFLAV)				
<b>2.1 ± 0.8 OUR AVERAGE</b>	Error includes scale factor of 5.4.				

$3.29 \pm 0.17^{+0.19}_{-0.20}$	3.2k	<sup>1</sup> LEE	10	BELL	$666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$1.58 \pm 0.13 \pm 0.12$	275	<sup>2</sup> AUBERT	08	BABR	$342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 3.7	90	BRIERE	03	CLE3	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
< 19	90	BARATE	98	ALEP	1991–1995 LEP runs

<sup>1</sup> See footnote to LEE 10  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements. Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  value.

<sup>2</sup> See footnote to AUBERT 08  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

 $\Gamma(K^- K^+ K^- \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  $\Gamma_{108}/\Gamma_{68}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$3.90 \pm 0.02^{+0.22}_{-0.23}$	3.2k	<sup>1</sup> LEE	10	BELL	$666 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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• • • Not independent of LEE 10  $\Gamma(\tau^- \rightarrow K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

 $\Gamma(K^- K^+ K^- \nu_\tau (\text{ex. } \phi)) / \Gamma_{\text{total}}$  $\Gamma_{109}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;2.5 × 10<sup>-6</sup></b>	90	AUBERT	08	BABR $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(K^- K^+ K^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{110}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.8 × 10<sup>-6</sup></b>	90	ARMS	05	CLE3 $7.6 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{111}/\Gamma$ 

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.25</b>	95	BAUER	94	TPC $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

 $\Gamma(e^- e^- e^+ \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{112}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.8±1.4±0.4</b>	5	ALAM	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{113}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.2</b>	90	ALAM	96	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\pi^- e^- e^+ \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{114}/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>seen</b>	400	<sup>1</sup> JIN	19	BELL $562 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.46 \pm 0.13 \pm 0.21$	400	<sup>1</sup> JIN	19	BELL axial-vector, $562 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$3.01 \pm 0.27 \pm 0.43$	400	<sup>1</sup> JIN	19	BELL vector, $562 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> JIN 19 measures  $B(\tau^- \rightarrow \pi^- e^- e^+ \nu_\tau (m_{\pi^- e^- e^+} > 1.05 \text{ GeV}/c^2)) = (5.90 \pm 0.53 \pm 0.86) \times 10^{-6}$ , which is only sensitive to the structure-dependent contribution, and assumes that the decay proceeds with either a pure axial-vector current or a pure vector current to obtain the two respective branching fraction measurements for this mode, which are 100% correlated.

NODE=S035C36

NODE=S035C36

NODE=S035C36

NODE=S035C36;LINKAGE=R1

NODE=S035C9

NODE=S035C9

NODE=S035C9;LINKAGE=LE

NODE=S035C9;LINKAGE=AU

NODE=S035C77

NODE=S035C77

NODE=S035C77;LINKAGE=LE

NODE=S035C67

NODE=S035C67

NODE=S035C62

NODE=S035C62

NODE=S035B36

NODE=S035B36

NODE=S035B65

NODE=S035B65

OCCUR=4

NODE=S035P01;LINKAGE=A

$\Gamma(\pi^- \mu^- \mu^+ \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{115}/\Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$<1.14 \times 10^{-5}$	90	JIN	19	BELL	$562 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035P00  
NODE=S035P00

$\Gamma(3h^- 2h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_s^0 \rightarrow \pi^- \pi^+ \text{) ("5-prong")})/\Gamma_{\text{total}}$					$\Gamma_{116}/\Gamma$
$\Gamma_{116}/\Gamma = (\Gamma_{117} + \Gamma_{123})/\Gamma$					

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.099 ± 0.004 OUR FIT</b>		(Produced by HFLAV)		

<b>0.107 ± 0.007 OUR AVERAGE</b>	Error includes scale factor of 1.1.			
0.170 ± 0.022 ± 0.026	1	ACHARD	01D	L3    1992–1995 LEP runs
0.097 ± 0.005 ± 0.011	419	GIBAUT	94B	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
0.102 ± 0.029	13	BYLSMA	87	HRS $E_{\text{cm}}^{ee} = 29 \text{ GeV}$
• • • We use the following data for averages but not for fits. • • •				

0.093 ± 0.009 ± 0.012	SCHAEL	05C	ALEP	1991–1995 LEP runs
0.115 ± 0.013 ± 0.006	112	2 ABREU	01M	DLPH    1992–1995 LEP runs
0.119 ± 0.013 ± 0.008	119	3 ACKERSTAFF	99E	OPAL    1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.26 ± 0.06 ± 0.05	ACTON	92H	OPAL	$E_{\text{cm}}^{ee} = 88.2\text{--}94.2 \text{ GeV}$
0.10 ± 0.05 ± 0.03	DECAMP	92C	ALEP	1989–1990 LEP runs
0.16 ± 0.13 ± 0.04	BEHREND	89B	CELL	$E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$
0.3 ± 0.1 ± 0.2	BARTEL	85F	JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
0.13 ± 0.04	10	BELTRAMI	85	HRS    Repl. by BYLSMA 87
0.16 ± 0.08 ± 0.04	4	BURCHAT	85	MRK2 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$
1.0 ± 0.4	10	BEHREND	82	CELL    Repl. by BEHREND 89B

1 The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau^- \rightarrow \text{"1-prong"})$  and  $B(\tau^- \rightarrow \text{"3-prong"})$  are  $-0.082$  and  $-0.19$  respectively.

2 The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau^- \rightarrow \text{1-prong})$  and  $B(\tau^- \rightarrow \text{3-prong})$  are  $-0.08$  and  $-0.08$  respectively.

3 Not independent of ACKERSTAFF 99E  $B(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau \text{ (ex. } K^0\text{)})$  and  $B(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau \text{ (ex. } K^0\text{)})$  measurements.

$\Gamma(3h^- 2h^+ \nu_\tau \text{ (ex. } K^0\text{)})/\Gamma_{\text{total}}$					$\Gamma_{117}/\Gamma = (\Gamma_{118} + \Gamma_{120} + 0.0153 \Gamma_{185})/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>8.29 ± 0.31 OUR FIT</b>		(Produced by HFLAV)			

NOTFITTED  
NOTFITTED  
NOTFITTED

$\Gamma_{117}/\Gamma = (\Gamma_{118} + \Gamma_{120} + 0.0153 \Gamma_{185})/\Gamma$				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>8.32 ± 0.35 OUR AVERAGE</b>				

NODE=S035R33;LINKAGE=CH  
NODE=S035R33;LINKAGE=M1  
NODE=S035R33;LINKAGE=A

9.7 ± 1.5 ± 0.5	96	1 ABDALLAH	06A	DLPH    1992–1995 LEP runs
7.2 ± 0.9 ± 1.2	165	2 SCHAEL	05C	ALEP    1991–1995 LEP runs
9.1 ± 1.4 ± 0.6	97	ACKERSTAFF	99E	OPAL    1991–1995 LEP runs
7.7 ± 0.5 ± 0.9	295	GIBAUT	94B	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
6.4 ± 2.3 ± 1.0	12	ALBRECHT	88B	ARG $E_{\text{cm}}^{ee} = 10 \text{ GeV}$
5.1 ± 2.0	7	BYLSMA	87	HRS $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

• • • We use the following data for averages but not for fits. • • •

8.56 ± 0.05 ± 0.42	34k	AUBERT,B	05W	BABR $232 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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NOTFITTED

• • • We do not use the following data for averages, fits, limits, etc. • • •

8.0 ± 1.1 ± 1.3	58	BUSKULIC	96	ALEP    Repl. by SCHAEL 05C
6.7 ± 3.0	5	3 BELTRAMI	85	HRS    Repl. by BYLSMA 87

NODE=S035R38;LINKAGE=AH  
NODE=S035R38;LINKAGE=SC  
NODE=S035R38;LINKAGE=B

1 See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

2 See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$  measurement for correlations with other measurements.

3 The error quoted is statistical only.

$\Gamma(3\pi^- 2\pi^+ \nu_\tau \text{ (ex. } K^0, \omega\text{)})/\Gamma_{\text{total}}$					$\Gamma_{118}/\Gamma = (\Gamma_{119} + \Gamma_{173})/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>8.27 ± 0.31 OUR FIT</b>		(Produced by HFLAV)			

NODE=S035C88  
NODE=S035C88

• • • We use the following data for averages but not for fits. • • •

<b>8.33 ± 0.04 ± 0.43</b>	<sup>1</sup> LEES	12X	BABR	$468 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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NOTFITTED

1 Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow f_1(1285) \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \nu_\tau)/\Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \nu_\tau \text{ (ex. } K^0, \omega, f_1(1285)\text{)})/\Gamma$  values.

$\Gamma(3\pi^- 2\pi^+ \nu_\tau \text{ (ex. } K^0, \omega, f_1(1285)\text{)})/\Gamma_{\text{total}}$					$\Gamma_{119}/\Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>7.75 ± 0.30 OUR FIT</b>		(Produced by HFLAV)			

NODE=S035C89  
NODE=S035C89

<b>7.68 ± 0.04 ± 0.40</b>	69k	LEES	12X	BABR	$468 \text{ fb}^{-1}, E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
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$\Gamma(K^- 2\pi^- 2\pi^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$					$\Gamma_{120} / \Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b>0.6±1.2 OUR FIT</b>	(Produced by HFLAV)				NODE=S035C94 NODE=S035C94
<b>0.6±0.5±1.1</b>	1 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$		NODE=S035C94;LINKAGE=LE

1 LEES 12X measurement corresponds to the lower limit of  $< 2.4 \times 10^{-6}$  at 90% CL.

$\Gamma(K^+ 3\pi^- \pi^+ \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{121} / \Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;5.0 × 10<sup>-6</sup></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	NODE=S035C95 NODE=S035C95

$\Gamma(K^+ K^- 2\pi^- \pi^+ \nu_\tau) / \Gamma_{\text{total}}$					$\Gamma_{122} / \Gamma$
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;4.5 × 10<sup>-7</sup></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	NODE=S035C96 NODE=S035C96

$\Gamma(3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$					$\Gamma_{123} / \Gamma = (\Gamma_{124} + \Gamma_{127}) / \Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.65±0.11 OUR FIT</b>	(Produced by HFLAV)				NODE=S035R39 NODE=S035R39

#### 1.74±0.27 OUR AVERAGE

1.6 ± 1.2 ± 0.6	13	1 ABDALLAH	06A DLPH	1992–1995 LEP runs	
2.1 ± 0.7 ± 0.9	95	2 SCHAEL	05C ALEP	1991–1995 LEP runs	
1.7 ± 0.2 ± 0.2	231	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	
2.7 ± 1.8 ± 0.9	23	ACKERSTAFF 99E	OPAL	1991–1995 LEP runs	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.8 ± 0.7 ± 1.2	18	BUSKULIC	96 ALEP	Repl. by SCHAEL 05C	
1.9 ± 0.4 ± 0.4	31	GIBAUT	94B CLEO	Repl. by ANASTASSOV 01	
5.1 ± 2.2	6	BYLSMA	87 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$	
6.7 ± 3.0	5	3 BELTRAMI	85 HRS	Repl. by BYLSMA 87	

1 See footnote to ABDALLAH 06A  $\Gamma(\tau^- \rightarrow h^- \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

2 SCHAEL 05C quote  $(1.4 \pm 0.7 \pm 0.9) \times 10^{-4}$ . We add  $0.7 \times 10^{-4}$  to remove their correction for  $\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  and  $\tau^- \rightarrow K^*(892)^- \eta \nu_\tau \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau$  decays. See footnote to SCHAEL 05C  $\Gamma(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$  measurement for correlations with other measurements.

3 The error quoted is statistical only.

$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$					$\Gamma_{124} / \Gamma$
$\Gamma_{124} / \Gamma = (\Gamma_{126} + 0.2302 \Gamma_{160} + 0.892 \Gamma_{185}) / \Gamma$					
<b>1.64±0.11 OUR FIT</b>	(Produced by HFLAV)				NODE=S035R39;LINKAGE=AH

• • • We use the following data for averages but not for fits. • • •

<b>1.65±0.05±0.09</b>	1 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$		NOTFITTED
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1 Not independent of LEES 12X measurements of  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0)) / \Gamma$ ,  $\Gamma(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma$ , and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma$ .

$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}}$					$\Gamma_{125} / \Gamma$
VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT		
<b>1.11±0.04±0.09</b>	1 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$		NODE=S035C91 NODE=S035C91

1 Not independent of LEES 12X  $\Gamma(\tau^- \rightarrow 2\pi^- \pi^+ \omega \nu_\tau (\text{ex. } K^0)) / \Gamma$  and  $\Gamma(\tau^- \rightarrow 3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma$  values.

$\Gamma(3\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0, \eta, \omega, f_1(1285))) / \Gamma_{\text{total}}$					$\Gamma_{126} / \Gamma$
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.38±0.09 OUR FIT</b>	(Produced by HFLAV)				NODE=S035C92 NODE=S035C92

<b>0.36±0.03±0.09</b>	7.3k	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	
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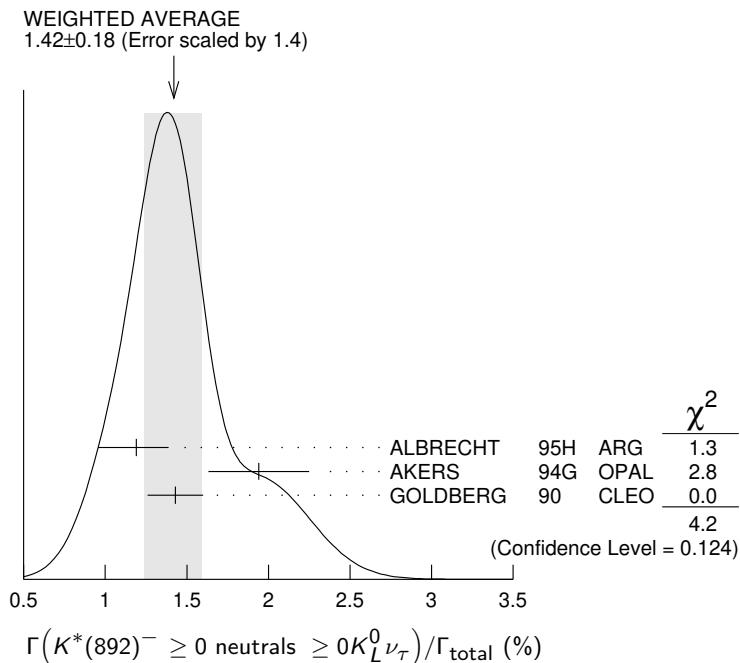
$\Gamma(K^- 2\pi^- 2\pi^+ \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$					$\Gamma_{127} / \Gamma$
VALUE (units $10^{-6}$ )	DOCUMENT ID	TECN	COMMENT		
<b>1.1±0.6 OUR FIT</b>	(Produced by HFLAV)				NODE=S035C97 NODE=S035C97

<b>1.1±0.4±0.4</b>	1 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	
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1 LEES 12X measurement corresponds to the lower limit of  $< 1.9 \times 10^{-6}$  at 90% CL.

NODE=S035C97;LINKAGE=LE

$\Gamma(K^+ 3\pi^- \pi^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{128}/\Gamma$	NODE=S035C98 NODE=S035C98
VALUE $<8 \times 10^{-7}$	CL% 90	DOCUMENT ID LEES
		TECN 12X BABR
		COMMENT $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\Gamma(3h^- 2h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{129}/\Gamma$	NODE=S035B44 NODE=S035B44
VALUE $<3.4 \times 10^{-6}$	CL% 90	DOCUMENT ID AUBERT,B
		TECN 06 BABR
		COMMENT $232 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$		
$<1.1 \times 10^{-4}$	90	GIBAUT 94B CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\Gamma((5\pi)^- \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{130}/\Gamma$	NODE=S035B45 NODE=S035B45 NODE=S035B45
		$\Gamma_{130}/\Gamma = (\Gamma_{30} + \frac{1}{2}\Gamma_{45} + \Gamma_{48} + \frac{1}{2}\Gamma_{61} + \Gamma_{85} + \Gamma_{117} + 0.5559\Gamma_{150} + 0.892\Gamma_{180})/\Gamma$
VALUE (%) <b>0.78 ± 0.05 OUR FIT</b>	DOCUMENT ID	TECN COMMENT
(Produced by HFLAV)		
$\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$		
<b>0.61 ± 0.06 ± 0.08</b>	1 GIBAUT	94B CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
1 Not independent of GIBAUT 94B B( $3h^- 2h^+ \nu_\tau$ ), PROCARIO 93 B( $h^- 4\pi^0 \nu_\tau$ ), and BORTOLETTO 93 B( $2h^- h^+ 2\pi^0 \nu_\tau$ )/B("3prong") measurements. Result is corrected for $\eta$ contributions.		
$\Gamma(4h^- 3h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ ("7-prong")})/\Gamma_{\text{total}}$	$\Gamma_{131}/\Gamma$	NODE=S035R36 NODE=S035R36
VALUE $<3.0 \times 10^{-7}$	CL% 90	DOCUMENT ID AUBERT,B
		TECN 05F BABR
		COMMENT $232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$		
$<1.8 \times 10^{-5}$	95	ACKERSTAFF 97J OPAL 1990–1995 LEP runs
$<2.4 \times 10^{-6}$	90	EDWARDS 97B CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.9 \times 10^{-4}$	90	BYLSMA 87 HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
$\Gamma(4h^- 3h^+ \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{132}/\Gamma$	NODE=S035C63 NODE=S035C63
VALUE $<4.3 \times 10^{-7}$	CL% 90	DOCUMENT ID AUBERT,B
		TECN 05F BABR
		COMMENT $232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\Gamma(4h^- 3h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{133}/\Gamma$	NODE=S035C57 NODE=S035C57
VALUE $<2.5 \times 10^{-7}$	CL% 90	DOCUMENT ID AUBERT,B
		TECN 05F BABR
		COMMENT $232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\Gamma(X^-(S=-1)\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{134}/\Gamma$	NODE=S035C47 NODE=S035C47
		$\Gamma_{134}/\Gamma = (\Gamma_{10} + \Gamma_{16} + \Gamma_{23} + \Gamma_{28} + \Gamma_{36} + \Gamma_{41} + \Gamma_{45} + \Gamma_{61} + \Gamma_{97} + \Gamma_{103} + \Gamma_{120} + \Gamma_{127} + \Gamma_{152} + \Gamma_{154} + \Gamma_{156} + 0.8312\Gamma_{170} + \Gamma_{179})/\Gamma$
VALUE (%) <b>2.92 ± 0.04 OUR FIT</b>	DOCUMENT ID	TECN COMMENT
(Produced by HFLAV)		
$\bullet \bullet \bullet$ We use the following data for averages but not for fits. $\bullet \bullet \bullet$		
<b>2.87 ± 0.12</b>	1 BARATE	99R ALEP 1991–1995 LEP runs
1 BARATE 99R perform a combined analysis of all ALEPH LEP 1 data on $\tau$ branching fraction measurements for decay modes having total strangeness equal to $-1$ .		
$\Gamma(K^*(892)^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{135}/\Gamma$	NODE=S035R45 NODE=S035R45
VALUE (%) <b>1.42 ± 0.18 OUR AVERAGE</b>	EVTS	DOCUMENT ID
		TECN COMMENT
Error includes scale factor of 1.4. See the ideogram below.		
1.19 ± 0.15 ± 0.13	104	ALBRECHT 95H ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
1.94 ± 0.27 ± 0.15	74	<sup>1</sup> AKERS 94G OPAL $E_{\text{cm}}^{\text{ee}} = 88\text{--}94 \text{ GeV}$
1.43 ± 0.11 ± 0.13	475	<sup>2</sup> GOLDBERG 90 CLEO $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9 \text{ GeV}$
1 AKERS 94G reject events in which a $K_S^0$ accompanies the $K^*(892)^-$ . We do not correct for them.		
2 GOLDBERG 90 estimates that 10% of observed $K^*(892)$ are accompanied by a $\pi^0$ .		



### $\Gamma(K^*(892)^- \nu_\tau) / \Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.20 ± 0.07 OUR AVERAGE</b>				Error includes scale factor of 1.8. See the ideogram below.
1.131±0.006±0.051	49k	<sup>1</sup> EPIFANOV 07	BELL 351 fb <sup>-1</sup>	$E_{\text{cm}}^{\text{ee}}=10.6$ GeV
1.326±0.063		BARATE 99R	ALEP	1991–1995 LEP runs
1.11 ± 0.12		<sup>2</sup> COAN 96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV
1.42 ± 0.22 ± 0.09		<sup>3</sup> ACCIARRI 95F	L3	1991–1993 LEP runs
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
1.39 ± 0.09 ± 0.10		<sup>4</sup> BUSKULIC 96	ALEP	Repl. by BARATE 99R
1.45 ± 0.13 ± 0.11	273	<sup>5</sup> BUSKULIC 94F	ALEP	Repl. by BUSKULIC 96
1.23 ± 0.21 ± 0.11	54	<sup>6</sup> ALBRECHT 88L	ARG	$E_{\text{cm}}^{\text{ee}}=10$ GeV
1.9 ± 0.3 ± 0.4	44	<sup>7</sup> TSCHIRHART 88	HRS	$E_{\text{cm}}^{\text{ee}}=29$ GeV
1.5 ± 0.4 ± 0.4	15	<sup>8</sup> AIHARA 87C	TPC	$E_{\text{cm}}^{\text{ee}}=29$ GeV
1.3 ± 0.3 ± 0.3	31	YELTON 86	MRK2	$E_{\text{cm}}^{\text{ee}}=29$ GeV
1.7 ± 0.7	11	DORFAN 81	MRK2	$E_{\text{cm}}^{\text{ee}}=4.2\text{--}6.7$ GeV

<sup>1</sup> EPIFANOV 07 quote  $B(\tau^- \rightarrow K^*(892)^- \nu_\tau) B(K^*(892)^- \rightarrow K_S^0 \pi^-) = (3.77 \pm 0.02(\text{stat}) \pm 0.12(\text{syst}) \pm 0.12(\text{mod})) \times 10^{-3}$ . We add the systematic and model uncertainties in quadrature and divide by  $B(K^*(892)^- \rightarrow K_S^0 \pi^-) = 0.3333$ .

<sup>2</sup> Not independent of COAN 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and BATTLE 94  $B(K^- \pi^0 \nu_\tau)$  measurements.  $K\pi$  final states are consistent with and assumed to originate from  $K^*(892)^-$  production.

<sup>3</sup> This result is obtained from their  $B(\pi^- \bar{K}^0 \nu_\tau)$  assuming all those decays originate in  $K^*(892)^-$  decays.

<sup>4</sup> Not independent of BUSKULIC 96  $B(\pi^- \bar{K}^0 \nu_\tau)$  and  $B(K^- \pi^0 \nu_\tau)$  measurements.

<sup>5</sup> BUSKULIC 94F obtain this result from BUSKULIC 94F  $B(\bar{K}^0 \pi^- \nu_\tau)$  and BUSKULIC 94E  $B(K^- \pi^0 \nu_\tau)$  assuming all of those decays originate in  $K^*(892)^-$  decays.

<sup>6</sup> The authors divide by  $\Gamma_2/\Gamma = 0.865$  to obtain this result.

<sup>7</sup> Not independent of TSCHIRHART 88  $\Gamma(\tau^- \rightarrow h^- \bar{K}^0 \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma$ .

<sup>8</sup> Decay  $\pi^-$  identified in this experiment, is assumed in the others.

### $\Gamma_{136}/\Gamma$

NODE=S035R9  
NODE=S035R9

NODE=S035R9;LINKAGE=EP

NODE=S035R9;LINKAGE=E

NODE=S035R9;LINKAGE=C

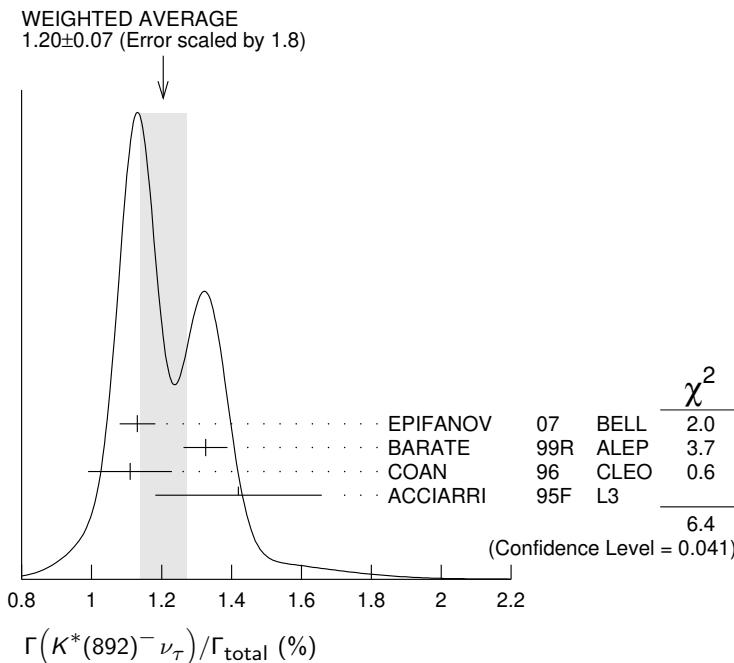
NODE=S035R9;LINKAGE=D

NODE=S035R9;LINKAGE=B

NODE=S035R9;LINKAGE=AL

NODE=S035R9;LINKAGE=A

NODE=S035R9;LINKAGE=AI



### $\Gamma(K^*(892)^- \nu_\tau)/\Gamma(\pi^- \pi^0 \nu_\tau)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{136}/\Gamma_{14}$
<b>0.075±0.027</b>	1 ABREU	94K	DLPH	LEP 1992 Z data	
1 ABREU 94K quote $B(\tau^- \rightarrow K^*(892)^- \nu_\tau)B(K^*(892)^- \rightarrow K^- \pi^0)/B(\tau^- \rightarrow \rho^- \nu_\tau) = 0.025 \pm 0.009$ . We divide by $B(K^*(892)^- \rightarrow K^- \pi^0) = 0.333$ to obtain this result.					

### $\Gamma(K^*(892)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \nu_\tau)/\Gamma(\pi^- \bar{K}^0 \nu_\tau)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{137}/\Gamma_{36}$
<b>0.933±0.027</b>	49k	EPIFANOV	07	BELL	$351 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

### $\Gamma(K^*(892)^0 K^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{138}/\Gamma$
<b>0.32±0.08±0.12</b>	119	GOLDBERG	90	CLEO	$E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$

### $\Gamma(K^*(892)^0 K^- \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{139}/\Gamma$
<b>0.21 ± 0.04 OUR AVERAGE</b>					
0.213±0.048 0.20 ± 0.05 ± 0.04					
47	1 BARATE	98	ALEP	1991–1995 LEP runs	
	ALBRECHT	95H	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$	

<sup>1</sup> BARATE 98 measure the  $K^- (\rho^0 \rightarrow \pi^+ \pi^-)$  fraction in  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  decays to be  $(35 \pm 11)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$  assuming the intermediate states are all  $K^- \rho$  and  $K^- K^*(892)^0$ .

### $\Gamma(\bar{K}^*(892)^0 \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{140}/\Gamma$
<b>0.38±0.11±0.13</b>	105	GOLDBERG	90	CLEO	$E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$

### $\Gamma(\bar{K}^*(892)^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	$\Gamma_{141}/\Gamma$
<b>0.22 ± 0.05 OUR AVERAGE</b>					
0.209±0.058 0.25 ± 0.10 ± 0.05					
27	1 BARATE	98	ALEP	1991–1995 LEP runs	
	ALBRECHT	95H	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$	

<sup>1</sup> BARATE 98 measure the  $K^- K^*(892)^0$  fraction in  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  decays to be  $(87 \pm 13)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$ .

NODE=S035B49  
NODE=S035B49

NODE=S035B49;LINKAGE=A

NODE=S035C68  
NODE=S035C68

NODE=S035B6  
NODE=S035B6

NODE=S035B60  
NODE=S035B60

NODE=S035B60;LINKAGE=B8

NODE=S035B7  
NODE=S035B7

NODE=S035B70  
NODE=S035B70

NODE=S035B70;LINKAGE=B8

$\Gamma((\bar{K}^*(892)\pi)^-\nu_\tau \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{142}/\Gamma$
VALUE (%)	DOCUMENT ID		TECN	COMMENT	
<b>0.10 ±0.04 OUR AVERAGE</b>					
0.097±0.044±0.036	1	BARATE	99K	ALEP	1991–1995 LEP runs
0.106±0.037±0.032	2	BARATE	98E	ALEP	1991–1995 LEP runs
1 BARATE 99K measure $K^0$ 's by detecting $K_L^0$ 's in their hadron calorimeter. They determine the $\bar{K}^0\rho^-$ fraction in $\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$ decays to be $(0.72 \pm 0.12 \pm 0.10)$ and multiply their $B(\pi^-\bar{K}^0\pi^0\nu_\tau)$ measurement by one minus this fraction to obtain the quoted result.					
2 BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+\pi^-$ decays. They determine the $\bar{K}^0\rho^-$ fraction in $\tau^- \rightarrow \pi^-\bar{K}^0\pi^0\nu_\tau$ decays to be $(0.64 \pm 0.09 \pm 0.10)$ and multiply their $B(\pi^-\bar{K}^0\pi^0\nu_\tau)$ measurement by one minus this fraction to obtain the quoted result.					

NODE=S035C11  
NODE=S035C11

NODE=S035C11;LINKAGE=9K

NODE=S035C11;LINKAGE=B9

$\Gamma(K_1(1270)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{143}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.47±0.11 OUR AVERAGE</b>					
0.48±0.11		BARATE	99R	ALEP	1991–1995 LEP runs
$0.41^{+0.41}_{-0.35} \pm 0.10$	5	1 BAUER	94	TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

NODE=S035B40  
NODE=S035B40

NODE=S035B40;LINKAGE=A

$\Gamma(K_1(1400)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{144}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.17±0.26 OUR AVERAGE</b>				Error includes scale factor of 1.7.	
0.05±0.17		BARATE	99R	ALEP	1991–1995 LEP runs
$0.76^{+0.40}_{-0.33} \pm 0.20$	11	1 BAUER	94	TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

NODE=S035B41  
NODE=S035B41

NODE=S035B41;LINKAGE=A

$[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]/\Gamma_{\text{total}}$					$(\Gamma_{143} + \Gamma_{144})/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>1.17<math>^{+0.41}_{-0.37}</math>±0.29</b>	16	1 BAUER	94	TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

NODE=S035B42  
NODE=S035B42

NODE=S035B42;LINKAGE=A

$\Gamma(K_1(1270)^-\nu_\tau)/[\Gamma(K_1(1270)^-\nu_\tau) + \Gamma(K_1(1400)^-\nu_\tau)]$					$\Gamma_{143}/(\Gamma_{143} + \Gamma_{144})$
VALUE	DOCUMENT ID	TECN	COMMENT		
<b>0.69±0.15 OUR AVERAGE</b>					

NODE=S035C41  
NODE=S035C41

0.71±0.16±0.11	1	ABBIENDI	00D	OPAL	1990–1995 LEP runs
0.66±0.19±0.13	2	ASNER	00B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
1 ABBIENDI 00D assume the resonance structure of $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$ decays is dominated by the $K_1(1270)^-$ and $K_1(1400)^-$ resonances.					
2 ASNER 00B assume the resonance structure of $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$ (ex. $K^0$ ) decays is dominated by $K_1(1270)^-$ and $K_1(1400)^-$ resonances.					

NODE=S035C41;LINKAGE=A

NODE=S035C41;LINKAGE=BA

$\Gamma(K_1^*(1410)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{145}/\Gamma$
VALUE (units $10^{-3}$ )	DOCUMENT ID	TECN	COMMENT		
<b>1.5<math>^{+1.4}_{-1.0}</math></b>	BARATE	99R	ALEP	1991–1995 LEP runs	

NODE=S035C45  
NODE=S035C45

$\Gamma(K_0^*(1430)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{146}/\Gamma$
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.5</b>	95	BARATE	99R	ALEP	1991–1995 LEP runs

NODE=S035C46  
NODE=S035C46

$\Gamma(K_2^*(1430)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{147}/\Gamma$
VALUE (%)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.3</b>	95		TSCHIRHART	88	HRS
					$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

NODE=S035R10  
NODE=S035R10

1 ACCIARRI 95F quote  $B(\tau^- \rightarrow K^*(1430)^- \rightarrow \pi^-\bar{K}^0\nu_\tau) < 0.11\%$ . We divide by  $B(K^*(1430)^- \rightarrow \pi^-\bar{K}^0) = 0.33$  to obtain the limit shown.

NODE=S035R10;LINKAGE=A

$\Gamma(a_0(980)^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \times B(a_0(980) \rightarrow K^0 K^-)$        $\Gamma_{148} / \Gamma \times B$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<2.8	90	GOLDBERG	90	CLEO $E_{\text{cm}}^{\text{ee}} = 9.4 - 10.9 \text{ GeV}$

NODE=S035B8  
NODE=S035B8 $\Gamma(\eta \pi^- \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{149} / \Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< 0.99	95	1	DEL-AMO-SA..11E	BABR	$470 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035R13  
NODE=S035R13

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 6.2	95	BUSKULIC	97C	ALEP	1991–1994 LEP runs
< 1.4	95	BARTEL	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
< 3.4	95	ARTUSO	92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
< 90	95	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$
< 140	90	BEHREND	88	CELL	$E_{\text{cm}}^{\text{ee}} = 14 - 46.8 \text{ GeV}$
< 180	95	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
< 250	90	COFFMAN	87	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$
510 $\pm 100 \pm 120$	65	DERRICK	87	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$
< 100	95	GAN	87B	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

1 DEL-AMO-SANCHEZ 11E also quote  $B(\tau^- \rightarrow \eta \pi^- \nu_\tau) = (3.4 \pm 3.4 \pm 2.1) \times 10^{-5}$ .

NODE=S035R13;LINKAGE=DE

 $\Gamma(\eta \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{150} / \Gamma$ 

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.39 <math>\pm 0.07</math> OUR FIT</b>	(Produced by HFLAV)				

NODE=S035R32  
NODE=S035R321.38  $\pm 0.09$  OUR AVERAGE Error includes scale factor of 1.2.

1.35 $\pm 0.03 \pm 0.07$	6.0k	INAMI	09	BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
1.8 $\pm 0.4 \pm 0.2$		BUSKULIC	97C	ALEP	1991–1994 LEP runs
1.7 $\pm 0.2 \pm 0.2$	125	ARTUSO	92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 11.0	95	ALBRECHT	88M	ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$
< 21.0	95	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
42.0 $^{+ 7.0}_{- 12.0} \pm 16.0$	1	GAN	87	MRK2	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

1 Highly correlated with GAN 87  $\Gamma(\pi^- 3\pi^0 \nu_\tau) / \Gamma(\text{total})$  value.

NODE=S035R32;LINKAGE=A

 $\Gamma(\eta \pi^- \pi^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{151} / \Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.9 <math>\pm 0.4</math> OUR FIT</b>	(Produced by HFLAV)				
<b>1.81 <math>\pm 0.31</math> OUR AVERAGE</b>					

NODE=S035R78  
NODE=S035R78

• • • We use the following data for averages but not for fits. • • •

1.5 $\pm 0.5$	30	1	ANASTASSOV	01	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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NOTFITTED

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.4 $\pm 0.6 \pm 0.3$	15	2	BERGFELD	97	CLEO   Repl. by ANAS-TASSOV 01
< 4.3	95		ARTUSO	92	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
< 120	95		ALBRECHT	88M	ARG $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

NODE=S035R78;LINKAGE=AN

1 Weighted average of BERGFELD 97 and ANASTASSOV 01 value of  $(1.5 \pm 0.6 \pm 0.3) \times 10^{-4}$  obtained using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+ \pi^- \pi^0$  decays.2 BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  decays. $\Gamma(\eta K^- \nu_\tau) / \Gamma_{\text{total}}$  $\Gamma_{152} / \Gamma$ 

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.55 <math>\pm 0.08</math> OUR FIT</b>	(Produced by HFLAV)				

NODE=S035B20  
NODE=S035B20**1.54  $\pm 0.08$  OUR AVERAGE**

1.42 $\pm 0.11 \pm 0.07$	690	DEL-AMO-SA..11E	BABR	$470 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
1.58 $\pm 0.05 \pm 0.09$	1.6k	INAMI	09	BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
2.9 $^{+ 1.3}_{- 1.2} \pm 0.7$		BUSKULIC	97C	ALEP	1991–1994 LEP runs
2.6 $\pm 0.5 \pm 0.5$	85	BARTEL	96	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
< 4.7	95	ARTUSO	92	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

$\Gamma(\eta K^*(892)^-\nu_\tau)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.38±0.15 OUR AVERAGE</b>				
1.34±0.12±0.09	245	<sup>1</sup> INAMI	09	BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
2.90±0.80±0.42	25	BISHAI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> Not independent of INAMI 09  $B(\tau^- \rightarrow \eta K^- \pi^0 \nu_\tau)$  and  $B(\tau^- \rightarrow \eta \bar{K}^0 \pi^- \nu_\tau)$  values.

 $\Gamma_{153}/\Gamma$ 

NODE=S035C26  
NODE=S035C26

 $\Gamma(\eta K^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.48±0.12 OUR FIT</b> (Produced by HFLAV)				
<b>0.48±0.12 OUR AVERAGE</b>				
0.46±0.11±0.04	270	INAMI	09	BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
1.77±0.56±0.71	36	BISHAI	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C26;LINKAGE=IN

 $\Gamma_{154}/\Gamma$ 

NODE=S035C27  
NODE=S035C27

 $\Gamma(\eta K^-\pi^0(\text{non-}K^*(892))\nu_\tau)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.5 \times 10^{-5}$	90	INAMI	09	BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C72  
NODE=S035C72

 $\Gamma(\eta \bar{K}^0 \pi^- \nu_\tau)/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.94±0.15 OUR FIT</b> (Produced by HFLAV)				
<b>0.93±0.15 OUR AVERAGE</b>				

NODE=S035C28  
NODE=S035C28

 $\Gamma(0.88\pm0.14\pm0.06)$ 

<sup>1</sup> INAMI 09     $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>2</sup> BISHAI 99     $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (0.44 \pm 0.07 \pm 0.03) \times 10^{-4}$  by 2 to obtain the listed value.

<sup>2</sup> We multiply the BISHAI 99 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (1.10 \pm 0.35 \pm 0.11) \times 10^{-4}$  by 2 to obtain the listed value.

 $\Gamma_{155}/\Gamma$ 

NODE=S035C72  
NODE=S035C72

 $\Gamma(\eta \bar{K}^0 \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.0 \times 10^{-5}$	90	<sup>1</sup> INAMI	09	BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C73  
NODE=S035C73

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \pi^0 \nu_\tau) < 2.5 \times 10^{-5}$  by 2 to obtain the listed value.

 $\Gamma_{157}/\Gamma$ 

NODE=S035C73  
NODE=S035C73

 $\Gamma(\eta K^-\bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<9.0 \times 10^{-6}$	90	<sup>1</sup> INAMI	09	BELL $490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C74  
NODE=S035C74

<sup>1</sup> We multiply the INAMI 09 measurement  $B(\tau^- \rightarrow \eta K^- \bar{K}^0 \nu_\tau) < 4.5 \times 10^{-6}$  by 2 to obtain the listed value.

 $\Gamma_{158}/\Gamma$ 

NODE=S035C74  
NODE=S035C74

 $\Gamma(\eta \pi^+ \pi^- \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ 

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<0.3	90	ABACHI	87B	HRS $E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

NODE=S035R50  
NODE=S035R50

 $\Gamma(\eta \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.20±0.13 OUR FIT</b> (Produced by HFLAV)				
<b>2.23±0.12 OUR AVERAGE</b>				

NODE=S035B89  
NODE=S035B89

 $\bullet \bullet \bullet$  We use the following data for averages but not for fits.  $\bullet \bullet \bullet$ 

2.3 ± 0.5      170      <sup>2</sup> ANASTASSOV 01      CLEO     $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\bullet \bullet \bullet$  We do not use the following data for averages, fits, limits, etc.  $\bullet \bullet \bullet$

1.60±0.05±0.11      1.8 k      AUBERT      08AE BABR    Repl. by LEES 12X

3.4  $^{+0.6}_{-0.5}$  ± 0.6      89      <sup>3</sup> BERGFELD      97      CLEO    Repl. by ANASTASSOV 01

OCCUR=2

OCCUR=3

OCCUR=4

NOTFITTED

<sup>1</sup> LEES 12X uses  $468 \text{ fb}^{-1}$  of data taken at  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$ . It gives the average of the three measurements listed here as  $(2.25 \pm 0.07 \pm 0.12) \times 10^{-4}$ .

<sup>2</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 measurements using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+ \pi^- \pi^0$  and  $\eta \rightarrow 3\pi^0$  decays.

<sup>3</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow 3\pi^0$  decays.

NODE=S035B89;LINKAGE=LE

NODE=S035B89;LINKAGE=AN

NODE=S035B89;LINKAGE=BF

$\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0,f_1(1285)))/\Gamma_{\text{total}}$   $\Gamma_{161}/\Gamma$ 

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	COMMENT
<b><math>0.99 \pm 0.09 \pm 0.13</math></b>	1 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

1 LEES 12X obtain this result by subtracting their  $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$  measurement from their  $B(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau(\text{ex.}K^0))$  measurement.

 $\Gamma(\eta a_1(1260)^-\nu_\tau \rightarrow \eta\pi^-\rho^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{162}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.9 \times 10^{-4}</math></b>	90	BERGFELD	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\eta\eta\pi^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{163}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;7.4 \times 10^{-6}</math></b>	90	INAMI	09 BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.1 \times 10^{-4}$	95	ARTUSO	92 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
$<8.3 \times 10^{-3}$	95	ALBRECHT	88M ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

 $\Gamma(\eta\eta\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{164}/\Gamma$ 

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.0</math></b>	95	ARTUSO	92 CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$<90$  95 ALBRECHT 88M ARG  $E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$

 $\Gamma(\eta\eta K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{165}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.0 \times 10^{-6}</math></b>	90	INAMI	09 BELL	$490 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\eta'(958)\pi^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{166}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;4.0 \times 10^{-6}</math></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<7.2 \times 10^{-6}$	90	AUBERT	08AE BABR	$384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$<7.4 \times 10^{-5}$  90 BERGFELD 97 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\eta'(958)\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{167}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.2 \times 10^{-5}</math></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$<8.0 \times 10^{-5}$  90 BERGFELD 97 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\eta'(958)K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{168}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.4 \times 10^{-6}</math></b>	90	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<20$  90 <sup>1</sup> AVERY 97 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$<35$  90 ALBRECHT 95H ARG  $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

1 AVERY 97 limit varies from  $(1.2\text{--}2.0) \times 10^{-4}$  depending on decay model assumptions.

 $\Gamma(\phi\pi^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{169}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>3.42 \pm 0.55 \pm 0.25</math></b>	344	AUBERT	08 BABR	$342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<20$  90 <sup>1</sup> AVERY 97 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$<35$  90 ALBRECHT 95H ARG  $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$

1 AVERY 97 limit varies from  $(1.2\text{--}2.0) \times 10^{-4}$  depending on decay model assumptions.

 $\Gamma(\phi K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{170}/\Gamma$ 

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>4.4 \pm 1.6 \text{ OUR FIT}</math></b>	(Produced by HFLAV)				

**$3.70 \pm 0.33 \text{ OUR AVERAGE}$**  Error includes scale factor of 1.3.

• • • We use the following data for averages but not for fits. • • •

$3.39 \pm 0.20 \pm 0.28$  274 AUBERT 08 BABR  $342 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$4.05 \pm 0.25 \pm 0.26$  551 INAMI 06 BELL  $401 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.7$  90 <sup>1</sup> AVERY 97 CLEO  $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

1 AVERY 97 limit varies from  $(5.4\text{--}6.7) \times 10^{-5}$  depending on decay model assumptions.

NODE=S035C81  
NODE=S035C81

NODE=S035C81;LINKAGE=LE

NODE=S035B90  
NODE=S035B90

NODE=S035R79  
NODE=S035R79

NODE=S035R80  
NODE=S035R80

NODE=S035C75  
NODE=S035C75

NODE=S035B91  
NODE=S035B91

NODE=S035B92  
NODE=S035B92

NODE=S035C93  
NODE=S035C93

NODE=S035B61  
NODE=S035B61

NODE=S035B61;LINKAGE=A

NODE=S035B82  
NODE=S035B82

NOTFITTED

NOTFITTED

NODE=S035B82;LINKAGE=A

$\Gamma(f_1(1285)\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{171}/\Gamma$	
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		NODE=S035B93 NODE=S035B93
<b>3.9 ± 0.5 OUR AVERAGE</b>		Error includes scale factor of 1.9.				OCCUR=2
4.73 ± 0.28 ± 0.45	3.7k	1 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
3.60 ± 0.18 ± 0.23	2.5k	2 LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
• • • We do not use the following data for averages, fits, limits, etc. • • •						
3.19 ± 0.18 ± 1.00	1.3 k	3 AUBERT	08AE BABR	Repl. by LEES 12X		NODE=S035B93;LINKAGE=LE
3.9 ± 0.7 ± 0.5	1.4 k	4 AUBERT,B	05W BABR	Repl. by LEES 12X		NODE=S035B93;LINKAGE=LS
5.8 $^{+1.4}_{-1.3}$ ± 1.8	54	5 BERGFELD	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		NODE=S035B93;LINKAGE=AB
1 LEES 12X obtain this value by dividing their $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^-2\pi^+\nu_\tau)$ measurement by the PDG 12 value of $B(f_1(1285) \rightarrow 2\pi^+2\pi^-) = 0.111^{+0.007}_{-0.006}$ .						
2 LEES 12X obtain this value by dividing their $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$ measurement by 2/3 of the PDG 12 value of $B(f_1(1285) \rightarrow \eta\pi\pi) = 0.524^{+0.019}_{-0.021}$ .						
3 AUBERT 08AE obtain this value by dividing their $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$ measurement by the PDG 06 value of $B(f_1(1285) \rightarrow \eta\pi^-\pi^+) = 0.35 \pm 0.11$ . The quote $(3.19 \pm 0.18 \pm 0.16 \pm 0.99) \times 10^{-4}$ where the final error is due to the uncertainty on $B(f_1(1285) \rightarrow \eta\pi^-\pi^+)$ . We combine the two systematic errors in quadrature.						
4 AUBERT,B 05W use the $f_1(1285) \rightarrow 2\pi^+2\pi^-$ decay mode and the PDG 04 value of $B(f_1(1285) \rightarrow 2\pi^+2\pi^-) = 0.110^{+0.007}_{-0.006}$ .						NODE=S035B93;LINKAGE=AU
5 BERGFELD 97 use the $f_1(1285) \rightarrow \eta\pi^+\pi^-$ decay mode.						NODE=S035B93;LINKAGE=BE
$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{172}/\Gamma$	
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		NODE=S035C69 NODE=S035C69
<b>1.18±0.07 OUR AVERAGE</b>		Error includes scale factor of 1.3.				
1.26 ± 0.06 ± 0.06	2.5k	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
1.11 ± 0.06 ± 0.05	1.3 k	AUBERT	08AE BABR	$384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0))$					$\Gamma_{172}/\Gamma_{160}$	
VALUE	DOCUMENT ID	TECN	COMMENT			NODE=S035B94 NODE=S035B94
<b>0.69±0.01±0.05</b>	1 AUBERT	08AE BABR	$384 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$			
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.55 ± 0.14	BERGFELD	97 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$			NODE=S035B94;LINKAGE=AU
1 Not independent of AUBERT 08AE $B(\tau^- \rightarrow f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)$ and $B(\tau^- \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau (\text{ex. } K^0))$ values.						
$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow 3\pi^-2\pi^+\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{173}/\Gamma$	
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT		NODE=S035C82 NODE=S035C82
<b>0.52 ± 0.04 OUR FIT</b>		(Produced by HFLAV)				
<b>0.520±0.031±0.037</b>	3.7k	LEES	12X BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
$\Gamma(\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{174}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		NODE=S035C42 NODE=S035C42
<b>&lt;1.0 × 10<sup>-4</sup></b>	90	ASNER	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
$\Gamma(\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_S\text{-wave}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{175}/\Gamma$	
VALUE	CL%	DOCUMENT ID	TECN	COMMENT		NODE=S035C43 NODE=S035C43
<b>&lt;1.9 × 10<sup>-4</sup></b>	90	ASNER	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$		
$\Gamma(h^-\omega \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{176}/\Gamma$	
$\Gamma_{176}/\Gamma = (\Gamma_{178} + \Gamma_{179} + \Gamma_{180})/\Gamma$						NODE=S035R15
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT		NODE=S035R15 NODE=S035R15
<b>2.40±0.08 OUR FIT</b>		(Produced by HFLAV)				
• • • We use the following data for averages but not for fits. • • •						
<b>1.65±0.3 ± 0.2</b>	1513	ALBRECHT	88M ARG	$E_{\text{cm}}^{\text{ee}} \approx 10 \text{ GeV}$		NOTFITTED

$\Gamma(h^-\omega\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{177}/\Gamma = (\Gamma_{178} + \Gamma_{179})/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.99±0.06 OUR FIT</b> (Produced by HFLAV)				
<b>1.92±0.07 OUR AVERAGE</b>				

1.91±0.07±0.06	5803	BUSKULIC	97C	ALEP	1991–1994 LEP runs
1.60±0.27±0.41	139	BARINGER	87	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5 \text{ GeV}$
<b>• • • We use the following data for averages but not for fits. • • •</b>					
1.95±0.07±0.11	2223	<sup>1</sup> BALEST	95C	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

<sup>1</sup> Not independent of BALEST 95C  $B(\tau^- \rightarrow h^-\omega\nu_\tau)/B(\tau^- \rightarrow h^-h^+\pi^0\nu_\tau)$  value.

$$[\Gamma(\pi^-\omega\nu_\tau) + \Gamma(K^-\omega\nu_\tau)]/\Gamma(h^-h^-h^+\pi^0\nu_\tau(\text{ex. } K^0)) = (\Gamma_{178} + \Gamma_{179})/\Gamma_{74}$$

$$(\Gamma_{178} + \Gamma_{179})/\Gamma_{74} = (\Gamma_{178} + \Gamma_{179})/(\Gamma_{78} + \Gamma_{103} + \Gamma_{107} + 0.2302\Gamma_{152} + 0.892\Gamma_{178} + 0.892\Gamma_{179} + 0.0153\Gamma_{180})$$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>43.5±1.4 OUR FIT</b> (Produced by HFLAV)				
<b>45.3±1.9 OUR AVERAGE</b>				

43.1±3.3	2350	<sup>1</sup> BUSKULIC	96	ALEP	LEP 1991–1993 data
46.4±1.6±1.7	2223	<sup>2</sup> BALEST	95C	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

**• • • We do not use the following data for averages, fits, limits, etc. • • •**

37 ± 5 ± 2	458	<sup>3</sup> ALBRECHT	91D	ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
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<sup>1</sup> BUSKULIC 96 quote the fraction of  $\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^-\omega$  final state =  $0.383 \pm 0.029$ . We divide this by the  $\omega(782) \rightarrow \pi^+\pi^-\pi^0$  branching fraction (0.888).

<sup>2</sup> BALEST 95C quote the fraction of  $\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^-\omega$  final state equals  $0.412 \pm 0.014 \pm 0.015$ . We divide this by the  $\omega(782) \rightarrow \pi^+\pi^-\pi^0$  branching fraction (0.888).

<sup>3</sup> ALBRECHT 91D quote the fraction of  $\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau$  decays which originate in a  $\pi^-\omega$  final state equals  $0.33 \pm 0.04 \pm 0.02$ . We divide this by the  $\omega(782) \rightarrow \pi^+\pi^-\pi^0$  branching fraction (0.888).

$\Gamma(\pi^-\omega\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{178}/\Gamma$		
VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>1.95±0.06 OUR FIT</b> (Produced by HFLAV)			

$\Gamma(K^-\omega\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{179}/\Gamma$				
VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>4.1±0.9 OUR FIT</b> (Produced by HFLAV)					
<b>4.1±0.6±0.7</b>	500	ARMS	05	CLE3	$7.6 \text{ fb}^{-1}$ , $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

$\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma_{\text{total}}$	$\Gamma_{180}/\Gamma$				
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.41±0.04 OUR FIT</b> (Produced by HFLAV)					
<b>0.43±0.06±0.05</b>	7283	BUSKULIC	97C	ALEP	1991–1994 LEP runs

$\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma(h^-h^-h^+ \geq 0 \text{ neutrals} \geq 0 K^0\nu_\tau)$	$\Gamma_{180}/\Gamma_{62}$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>(2.69 ± 0.28) × 10<sup>-2</sup> OUR FIT</b> (Produced by HFLAV)				
<b>0.028±0.003±0.003</b>	430	<sup>1</sup> BORTOLETTO 93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

<sup>1</sup> Not independent of BORTOLETTO 93  $\Gamma(\tau^- \rightarrow h^-\omega\pi^0\nu_\tau)/\Gamma(\tau^- \rightarrow h^-h^-h^+\pi^0\nu_\tau)$  (ex.  $K^0$ ) value.

$\Gamma(h^-\omega\pi^0\nu_\tau)/\Gamma(h^-h^-h^+ 2\pi^0\nu_\tau(\text{ex. } K^0))$	$\Gamma_{180}/\Gamma_{84}$			
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>82±8 OUR FIT</b> (Produced by HFLAV)				
<b>81±6±6</b>		BORTOLETTO 93	CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

NODE=S035R23  
NODE=S035R23

NOTFITTED  
NODE=S035R23;LINKAGE=A

NODE=S035R14  
NODE=S035R14

NODE=S035R14

NODE=S035R14;LINKAGE=D

NODE=S035R14;LINKAGE=B

NODE=S035R14;LINKAGE=C

NODE=S035R00  
NODE=S035R00

NODE=S035C61  
NODE=S035C61

NODE=S035B58  
NODE=S035B58

NODE=S035B26

NOTFITTED  
NODE=S035B26;LINKAGE=A

NODE=S035B27  
NODE=S035B27  
NODE=S035B27

$\Gamma(h^-\omega 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{181}/\Gamma$
<b>1.4 ± 0.4 ± 0.3</b>	53	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035B96 NODE=S035B96
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$1.89^{+0.74}_{-0.67} \pm 0.40$	19	ANDERSON	97	CLEO Repl. by ANASTASSOV 01	

 $\Gamma(\pi^-\omega 2\pi^0\nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{182}/\Gamma$
<b>0.72±0.16 OUR FIT</b>		(Produced by HFLAV)			NODE=S035C84 NODE=S035C84
<b>0.73±0.12±0.12</b>	1.1k	LEES	12x BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

 $\Gamma(h^-2\omega\nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{183}/\Gamma$
$< 5.4 \times 10^{-7}$	90	AUBERT,B	06	BABR $232 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C66 NODE=S035C66

 $\Gamma(2h^- h^+\omega\nu_\tau)/\Gamma_{\text{total}}$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{184}/\Gamma$
<b>1.2±0.2±0.1</b>	110	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C53 NODE=S035C53

 $\Gamma(2\pi^-\pi^+\omega\nu_\tau(\text{ex. } K^0))/\Gamma_{\text{total}}$ 

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{185}/\Gamma$
<b>0.84±0.06 OUR FIT</b>		(Produced by HFLAV)			NODE=S035C83 NODE=S035C83
<b>0.84±0.04±0.06</b>	2.4k	LEES	12x BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

 $\Gamma(e^-\gamma)/\Gamma_{\text{total}}$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{186}/\Gamma$
$< 3.3 \times 10^{-8}$	90	AUBERT	10B BABR	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R54 NODE=S035R54 NODE=S035R54
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 5.6 \times 10^{-8}$	90	UNO	21 BELL	$988 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 1.2 \times 10^{-7}$	90	HAYASAKA	08 BELL	$535 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 1.1 \times 10^{-7}$	90	AUBERT	06C BABR	$232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 3.9 \times 10^{-7}$	90	HAYASAKA	05 BELL	$86.7 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 2.7 \times 10^{-6}$	90	EDWARDS	97 CLEO		
$< 1.1 \times 10^{-4}$	90	ABREU	95U DLPH	1990–1993 LEP runs	
$< 1.2 \times 10^{-4}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	
$< 2.0 \times 10^{-4}$	90	KEH	88 CBAL	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	
$< 6.4 \times 10^{-4}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$	

 $\Gamma(e^-\gamma\gamma)/\Gamma_{\text{total}}$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{187}/\Gamma$
$< 2.5 \times 10^{-4}$	90	<sup>1</sup> BRYMAN	21 RVUE	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035P06 NODE=S035P06

<sup>1</sup> BRYMAN 21 reinterprets the upper limit result on  $B(\tau^- \rightarrow e^-\gamma)$  and  $B(\tau^- \rightarrow \mu^-\gamma)$  by AUBERT 10B, estimating with a simulation the efficiency for this decay mode to be detected as the corresponding AUBERT 10B decay mode.

 $\Gamma(\mu^-\gamma)/\Gamma_{\text{total}}$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	$\Gamma_{188}/\Gamma$
$< 4.2 \times 10^{-8}$	90	UNO	21 BELL	$988 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R53 NODE=S035R53 NODE=S035R53
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$< 4.4 \times 10^{-8}$	90	AUBERT	10B BABR	$516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 4.5 \times 10^{-8}$	90	HAYASAKA	08 BELL	$535 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 6.8 \times 10^{-8}$	90	AUBERT,B	05A BABR	$232 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 3.1 \times 10^{-7}$	90	ABE	04B BELL	$86.3 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 1.1 \times 10^{-6}$	90	AHMED	00 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 3.0 \times 10^{-6}$	90	EDWARDS	97 CLEO		
$< 6.2 \times 10^{-5}$	90	ABREU	95U DLPH	1990–1993 LEP runs	
$< 0.42 \times 10^{-5}$	90	BEAN	93 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 3.4 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	
$< 55 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$	

$\Gamma(\mu^-\gamma\gamma)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 5.8 \times 10^{-4}$	90	1 BRYMAN	21	RVUE $516 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

1 BRYMAN 21 reinterprets the upper limit result on  $B(\tau^- \rightarrow e^-\gamma)$  and  $B(\tau^- \rightarrow \mu^-\gamma)$  by AUBERT 10B, estimating with a simulation the efficiency for this decay mode to be detected as the corresponding AUBERT 10B decay mode.

 $\Gamma_{189}/\Gamma$ 

NODE=S035P07  
NODE=S035P07

 $\Gamma(e^-\pi^0)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.0 \times 10^{-8}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.3 \times 10^{-7}$	90	AUBERT	07I	BABR $339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.7 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 17 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 14 \times 10^{-5}$	90	KEH	88	CBAL $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 210 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

 $\Gamma_{190}/\Gamma$ 

NODE=S035R60  
NODE=S035R60  
NODE=S035R60

 $\Gamma(\mu^-\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{191}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.1 \times 10^{-7}$	90	AUBERT	07I	BABR $339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.2 \times 10^{-7}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.1 \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.0 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 82 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

NODE=S035R59  
NODE=S035R59  
NODE=S035R59

 $\Gamma(e^-\kappa_S^0)/\Gamma_{\text{total}}$  $\Gamma_{192}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.6 \times 10^{-8}$	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 3.3 \times 10^{-8}$	90	AUBERT	09D	BABR $469 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.6 \times 10^{-8}$	90	MIYAZAKI	06A	BELL $281 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 9.1 \times 10^{-7}$	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.3 \times 10^{-3}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

NODE=S035R62  
NODE=S035R62  
NODE=S035R62

 $\Gamma(\mu^-\kappa_S^0)/\Gamma_{\text{total}}$  $\Gamma_{193}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.3 \times 10^{-8}$	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 4.0 \times 10^{-8}$	90	AUBERT	09D	BABR $469 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.9 \times 10^{-8}$	90	MIYAZAKI	06A	BELL $281 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 9.5 \times 10^{-7}$	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.0 \times 10^{-3}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

NODE=S035R61  
NODE=S035R61  
NODE=S035R61

 $\Gamma(e^-\eta)/\Gamma_{\text{total}}$  $\Gamma_{194}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 9.2 \times 10^{-8}$	90	MIYAZAKI	07	BELL $401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.6 \times 10^{-7}$	90	AUBERT	07I	BABR $339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.4 \times 10^{-7}$	90	ENARI	05	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 8.2 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 24 \times 10^{-5}$	90	KEH	88	CBAL $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

NODE=S035R77  
NODE=S035R77  
NODE=S035R77

$\Gamma(\mu^- \eta)/\Gamma_{\text{total}}$  $\Gamma_{195}/\Gamma$ 

Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 6.5 \times 10^{-8}$	90	MIYAZAKI	07	BELL	$401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.5 \times 10^{-7}$	90	AUBERT	07I	BABR	$339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-7}$	90	ENARI	05	BELL	$154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.4 \times 10^{-7}$	90	ENARI	04	BELL	$84.3 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 9.6 \times 10^{-6}$	90	BONVICINI	97	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 7.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

 $\Gamma(e^- \rho^0)/\Gamma_{\text{total}}$  $\Gamma_{196}/\Gamma$ 

Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 2.2 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23	BELL	$980 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.8 \times 10^{-8}$	90	MIYAZAKI	11	BELL	$854 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.6 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.5 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.2 \times 10^{-6}$	90	<sup>2</sup> BARTEL	94	CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 37 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTEL 94 assume phase space decays. $\Gamma(\mu^- \rho^0)/\Gamma_{\text{total}}$  $\Gamma_{197}/\Gamma$ 

Test of lepton family number conservation.					
VALUE	CL%	DOCUMENT ID	TECN	COMMENT	
$< 1.7 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23	BELL	$980 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.2 \times 10^{-8}$	90	MIYAZAKI	11	BELL	$854 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.6 \times 10^{-8}$	90	AUBERT	09W	BABR	$451 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.8 \times 10^{-8}$	90	NISHIO	08	BELL	$543 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 6.3 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.7 \times 10^{-6}$	90	<sup>2</sup> BARTEL	94	CLEO	Repl. by BLISS 98
$< 2.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 44 \times 10^{-5}$	90	HAYES	82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTEL 94 assume phase space decays. $\Gamma(e^- \omega)/\Gamma_{\text{total}}$  $\Gamma_{198}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.4 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23	$980 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$< 4.8 \times 10^{-8}$	90	MIYAZAKI	11	BELL
$< 1.1 \times 10^{-7}$	90	AUBERT	08K	BABR
$< 1.8 \times 10^{-7}$	90	NISHIO	08	BELL

<sup>1</sup> Supersedes MIYAZAKI 11. $\Gamma(\mu^- \omega)/\Gamma_{\text{total}}$  $\Gamma_{199}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.9 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI	23	$980 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$< 4.7 \times 10^{-8}$	90	MIYAZAKI	11	BELL
$< 1.0 \times 10^{-7}$	90	AUBERT	08K	BABR
$< 8.9 \times 10^{-8}$	90	NISHIO	08	BELL

<sup>1</sup> Supersedes MIYAZAKI 11.

NODE=S035B14

NODE=S035B14

NODE=S035B14

NODE=S035R64

NODE=S035R64

NODE=S035R64

NODE=S035R64;LINKAGE=A

NODE=S035R64;LINKAGE=B9

NODE=S035R63

NODE=S035R63

NODE=S035R63

NODE=S035C04;LINKAGE=A

NODE=S035C04;LINKAGE=B9

NODE=S035C04

NODE=S035C04

NODE=S035C05;LINKAGE=A

$\Gamma(e^- K^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{200}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<1.9 \times 10^{-8}$	90	1 TSUZUKI	23	BELL $980 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.2 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<5.9 \times 10^{-8}$	90	AUBERT	09w	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.8 \times 10^{-8}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<5.1 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.3 \times 10^{-6}$	90	2 BARTEL	94	CLEO Repl. by BLISS 98
$<3.8 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTEL 94 assume phase space decays. $\Gamma(\mu^- K^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{201}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<2.9 \times 10^{-8}$	90	1 TSUZUKI	23	BELL $980 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<7.2 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.7 \times 10^{-7}$	90	AUBERT	09w	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<5.9 \times 10^{-8}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.9 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<9.4 \times 10^{-6}$	90	2 BARTEL	94	CLEO Repl. by BLISS 98
$<4.5 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTEL 94 assume phase space decays. $\Gamma(e^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{202}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<1.7 \times 10^{-8}$	90	1 TSUZUKI	23	BELL $980 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.4 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.6 \times 10^{-8}$	90	AUBERT	09w	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.7 \times 10^{-8}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.1 \times 10^{-5}$	90	2 BARTEL	94	CLEO Repl. by BLISS 98

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTEL 94 assume phase space decays. $\Gamma(\mu^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$  $\Gamma_{203}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<4.3 \times 10^{-8}$	90	1 TSUZUKI	23	BELL $980 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<7.0 \times 10^{-8}$	90	MIYAZAKI	11	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.3 \times 10^{-8}$	90	AUBERT	09w	BABR $451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.0 \times 10^{-7}$	90	NISHIO	08	BELL $543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<8.7 \times 10^{-6}$	90	2 BARTEL	94	CLEO Repl. by BLISS 98

<sup>1</sup> Supersedes MIYAZAKI 11.<sup>2</sup> BARTEL 94 assume phase space decays.

NODE=S035R73

NODE=S035R73

NODE=S035R73

NODE=S035R73;LINKAGE=A

NODE=S035R73;LINKAGE=B9

NODE=S035R74

NODE=S035R74

NODE=S035R74

NODE=S035R74;LINKAGE=A

NODE=S035R74;LINKAGE=B9

NODE=S035B47

NODE=S035B47

NODE=S035B47

NODE=S035B47;LINKAGE=A

NODE=S035B47;LINKAGE=B9

NODE=S035B48

NODE=S035B48

NODE=S035B48

NODE=S035B48;LINKAGE=A

NODE=S035B48;LINKAGE=B9

$\Gamma(e^- \eta'(958))/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{204}/\Gamma$
$< 1.6 \times 10^{-7}$	90	MIYAZAKI 07	BELL	$401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C58
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 2.4 \times 10^{-7}$	90	AUBERT 07I	BABR	$339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C58
$< 10. \times 10^{-7}$	90	ENARI 05	BELL	$154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

 $\Gamma(\mu^- \eta'(958))/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{205}/\Gamma$
$< 1.3 \times 10^{-7}$	90	MIYAZAKI 07	BELL	$401 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C59
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 1.4 \times 10^{-7}$	90	AUBERT 07I	BABR	$339 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C59
$< 4.7 \times 10^{-7}$	90	ENARI 05	BELL	$154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

 $\Gamma(e^- f_0(980) \rightarrow e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{206}/\Gamma$
$< 3.2 \times 10^{-8}$	90	MIYAZAKI 09	BELL	$671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C70

 $\Gamma(\mu^- f_0(980) \rightarrow \mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{207}/\Gamma$
$< 3.4 \times 10^{-8}$	90	MIYAZAKI 09	BELL	$671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C71

 $\Gamma(e^- \phi)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{208}/\Gamma$
Test of lepton family number conservation.					
$< 2.0 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI 23	BELL	$980 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C16
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 3.1 \times 10^{-8}$	90	MIYAZAKI 11	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C16
$< 3.1 \times 10^{-8}$	90	AUBERT 09W	BABR	$451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C16
$< 7.3 \times 10^{-8}$	90	NISHIO 08	BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C16
$< 7.3 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 6.9 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

<sup>1</sup> Supersedes MIYAZAKI 11.

 $\Gamma(\mu^- \phi)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{209}/\Gamma$
$< 2.3 \times 10^{-8}$	90	<sup>1</sup> TSUZUKI 23	BELL	$980 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C17
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 8.4 \times 10^{-8}$	90	MIYAZAKI 11	BELL	$854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C17
$< 1.9 \times 10^{-7}$	90	AUBERT 09W	BABR	$451 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C17
$< 1.3 \times 10^{-7}$	90	NISHIO 08	BELL	$543 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035C17
$< 7.7 \times 10^{-7}$	90	YUSA 06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 7.0 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	

<sup>1</sup> Supersedes MIYAZAKI 11.

 $\Gamma(e^- e^+ e^-)/\Gamma_{\text{total}}$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT	$\Gamma_{210}/\Gamma$
$< 2.7 \times 10^{-8}$	90	HAYASAKA 10	BELL	$782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R58
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$< 2.9 \times 10^{-8}$	90	LEES 10A	BABR	$468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R58
$< 3.6 \times 10^{-8}$	90	MIYAZAKI 08	BELL	$535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R58
$< 4.3 \times 10^{-8}$	90	AUBERT 07BK	BABR	$376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R58
$< 2.0 \times 10^{-7}$	90	AUBERT 04J	BABR	$91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	NODE=S035R58
$< 3.5 \times 10^{-7}$	90	YUSA 04	BELL	$87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 2.9 \times 10^{-6}$	90	BLISS 98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$	
$< 0.33 \times 10^{-5}$	90	<sup>1</sup> BARTELTT 94	CLEO	Repl. by BLISS 98	
$< 1.3 \times 10^{-5}$	90	ALBRECHT 92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$	
$< 2.7 \times 10^{-5}$	90	BOWCOCK 90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$	
$< 40 \times 10^{-5}$	90	HAYES 82	MRK2	$E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$	

<sup>1</sup> BARTELTT 94 assume phase space decays.

$\Gamma(e^- \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{211}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.7 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 3.2 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 4.1 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.7 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 3.3 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.36 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$
$< 33 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTEL 94 assume phase space decays.

NODE=S035R56

NODE=S035R56

NODE=S035R56

 $\Gamma(e^+ \mu^- \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{212}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.7 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 2.6 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.3 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.6 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.3 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.35 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$< 1.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

NODE=S035R56;LINKAGE=B9

NODE=S035R75

NODE=S035R75

NODE=S035R75

 $\Gamma(\mu^- e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{213}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.8 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 2.2 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.7 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 8.0 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.7 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.7 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 0.34 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$< 1.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$
$< 44 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{\text{ee}} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> BARTEL 94 assume phase space decays.

NODE=S035R75;LINKAGE=B9

NODE=S035R57

NODE=S035R57

NODE=S035R57

 $\Gamma(\mu^+ e^- e^-)/\Gamma_{\text{total}}$  $\Gamma_{214}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.5 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 1.8 \times 10^{-8}$	90	LEES	10A	BABR $468 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 5.8 \times 10^{-8}$	90	AUBERT	07BK	BABR $376 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 1.1 \times 10^{-7}$	90	AUBERT	04J	BABR $91.5 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035R76;LINKAGE=B9

NODE=S035R76

NODE=S035R76

$<1.5 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
$<0.34 \times 10^{-5}$	90	<sup>1</sup> BARTELT	94	CLEO	Repl. by BLISS 98
$<1.4 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{cm}^{ee} = 10 \text{ GeV}$
$<1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{cm}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

NODE=S035R76;LINKAGE=B9

### $\Gamma(\mu^-\mu^+\mu^-)/\Gamma_{\text{total}}$

### $\Gamma_{215}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.9 \times 10^{-8}$	90	ADACHI	24R	BELL $424 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 2.9 \times 10^{-8}$	90	<sup>1</sup> HAYRAPETYAN	24J	CMS $131 \text{ fb}^{-1}, pp \text{ at } 13 \text{ TeV}$
$< 8.0 \times 10^{-8}$	90	<sup>2</sup> SIRUNYAN	21D	CMS $33.2 \text{ fb}^{-1}, pp \text{ at } 13 \text{ TeV}$
$< 3.8 \times 10^{-7}$	90	AAD	16BA ATLAS	$20.3 \text{ fb}^{-1}, pp \text{ at } 8 \text{ TeV}$
$< 4.6 \times 10^{-8}$	90	AAIJ	15AI LHCb	$3.0 \text{ fb}^{-1}, pp \text{ at } 7, 8 \text{ TeV}$
$< 8.0 \times 10^{-8}$	90	<sup>3</sup> AAIJ	13AH LHCb	$1.0 \text{ fb}^{-1}, pp \text{ at } 7 \text{ TeV}$
$< 2.1 \times 10^{-8}$	90	HAYASAKA	10	BELL $782 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 3.3 \times 10^{-8}$	90	LEES	10A BABR	$468 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 3.2 \times 10^{-8}$	90	MIYAZAKI	08	BELL $535 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 5.3 \times 10^{-8}$	90	AUBERT	07BK BABR	$376 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-7}$	90	AUBERT	04J BABR	$91.5 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 2.0 \times 10^{-7}$	90	YUSA	04	BELL $87.1 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 0.43 \times 10^{-5}$	90	<sup>4</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{cm}^{ee} = 10 \text{ GeV}$
$< 1.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{cm}^{ee} = 10.4\text{--}10.9$
$< 49 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{cm}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>1</sup> HAYRAPETYAN 24J use  $33.2 \text{ fb}^{-1}$  of data taken in 2016 and used for the same search in SIRUNYAN 21D and  $97.7 \text{ fb}^{-1}$  of data taken in 2017-2018.

<sup>2</sup> Superseded by HAYRAPETYAN 24J.

<sup>3</sup> Repl. by AAIJ 15AI.

<sup>4</sup> BARTELT 94 assume phase space decays.

NODE=S035R55;LINKAGE=B

NODE=S035R55;LINKAGE=C

NODE=S035R55;LINKAGE=A

NODE=S035R55;LINKAGE=B9

### $\Gamma(e^-\pi^+\pi^-)/\Gamma_{\text{total}}$

### $\Gamma_{216}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.3 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 4.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 7.3 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 1.2 \times 10^{-7}$	90	AUBERT,BE	05D BABR	$221 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 2.2 \times 10^{-6}$	90	BLISS	98	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 4.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 2.7 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{cm}^{ee} = 10 \text{ GeV}$
$< 6.0 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{cm}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

NODE=S035R65;LINKAGE=B9

### $\Gamma(e^+\pi^-\pi^-)/\Gamma_{\text{total}}$

### $\Gamma_{217}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.0 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$< 8.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$< 2.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 2.7 \times 10^{-7}$	90	AUBERT,BE	05D BABR	$221 \text{ fb}^{-1}, E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 1.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$
$< 4.4 \times 10^{-6}$	90	<sup>1</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 1.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{cm}^{ee} = 10 \text{ GeV}$
$< 1.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{cm}^{ee} = 10.4\text{--}10.9$

<sup>1</sup> BARTELT 94 assume phase space decays.

NODE=S035R66;LINKAGE=B9

$\Gamma(\mu^-\pi^+\pi^-)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.1 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.3 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<4.8 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.9 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<8.2 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<3.6 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma_{218}/\Gamma$ 

NODE=S035R67

NODE=S035R67

NODE=S035R67

 $\Gamma(\mu^+\pi^-\pi^-)/\Gamma_{\text{total}}$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.9 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.7 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<3.4 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7 \times 10^{-8}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.9 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<6.3 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma_{219}/\Gamma$ 

NODE=S035R68

NODE=S035R68

NODE=S035R68

 $\Gamma(e^-\pi^+K^-)/\Gamma_{\text{total}}$  $\Gamma_{220}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.7 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<5.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<7.2 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.7 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<2.9 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<5.8 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

NODE=S035R68;LINKAGE=B9

 $\Gamma(e^-\pi^-K^+)/\Gamma_{\text{total}}$  $\Gamma_{221}/\Gamma$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.1 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<5.2 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<1.6 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.7 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.6 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

NODE=S035R69;LINKAGE=B9

 $\Gamma(e^+\pi^-K^-)/\Gamma_{\text{total}}$  $\Gamma_{222}/\Gamma$ 

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.2 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035R86

NODE=S035R86

NODE=S035R86

NODE=S035R86;LINKAGE=B9

NODE=S035R70

NODE=S035R70

NODE=S035R70

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.7 \times 10^{-8}$	90	MIYAZAKI	10	BELL	Repl. by MIYAZAKI 13
$<1.9 \times 10^{-7}$	90	YUSA	06	BELL	$158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.8 \times 10^{-7}$	90	AUBERT,BE	05D	BABR	$221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.1 \times 10^{-6}$	90	BLISS	98	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.5 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO	Repl. by BLISS 98
$<2.0 \times 10^{-5}$	90	ALBRECHT	92K	ARG	$E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<4.9 \times 10^{-5}$	90	BOWCOCK	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

### $\Gamma(e^- K_S^0 K_S^0)/\Gamma_{\text{total}}$

$\Gamma_{223}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.1 \times 10^{-8}$	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<2.2 \times 10^{-6}$	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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### $\Gamma(e^- K^+ K^-)/\Gamma_{\text{total}}$

$\Gamma_{224}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<5.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<3.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.4 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(e^+ K^- K^-)/\Gamma_{\text{total}}$

$\Gamma_{225}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.3 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.0 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<3.1 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.5 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\Gamma(\mu^- \pi^+ K^-)/\Gamma_{\text{total}}$

$\Gamma_{226}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8.6 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.6 \times 10^{-7}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<2.7 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.6 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<8.7 \times 10^{-6}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<11 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<7.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

### $\Gamma(\mu^- \pi^- K^+)/\Gamma_{\text{total}}$

$\Gamma_{227}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.5 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.0 \times 10^{-7}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<7.3 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<3.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<1.5 \times 10^{-5}$	90	<sup>1</sup> BARTEL	94	CLEO Repl. by BLISS 98
$<7.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays.

NODE=S035R70;LINKAGE=B9

NODE=S035C55

NODE=S035C55

NODE=S035C55

NODE=S035C12

NODE=S035C12

NODE=S035C12

NODE=S035C13

NODE=S035C13

NODE=S035C13

NODE=S035R71

NODE=S035R71

NODE=S035R71

NODE=S035R87;LINKAGE=B9

NODE=S035R87

NODE=S035R87

NODE=S035R87;LINKAGE=B9

$\Gamma(\mu^+\pi^-K^-)/\Gamma_{\text{total}}$  $\Gamma_{228}/\Gamma$ 

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<4.8 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<9.4 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<2.9 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.2 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<7.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.0 \times 10^{-5}$	90	BARTEL	94	CLEO Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	ALBRECHT	92K	ARG $E_{\text{cm}}^{\text{ee}} = 10 \text{ GeV}$
$<4.0 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{\text{ee}} = 10.4\text{--}10.9$

<sup>1</sup> BARTEL 94 assume phase space decays. $\Gamma(\mu^-K_S^0K_S^0)/\Gamma_{\text{total}}$  $\Gamma_{229}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<8.0 \times 10^{-8}$	90	MIYAZAKI	10A	BELL $671 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<3.4 \times 10^{-6}$	90	CHEN	02C	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^-K^+K^-)/\Gamma_{\text{total}}$  $\Gamma_{230}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<4.4 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<6.8 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<8.0 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<2.5 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<15 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^+K^-K^-)/\Gamma_{\text{total}}$  $\Gamma_{231}/\Gamma$ 

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<4.7 \times 10^{-8}$	90	MIYAZAKI	13	BELL $854 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
$<9.6 \times 10^{-8}$	90	MIYAZAKI	10	BELL Repl. by MIYAZAKI 13
$<4.4 \times 10^{-7}$	90	YUSA	06	BELL $158 \text{ fb}^{-1} E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<4.8 \times 10^{-7}$	90	AUBERT,BE	05D	BABR $221 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
$<6.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^-\pi^0\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{232}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<6.5 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^-\pi^0\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{233}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<14 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(e^-\eta\eta)/\Gamma_{\text{total}}$  $\Gamma_{234}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<35 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma(\mu^-\eta\eta)/\Gamma_{\text{total}}$  $\Gamma_{235}/\Gamma$ 

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<60 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035R72

NODE=S035R72

NODE=S035R72

NODE=S035R72;LINKAGE=B9

NODE=S035C56

NODE=S035C56

NODE=S035C14

NODE=S035C14

NODE=S035C14

NODE=S035C15

NODE=S035C15

NODE=S035C15

NODE=S035B84

NODE=S035B84

NODE=S035B84

NODE=S035B85

NODE=S035B85

NODE=S035B85

NODE=S035B86

NODE=S035B86

NODE=S035B86

$\Gamma(e^- \pi^0 \eta)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<24 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma_{236}/\Gamma$ 

NODE=S035B87  
NODE=S035B87  
NODE=S035B87

 $\Gamma(\mu^- \pi^0 \eta)/\Gamma_{\text{total}}$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<22 \times 10^{-6}$	90	BONVICINI	97	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

 $\Gamma_{237}/\Gamma$ 

NODE=S035B88  
NODE=S035B88  
NODE=S035B88

 $\Gamma(p e^- e^-)/\Gamma_{\text{total}}$  $\Gamma_{238}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.0 \times 10^{-8}$	90	SAHOO	20	BELL $921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035P02  
NODE=S035P02

 $\Gamma(\bar{p} e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{239}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.0 \times 10^{-8}$	90	SAHOO	20	BELL $921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035P03  
NODE=S035P03

 $\Gamma(\bar{p} e^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{240}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.0 \times 10^{-8}$	90	SAHOO	20	BELL $921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035P04  
NODE=S035P04

 $\Gamma(\bar{p} e^- \mu^+)/\Gamma_{\text{total}}$  $\Gamma_{241}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-8}$	90	SAHOO	20	BELL $921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035P05  
NODE=S035P05

 $\Gamma(p \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{242}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.0 \times 10^{-8}$	90	SAHOO	20	BELL $921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035D02  
NODE=S035D02

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<4.4 \times 10^{-7}$  90 AAIJ 13AH LHCb  $1.0 \text{ fb}^{-1}$ ,  $pp$  at 7 TeV

 $\Gamma(\bar{p} \mu^+ \mu^-)/\Gamma_{\text{total}}$  $\Gamma_{243}/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.8 \times 10^{-8}$	90	SAHOO	20	BELL $921 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035D03  
NODE=S035D03

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<3.3 \times 10^{-7}$  90 AAIJ 13AH LHCb  $1.0 \text{ fb}^{-1}$ ,  $pp$  at 7 TeV

 $\Gamma(\bar{p} \gamma)/\Gamma_{\text{total}}$  $\Gamma_{244}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.5 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035B15  
NODE=S035B15  
NODE=S035B15

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<29 \times 10^{-5}$  90 ALBRECHT 92K ARG  $E_{\text{cm}}^{ee} = 10 \text{ GeV}$

 $\Gamma(\bar{p} \pi^0)/\Gamma_{\text{total}}$  $\Gamma_{245}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<15 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035B16  
NODE=S035B16  
NODE=S035B16

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<66 \times 10^{-5}$  90 ALBRECHT 92K ARG  $E_{\text{cm}}^{ee} = 10 \text{ GeV}$

 $\Gamma(\bar{p} 2\pi^0)/\Gamma_{\text{total}}$  $\Gamma_{246}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<33 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035C29  
NODE=S035C29  
NODE=S035C29

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<130 \times 10^{-5}$  90 ALBRECHT 92K ARG  $E_{\text{cm}}^{ee} = 10 \text{ GeV}$

 $\Gamma(\bar{p} \eta)/\Gamma_{\text{total}}$  $\Gamma_{247}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.9 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

NODE=S035B17  
NODE=S035B17  
NODE=S035B17

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<130 \times 10^{-5}$  90 ALBRECHT 92K ARG  $E_{\text{cm}}^{ee} = 10 \text{ GeV}$

$\Gamma(\bar{p}\pi^0\eta)/\Gamma_{\text{total}}$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<27 \times 10^{-6}$	90	GODANG	99	CLEO $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

 $\Gamma_{248}/\Gamma$ 

NODE=S035C30

NODE=S035C30

NODE=S035C30

 $\Gamma(\Lambda\pi^-)/\Gamma_{\text{total}}$  $\Gamma_{249}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.7 \times 10^{-8}$	90	ADACHI	24Q	BELL $364 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C64

NODE=S035C64

NODE=S035C64

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<7.2 \times 10^{-8}$	90	MIYAZAKI	06	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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 $\Gamma(\bar{\Lambda}\pi^-)/\Gamma_{\text{total}}$  $\Gamma_{250}/\Gamma$ 

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<4.3 \times 10^{-8}$	90	ADACHI	24Q	BELL $364 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

NODE=S035C65

NODE=S035C65

NODE=S035C65

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<14 \times 10^{-8}$	90	MIYAZAKI	06	BELL $154 \text{ fb}^{-1}, E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$
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 $\Gamma(e^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  $\Gamma_{251}/\Gamma_5$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<5.3 \times 10^{-3}$	95	1 ADACHI	23A	BELL $E_{\text{cm}}^{\text{ee}} = 10.58 \text{ GeV}$

NODE=S035B11

NODE=S035B11

NODE=S035B11

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.008	95	2 BRYMAN	21	RVUE
<0.015	95	3 ALBRECHT	95G	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
<0.018	95	4 ALBRECHT	90E	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
<0.040	95	5 BALTRUSAIT..85	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$

1 ADACHI 23A limit holds for bosons with mass = 0.0 GeV. The limit rises to  $9.7 \times 10^{-3}$  for a mass of 1.0 GeV, then falls to  $1.1 \times 10^{-3}$  at the upper mass limit of 1.6 GeV.2 BRYMAN 21 reports indirect limits obtained from the consistency of the world averages of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  with their Standard Model predictions, without a simulation of the efficiency as a function of the  $X$  mass for the searched decay modes to be detected as the corresponding Standard Model decay modes.

3 ALBRECHT 95G limit holds for bosons with mass &lt; 0.4 GeV. The limit rises to 0.036 for a mass of 1.0 GeV, then falls to 0.006 at the upper mass limit of 1.6 GeV.

4 ALBRECHT 90E limit applies for spinless boson with mass &lt; 100 MeV, and rises to 0.050 for mass = 500 MeV.

5 BALTRUSAITIS 85 limit applies for spinless boson with mass &lt; 100 MeV.

NODE=S035B11;LINKAGE=G

NODE=S035B11;LINKAGE=E

 $\Gamma(\mu^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  $\Gamma_{252}/\Gamma_5$ 

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.4 \times 10^{-3}$	95	1 ADACHI	23A	BELL $E_{\text{cm}}^{\text{ee}} = 10.58 \text{ GeV}$

NODE=S035B12

NODE=S035B12

NODE=S035B12

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.011	95	2 BRYMAN	21	RVUE
<0.026	95	3 ALBRECHT	95G	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
<0.033	95	4 ALBRECHT	90E	ARG $E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6 \text{ GeV}$
<0.125	95	5 BALTRUSAIT..85	MRK3	$E_{\text{cm}}^{\text{ee}} = 3.77 \text{ GeV}$

1 ADACHI 23A limit holds for bosons with mass = 0.0 GeV. The limit rises to  $12.2 \times 10^{-3}$  for a mass of 1.0 GeV, then falls to  $0.7 \times 10^{-3}$  at the upper mass limit of 1.6 GeV.2 BRYMAN 21 reports indirect limits obtained from the consistency of the world averages of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  with their Standard Model predictions, without a simulation of the efficiency as a function of the  $X$  mass for the searched decay modes to be detected as the corresponding Standard Model decay modes.

3 ALBRECHT 95G limit holds for bosons with mass &lt; 1.3 GeV. The limit rises to 0.034 for a mass of 1.4 GeV, then falls to 0.003 at the upper mass limit of 1.6 GeV.

4 ALBRECHT 90E limit applies for spinless boson with mass &lt; 100 MeV, and rises to 0.071 for mass = 500 MeV.

5 BALTRUSAITIS 85 limit applies for spinless boson with mass &lt; 100 MeV.

NODE=S035B12;LINKAGE=F

NODE=S035B12;LINKAGE=E

NODE=S035B12;LINKAGE=C

NODE=S035B12;LINKAGE=B

NODE=S035B12;LINKAGE=A

## $\tau$ -DECAY PARAMETERS

See the related review(s):

[τ-Lepton Decay Parameters](#)

### $\rho(e \text{ or } \mu)$ PARAMETER

(V-A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.745±0.008 OUR FIT</b>				
<b>0.749±0.008 OUR AVERAGE</b>				

0.742±0.014±0.006	81k	HEISTER	01E	ALEP	1991–1995 LEP runs
0.775±0.023±0.020	36k	ABREU	00L	DLPH	1992–1995 runs
0.781±0.028±0.018	46k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
0.762±0.035	54k	ACCIARRI	98R	L3	1991–1995 LEP runs
0.731±0.031		1 ALBRECHT	98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.72 ± 0.09 ± 0.03		2 ABE	970	SLD	1993–1995 SLC runs
0.747±0.010±0.006	55k	ALEXANDER	97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
0.79 ± 0.10 ± 0.10	3732	FORD	87B	MAC	$E_{cm}^{ee} = 29 \text{ GeV}$
0.71 ± 0.09 ± 0.03	1426	BEHRENDS	85	CLEO	$e^+ e^-$ near $\gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.735±0.013±0.008	31k	AMMAR	97B	CLEO	Repl. by ALEXANDER 97F
0.794±0.039±0.031	18k	ACCIARRI	96H	L3	Repl. by ACCIARRI 98R
0.732±0.034±0.020	8.2k	3 ALBRECHT	95	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.738±0.038		4 ALBRECHT	95C	ARG	Repl. by ALBRECHT 98
0.751±0.039±0.022		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
0.742±0.035±0.020	8000	ALBRECHT	90E	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 970 assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\rho$  value of  $0.69 \pm 0.13 \pm 0.05$ .

<sup>3</sup> Value is from a simultaneous fit for the  $\rho$  and  $\eta$  decay parameters to the lepton energy spectrum. Not independent of ALBRECHT 90E  $\rho(e \text{ or } \mu)$  value which assumes  $\eta = 0$ . Result is strongly correlated with ALBRECHT 95C.

<sup>4</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

### $\rho(e)$ PARAMETER

(V-A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.747±0.010 OUR FIT</b>				
<b>0.744±0.010 OUR AVERAGE</b>				

0.747±0.019±0.014	44k	HEISTER	01E	ALEP	1991–1995 LEP runs
0.744±0.036±0.037	17k	ABREU	00L	DLPH	1992–1995 runs
0.779±0.047±0.029	25k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
0.68 ± 0.04 ± 0.07		1 ALBRECHT	98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.71 ± 0.14 ± 0.05		ABE	970	SLD	1993–1995 SLC runs
0.747±0.012±0.004	34k	ALEXANDER	97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
0.735±0.036±0.020	4.7k	2 ALBRECHT	95	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.79 ± 0.08 ± 0.06	3230	3 ALBRECHT	93G	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.64 ± 0.06 ± 0.07	2753	JANSSEN	89	CBAL	$E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.62 ± 0.17 ± 0.14	1823	FORD	87B	MAC	$E_{cm}^{ee} = 29 \text{ GeV}$
0.60 ± 0.13	699	BEHRENDS	85	CLEO	$e^+ e^-$ near $\gamma(4S)$
0.72 ± 0.10 ± 0.11	594	BACINO	79B	DLCO	$E_{cm}^{ee} = 3.5\text{--}7.4 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.732±0.014±0.009	19k	AMMAR	97B	CLEO	Repl. by ALEXANDER 97F
0.793±0.050±0.025		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
0.747±0.045±0.028	5106	ALBRECHT	90E	ARG	Repl. by ALBRECHT 95

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(\pi^+\pi^0\bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ALBRECHT 95 use tau pair events of the type  $\tau^-\tau^+ \rightarrow (\ell^-\bar{\nu}_\ell\nu_\tau)(h^+h^-(\pi^0)\bar{\nu}_\tau)$  and their charged conjugates.

<sup>3</sup> ALBRECHT 93G use tau pair events of the type  $\tau^-\tau^+ \rightarrow (\mu^-\bar{\nu}_\mu\nu_\tau)(e^+\nu_e\bar{\nu}_\tau)$  and their charged conjugates.

NODE=S035225

NODE=S035RHO

NODE=S035RHO

NODE=S035RHO

NODE=S035RHO;LINKAGE=RC

NODE=S035RHO;LINKAGE=D

NODE=S035RHO;LINKAGE=A

NODE=S035RHO;LINKAGE=C

NODE=S035RHE

NODE=S035RHE

NODE=S035RHE

NODE=S035RHE;LINKAGE=RC

NODE=S035RHE;LINKAGE=A

NODE=S035RHE;LINKAGE=B

## $\rho(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.763±0.020 OUR FIT</b>				
<b>0.770±0.022 OUR AVERAGE</b>				
0.776±0.045±0.019	46k	HEISTER	01E	ALEP 1991–1995 LEP runs
0.999±0.098±0.045	22k	ABREU	00L	DLPH 1992–1995 runs
0.777±0.044±0.016	27k	ACKERSTAFF	99D	OPAL 1990–1995 LEP runs
0.69 ± 0.06 ± 0.06		<sup>1</sup> ALBRECHT	98	ARG $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.54 ± 0.28 ± 0.14		ABE	970	SLD 1993–1995 SLC runs
0.750±0.017±0.045	22k	ALEXANDER	97F	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$
0.76 ± 0.07 ± 0.08	3230	ALBRECHT	93G	ARG $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.734±0.055±0.027	3041	ALBRECHT	90E	ARG $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
0.89 ± 0.14 ± 0.08	1909	FORD	87B	MAC $E_{cm}^{ee} = 29 \text{ GeV}$
0.81 ± 0.13	727	BEHRENDS	85	CLEO $e^+ e^-$ near $\gamma(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.747±0.048±0.044	13k	AMMAR	97B	CLEO Repl. by ALEXANDER 97F
0.693±0.057±0.028		BUSKULIC	95D	ALEP Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

NODE=S035RHM

NODE=S035RHM

NODE=S035RHM

## $\xi(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.985±0.030 OUR FIT</b>				
<b>0.981±0.031 OUR AVERAGE</b>				
0.986±0.068±0.031	81k	HEISTER	01E	ALEP 1991–1995 LEP runs
0.929±0.070±0.030	36k	ABREU	00L	DLPH 1992–1995 runs
0.98 ± 0.22 ± 0.10	46k	ACKERSTAFF	99D	OPAL 1990–1995 LEP runs
0.70 ± 0.16	54k	ACCIARRI	98R	L3 1991–1995 LEP runs
1.03 ± 0.11		<sup>1</sup> ALBRECHT	98	ARG $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
1.05 ± 0.35 ± 0.04		<sup>2</sup> ABE	970	SLD 1993–1995 SLC runs
1.007±0.040±0.015	55k	ALEXANDER	97F	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.94 ± 0.21 ± 0.07	18k	ACCIARRI	96H	L3 Repl. by ACCIARRI 98R
0.97 ± 0.14		<sup>3</sup> ALBRECHT	95C	ARG Repl. by ALBRECHT 98
1.18 ± 0.15 ± 0.16		BUSKULIC	95D	ALEP Repl. by HEISTER 01E
0.90 ± 0.15 ± 0.10	3230	<sup>4</sup> ALBRECHT	93G	ARG $E_{cm}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

<sup>1</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>2</sup> ABE 970 assume  $\eta = 0$  in their fit. Letting  $\eta$  vary in the fit gives a  $\xi$  value of  $1.02 \pm 0.36 \pm 0.05$ .

<sup>3</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

<sup>4</sup> ALBRECHT 93G measurement determines  $|\xi|$  for the case  $\xi(e) = \xi(\mu)$ , but the authors point out that other LEP experiments determine the sign to be positive.

NODE=S035RHM;LINKAGE=RC

NODE=S035XI

NODE=S035XI

NODE=S035XI

OCCUR=2

NODE=S035XI;LINKAGE=RC

NODE=S035XI;LINKAGE=B

NODE=S035XI;LINKAGE=FF

NODE=S035XI;LINKAGE=A

NODE=S035XE

NODE=S035XE

NODE=S035XE

## $\xi(e)$ PARAMETER

( $V-A$ ) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.994±0.040 OUR FIT</b>				
<b>1.00 ± 0.04 OUR AVERAGE</b>				
1.011±0.094±0.038	44k	HEISTER	01E	ALEP 1991–1995 LEP runs
1.01 ± 0.12 ± 0.05	17k	ABREU	00L	DLPH 1992–1995 runs
1.13 ± 0.39 ± 0.14	25k	ACKERSTAFF	99D	OPAL 1990–1995 LEP runs
1.11 ± 0.20 ± 0.08		<sup>1</sup> ALBRECHT	98	ARG $E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
1.16 ± 0.52 ± 0.06		ABE	970	SLD 1993–1995 SLC runs
0.979±0.048±0.016	34k	ALEXANDER	97F	CLEO $E_{cm}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.03 ± 0.23 ± 0.09		BUSKULIC	95D	ALEP Repl. by HEISTER 01E

<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

NODE=S035XE;LINKAGE=RC

## $\xi(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.030±0.059 OUR FIT</b>				

### 1.06 ± 0.06 OUR AVERAGE

1.030 ± 0.120 ± 0.050	46k	HEISTER	01E	ALEP	1991–1995 LEP runs
1.16 ± 0.19 ± 0.06	22k	ABREU	00L	DLPH	1992–1995 runs
0.79 ± 0.41 ± 0.09	27k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
1.26 ± 0.27 ± 0.14		<sup>1</sup> ALBRECHT	98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.75 ± 0.50 ± 0.14		ABE	970	SLD	1993–1995 SLC runs
1.054 ± 0.069 ± 0.047	22k	ALEXANDER	97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.23 ± 0.22 ± 0.10		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E
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<sup>1</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

NODE=S035XM

NODE=S035XM

NODE=S035XM

## $\eta(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.013±0.020 OUR FIT</b>				

### 0.015±0.021 OUR AVERAGE

0.012 ± 0.026 ± 0.004	81k	HEISTER	01E	ALEP	1991–1995 LEP runs
-0.005 ± 0.036 ± 0.037		ABREU	00L	DLPH	1992–1995 runs
0.027 ± 0.055 ± 0.005	46k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
0.27 ± 0.14	54k	ACCIARRI	98R	L3	1991–1995 LEP runs
-0.13 ± 0.47 ± 0.15		ABE	970	SLD	1993–1995 SLC runs
-0.015 ± 0.061 ± 0.062	31k	AMMAR	97B	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$
0.03 ± 0.18 ± 0.12	8.2k	ALBRECHT	95	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.25 ± 0.17 ± 0.11	18k	ACCIARRI	96H	L3	Repl. by ACCIARRI 98R
-0.04 ± 0.15 ± 0.11		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

NODE=S035XM;LINKAGE=RC

NODE=S035ETA

NODE=S035ETA

NODE=S035ETA

## $\eta(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.094±0.073 OUR FIT</b>				

### 0.17 ± 0.15 OUR AVERAGE

0.160 ± 0.150 ± 0.060	46k	HEISTER	01E	ALEP	1991–1995 LEP runs
0.72 ± 0.32 ± 0.15		ABREU	00L	DLPH	1992–1995 runs
-0.59 ± 0.82 ± 0.45		<sup>1</sup> ABE	970	SLD	1993–1995 SLC runs
0.010 ± 0.149 ± 0.171	13k	<sup>2</sup> AMMAR	97B	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.010 ± 0.065 ± 0.001	27k	<sup>3</sup> ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
-0.24 ± 0.23 ± 0.18		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

NODE=S035ETM

NODE=S035ETM

NODE=S035ETM

1 Highly correlated (corr. = 0.92) with ABE 970  $\rho(\mu)$  measurement.

2 Highly correlated (corr. = 0.949) with AMMAR 97B  $\rho(\mu)$  value.

3 ACKERSTAFF 99D result is dominated by a constraint on  $\eta$  from the OPAL measurements of the  $\tau$  lifetime and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  assuming lepton universality for the total coupling strength.

NODE=S035ETM;LINKAGE=B

NODE=S035ETM;LINKAGE=A

NODE=S035ETM;LINKAGE=K9

## $(\delta\xi)(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.746±0.021 OUR FIT</b>				

### 0.744±0.022 OUR AVERAGE

0.776 ± 0.045 ± 0.024	81k	HEISTER	01E	ALEP	1991–1995 LEP runs
0.779 ± 0.070 ± 0.028	36k	ABREU	00L	DLPH	1992–1995 runs
0.65 ± 0.14 ± 0.07	46k	ACKERSTAFF	99D	OPAL	1990–1995 LEP runs
0.70 ± 0.11	54k	ACCIARRI	98R	L3	1991–1995 LEP runs
0.63 ± 0.09		<sup>1</sup> ALBRECHT	98	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6 \text{ GeV}$
0.88 ± 0.27 ± 0.04		<sup>2</sup> ABE	970	SLD	1993–1995 SLC runs
0.745 ± 0.026 ± 0.009	55k	ALEXANDER	97F	CLEO	$E_{cm}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.81 ± 0.14 ± 0.06	18k	ACCIARRI	96H	L3	Repl. by ACCIARRI 98R
0.65 ± 0.12		<sup>3</sup> ALBRECHT	95C	ARG	Repl. by ALBRECHT 98
0.88 ± 0.11 ± 0.07		BUSKULIC	95D	ALEP	Repl. by HEISTER 01E

OCCUR=2



## $\xi(a_1)$ PARAMETER

(V-A) theory predicts  $\xi(a_1) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.001±0.027 OUR FIT</b>				
<b>1.002±0.028 OUR AVERAGE</b>				
1.000±0.016±0.024	35k	1 HEISTER 01E	ALEP	1991–1995 LEP runs
1.02 ± 0.13 ± 0.03	17.2k	2 ASNER 00	CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.29 ± 0.26 ± 0.11	7.4k	ACKERSTAFF 97R	OPAL	1992–1994 LEP runs
0.85 ± 0.15 ± 0.05		ALBRECHT 95C	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
1.25 ± 0.23 ± 0.15	7.5k	ALBRECHT 93C	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.08 ± 0.46 ± 0.14	2.6k	3 AKERS 95P	OPAL	Repl. by ACKERSTAFF 97R
0.937±0.116±0.064		BUSKULIC 95D	ALEP	Repl. by HEISTER 01E

1 HEISTER 01E quote  $1.000 \pm 0.016 \pm 0.013 \pm 0.020$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty.

2 ACKERSTAFF 97R obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.16 \pm 0.04$ , and with the model of Isgur *et al.* (PR **D39**, 1357 (1989)) they obtain  $1.20 \pm 0.21 \pm 0.14$ .

3 AKERS 95P obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.27^{+0.05}_{-0.06}$ , and with the model of Isgur *et al.* (PR **D39**, 1357 (1989)) they obtain  $1.10 \pm 0.31^{+0.13}_{-0.14}$ .

## $\xi(\text{all hadronic modes})$ PARAMETER

(V-A) theory predicts  $\xi = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.995±0.007 OUR FIT</b>				
<b>0.997±0.007 OUR AVERAGE</b>				

0.992±0.007±0.008	102k	1 HEISTER 01E	ALEP	1991–1995 LEP runs
0.997±0.027±0.011	39k	2 ABREU 00L	DLPH	1992–1995 runs
1.02 ± 0.13 ± 0.03	17.2k	3 ASNER 00	CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.032±0.031	37k	4 ACCIARRI 98R	L3	1991–1995 LEP runs
0.93 ± 0.10 ± 0.04		ABE 970	SLD	1993–1995 SLC runs
1.29 ± 0.26 ± 0.11	7.4k	5 ACKERSTAFF 97R	OPAL	1992–1994 LEP runs
0.995±0.010±0.003	66k	6 ALEXANDER 97F	CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.03 ± 0.06 ± 0.04	2.0k	7 COAN 97	CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.017±0.039		8 ALBRECHT 95C	ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
1.25 ± 0.23 ± 0.15	7.5k	9 ALBRECHT 93C	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.970±0.053±0.011	14k	10 ACCIARRI 96H	L3	Repl. by ACCIARRI 98R
1.08 ± 0.46 ± 0.14	2.6k	11 AKERS 95P	OPAL	Repl. by ACKERSTAFF 97R
1.006±0.032±0.019		12 BUSKULIC 95D	ALEP	Repl. by HEISTER 01E
1.022±0.028±0.030	1.7k	13 ALBRECHT 94E	ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
0.99 ± 0.07 ± 0.04		14 BUSKULIC 94D	ALEP	1990+1991 LEP run
1.14 ± 0.34 ± 0.34	3.9k	9 ALBRECHT 90I	ARG	Repl. by ALBRECHT 93C

1 HEISTER 01E quote  $0.992 \pm 0.007 \pm 0.006 \pm 0.005$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty. They use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ ,  $\tau \rightarrow \rho\nu_\tau$ , and  $\tau \rightarrow a_1\nu_\tau$  decays.

2 ABREU 00L use  $\tau^- \rightarrow h^- \geq 0\pi^0\nu_\tau$  decays.

3 ASNER 00 use  $\tau^- \rightarrow \pi^- 2\pi^0\nu_\tau$  decays.

4 ACCIARRI 98R use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ , and  $\tau \rightarrow \rho\nu_\tau$  decays.

5 ACKERSTAFF 97R use  $\tau \rightarrow a_1\nu_\tau$  decays.

6 ALEXANDER 97F use  $\tau \rightarrow \rho\nu_\tau$  decays.

7 COAN 97 use  $h^+ h^-$  energy correlations.

8 Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

9 Uses  $\tau \rightarrow a_1\nu_\tau$  decays. Replaced by ALBRECHT 95C.

10 ACCIARRI 96H use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ , and  $\tau \rightarrow \rho\nu_\tau$  decays.

11 AKERS 95P use  $\tau \rightarrow a_1\nu_\tau$  decays.

NODE=S035XA1

NODE=S035XA1

NODE=S035XA1

OCCUR=2

NODE=S035XA1;LINKAGE=AA

NODE=S035XA1;LINKAGE=C

NODE=S035XA1;LINKAGE=D

NODE=S035XAL

NODE=S035XAL

NODE=S035XAL

NODE=S035XAL;LINKAGE=AA

NODE=S035XAL;LINKAGE=LO

NODE=S035XAL;LINKAGE=SA

NODE=S035XAL;LINKAGE=BX

NODE=S035XAL;LINKAGE=H

NODE=S035XAL;LINKAGE=I

NODE=S035XAL;LINKAGE=J

NODE=S035XAL;LINKAGE=F

NODE=S035XAL;LINKAGE=B

NODE=S035XAL;LINKAGE=G

NODE=S035XAL;LINKAGE=D

12 BUSKULIC 95D use  $\tau \rightarrow \pi \nu_\tau$ ,  $\tau \rightarrow \rho \nu_\tau$ , and  $\tau \rightarrow a_1 \nu_\tau$  decays.

13 ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result. Uses  $\tau \rightarrow a_1 \nu_\tau$  decays. Replaced by ALBRECHT 95C.

14 BUSKULIC 94D use  $\tau \rightarrow \pi \nu_\tau$  and  $\tau \rightarrow \rho \nu_\tau$  decays. Superseded by BUSKULIC 95D.

### $\xi'(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\xi' = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22±0.94±0.42</b>	165	BODROV	23	BELL $E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

### $\bar{\eta}(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\bar{\eta}(\mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-1.3±1.5±0.8</b>	71k	1 SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(\mu)$ ,  $\xi\kappa(\mu)$  and  $\eta''(\mu)$ , floating  $\bar{\eta}(\mu)$  and  $\xi\kappa(\mu)$  and fixing  $\eta''(\mu) = 0$ . The contribution of  $\eta''(\mu)$  is suppressed by  $m_\mu/m_\tau$ .

### $(\xi\kappa)(e \text{ or } \mu)$ PARAMETER

( $V-A$ ) theory predicts  $(\xi\kappa)(e \text{ or } \mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.5±0.4±0.2</b>	149k	1,2 SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e \gamma$ and $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution of radiative tau decays into both electrons and muons affected by  $\bar{\eta}(e \text{ or } \mu)$ ,  $\xi\kappa(e \text{ or } \mu)$  and  $\eta''(e \text{ or } \mu)$ , floating  $\bar{\eta}(e \text{ or } \mu)$  and  $\xi\kappa(e \text{ or } \mu)$  and fixing  $\eta''(e \text{ or } \mu) = 0$ . The contribution of  $\eta''(e \text{ or } \mu)$  is suppressed by  $m_e/m_\tau$  for tau decaying to electrons and by  $m_\mu/m_\tau$  for tau decaying to muons.

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e)$  and  $(\xi\kappa)(\mu)$  values.

### $(\xi\kappa)(e)$ PARAMETER

( $V-A$ ) theory predicts  $(\xi\kappa)(e) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>-0.4±0.8±0.9</b>	78k	1,2 SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau e^- \bar{\nu}_e \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(e)$ ,  $(\xi\kappa)(e)$  and  $\eta''(e)$ , floating  $(\xi\kappa)(e)$  and fixing  $\bar{\eta}(e) = 0$  and  $\eta''(e) = 0$ . The contribution of  $\eta''(e)$  is suppressed by  $m_e/m_\tau$ .

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e \text{ or } \mu)$  value.

### $(\xi\kappa)(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $(\xi\kappa)(\mu) = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.8±0.5±0.3</b>	71k	1,2 SHIMIZU	18A	BELL $\tau^- \rightarrow \nu_\tau \mu^- \bar{\nu}_\mu \gamma$

<sup>1</sup> SHIMIZU 18A measurement procedure fits a distribution affected by  $\bar{\eta}(\mu)$ ,  $\xi\kappa(\mu)$  and  $\eta''(\mu)$ , floating  $\bar{\eta}(\mu)$  and  $\xi\kappa(\mu)$  and fixing  $\eta''(\mu) = 0$ . The contribution of  $\eta''(\mu)$  is suppressed by  $m_\mu/m_\tau$ .

<sup>2</sup> Error correlated with SHIMIZU 18A  $(\xi\kappa)(e \text{ or } \mu)$  value.

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PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	REFID=51004
YUSA	06	PL B640 138	Y. Yusa <i>et al.</i>	REFID=51378
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ABDALLAH	04T	EPJ C36 283	J. Abdallah <i>et al.</i>	REFID=50145
ABE	04B	PRL 92 171802	K. Abe <i>et al.</i>	REFID=49938
ACHARD	04G	PL B585 53	P. Achard <i>et al.</i>	REFID=50105
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ENARI	04	PRL 93 081803	Y. Enari <i>et al.</i>	REFID=50073
PDG	04	PL B592 1	S. Eidelman <i>et al.</i>	REFID=49653
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ABBIENDI	03	PL B551 35	G. Abbiendi <i>et al.</i>	REFID=49150
BRIERE	03	PRL 90 181802	R. A. Briere <i>et al.</i>	REFID=49360
HEISTER	03F	EPJ C30 291	A. Heister <i>et al.</i>	REFID=49552
INAMI	03	PL B551 16	K. Inami <i>et al.</i>	REFID=49149
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REGAN	02	PRL 88 071805	B.C. Regan <i>et al.</i>	REFID=48608
ABBIENDI	01J	EPJ C19 653	G. Abbiendi <i>et al.</i>	REFID=48146
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ACCIARRI	01F	PL B507 47	M. Acciari <i>et al.</i>	REFID=48143
ACHARD	01D	PL B519 189	P. Achard <i>et al.</i>	REFID=48384
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HEISTER	01E	EPJ C22 217	A. Heister <i>et al.</i>	REFID=48509
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ACKERSTAFF	99D	EPJ C8 3	K. Ackerstaff <i>et al.</i>	REFID=47012
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BARATE	99K	EPJ C10 1	R. Barate <i>et al.</i>	REFID=47181
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BISHAI	99	PRL 82 281	M. Bishai <i>et al.</i>	REFID=46549
GODANG	99	PR D59 091303	R. Godang <i>et al.</i>	REFID=46998
RICHICHI	99	PR D60 112002	S.J. Richichi <i>et al.</i>	REFID=47268

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ACCIARRI	98R	PL B438 405	M. Acciarri <i>et al.</i>	(L3 Collab.)	REFID=46502
ACKERSTAFF	98M	EPJ C4 193	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)	REFID=45944
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ALBRECHT	98	PL B431 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=46067
BARATE	98	EPJ C1 65	R. Barate <i>et al.</i>	(ALEPH Collab.)	REFID=45809
BARATE	98E	EPJ C4 29	R. Barate <i>et al.</i>	(ALEPH Collab.)	REFID=45917
BLISS	98	PR D57 5903	D.W. Bliss <i>et al.</i>	(CLEO Collab.)	REFID=45943
ABE	97O	PRL 78 4691	K. Abe <i>et al.</i>	(SLD Collab.)	REFID=45461
ACKERSTAFF	97J	PL B404 213	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)	REFID=45455
ACKERSTAFF	97L	ZPHY C74 403	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)	REFID=45486
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ALEXANDER	97F	PR D56 5320	J.P. Alexander <i>et al.</i>	(CLEO Collab.)	REFID=45677
AMMAR	97B	PRL 78 4686	R. Ammar <i>et al.</i>	(CLEO Collab.)	REFID=45460
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BARATE	97I	ZPHY C74 387	R. Barate <i>et al.</i>	(ALEPH Collab.)	REFID=45484
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BERGFELD	97	PRL 79 2406	T. Bergfeld <i>et al.</i>	(CLEO Collab.)	REFID=45695
BONVICINI	97	PRL 79 1221	G. Bonvicini <i>et al.</i>	(CLEO Collab.)	REFID=45591
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ALBRECHT	96E	PRPL 276 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=44953
ALEXANDER	96D	PL B369 163	G. Alexander <i>et al.</i>	(OPAL Collab.)	REFID=44663
ALEXANDER	96E	PL B374 341	G. Alexander <i>et al.</i>	(OPAL Collab.)	REFID=44666
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BARTELT	96	PRL 76 4119	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)	REFID=44665
BUSKULIC	96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=44588
BUSKULIC	96C	ZPHY C70 561	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=44658
COAN	96	PR D53 6037	T.E. Coan <i>et al.</i>	(CLEO Collab.)	REFID=44667
ABE	95Y	PR D52 4828	K. Abe <i>et al.</i>	(SLD Collab.)	REFID=44479
ABREU	95T	PL B357 715	P. Abreu <i>et al.</i>	(DELPHI Collab.)	REFID=44470
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ACCIARRI	95	PL B345 93	M. Acciarri <i>et al.</i>	(L3 Collab.)	REFID=44125
ACCIARRI	95F	PL B352 487	M. Acciarri <i>et al.</i>	(L3 Collab.)	REFID=44280
AKERS	95F	ZPHY C66 31	R. Akers <i>et al.</i>	(OPAL Collab.)	REFID=44183
AKERS	95I	ZPHY C66 543	R. Akers <i>et al.</i>	(OPAL Collab.)	REFID=44288
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)	REFID=44366
AKERS	95Y	ZPHY C68 555	R. Akers <i>et al.</i>	(OPAL Collab.)	REFID=44539
ALBRECHT	95	PL B341 441	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=44114
ALBRECHT	95C	PL B349 576	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=44267
ALBRECHT	95G	ZPHY C68 25	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=44461
ALBRECHT	95H	ZPHY C68 215	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=44462
BALEST	95C	PRL 75 3809	R. Balest <i>et al.</i>	(CLEO Collab.)	REFID=44557
BERNABEU	95	NP B436 474	J. Bernabeu <i>et al.</i>		REFID=48066
BUSKULIC	95C	PL B346 371	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=44206
BUSKULIC	95D	PL B346 379	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=44207
		Also PL B363 265 (errat.)	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=44559
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AKERS	94E	PL B328 207	R. Akers <i>et al.</i>	(OPAL Collab.)	REFID=43802
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BATTLE	94	PRL 73 1079	M. Battle <i>et al.</i>	(CLEO Collab.)	REFID=43918
BAUER	94	PR D50 13	D.A. Bauer <i>et al.</i>	(TPC/2gamma Collab.)	REFID=43861
BUSKULIC	94D	PL B321 168	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=43885
BUSKULIC	94E	PL B332 209	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=43898
BUSKULIC	94F	PL B332 219	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=43899
GIBAUT	94B	PRL 73 934	D. Gibaut <i>et al.</i>	(CLEO Collab.)	REFID=43917
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)	REFID=43644
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=43310
ALBRECHT	93G	PL B316 608	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=43627
BALEST	93	PR D47 3671	R. Balest <i>et al.</i>	(CLEO Collab.)	REFID=43373
BEAN	93	PRL 70 138	A. Bean <i>et al.</i>	(CLEO Collab.)	REFID=43203
BORTOLETTO	93	PRL 71 1791	D. Bortolotto <i>et al.</i>	(CLEO Collab.)	REFID=43527
ESCRIBANO	93	PL B301 419	R. Escrivano, E. Masso	(BARC)	REFID=43289
PROCARIO	93	PRL 70 1207	M. Procario <i>et al.</i>	(CLEO Collab.)	REFID=43211
ABREU	92N	ZPHY C55 555	P. Abreu <i>et al.</i>	(DELPHI Collab.)	REFID=42204
ACTON	92F	PL B281 405	D.P. Acton <i>et al.</i>	(OPAL Collab.)	REFID=42082
ACTON	92H	PL B288 373	P.D. Acton <i>et al.</i>	(OPAL Collab.)	REFID=42152
AKERIB	92	PRL 69 3610	D.S. Akerib <i>et al.</i>	(CLEO Collab.)	REFID=43125
		Also PRL 71 3395 (errat.)	D.S. Akerib <i>et al.</i>	(CLEO Collab.)	REFID=43574
ALBRECHT	92D	ZPHY C53 367	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=41955
ALBRECHT	92K	ZPHY C55 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=42198
ALBRECHT	92M	PL B292 221	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=42211
ALBRECHT	92Q	ZPHY C56 339	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=43150
AMMAR	92	PR D45 3976	R. Ammar <i>et al.</i>	(CLEO Collab.)	REFID=41904
ARTUSO	92	PRL 69 3278	M. Artuso <i>et al.</i>	(CLEO Collab.)	REFID=43120
BAI	92	PRL 69 3021	J.Z. Bai <i>et al.</i>	(BES Collab.)	REFID=43117
BATTLE	92	PL B291 488	M. Battle <i>et al.</i>	(CLEO Collab.)	REFID=42208
BUSKULIC	92J	PL B297 459	D. Buskulic <i>et al.</i>	(ALEPH Collab.)	REFID=43222

DECAMP	92C	ZPHY C54 211	D. Decamp <i>et al.</i>	(ALEPH Collab.)	REFID=41902
ADEVA	91F	PL B265 451	B. Adeva <i>et al.</i>	(L3 Collab.)	REFID=41610
ALBRECHT	91D	PL B260 259	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=41529
ALEXANDER	91D	PL B266 201	G. Alexander <i>et al.</i>	(OPAL Collab.)	REFID=41613
ANTREASYAN	91	PL B259 216	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)	REFID=41486
GRIFOLS	91	PL B255 611	J.A. Grifols, A. Mendez	(BARC)	REFID=41442
ABACHI	90	PR D41 1414	S. Abachi <i>et al.</i>	(HRS Collab.)	REFID=41243
ALBRECHT	90E	PL B246 278	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=41335
ALBRECHT	90I	PL B250 164	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=41414
BEHREND	90	ZPHY C46 537	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	REFID=40983
BOWCOCK	90	PR D41 805	T.J.V. Bowcock <i>et al.</i>	(CLEO Collab.)	REFID=41239
DELAGUILA	90	PL B252 116	F. del Aguila, M. Sher	(BARC, WILL)	REFID=41420
GOLDBERG	90	PL B251 223	M. Goldberg <i>et al.</i>	(CLEO Collab.)	REFID=41457
WU	90	PR D41 2339	D.Y. Wu <i>et al.</i>	(Mark II Collab.)	REFID=41246
ABACHI	89B	PR D40 902	S. Abachi <i>et al.</i>	(HRS Collab.)	REFID=40847
BEHREND	89B	PL B222 163	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	REFID=40817
JANSSEN	89	PL B228 273	H. Janssen <i>et al.</i>	(Crystal Ball Collab.)	REFID=40827
KLEINWORT	89	ZPHY C42 7	C. Kleinwort <i>et al.</i>	(JADE Collab.)	REFID=40854
ADEVA	88	PR D38 2665	B. Adeva <i>et al.</i>	(Mark-J Collab.)	REFID=40686
ALBRECHT	88B	PL B202 149	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=40506
ALBRECHT	88L	ZPHY C41 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=40861
ALBRECHT	88M	ZPHY C41 405	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=40862
AMIDEI	88	PR D37 1750	D. Amidei <i>et al.</i>	(Mark II Collab.)	REFID=40433
BEHREND	88	PL B200 226	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	REFID=40441
BRAUNSCH...	88C	ZPHY C39 331	W. Braunschweig <i>et al.</i>	(TASSO Collab.)	REFID=40696
KEH	88	PL B212 123	S. Keh <i>et al.</i>	(Crystal Ball Collab.)	REFID=40655
TSCHIRHART	88	PL B205 407	R. Tschirhart <i>et al.</i>	(HRS Collab.)	REFID=40640
ABACHI	87B	PL B197 291	S. Abachi <i>et al.</i>	(HRS Collab.)	REFID=40432
ABACHI	87C	PRL 59 2519	S. Abachi <i>et al.</i>	(HRS Collab.)	REFID=40439
ADLER	87B	PRL 59 1527	J. Adler <i>et al.</i>	(Mark III Collab.)	REFID=40431
AIHARA	87B	PR D35 1553	H. Aihara <i>et al.</i>	(TPC Collab.)	REFID=40423
AIHARA	87C	PRL 59 751	H. Aihara <i>et al.</i>	(TPC Collab.)	REFID=40428
ALBRECHT	87L	PL B185 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=40418
ALBRECHT	87P	PL B199 580	H. Albrecht <i>et al.</i>	(ARGUS Collab.)	REFID=40440
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)	REFID=40263
BAND	87B	PRL 59 415	H.R. Band <i>et al.</i>	(MAC Collab.)	REFID=40427
BARINGER	87	PR D35 1993	P. Baringer <i>et al.</i>	(CLEO Collab.)	REFID=40436
BEBEK	87C	PR D36 690	C. Bebek <i>et al.</i>	(CLEO Collab.)	REFID=40429
BURCHAT	87	PR D35 27	P.R. Burchat <i>et al.</i>	(Mark II Collab.)	REFID=40422
BYLSMA	87	PR D35 2269	B.G. Bylsma <i>et al.</i>	(HRS Collab.)	REFID=40424
COFFMAN	87	PR D36 2185	D.M. Coffman <i>et al.</i>	(Mark III Collab.)	REFID=40430
DERRICK	87	PL B189 260	M. Derrick <i>et al.</i>	(HRS Collab.)	REFID=40421
FORD	87	PR D35 408	W.T. Ford <i>et al.</i>	(MAC Collab.)	REFID=40434
FORD	87B	PR D36 1971	W.T. Ford <i>et al.</i>	(MAC Collab.)	REFID=40438
GAN	87	PRL 59 411	K.K. Gan <i>et al.</i>	(Mark II Collab.)	REFID=40426
GAN	87B	PL B197 561	K.K. Gan <i>et al.</i>	(Mark II Collab.)	REFID=40437
AIHARA	86E	PRL 57 1836	H. Aihara <i>et al.</i>	(TPC Collab.)	REFID=40420
BARTEL	86D	PL B182 216	W. Bartel <i>et al.</i>	(JADE Collab.)	REFID=40417
PDG	86	PL 170B 1	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)	REFID=40122
RUCKSTUHL	86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)	REFID=10349
SCHMIDKE	86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)	REFID=10350
YELTON	86	PRL 56 812	J.M. Yelton <i>et al.</i>	(Mark II Collab.)	REFID=10351
ALTHOFF	85	ZPHY C26 521	M. Althoff <i>et al.</i>	(TASSO Collab.)	REFID=10339
ASH	85B	PR D55 2118	W.W. Ash <i>et al.</i>	(MAC Collab.)	REFID=10340
BALTRUSAIT...	85	PRL 55 1842	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)	REFID=10341
BARTEL	85F	PL 161B 188	W. Bartel <i>et al.</i>	(JADE Collab.)	REFID=10342
BEHREND	85	PR D32 2468	S. Behrends <i>et al.</i>	(CLEO Collab.)	REFID=10343
BELTRAMI	85	PRL 54 1775	I. Beltrami <i>et al.</i>	(HRS Collab.)	REFID=10344
BERGER	85	ZPHY C28 1	C. Berger <i>et al.</i>	(PLUTO Collab.)	REFID=10345
BURCHAT	85	PRL 54 2489	P.R. Burchat <i>et al.</i>	(Mark II Collab.)	REFID=10290
FERNANDEZ	85	PRL 54 1624	E. Fernandez <i>et al.</i>	(MAC Collab.)	REFID=10347
MILLS	85	PRL 54 624	G.B. Mills <i>et al.</i>	(DELCO Collab.)	REFID=10292
AIHARA	84C	PR D30 2436	H. Aihara <i>et al.</i>	(TPC Collab.)	REFID=10334
BEHREND	84	ZPHY C23 103	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	REFID=10336
MILLS	84	PRL 52 1944	G.B. Mills <i>et al.</i>	(DELCO Collab.)	REFID=10337
BEHREND	83C	PL 127B 270	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	REFID=10331
SILVERMAN	83	PR D27 1196	D.J. Silverman, G.L. Shaw	(UCI)	REFID=43784
BEHREND	82	PL 114B 282	H.J. Behrend <i>et al.</i>	(CELLO Collab.)	REFID=10324
BLOCKER	82B	PRL 48 1586	C.A. Blocker <i>et al.</i>	(Mark II Collab.)	REFID=10326
BLOCKER	82D	PL 109B 119	C.A. Blocker <i>et al.</i>	(Mark II Collab.)	REFID=10325
FELDMAN	82	PRL 48 66	G.J. Feldman <i>et al.</i>	(Mark II Collab.)	REFID=10328
HAYES	82	PR D25 2869	K.G. Hayes <i>et al.</i>	(Mark II Collab.)	REFID=10330
BERGER	81B	PL 99B 489	C. Berger <i>et al.</i>	(PLUTO Collab.)	REFID=10322
DORFAN	81	PRL 46 215	J.M. Dorfan <i>et al.</i>	(Mark II Collab.)	REFID=10323
BRANDELIK	80	PL 92B 199	R. Brandelik <i>et al.</i>	(TASSO Collab.)	REFID=10318
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)	REFID=10320
Also		SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)	REFID=10321
BACINO	79B	PRL 42 749	Translated from YAF 34 1471.	(DELCO Collab.)	REFID=10316
KIRKBY	79	SLAC-PUB-2419	W.J. Bacino <i>et al.</i>	(SLAC)	REFID=10283
Batavia Lepton Photon Conference.			J. Kirkby	(DELCO Collab.)	REFID=10316
BACINO	78B	PRL 41 13	W.J. Bacino <i>et al.</i>	(SLAC)	REFID=10283
Also		Tokyo Conf. 249	J. Kirz	(STON)	REFID=10304
Also		PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)	REFID=10305
BRANDELIK	78	PL 73B 109	R. Brandelik <i>et al.</i>	(DASP Collab.)	REFID=10320
FELDMAN	78	Tokyo Conf. 777	G.J. Feldman	(SLAC)	REFID=10280
JAROS	78	PRL 40 1120	J. Jaros <i>et al.</i>	(LGW Collab.)	REFID=10355
PERL	75	PRL 35 1489	M.L. Perl <i>et al.</i>	(LBL, SLAC)	REFID=10310

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PERL	80	ARNPS 30 299	M.L. Perl	(SLAC)	REFID=10359