Extra Dimensions

NODE=S071

NODE=S071

For explanation of terms used and discussion of significant model dependence of following limits, see the "Extra Dimensions" review. Footnotes describe originally quoted limit. δ indicates the number of extra dimensions.

Limits not encoded here are summarized in the "Extra Dimensions" review, where the latest unpublished results are also described.

See the related review(s):

Extra Dimensions

CONTENTS:

NODE=S071CNT

NODE=S071CNT

NODE=S071CNT

Limits on R from Deviations in Gravitational Force Law

NODE=S071DGF

NODE=S071DGF

This section includes limits on the size of extra dimensions from deviations in the Newtonian ($1/r^2$) gravitational force law at short distances. Deviations are parametrized by a gravitational potential of the form $V = -(G\ m\ m'/r)\ [1 + \alpha\ \exp(-r/R)]$. For δ toroidal extra dimensions of equal size, $\alpha = 8\delta/3$. Quoted bounds are for $\delta = 2$ unless otherwise noted.

NODE=S071DGF

VALUE (μm)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
		[1] BLAKEMORE 21		Optical levitation
		[2] HEACOCK 21		Neutron scattering
		[3] LEE 20		Torsion pendulum
< 37	95	[4] TAN 20A		Torsion pendulum
		[5] BERGE 18	MICR	Space accelerometer
		[6] FAYET 18A	MICR	Space accelerometer
		[7] KLIMCHITSK... 17A		Torsion oscillator
		[8] XU 13		Nuclei properties
		[9] BEZERRA 11		Torsion oscillator
		[10] SUSHKOV 11		Torsion pendulum
		[11] BEZERRA 10		Microcantilever
		[12] MASUDA 09		Torsion pendulum
		[13] GERACI 08		Microcantilever
		[14] TRENKEL 08		Newton's constant
		[15] DECCA 07A		Torsion oscillator
< 37	95	[16] KAPNER 07		Torsion pendulum
< 47	95	[17] TU 07		Torsion pendulum
		[18] SMULLIN 05		Microcantilever
<130	95	[19] HOYLE 04		Torsion pendulum
		[20] CHIAVERINI 03		Microcantilever
$\lesssim 200$	95	[21] LONG 03		Microcantilever
<190	95	[22] HOYLE 01		Torsion pendulum
		[23] HOSKINS 85		Torsion pendulum

OCCUR=2

[1] BLAKEMORE 21 obtain constraints on non-Newtonian forces with strengths $|\alpha| \gtrsim 10^8$ and length scales $R > 10\ \mu$m. See their Fig. 4 for more details including comparison with previous searches. NODE=S071DGF;LINKAGE=K

[2] HEACOCK 21 obtain constraints on non-Newtonian forces with strengths $10^{18} \lesssim |\alpha| \lesssim 10^{25}$ and length scales $R \simeq$ 0.02–10 nm. See their Figure 3 for more details. This improves the results of HADDOCK 18. These constraints do not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=M

[3] LEE 20 search for new forces probing a range of $|\alpha| \simeq 0.1$–10^5 and length scales $R \simeq$ 7–90 μm. For $\delta = 1$ the bound on R is 30 μm. See their Fig. 5 for details on the bound. NODE=S071DGF;LINKAGE=J

[4] TAN 20A search for new forces probing a range of $|\alpha| \simeq 4 \times 10^{-3}$–$1 \times 10^2$ and length scales $R \simeq$ 40–350 μm. See their Fig. 6 for details on the bound. NODE=S071DGF;LINKAGE=I

[5] BERGE 18 uses results from the MICROSCOPE experiment to obtain constraints on non-Newtonian forces with strengths $10^{-11} \lesssim |\alpha| \lesssim 10^{-7}$ and length scales $R \gtrsim 10^{5}$ m. See their Figure 1 for more details. These constraints do not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=E

[6] FAYET 18A uses results from the MICROSCOPE experiment to obtain constraints on an EP-violating force possibly arising from a new U(1) gauge boson. For $R \gtrsim 10^{7}$ m the limits are $|\alpha| \lesssim$ a few 10^{-13} to a few 10^{-11} depending on the coupling, corresponding to $|\epsilon| \lesssim 10^{-24}$ for the coupling of the new spin-1 or spin-0 mediator. These constraints do not place limits on the size of extra flat dimensions. This extends the results of FAYET 18. NODE=S071DGF;LINKAGE=H

[7] KLIMCHITSKAYA 17A uses an experiment that measures the difference of Casimir forces to obtain bounds on non-Newtonian forces with strengths $|\alpha| \simeq 10^{5}$–10^{17} and length scales R = 0.03–10 μm. See their Fig. 3. These constraints do not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=B

[8] XU 13 obtain constraints on non-Newtonian forces with strengths $|\alpha| \simeq 10^{34}$–10^{36} and length scales $R \simeq$ 1–10 fm. See their Fig. 4 for more details. These constraints do not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=A

[9] BEZERRA 11 obtain constraints on non-Newtonian forces with strengths $10^{11} \lesssim |\alpha| \lesssim 10^{18}$ and length scales R = 30–1260 nm. See their Fig. 2 for more details. These constraints do not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=BZ

[10] SUSHKOV 11 obtain improved limits on non-Newtonian forces with strengths $10^{7} \lesssim |\alpha| \lesssim 10^{11}$ and length scales 0.4 μm < R < 4 μm (95% CL). See their Fig. 2. These bounds do not place limits on the size of extra flat dimensions. However, a model dependent bound of M_{*} > 70 TeV is obtained assuming gauge bosons that couple to baryon number also propagate in $(4 + \delta)$ dimensions. NODE=S071DGF;LINKAGE=SU

[11] BEZERRA 10 obtain improved constraints on non-Newtonian forces with strengths $10^{19} \lesssim |\alpha| \lesssim 10^{29}$ and length scales R = 1.6–14 nm (95% CL). See their Fig. 1. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=BE

[12] MASUDA 09 obtain improved constraints on non-Newtonian forces with strengths $10^{9} \lesssim |\alpha| \lesssim 10^{11}$ and length scales R = 1.0–2.9 μm (95% CL). See their Fig. 3. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=MA

[13] GERACI 08 obtain improved constraints on non-Newtonian forces with strengths $|\alpha| >$ 14,000 and length scales R = 5–15 μm. See their Fig. 9. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=GE

[14] TRENKEL 08 uses two independent measurements of Newton's constant G to constrain new forces with strength $|\alpha| \simeq 10^{-4}$ and length scales R = 0.02–1 m. See their Fig. 1. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=TR

[15] DECCA 07A search for new forces and obtain bounds in the region with strengths $|\alpha| \simeq 10^{13}$–10^{18} and length scales R = 20–86 nm. See their Fig. 6. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=DE

[16] KAPNER 07 search for new forces, probing a range of $|\alpha| \simeq 10^{-3}$–10^{5} and length scales $R \simeq$ 10–1000 μm. For $\delta = 1$ the bound on R is 44 μm. For $\delta = 2$, the bound is expressed in terms of M_{*}, here translated to a bound on the radius. See their Fig. 6 for details on the bound. NODE=S071DGF;LINKAGE=KA

[17] TU 07 search for new forces probing a range of $|\alpha| \simeq 10^{-1}$–10^{5} and length scales $R \simeq$ 20–1000 μm. For $\delta = 1$ the bound on R is 53 μm. See their Fig. 3 for details on the bound. NODE=S071DGF;LINKAGE=TU

[18] SMULLIN 05 search for new forces, and obtain bounds in the region with strengths $\alpha \simeq 10^{3}$–10^{8} and length scales R = 6–20 μm. See their Figs. 1 and 16 for details on the bound. This work does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=SM

[19] HOYLE 04 search for new forces, probing α down to 10^{-2} and distances down to 10μm. Quoted bound on R is for $\delta = 2$. For $\delta = 1$, bound goes to 160 μm. See their Fig. 34 for details on the bound. NODE=S071DGF;LINKAGE=HO

[20] CHIAVERINI 03 search for new forces, probing α above 10^{4} and λ down to 3μm, finding no signal. See their Fig. 4 for details on the bound. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=C

[21] LONG 03 search for new forces, probing α down to 3, and distances down to about 10μm. See their Fig. 4 for details on the bound. NODE=S071DGF;LINKAGE=L

[22] HOYLE 01 search for new forces, probing α down to 10^{-2} and distances down to 20μm. See their Fig. 4 for details on the bound. The quoted bound is for $\alpha \geq 3$. NODE=S071DGF;LINKAGE=HL

[23] HOSKINS 85 search for new forces, probing distances down to 4 mm. See their Fig. 13 for details on the bound. This bound does not place limits on the size of extra flat dimensions. NODE=S071DGF;LINKAGE=HK

Limits on R from On-Shell Production of Gravitons: $\delta = 2$

NODE=S071OS6
NODE=S071OS6

This section includes limits on on-shell production of gravitons in collider and astrophysical processes. Bounds quoted are on R, the assumed common radius of the flat extra dimensions, for $\delta = 2$ extra dimensions. Studies often quote bounds in terms of derived parameter; experiments are actually sensitive to the masses of the KK gravitons: $m_{\vec{n}} = |\vec{n}|/R$. See the Review on "Extra Dimensions" for details. Bounds are given in μm for $\delta = 2$.

VALUE (μm)	*CL%*	*DOCUMENT ID*		*TECN*	*COMMENT*	
< **3.8**	95	[1] AAD	21F	ATLS	$pp \rightarrow jG$	NODE=S071OS6;CHECK LIMITS
< **0.00016**	95	[2] HANNESTAD	03		Neutron star heating	OCCUR=4

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<$ 56	95	[3] SIRUNYAN	21A	CMS	$pp \to ZG$	
$<$ 4.1	95	[4] TUMASYAN	21D	CMS	$pp \to jG$	
		[5] SIRUNYAN	17AQ	CMS	$pp \to \gamma G$	
$<$ 90	95	[6] AABOUD	16F	ATLS	$pp \to \gamma G$	
		[7] KHACHATRY...	16N	CMS	$pp \to \gamma G$	
		[8] AAD	15CS	ATLS	$pp \to \gamma G$	
$<$ 127	95	[9] AAD	13C	ATLS	$pp \to \gamma G$	
$<$ 34.4	95	[10] AAD	13D	ATLS	$pp \to jj$	
$<$ 0.0087	95	[11] AJELLO	12	FLAT	Neutron star γ sources	
$<$ 245	95	[12] AALTONEN	08AC	CDF	$p\overline{p} \to \gamma G$, jG	
$<$ 615	95	[13] ABAZOV	08S	D0	$p\overline{p} \to \gamma G$	
$<$ 0.916	95	[14] DAS	08		Supernova cooling	
$<$ 350	95	[15] ABULENCIA,A	06	CDF	$p\overline{p} \to jG$	
$<$ 270	95	[16] ABDALLAH	05B	DLPH	$e^+e^- \to \gamma G$	
$<$ 210	95	[17] ACHARD	04E	L3	$e^+e^- \to \gamma G$	
$<$ 480	95	[18] ACOSTA	04C	CDF	$\overline{p}p \to jG$	
$<$ 0.00038	95	[19] CASSE	04		Neutron star γ sources	
$<$ 610	95	[20] ABAZOV	03	D0	$\overline{p}p \to jG$	
$<$ 0.96	95	[21] HANNESTAD	03		Supernova cooling	
$<$ 0.096	95	[22] HANNESTAD	03		Diffuse γ background	OCCUR=2
$<$ 0.051	95	[23] HANNESTAD	03		Neutron star γ sources	OCCUR=3
$<$ 300	95	[24] HEISTER	03C	ALEP	$e^+e^- \to \gamma G$	
		[25] FAIRBAIRN	01		Cosmology	
$<$ 0.66	95	[26] HANHART	01		Supernova cooling	
		[27] CASSISI	00		Red giants	
$<$1300	95	[28] ACCIARRI	99S	L3	$e^+e^- \to ZG$	

[1] AAD 21F search for $pp \to jG$, using 139 fb^{-1} of data at $\sqrt{s} = 13$ TeV to place lower limits on M_D for two to six extra dimensions (see their Table X), from which this bound on R is derived. This limit supersedes that in AABOUD 18I. NODE=S071OS6;LINKAGE=M

[2] HANNESTAD 03 obtain a limit on R from the heating of old neutron stars by the surrounding cloud of trapped KK gravitons. Limits for all $\delta \le 7$ are given in their Tables V and VI. These limits supersede those in HANNESTAD 02. NODE=S071OS;LINKAGE=HD

[3] SIRUNYAN 21A search for $pp \to ZG$, using 137 fb^{-1} of data at $\sqrt{s} = 13$ TeV to place lower limits on M_D for two to seven extra dimensions (see their Figure 12), from which this bound on R is derived. These limits supersede those obtained in SIRUNYAN 18BV. NODE=S071OS6;LINKAGE=L

[4] TUMASYAN 21D search for $pp \to jG$, using 137 fb^{-1} of data at $\sqrt{s} = 13$ TeV to place lower limits on M_D for two to seven extra dimensions (see their Table 3), from which this bound on R is derived. This limit supersedes that in SIRUNYAN 18S. NODE=S071OS6;LINKAGE=N

[5] SIRUNYAN 17AQ search for $pp \to \gamma G$, using 12.9 fb^{-1} of data at $\sqrt{s} = 13$ TeV to place limits on M_D for three to six extra dimensions (see their Table 3). NODE=S071OS6;LINKAGE=H

[6] AABOUD 16F search for $pp \to \gamma G$, using 3.2 fb^{-1} of data at $\sqrt{s} = 13$ TeV to place limits on M_D for two to six extra dimensions (see their Figure 9), from which this bound on R is derived. NODE=S071OS6;LINKAGE=E

[7] KHACHATRYAN 16N search for $pp \to \gamma G$, using 19.6 fb^{-1} of data at $\sqrt{s} = 8$ TeV to place limits on M_D for three to six extra dimensions (see their Table 5). NODE=S071OS6;LINKAGE=G

[8] AAD 15CS search for $pp \to \gamma G$, using 20.3 fb^{-1} of data at $\sqrt{s} = 8$ TeV to place lower limits on M_D for two to six extra dimensions (see their Fig. 18). NODE=S071OS6;LINKAGE=C

[9] AAD 13C search for $pp \to \gamma G$, using 4.6 fb^{-1} of data at $\sqrt{s} = 7$ TeV to place bounds on M_D for two to six extra dimensions, from which this bound on R is derived. NODE=S071OS6;LINKAGE=GA

[10] AAD 13D search for the dijet decay of quantum black holes in 4.8 fb^{-1} of data produced in pp collisions at $\sqrt{s} = 7$ TeV to place bounds on M_D for two to seven extra dimensions, from which these bounds on R are derived. Limits on M_D for all $\delta \le 7$ are given in their Table 3. NODE=S071OS6;LINKAGE=TA

[11] AJELLO 12 obtain a limit on R from the gamma-ray emission of point γ sources that arise from the photon decay of KK gravitons which are gravitationally bound around neutron stars. Limits for all $\delta \le 7$ are given in their Table 7. NODE=S071OS6;LINKAGE=AJ

[12] AALTONEN 08AC search for $p\overline{p} \to \gamma G$ and $p\overline{p} \to jG$ at $\sqrt{s} = 1.96$ TeV with 2.0 fb^{-1} and 1.1 fb^{-1} respectively, in order to place bounds on the fundamental scale and size of the extra dimensions. See their Table III for limits on all $\delta \le 6$. NODE=S071OS;LINKAGE=LO

[13] ABAZOV 08S search for $p\overline{p} \to \gamma G$, using 1 fb^{-1} of data at $\sqrt{s} = 1.96$ TeV to place bounds on M_D for two to eight extra dimensions, from which these bounds on R are derived. See their paper for intermediate values of δ. NODE=S071OS;LINKAGE=BA

[14] DAS 08 obtain a limit on R from Kaluza-Klein graviton cooling of SN1987A due to plasmon-plasmon annihilation. NODE=S071OS;LINKAGE=DA

[15] ABULENCIA,A 06 search for $p\overline{p} \to jG$ using 368 pb^{-1} of data at $\sqrt{s} = 1.96$ TeV. See their Table II for bounds for all $\delta \le 6$. NODE=S071OS6;LINKAGE=LE

[16] ABDALLAH 05B search for $e^+e^- \to \gamma G$ at $\sqrt{s} = 180$–209 GeV to place bounds on the size of extra dimensions and the fundamental scale. Limits for all $\delta \le 6$ are given in their Table 6. These limits supersede those in ABREU 00Z. NODE=S071OS;LINKAGE=AB

[17] ACHARD 04E search for $e^+ e^- \to \gamma G$ at $\sqrt{s}$ = 189–209 GeV to place bounds on the size of extra dimensions and the fundamental scale. See their Table 8 for limits with $\delta \leq 8$. These limits supersede those in ACCIARRI 99R. NODE=S071OS;LINKAGE=AR

[18] ACOSTA 04C search for $\overline{p}p \to jG$ at $\sqrt{s}$ = 1.8 TeV to place bounds on the size of extra dimensions and the fundamental scale. See their paper for bounds on δ = 4, 6. NODE=S071OS;LINKAGE=AC

[19] CASSE 04 obtain a limit on R from the gamma-ray emission of point γ sources that arises from the photon decay of gravitons around newly born neutron stars, applying the technique of HANNESTAD 03 to neutron stars in the galactic bulge. Limits for all $\delta \leq 7$ are given in their Table I. NODE=S071OS;LINKAGE=CA

[20] ABAZOV 03 search for $p\overline{p} \to jG$ at $\sqrt{s}$=1.8 TeV to place bounds on M_D for 2 to 7 extra dimensions, from which these bounds on R are derived. See their paper for bounds on intermediate values of δ. We quote results without the approximate NLO scaling introduced in the paper. NODE=S071OS;LINKAGE=ZB

[21] HANNESTAD 03 obtain a limit on R from graviton cooling of supernova SN1987a. Limits for all $\delta \leq 7$ are given in their Tables V and VI. NODE=S071OS;LINKAGE=HA

[22] HANNESTAD 03 obtain a limit on R from gravitons emitted in supernovae and which subsequently decay, contaminating the diffuse cosmic γ background. Limits for all $\delta \leq 7$ are given in their Tables V and VI. These limits supersede those in HANNESTAD 02. NODE=S071OS;LINKAGE=HB

[23] HANNESTAD 03 obtain a limit on R from gravitons emitted in two recent supernovae and which subsequently decay, creating point γ sources. Limits for all $\delta \leq 7$ are given in their Tables V and VI. These limits are corrected in the published erratum. NODE=S071OS;LINKAGE=HC

[24] HEISTER 03C use the process $e^+ e^- \to \gamma G$ at $\sqrt{s}$ = 189–209 GeV to place bounds on the size of extra dimensions and the scale of gravity. See their Table 4 for limits with $\delta \leq 6$ for derived limits on M_D. NODE=S071OS;LINKAGE=3H

[25] FAIRBAIRN 01 obtains bounds on R from over production of KK gravitons in the early universe. Bounds are quoted in paper in terms of fundamental scale of gravity. Bounds depend strongly on temperature of QCD phase transition and range from $R<$ 0.13 μm to 0.001 μm for δ=2; bounds for δ=3,4 can be derived from Table 1 in the paper. NODE=S071OS;LINKAGE=F

[26] HANHART 01 obtain bounds on R from limits on graviton cooling of supernova SN 1987a using numerical simulations of proto-neutron star neutrino emission. NODE=S071OS;LINKAGE=HT

[27] CASSISI 00 obtain rough bounds on M_D (and thus R) from red giant cooling for δ=2,3. See their paper for details. NODE=S071OS;LINKAGE=CS

[28] ACCIARRI 99S search for $e^+ e^- \to ZG$ at $\sqrt{s}$=189 GeV. Limits on the gravity scale are found in their Table 2, for $\delta \leq 4$. NODE=S071OS;LINKAGE=S9

Mass Limits on M_{TT}

NODE=S071GEX

This section includes limits on the cut-off mass scale, M_{TT}, of dimension-8 operators from KK graviton exchange in models of large extra dimensions. Ambiguities in the UV-divergent summation are absorbed into the parameter λ, which is taken to be $\lambda = \pm1$ in the following analyses. Bounds for $\lambda = -1$ are shown in parenthesis after the bound for $\lambda = +1$, if appropriate. Different papers use slightly different definitions of the mass scale. The definition used here is related to another popular convention by $M^4_{TT} = (2/\pi)\, \Lambda^4_T$, as discussed in the above Review on "Extra Dimensions." NODE=S071GEX

NODE=S071GEX

VALUE (TeV)		*CL%*	*DOCUMENT ID*		*TECN*	*COMMENT*
> 9.02		95	[1] SIRUNYAN	18DD	CMS	$pp \to$ dijet, ang. distrib.
>20.6	**(> 15.7)**	95	[2] GIUDICE	03	RVUE	Dim-6 operators
• • • We do not use the following data for averages, fits, limits, etc. • • •						
> 6.7		95	[3] SIRUNYAN	21N	CMS	$pp \to e^+e^-, \mu^+\mu^-$
> 6.9		95	[4] SIRUNYAN	19AC	CMS	$pp \to e^+e^-, \mu^+\mu^-, \gamma\gamma$
> 7.0	(>5.6)	95	[5] SIRUNYAN	18DU	CMS	$pp \to \gamma\gamma$
> 6.5		95	[6] AABOUD	17AP	ATLS	$pp \to \gamma\gamma$
> 3.8		95	[7] AAD	14BE	ATLS	$pp \to e^+e^-, \mu^+\mu^-$
> 3.2		95	[8] AAD	13E	ATLS	$pp \to e^+e^-, \mu^+\mu^-, \gamma\gamma$
			[9] BAAK	12	RVUE	Electroweak
> 0.90	(>0.92)	95	[10] AARON	11C	H1	$e^\pm p \to e^\pm X$
> 1.48		95	[11] ABAZOV	09AE	D0	$p\overline{p} \to$ dijet, ang. distrib.
> 1.45		95	[12] ABAZOV	09D	D0	$p\overline{p} \to e^+e^-, \gamma\gamma$
> 1.1	(> 1.0)	95	[13] SCHAEL	07A	ALEP	$e^+e^- \to e^+e^-$
> 0.898	(> 0.998)	95	[14] ABDALLAH	06C	DLPH	$e^+e^- \to \ell^+\ell^-$
> 0.853	(> 0.939)	95	[15] GERDES	06		$p\overline{p} \to e^+e^-, \gamma\gamma$
> 0.96	(> 0.93)	95	[16] ABAZOV	05V	D0	$p\overline{p} \to \mu^+\mu^-$
> 0.78	(> 0.79)	95	[17] CHEKANOV	04B	ZEUS	$e^\pm p \to e^\pm X$
> 0.805	(> 0.956)	95	[18] ABBIENDI	03D	OPAL	$e^+e^- \to \gamma\gamma$
> 0.7	(> 0.7)	95	[19] ACHARD	03D	L3	$e^+e^- \to ZZ$
> 0.82	(> 0.78)	95	[20] ADLOFF	03	H1	$e^\pm p \to e^\pm X$
> 1.28	(> 1.25)	95	[21] GIUDICE	03	RVUE	
> 0.80	(> 0.85)	95	[22] HEISTER	03C	ALEP	$e^+e^- \to \gamma\gamma$
> 0.84	(> 0.99)	95	[23] ACHARD	02D	L3	$e^+e^- \to \gamma\gamma$
> 1.2	(> 1.1)	95	[24] ABBOTT	01	D0	$p\overline{p} \to e^+e^-, \gamma\gamma$
> 0.60	(> 0.63)	95	[25] ABBIENDI	00R	OPAL	$e^+e^- \to \mu^+\mu^-$

OCCUR=2

$>$ 0.63	($>$ 0.50)	95	[25] ABBIENDI	00R	OPAL	$e^+e^- \rightarrow \tau^+\tau^-$	OCCUR=3
$>$ 0.68	($>$ 0.61)	95	[25] ABBIENDI	00R	OPAL	$e^+e^- \rightarrow \mu^+\mu^-, \tau^+\tau^-$	OCCUR=5
			[26] ABREU	00A	DLPH	$e^+e^- \rightarrow \gamma\gamma$	
$>$ 0.680	($>$ 0.542)	95	[27] ABREU	00S	DLPH	$e^+e^- \rightarrow \mu^+\mu^-, \tau^+\tau^-$	OCCUR=6
$>$ 15–28		99.7	[28] CHANG	00B	RVUE	Electroweak	
$>$ 0.98		95	[29] CHEUNG	00	RVUE	$e^+e^- \rightarrow \gamma\gamma$	
$>$ 0.29–0.38		95	[30] GRAESSER	00	RVUE	$(g-2)_\mu$	
$>$ 0.50–1.1		95	[31] HAN	00	RVUE	Electroweak	
$>$ 2.0	($>$ 2.0)	95	[32] MATHEWS	00	RVUE	$\overline{p}p \rightarrow jj$	
$>$ 1.0	($>$ 1.1)	95	[33] MELE	00	RVUE	$e^+e^- \rightarrow VV$	
			[34] ABBIENDI	99P	OPAL		
			[35] ACCIARRI	99M	L3		
			[36] ACCIARRI	99S	L3		
$>$ 1.412	($>$ 1.077)	95	[37] BOURILKOV	99		$e^+e^- \rightarrow e^+e^-$	

[1] SIRUNYAN 18DD use dijet angular distributions in 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to place a lower bound on Λ_T, here converted to M_{TT}. This updates the results of SIRUNYAN 17F. NODE=S071GEX;LINKAGE=K

[2] GIUDICE 03 place bounds on Λ_6, the coefficient of the gravitationally-induced dimension-6 operator $(2\pi\lambda/\Lambda_6^2)(\sum \overline{f}\gamma_\mu\gamma^5 f)(\sum \overline{f}\gamma^\mu\gamma^5 f)$, using data from a variety of experiments. Results are quoted for $\lambda=\pm1$ and are independent of δ. NODE=S071GEX;LINKAGE=GB

[3] SIRUNYAN 21N use 137 (140) fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV in the dielectron (dimuon) channels to place a lower limit on Λ_T, here converted to M_{TT}. Bounds on individual channels can be found in their Table 7. NODE=S071GEX;LINKAGE=N

[4] SIRUNYAN 19AC use 35.9 (36.3) fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV in the dielectron (dimuon) channels to place a lower limit on Λ_T, here converted to M_{TT}. The dielectron and dimuon channels are combined with previous results in the diphoton channel to set the best limit. Bounds on individual channels and different priors can be found in their Table 2. This updates the results in KHACHATRYAN 15AE. NODE=S071GEX;LINKAGE=M

[5] SIRUNYAN 18DU use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to place lower limits on M_{TT} (equivalent to their M_S). This updates the results of CHATRCHYAN 12R. NODE=S071GEX;LINKAGE=L

[6] AABOUD 17AP use 36.7 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to place lower limits on M_{TT} (equivalent to their M_S). This updates the results of AAD 13AS. NODE=S071GEX;LINKAGE=J

[7] AAD 14BE use 20 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV in the dilepton channel to place lower limits on M_{TT} (equivalent to their M_S). NODE=S071GEX;LINKAGE=F

[8] AAD 13E use 4.9 and 5.0 fb^{-1} of data from pp collisions at $\sqrt{s} = 7$ TeV in the dielectron and dimuon channels, respectively, to place lower limits on M_{TT} (equivalent to their M_S). The dielectron and dimuon channels are combined with previous results in the diphoton channel to set the best limit. Bounds on individual channels and different priors can be found in their Table VIII. NODE=S071GEX;LINKAGE=AD

[9] BAAK 12 use electroweak precision observables to place bounds on the ratio Λ_T/M_D as a function of M_D. See their Fig. 22 for constraints with a Higgs mass of 120 GeV. NODE=S071GEX;LINKAGE=BK

[10] AARON 11C search for deviations in the differential cross section of $e^\pm p \rightarrow e^\pm X$ in 446 pb^{-1} of data taken at $\sqrt{s} = 301$ and 319 GeV to place a bound on M_{TT}. NODE=S071GEX;LINKAGE=AA

[11] ABAZOV 09AE use dijet angular distributions in 0.7 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s} = 1.96$ TeV to place lower bounds on Λ_T (equivalent to their M_S), here converted to M_{TT}. NODE=S071GEX;LINKAGE=ZO

[12] ABAZOV 09D use 1.05 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s} = 1.96$ TeV to place lower bounds on Λ_T (equivalent to their M_s), here converted to M_{TT}. NODE=S071GEX;LINKAGE=BA

[13] SCHAEL 07A use e^+e^- collisions at $\sqrt{s} = 189$–209 GeV to place lower limits on Λ_T, here converted to limits on M_{TT}. NODE=S071GEX;LINKAGE=SC

[14] ABDALLAH 06C use e^+e^- collisions at $\sqrt{s} \sim 130$–207 GeV to place lower limits on M_{TT}, which is equivalent to their definition of M_s. Bound shown includes all possible final state leptons, $\ell = e, \mu, \tau$. Bounds on individual leptonic final states can be found in their Table 31. NODE=S071GEX;LINKAGE=BD

[15] GERDES 06 use 100 to 110 pb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s} = 1.8$ TeV, as recorded by the CDF Collaboration during Run I of the Tevatron. Bound shown includes a K-factor of 1.3. Bounds on individual e^+e^- and $\gamma\gamma$ final states are found in their Table I. NODE=S071GEX;LINKAGE=GE

[16] ABAZOV 05V use 246 pb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s} = 1.96$ TeV to search for deviations in the differential cross section to $\mu^+\mu^-$ from graviton exchange. NODE=S071GEX;LINKAGE=AZ

[17] CHEKANOV 04B search for deviations in the differential cross section of $e^\pm p \rightarrow e^\pm X$ with 130 pb^{-1} of combined data and Q^2 values up to 40,000 GeV^2 to place a bound on M_{TT}. NODE=S071GEX;LINKAGE=CE

[18] ABBIENDI 03D use e^+e^- collisions at $\sqrt{s}$=181–209 GeV to place bounds on the ultraviolet scale M_{TT}, which is equivalent to their definition of M_s. NODE=S071GEX;LINKAGE=GD

[19] ACHARD 03D look for deviations in the cross section for $e^+e^- \rightarrow ZZ$ from $\sqrt{s} =$ 200–209 GeV to place a bound on M_{TT}. NODE=S071GEX;LINKAGE=DH

[20] ADLOFF 03 search for deviations in the differential cross section of $e^\pm p \rightarrow e^\pm X$ at $\sqrt{s}$=301 and 319 GeV to place bounds on M_{TT}. NODE=S071GEX;LINKAGE=LF

[21] GIUDICE 03 review existing experimental bounds on M_{TT} and derive a combined limit. NODE=S071GEX;LINKAGE=GA

[22] HEISTER 03C use e^+e^- collisions at $\sqrt{s}$= 189–209 GeV to place bounds on the scale of dim-8 gravitational interactions. Their $M_s^{\pm}$ is equivalent to our M_{TT} with $\lambda=\pm1$. NODE=S071GEX;LINKAGE=3H

[23] ACHARD 02 search for s-channel graviton exchange effects in $e^+e^- \rightarrow \gamma\gamma$ at E_{cm} = 192–209 GeV. NODE=S071GEX;LINKAGE=CD

[24] ABBOTT 01 search for variations in differential cross sections to e^+e^- and $\gamma\gamma$ final states at the Tevatron. NODE=S071GEX;LINKAGE=AB

[25] ABBIENDI 00R uses e^+e^- collisions at $\sqrt{s}$= 189 GeV. NODE=S071GEX;LINKAGE=A

[26] ABREU 00A search for s-channel graviton exchange effects in $e^+e^- \rightarrow \gamma\gamma$ at E_{cm}= 189–202 GeV. NODE=S071GEX;LINKAGE=EA

[27] ABREU 00S uses e^+e^- collisions at $\sqrt{s}$=183 and 189 GeV. Bounds on μ and τ individual final states given in paper. NODE=S071GEX;LINKAGE=S

[28] CHANG 00B derive 3σ limit on M_{TT} of (28,19,15) TeV for δ=(2,4,6) respectively assuming the presence of a torsional coupling in the gravitational action. Highly model dependent. NODE=S071GEX;LINKAGE=CG

[29] CHEUNG 00 obtains limits from anomalous diphoton production at OPAL due to graviton exchange. Original limit for δ=4. However, unknown UV theory renders δ dependence unreliable. Original paper works in HLZ convention. NODE=S071GEX;LINKAGE=CH

[30] GRAESSER 00 obtains a bound from graviton contributions to $g-2$ of the muon through loops of 0.29 TeV for δ=2 and 0.38 TeV for δ=4,6. Limits scale as $\lambda^{1/2}$. However calculational scheme not well-defined without specification of high-scale theory. See the "Extra Dimensions Review." NODE=S071GEX;LINKAGE=GR

[31] HAN 00 calculates corrections to gauge boson self-energies from KK graviton loops and constrain them using S and T. Bounds on M_{TT} range from 0.5 TeV (δ=6) to 1.1 TeV (δ=2); see text. Limits have strong dependence, $\lambda^{\delta+2}$, on unknown λ coefficient. NODE=S071GEX;LINKAGE=HN

[32] MATHEWS 00 search for evidence of graviton exchange in CDF and DØ dijet production data. See their Table 2 for slightly stronger δ-dependent bounds. Limits expressed in terms of $\widetilde{M}_S^4 = M_{TT}^4/8$. NODE=S071GEX;LINKAGE=MT

[33] MELE 00 obtains bound from KK graviton contributions to $e^+e^- \rightarrow VV$ (V=γ,W,Z) at LEP. Authors use Hewett conventions. NODE=S071GEX;LINKAGE=ME

[34] ABBIENDI 99P search for s-channel graviton exchange effects in $e^+e^- \rightarrow \gamma\gamma$ at E_{cm}=189 GeV. The limits $G_+ > 660$ GeV and $G_- > 634$ GeV are obtained from combined E_{cm}=183 and 189 GeV data, where $G_{\pm}$ is a scale related to the fundamental gravity scale. NODE=S071GEX;LINKAGE=AP

[35] ACCIARRI 99M search for the reaction $e^+e^- \rightarrow \gamma G$ and s-channel graviton exchange effects in $e^+e^- \rightarrow \gamma\gamma$, W^+W^-, ZZ, e^+e^-, $\mu^+\mu^-$, $\tau^+\tau^-$, $q\overline{q}$ at E_{cm}=183 GeV. Limits on the gravity scale are listed in their Tables 1 and 2. NODE=S071GEX;LINKAGE=B3

[36] ACCIARRI 99S search for the reaction $e^+e^- \rightarrow ZG$ and s-channel graviton exchange effects in $e^+e^- \rightarrow \gamma\gamma$, W^+W^-, ZZ, e^+e^-, $\mu^+\mu^-$, $\tau^+\tau^-$, $q\overline{q}$ at E_{cm}=189 GeV. Limits on the gravity scale are listed in their Tables 1 and 2. NODE=S071GEX;LINKAGE=DQ

[37] BOURILKOV 99 performs global analysis of LEP data on e^+e^- collisions at $\sqrt{s}$=183 and 189 GeV. Bound is on Λ_T. NODE=S071GEX;LINKAGE=C

Limits on $1/R = M_c$

NODE=S071KK

This section includes limits on $1/R = M_c$, the compactification scale in models with one TeV-sized extra dimension, due to exchange of Standard Model KK excitations. Bounds assume fermions are not in the bulk, unless stated otherwise. See the "Extra Dimensions" review for discussion of model dependence. NODE=S071KK

NODE=S071KK;CHECK LIMITS

VALUE (TeV)	*CL%*	*DOCUMENT ID*		*TECN*	*COMMENT*
>4.16	95	[1] AAD	12CC	ATLS	$pp \rightarrow \ell\overline{\ell}$
>6.1		[2] BARBIERI	04	RVUE	Electroweak
• • • We do not use the following data for averages, fits, limits, etc. • • •					
		[3] FLORES	23	RVUE	minimal universal extra dims
		[4] AVNISH	21	RVUE	$pp \rightarrow$ multijet
		[5] AABOUD	18AV	ATLS	$pp \rightarrow t\overline{t}t\overline{t}$
		[6] AABOUD	18CE	ATLS	$pp \rightarrow t\overline{t}t\overline{t}$
>3.8	95	[7] ACCOMANDO	15	RVUE	Electroweak
>3.40	95	[8] KHACHATRY...	15T	CMS	$pp \rightarrow \ell X$
		[9] CHATRCHYAN	13AQ	CMS	$pp \rightarrow \ell X$
>1.38	95	[10] CHATRCHYAN	13W	CMS	$pp \rightarrow \gamma\gamma$, δ=6, M_D=5 TeV
>0.715	95	[11] EDELHAUSER	13	RVUE	$pp \rightarrow \ell\overline{\ell} + X$
>1.40	95	[12] AAD	12CP	ATLS	$pp \rightarrow \gamma\gamma$, δ=6, M_D=5 TeV
>1.23	95	[13] AAD	12X	ATLS	$pp \rightarrow \gamma\gamma$, δ=6, M_D=5 TeV
>0.26	95	[14] ABAZOV	12M	D0	$p\overline{p} \rightarrow \mu\mu$
>0.75	95	[15] BAAK	12	RVUE	Electroweak
		[16] FLACKE	12	RVUE	Electroweak
>0.43	95	[17] NISHIWAKI	12	RVUE	$H \rightarrow WW$, $\gamma\gamma$

>0.729	95	[18] AAD	11F	ATLS	$pp \to \gamma\gamma$, δ=6, M_D=5 TeV
>0.961	95	[19] AAD	11X	ATLS	$pp \to \gamma\gamma$, δ=6, M_D=5 TeV
>0.477	95	[20] ABAZOV	10P	D0	$p\overline{p} \to \gamma\gamma$, δ=6, M_D=5 TeV
>1.59	95	[21] ABAZOV	09AE	D0	$p\overline{p} \to$ dijet, angular dist.
>0.6	95	[22] HAISCH	07	RVUE	$\overline{B} \to X_s\gamma$
>0.6	90	[23] GOGOLADZE	06	RVUE	Electroweak
>3.3	95	[24] CORNET	00	RVUE	Electroweak
> 3.3–3.8	95	[25] RIZZO	00	RVUE	Electroweak

[1] AAD 12CC use 4.9 and 5.0 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 7 TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK Z/γ boson (equivalent to $1/R = M_c$). The limit quoted here assumes a flat prior corresponding to when the pure Z/γ KK cross section term dominates. See their Section 15 for more details. NODE=S071KK;LINKAGE=GA

[2] BARBIERI 04 use electroweak precision observables to place a lower bound on the compactification scale $1/R$. Both the gauge bosons and the Higgs boson are assumed to propagate in the bulk. NODE=S071KK;LINKAGE=F

[3] FLORES 23 use a number of 13 TeV Run 2 searches at the LHC to place constraints on the compactification scale $1/R$ and cutoff scale Λ in the minimal universal extra dimension model with Standard Model fields propagating in the bulk (see their Fig.6). NODE=S071KK;LINKAGE=N

[4] AVNISH 21 perform a study on the ATLAS collaboration search for multiple jets plus missing transverse energy from pp collisions at $\sqrt{s}$ = 13 TeV and integrated luminosity of 139 fb^{-1}, to place constraints on the compactification scale and cutoff scale Λ in universal extra dimension models with Standard Model fields propagating in the bulk. NODE=S071KK;LINKAGE=M

[5] AABOUD 18AV use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV in final states with multiple b-jets, to place a lower bound on the compactification scale in a model with two universal extra dimensions. Assuming the radii of the two extra dimensions are equal, a lower limit of 1.8 TeV for the Kaluza-Klein mass is obtained. NODE=S071KK;LINKAGE=J

[6] AABOUD 18CE use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV in final states with same-charge leptons and b-jets, to place a lower bound on the compactification scale in a model with two universal extra dimensions. Assuming the radii of the two extra dimensions are equal, a lower limit of 1.45 TeV for the Kaluza-Klein mass is obtained. NODE=S071KK;LINKAGE=L

[7] ACCOMANDO 15 use electroweak precision observables to place a lower bound on the compactification scale $1/R$. See their Fig. 2 for the bound as a function of sinβ, which parametrizes the VEV contribution from brane and bulk Higgs fields. The quoted value is for the minimum bound which occurs at sinβ = 0.45. NODE=S071KK;LINKAGE=H

[8] KHACHATRYAN 15T use 19.7 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 8 TeV to place a lower bound on the compactification scale $1/R$. NODE=S071KK;LINKAGE=G

[9] CHATRCHYAN 13AQ use 5.0 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 7 TeV and a further 3.7 fb^{-1} of data at $\sqrt{s}$ = 8 TeV to place a lower bound on the compactification scale $1/R$, in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 5 for the bound as a function of the universal bulk fermion mass parameter μ. NODE=S071KK;LINKAGE=E

[10] CHATRCHYAN 13W use diphoton events with large missing transverse momentum in 4.93 fb^{-1} of data produced from pp collisions at $\sqrt{s}$ = 7 TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies Λ/M_c = 20. The model parameters are chosen such that the decay $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction. NODE=S071KK;LINKAGE=C

[11] EDELHAUSER 13 use 19.6 and 20.6 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 8 TeV analyzed by the CMS Collaboration in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the second lightest Kaluza-Klein Z/γ boson (converted to a limit on $1/R = M_c$). The bound assumes Standard Model fields propagating in the bulk and that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies Λ/M_c = 20. NODE=S071KK;LINKAGE=D

[12] AAD 12CP use diphoton events with large missing transverse momentum in 4.8 fb^{-1} of data produced from pp collisions at $\sqrt{s}$ = 7 TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies Λ/M_c = 20. The model parameters are chosen such that the decay $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction. NODE=S071KK;LINKAGE=GD

[13] AAD 12X use diphoton events with large missing transverse momentum in 1.07 fb^{-1} of data produced from pp collisions at $\sqrt{s}$ = 7 TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies Λ/M_c = 20. The model parameters are chosen such that the decay $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction. NODE=S071KK;LINKAGE=DG

[14] ABAZOV 12M use same-sign dimuon events in 7.3 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}$ = 1.96 TeV to place a lower bound on the compactification scale $1/R$, in models with universal extra dimensions where all Standard Model fields propagate in the bulk. NODE=S071KK;LINKAGE=AZ

[15] BAAK 12 use electroweak precision observables to place a lower bound on the compactification scale $1/R$, in models with universal extra dimensions and Standard Model fields propagating in the bulk. Bound assumes a 125 GeV Higgs mass. See their Fig. 25 for the bound as a function of the Higgs mass. NODE=S071KK;LINKAGE=BK

[16] FLACKE 12 use electroweak precision observables to place a lower bound on the compactification scale $1/R$, in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 1 for the bound as a function of the universal bulk fermion mass parameter μ. NODE=S071KK;LINKAGE=FL

[17] NISHIWAKI 12 use up to 2 fb^{-1} of data from the ATLAS and CMS experiments that constrains the production cross section of a Higgs-like particle to place a lower bound on the compactification scale $1/R$ in universal extra dimension models. The quoted bound assumes Standard Model fields propagating in the bulk and a 125 GeV Higgs mass. See their Fig. 1 for the bound as a function of the Higgs mass. NODE=S071KK;LINKAGE=NI

[18] AAD 11F use diphoton events with large missing transverse energy in 3.1 pb^{-1} of data produced from pp collisions at $\sqrt{s}=7$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies $\Lambda/M_c = 20$. The model parameters are chosen such that the decay $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction. NODE=S071KK;LINKAGE=AA

[19] AAD 11X use diphoton events with large missing transverse energy in 36 pb^{-1} of data produced from pp collisions at $\sqrt{s}=7$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies $\Lambda/M_c = 20$. The model parameters are chosen such that the decay $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction. NODE=S071KK;LINKAGE=AD

[20] ABAZOV 10P use diphoton events with large missing transverse energy in 6.3 fb^{-1} of data produced from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ, for the radiative corrections to the Kaluza-Klein masses, satisfies Λ/M_c=20. The model parameters are chosen such that the decay $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction. NODE=S071KK;LINKAGE=AB

[21] ABAZOV 09AE use dijet angular distributions in 0.7 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the compactification scale. NODE=S071KK;LINKAGE=ZO

[22] HAISCH 07 use inclusive $\overline{B}$-meson decays to place a Higgs mass independent bound on the compactification scale $1/R$ in the minimal universal extra dimension model. NODE=S071KK;LINKAGE=HA

[23] GOGOLADZE 06 use electroweak precision observables to place a lower bound on the compactification scale in models with universal extra dimensions. Bound assumes a 115 GeV Higgs mass. See their Fig. 3 for the bound as a function of the Higgs mass. NODE=S071KK;LINKAGE=GO

[24] CORNET 00 translates a bound on the coefficient of the 4-fermion operator $(\overline{\ell}\gamma_\mu \tau^a \ell)(\overline{\ell}\gamma^\mu \tau^a \ell)$ derived by Hagiwara and Matsumoto into a limit on the mass scale of KK W bosons. NODE=S071KK;LINKAGE=B

[25] RIZZO 00 obtains limits from global electroweak fits in models with a Higgs in the bulk (3.8 TeV) or on the standard brane (3.3 TeV). NODE=S071KK;LINKAGE=A

Limits on Kaluza-Klein Gravitons in Warped Extra Dimensions

NODE=S071RSG

NODE=S071RSG

This section places limits on the mass of the first Kaluza-Klein (KK) excitation of the graviton in the warped extra dimension model of Randall and Sundrum. Bounds in parenthesis assume Standard Model fields propagate in the bulk. Experimental bounds depend strongly on the warp parameter, k. See the "Extra Dimensions" review for a full discussion.

Here we list limits for the value of the warp parameter $k/\overline{M}_P = 0.1$.

NODE=S071RSG

VALUE (TeV)	*CL%*	*DOCUMENT ID*		*TECN*	*COMMENT*
>4.78	95	[1] SIRUNYAN	21N	CMS	$pp \to G \to e^+e^-, \mu^+\mu^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
		[2] TUMASYAN	23AP	CMS	$pp \to G \to WW,\ ZZ$
		[3] AAD	22F	ATLS	$pp \to G \to HH$
		[4] TUMASYAN	22D	CMS	$pp \to G \to WW$
		[5] TUMASYAN	22J	CMS	$pp \to G \to ZZ$
		[6] TUMASYAN	22R	CMS	$pp \to G \to ZZ$
		[7] TUMASYAN	22U	CMS	$pp \to G \to HH$
		[8] AAD	21AF	ATLS	$pp \to G \to ZZ$
>4.5	95	[9] AAD	21AY	ATLS	$pp \to G \to \gamma\gamma$
		[10] AAD	20AT	ATLS	$pp \to G \to WW,\ ZZ$
		[11] AAD	20C	ATLS	$pp \to G \to HH$
		[12] AAD	20T	ATLS	$pp \to G \to b\overline{b}$
>2.6	95	[13] SIRUNYAN	20AI	CMS	$pp \to G \to jj$
		[14] SIRUNYAN	20F	CMS	$pp \to G \to HH$
		[15] AABOUD	19O	ATLS	$pp \to G \to HH$
		[16] AAD	19D	ATLS	$pp \to G \to WW,\ ZZ$
		[17] SIRUNYAN	19	CMS	$pp \to G \to HH$
		[18] SIRUNYAN	19BE	CMS	$pp \to G \to HH$
		[19] AABOUD	18BI	ATLS	$pp \to G \to t\overline{t}$
		[20] AABOUD	18CJ	ATLS	$pp \to G \to VV, VH, \ell\overline{\ell}$
		[21] AABOUD	18CQ	ATLS	$pp \to G \to HH$

		22 AABOUD	18CW	ATLS	$pp \to G \to HH$
		23 SIRUNYAN	18AF	CMS	$pp \to G \to HH$
		24 SIRUNYAN	18AS	CMS	$pp \to G \to ZZ$
		25 SIRUNYAN	18CW	CMS	$pp \to G \to HH$
>4.1	95	26 SIRUNYAN	18DU	CMS	$pp \to G \to \gamma\gamma$
		27 SIRUNYAN	18I	CMS	$pp \to G \to b\overline{b}$
		28 AAD	16R	ATLS	$pp \to G \to WW,ZZ$
		29 AAD	15AZ	ATLS	$pp \to G \to WW$
		30 AAD	15CP	ATLS	$pp \to G \to WW,ZZ$
>2.68	95	31 AAD	14V	ATLS	$pp \to G \to e^+e^-, \mu^+\mu^-$
>1.23 (>0.84)	95	32 AAD	13A	ATLS	$pp \to G \to WW$
>0.94 (>0.71)	95	33 AAD	13AO	ATLS	$pp \to G \to WW$
>2.23	95	34 AAD	13AS	ATLS	$pp \to \gamma\gamma, e^+e^-, \mu^+\mu^-$
>0.845	95	35 AAD	12AD	ATLS	$pp \to G \to ZZ$
		36 AALTONEN	12V	CDF	$p\overline{p} \to G \to ZZ$
		37 BAAK	12	RVUE	Electroweak
		38 AALTONEN	11G	CDF	$p\overline{p} \to G \to ZZ$
>1.058	95	39 AALTONEN	11R	CDF	$p\overline{p} \to G \to e^+e^-, \gamma\gamma$
>0.754	95	40 ABAZOV	11H	D0	$p\overline{p} \to G \to WW$
>0.607		41 AALTONEN	10N	CDF	$p\overline{p} \to G \to WW$
>1.05		42 ABAZOV	10F	D0	$p\overline{p} \to G \to e^+e^-, \gamma\gamma$
		43 AALTONEN	08S	CDF	$p\overline{p} \to G \to ZZ$
>0.90		44 ABAZOV	08J	D0	$p\overline{p} \to G \to e^+e^-, \gamma\gamma$
		45 AALTONEN	07G	CDF	$p\overline{p} \to G \to \gamma\gamma$
>0.889		46 AALTONEN	07H	CDF	$p\overline{p} \to G \to e\overline{e}$
>0.785		47 ABAZOV	05N	D0	$p\overline{p} \to G \to \ell\ell, \gamma\gamma$
>0.71		48 ABULENCIA	05A	CDF	$p\overline{p} \to G \to \ell\overline{\ell}$

[1] SIRUNYAN 21N use 137 (140) fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for dilepton resonances in the dielectron (dimuon) channel. See Table 6 for other limits with warp parameter values $k/\overline{M}_P$ = 0.01 and 0.05. This updates the results of SIRUNYAN 18BB. NODE=S071RSG;LINKAGE=KB

[2] TUMASYAN 23AP use 138 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for WW, ZZ diboson resonances in $q\overline{q}q\overline{q}$ final states. See their Figure 7 for the limit on the cross section times branching fraction as a function of the KK graviton mass. Assuming $k/\overline{M}_P$ = 0.5, a graviton mass is excluded below 1400 GeV. This updates the result of SIRUNYAN 20Q. NODE=S071RSG;LINKAGE=RB

[3] AAD 22F use 126–139 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for Higgs boson pair production in the $b\overline{b}b\overline{b}$ final state. See their Figure 14 for limits on the cross section times branching fraction as a function of the KK graviton mass. Assuming $k/\overline{M}_P$ = 1, gravitons in the mass range 298–1460 GeV are excluded. This updates the results of AABOUD 19A. NODE=S071RSG;LINKAGE=MB

[4] TUMASYAN 22D use 137 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for WW resonances in $\ell\nu qq$ final states ($\ell = e, \mu$). See their Figure 6 for the limit on the KK graviton mass as a function of the cross section times branching fraction, including theoretical values for $k/\overline{M}_P$ = 0.5. This updates the results of SIRUNYAN 18AX. NODE=S071RSG;LINKAGE=NB

[5] TUMASYAN 22J use 137 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for ZZ resonances in the $\nu\overline{\nu}q\overline{q}$ final state. See their Figure 10 for the limit on the KK graviton mass as a function of the cross section times branching fraction, assuming $k/\overline{M}_P$ = 0.5 . This updates the result of SIRUNYAN 18BK. NODE=S071RSG;LINKAGE=OB

[6] TUMASYAN 22R use 138 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for ZZ resonances in $2\ell 2q$ final states ($\ell = e, \mu$). See their Figure 8 for the limit on the KK graviton mass as a function of the cross section times branching fraction. Assuming $k/\overline{M}_P$ = 0.5, a graviton mass is excluded below 1200 GeV. This updates the result of SIRUNYAN 18DJ. NODE=S071RSG;LINKAGE=PB

[7] TUMASYAN 22U use 138 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for Higgs boson pair production in the $b\overline{b}q\overline{q}'\ell\nu$, $b\overline{b}\ell\nu\ell\nu$ and $b\overline{b}\ell\nu\nu\ell\nu\nu$ final states ($\ell = e, \mu$). See their Figure 7 for limits on the cross section times branching fraction as a function of the KK graviton mass, including theoretical values for $k/\overline{M}_P$ = 0.3 and 0.5. This updates the results of SIRUNYAN 19CF and SIRUNYAN 18F. NODE=S071RSG;LINKAGE=QB

[8] AAD 21AF use 139 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for ZZ resonances in the $\ell\ell\ell\ell$ and $\ell\ell\nu\overline{\nu}$ final states (ℓ=e, μ). See their Figure 8 for the limit on the cross section times branching fraction as a function of the KK graviton mass, including theoretical values for $k/\overline{M}_P$ = 1. This updates the results of AAD 15AU and AABOUD 18BF. NODE=S071RSG;LINKAGE=JB

[9] AAD 21AY use 139 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV in the diphoton channel to place a lower limit on the mass of the lightest KK graviton. This updates the results of AABOUD 17AP. NODE=S071RSG;LINKAGE=LB

[10] AAD 20AT use 139 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for diboson resonances in semileptonic final states ($\ell\nu qq$,$\ell\ell qq$, $\nu\nu qq$). See their Figure 15 for the limit on the cross section times branching fraction as a function of the KK graviton mass. Lower limits on the graviton mass are also given for $k/\overline{M}_P$ = 1. This updates the results of AABOUD 18AK and AABOUD 18AL. NODE=S071RSG;LINKAGE=IB

[11] AAD 20C use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 13 TeV to search for Higgs boson pair production in the $b\overline{b}b\overline{b}$, $b\overline{b}W^+W^-$, and $b\overline{b}\tau^+\tau^-$ final states. See their NODE=S071RSG;LINKAGE=CB

Figure 5(b)(c) for limits on the cross section as a function of the KK graviton mass. In the case of $k/\overline{M}_P = 1$ and 2, gravitons are excluded in the mass range 260–3000 GeV and 260–1760 GeV, respectively.

[12] AAD 20T use 139 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for narrow resonances decaying to bottom quark pairs. See their Figure 7 for the limit on the product of the cross section, branching fraction, acceptance and b-tagging efficiency as a function of the KK graviton mass. In the case of $k/\overline{M}_P = 0.2$, KK gravitons in the mass range 1.25–2.8 TeV are excluded. NODE=S071RSG;LINKAGE=EB

[13] SIRUNYAN 20AI use 137 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for dijet resonances. See their Figure 6 for the limit on the product of the cross section, branching fraction and acceptance as a function of the KK graviton mass. This updates the results of SIRUNYAN 18BO. NODE=S071RSG;LINKAGE=FB

[14] SIRUNYAN 20F use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $b\overline{b}ZZ$ final state. See their Figure 4 for limits on the cross section times branching fraction as a function of the KK graviton mass, and Figure 5 for limits as a function of $k/\overline{M}_P$. NODE=S071RSG;LINKAGE=GB

[15] AABOUD 19O use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $b\overline{b}WW$ final state. See their Figure 12 for limits on the cross section times branching fraction as a function of the KK graviton mass for $k/\overline{M}_P = 1$ and $k/\overline{M}_P = 2$. NODE=S071RSG;LINKAGE=XA

[16] AAD 19D use 139 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for diboson resonances in the all-hadronic final state. See their Figure 9(b) for the limit on the cross section times branching fraction as a function of the KK graviton mass, including theoretical values for $k/\overline{M}_P = 1$. This updates the results of AABOUD 18F. NODE=S071RSG;LINKAGE=BB

[17] SIRUNYAN 19 use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $\gamma\gamma b\overline{b}$ final state. See their Figure 9 for limits on the cross section times branching fraction as a function of the KK graviton mass. Assuming $k/\overline{M}_P = 1$, gravitons in the mass range 290–810 GeV are excluded. This updates the result of KHACHATRYAN 16BQ. NODE=S071RSG;LINKAGE=WA

[18] SIRUNYAN 19BE use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production by combining the results from four final states: $b\overline{b}\gamma\gamma$, $b\overline{b}\tau\overline{\tau}$, $b\overline{b}b\overline{b}$, and $b\overline{b}VV$. See their Figure 7 for limits on the cross section times branching fraction as a function of the KK graviton mass. NODE=S071RSG;LINKAGE=ZA

[19] AABOUD 18BI use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for top-quark pairs decaying into the lepton-plus jets topology. See their Figure 16 for the limit on the KK graviton mass as a function of the cross section times branching fraction, including theoretical values for $k/\overline{M}_P = 1$. NODE=S071RSG;LINKAGE=CA

[20] AABOUD 18CJ combine the searches for heavy resonances decaying into bosonic and leptonic final states from 36.1 fb^{-1} of pp collision data at $\sqrt{s} = 13$ TeV. The lower limit on the KK graviton mass, with $k/\overline{M}_P = 1$, is 2.3 TeV. NODE=S071RSG;LINKAGE=QA

[21] AABOUD 18CQ use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $b\overline{b}\tau^+\tau^-$ final state. See their Figure 2 for limits on the cross section times branching fraction as a function of the KK graviton mass. Assuming $k/\overline{M}_P = 1$, gravitons in the mass range 325–885 GeV are excluded. NODE=S071RSG;LINKAGE=RA

[22] AABOUD 18CW use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $\gamma\gamma b\overline{b}$ final state. See their Figure 7 for limits on the cross section times branching fraction as a function of the KK graviton mass. NODE=S071RSG;LINKAGE=SA

[23] SIRUNYAN 18AF use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $b\overline{b}b\overline{b}$ final state. See their Figure 9 for limits on the cross section times branching fraction as a function of the KK graviton mass, including theoretical values for $k/\overline{M}_P = 0.5$. This updates the results of KHACHATRYAN 15R. NODE=S071RSG;LINKAGE=HA

[24] SIRUNYAN 18AS use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for ZZ resonances in the $\ell\ell\nu\overline{\nu}$ final state (ℓ=e, μ). See their Figure 5 for the limit on the KK graviton mass as a function of the cross section times branching fraction, including theoretical values for $k/\overline{M}_P = 0.1$, 0.5, and 1.0. NODE=S071RSG;LINKAGE=IA

[25] SIRUNYAN 18CW use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for Higgs boson pair production in the $b\overline{b}b\overline{b}$ final state. See their Figure 8 for limits on the cross section times branching fraction as a function of the KK graviton mass, including theoretical values for $k/\overline{M}_P = 0.5$. NODE=S071RSG;LINKAGE=NA

[26] SIRUNYAN 18DU use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV, in the diphoton channel to place a lower limit on the mass of the lightest KK graviton. See their paper for limits with other warp parameter values $k/\overline{M}_P = 0.01$ and 0.2. This updates the results of KHACHATRYAN 16M. NODE=S071RSG;LINKAGE=VA

[27] SIRUNYAN 18I use 19.7 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV to search for narrow resonances decaying to bottom quark pairs. See their Figure 3 for the limit on the KK graviton mass as a function of the cross section times branching fraction in the mass range of 325–1200 GeV. NODE=S071RSG;LINKAGE=OA

[28] AAD 16R use 20.3 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV to place a lower bound on the mass of the lightest KK graviton. See their Figure 4 for the limit on the KK graviton mass as a function of the cross section times branching fraction. NODE=S071RSG;LINKAGE=M

[29] AAD 15AZ use 20.3 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV to place a lower bound on the mass of the lightest KK graviton. See their Figure 2 for limits on the KK graviton mass as a function of the cross section times branching ratio. NODE=S071RSG;LINKAGE=K

[30] AAD 15CP use 20.3 fb^{-1} of data from pp collisions at $\sqrt{s}=8$ TeV to place a lower bound on the mass of the lightest KK graviton. See their Figures 6b and 6c for the limit on the KK graviton mass as a function of the cross section times branching fraction. NODE=S071RSG;LINKAGE=L

[31] AAD 14V use 20.3 (20.5) fb^{-1} of data from pp collisions at $\sqrt{s}=8$ TeV in the dielectron (dimuon) channels to place a lower bound on the mass of the lightest KK graviton. This updates the results of AAD 12CC . NODE=S071RSG;LINKAGE=B

[32] AAD 13A use 4.7 fb^{-1} of data from pp collisions at $\sqrt{s}=7$ TeV in the $\ell\nu\ell\nu$ channel, to place a lower bound on the mass of the lightest KK graviton. NODE=S071RSG;LINKAGE=DD

[33] AAD 13AO use 4.7 fb^{-1} of data from pp collisions at $\sqrt{s}=7$ TeV in the $\ell\nu jj$ channel, to place a lower bound on the mass of the lightest KK graviton. NODE=S071RSG;LINKAGE=N

[34] AAD 13AS use 4.9 fb^{-1} of data from pp collisions at $\sqrt{s}=7$ TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. The diphoton channel is combined with previous results in the dielectron and dimuon channels to set the best limit. See their Table 2 for warp parameter values $k/\overline{M}_P$ between 0.01 and 0.1. This updates the results of AAD 12Y . NODE=S071RSG;LINKAGE=A

[35] AAD 12AD use 1.02 fb^{-1} of data from pp collisions at $\sqrt{s}=7$ TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons in the $lljj$ and $llll$ channels (ℓ=e, μ). The limit is quoted for the combined $lljj+llll$ channels. See their Figure 5 for limits on the cross section $\sigma(G \rightarrow ZZ)$ as a function of the graviton mass. NODE=S071RSG;LINKAGE=AD

[36] AALTONEN 12V use 6 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons in the $lljj$ and $llll$ channels (ℓ=e, μ). It provides improved limits over the previous analysis in AALTONEN 11G. See their Figure 16 for limits from all channels combined on the cross section times branching ratio $\sigma(p\overline{p} \rightarrow G^* \rightarrow ZZ)$ as a function of the graviton mass. NODE=S071RSG;LINKAGE=EN

[37] BAAK 12 use electroweak precision observables to place a lower bound on the compactification scale $k\, e^{-\pi k R}$, assuming Standard Model fields propagate in the bulk and the Higgs is confined to the IR brane. See their Fig. 27 for more details. NODE=S071RSG;LINKAGE=BK

[38] AALTONEN 11G use 2.5–2.9 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons via the $eeee$, $ee\mu\mu$, $\mu\mu\mu\mu$, $eejj$, and $\mu\mu jj$ channels. See their Fig. 20 for limits on the cross section $\sigma(G \rightarrow ZZ)$ as a function of the graviton mass. NODE=S071RSG;LINKAGE=LN

[39] AALTONEN 11R uses 5.7 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV in the dielectron channel to place a lower bound on the mass of the lightest graviton. It provides combined limits with the diphoton channel analysis of AALTONEN 11U. For warp parameter values $k/\overline{M}_P$ between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 612 and 1058 GeV. See their Table I for more details. NODE=S071RSG;LINKAGE=AT

[40] ABAZOV 11H use 5.4 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the mass of the lightest graviton. Their 95% C.L. exclusion limit does not include masses less than 300 GeV. NODE=S071RSG;LINKAGE=BA

[41] AALTONEN 10N use 2.9 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the mass of the lightest graviton. NODE=S071RSG;LINKAGE=LO

[42] ABAZOV 10F use 5.4 fb^{-1} of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the mass of the lightest graviton. For warp parameter values of $k/\overline{M}_P$ between 0.01 and 0.1 the lower limit on the mass of the lightest graviton is between 560 and 1050 GeV. See their Fig. 3 for more details. NODE=S071RSG;LINKAGE=AO

[43] AALTONEN 08S use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to four electrons via two Z bosons using 1.1 fb^{-1} of data. See their Fig. 8 for limits on $\sigma \cdot \mathrm{B}(G \rightarrow ZZ)$ versus the graviton mass. NODE=S071RSG;LINKAGE=LT

[44] ABAZOV 08J use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons and photons using 1 fb^{-1} of data. For warp parameter values of $k/\overline{M}_P$ between 0.01 and 0.1 the lower limit on the mass of the lightest excitation is between 300 and 900 GeV. See their Fig. 4 for more details. NODE=S071RSG;LINKAGE=AZ

[45] AALTONEN 07G use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to photons using 1.2 fb^{-1} of data. For warp parameter values of $k/\overline{M}_P=0.1$, 0.05, and 0.01 the bounds on the graviton mass are 850, 694, and 230 GeV, respectively. See their Fig. 3 for more details. See also AALTONEN 07H. NODE=S071RSG;LINKAGE=AA

[46] AALTONEN 07H use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons using 1.3 fb^{-1} of data. For a warp parameter value of $k/\overline{M}_P=0.1$ the bound on the graviton mass is 807 GeV. See their Fig. 4 for more details. A combined analysis with the diphoton data of AALTONEN 07G yields for $k/\overline{M}_P=0.1$ a graviton mass lower bound of 889 GeV. NODE=S071RSG;LINKAGE=AL

[47] ABAZOV 05N use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons, electrons or photons, using 260 pb^{-1} of data. For warp parameter values of $k/\overline{M}_P=0.1$, 0.05, and 0.01, the bounds on the graviton mass are 785, 650 and 250 GeV respectively. See their Fig. 3 for more details. NODE=S071RSG;LINKAGE=AB

[48] ABULENCIA 05A use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons or electrons, using 200 pb^{-1} of data. For warp parameter values of $k/\overline{M}_P=0.1$, 0.05, and 0.01, the bounds on the graviton mass are 710, 510 and 170 GeV respectively. NODE=S071RSG;LINKAGE=AU

Limits on Kaluza-Klein Gluons in Warped Extra Dimensions

NODE=S071KKG

This section places limits on the mass of the first Kaluza-Klein (KK) excitation of the gluon in warped extra dimension models with Standard Model fields propagating in the bulk. Bounds are given for a specific benchmark model with $\Gamma/m = 15.3\%$ where Γ is the width and m the mass of the KK gluon. See the "Extra Dimensions" review for more discussion.

NODE=S071KKG

VALUE (TeV)	CL%	DOCUMENT ID		TECN	COMMENT
>3.8	95	[1] AABOUD	18BI	ATLS	$g_{KK} \to t\bar{t} \to \ell j$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
		[2] TUMASYAN	22C	CMS	$g_{KK} \to Rj \to jjj$
		[3] AABOUD	19AS	ATLS	$g_{KK} \to t\bar{t} \to jj$
		[4] SIRUNYAN	19AL	CMS	$g_{KK} \to tT$
>2.5	95	[5] CHATRCHYAN	13BM	CMS	$g_{KK} \to t\bar{t}$
		[6] CHEN	13A		$\overline{B} \to X_s \gamma$
>1.5	95	[7] AAD	12BV	ATLS	$g_{KK} \to t\bar{t} \to \ell j$

NODE=S071KKG

[1] AABOUD 18BI use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV. This result updates AAD 13AQ. NODE=S071KKG;LINKAGE=E

[2] TUMASYAN 22C use 138 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to place limits on a KK gluon decaying to gluons via a spin-0 radion, R. See their Figure 5 for limits on the cross section times branching fraction as a function of the KK gluon mass and various values of the radion mass. NODE=S071KKG;LINKAGE=H

[3] AABOUD 19AS use 36.1 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV. An upper bound of 3.4 TeV is placed on the KK gluon mass for $\Gamma/m = 30\%$. NODE=S071KKG;LINKAGE=F

[4] SIRUNYAN 19AL use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to place limits on a KK gluon decaying to a top quark and a heavy vector-like fermion, T. KK gluon masses between 1.5 and 2.3 TeV and between 2.0 and 2.4 TeV are excluded for T masses of 1.2 and 1.5 TeV, respectively. NODE=S071KKG;LINKAGE=G

[5] CHATRCHYAN 13BM use 19.7 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV. Bound is for a width of approximately 15–20% of the KK gluon mass. NODE=S071KKG;LINKAGE=D

[6] CHEN 13A place limits on the KK mass scale for a specific warped model with custodial symmetry and bulk fermions. See their Figures 4 and 5. NODE=S071KKG;LINKAGE=C

[7] AAD 12BV use 2.05 fb^{-1} of data from pp collisions at $\sqrt{s} = 7$ TeV. NODE=S071KKG;LINKAGE=A

Black Hole Production Limits

NODE=S071405

Semiclassical Black Holes

NODE=S071BGR

NODE=S071BGR

VALUE (GeV)	DOCUMENT ID		TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
	[1] SIRUNYAN	18DA	CMS	$pp \to$ multijet
	[2] AAD	16N	ATLS	$pp \to$ multijet
	[3] AAD	16O	ATLS	$pp \to \ell + (\ell\ell/\ell j/jj)$
	[4] AAD	13AW	ATLS	$pp \to \mu\mu$

[1] SIRUNYAN 18DA use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for semiclassical black holes decaying to multijet final states. No excess of events above the expected level of standard model background was observed. Exclusions at 95% CL are set on the mass threshold for black hole production as a function of the higher-dimensional Planck scale for rotating and nonrotating black holes under several model assumptions (ADD, 2, 4, 6 extra dimensions model) in the 7.1–10.3 TeV range. These limits supersede those in SIRUNYAN 17CP. NODE=S071BGR;LINKAGE=E

[2] AAD 16N use 3.6 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for semiclassical black hole decays to multijet final states. No excess of events above the expected level of Standard Model background was observed. Exclusion contours at 95% C.L. are set on the mass threshold for black hole production versus higher-dimensional Planck scale for rotating black holes (ADD, 6 extra dimensions model). NODE=S071BGR;LINKAGE=A

[3] AAD 16O use 3.2 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for semiclassical black hole decays to high-mass final states with leptons and jets. No excess of events above the expected level of Standard Model background was observed. Exclusion contours at 95% C.L. are set on the mass threshold for black hole production versus higher-dimensional Planck scale for rotating black holes (ADD, 2 to 6 extra dimensions). NODE=S071BGR;LINKAGE=B

[4] AAD 13AW use 20.3 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV to search for semiclassical black hole decays to like-sign dimuon final states using large track multiplicity. No excess of events above the expected level of Standard Model background was observed. Exclusion contours at 95% C.L. are set on the mass threshold for black hole production versus higher-dimensional Planck scale in various extra dimensions, rotating and non-rotating models. NODE=S071BGR;LINKAGE=C

Quantum Black Holes

NODE=S071BHQ
NODE=S071BHQ

VALUE (GeV)	DOCUMENT ID		TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
	[1] AAD	23CB	ATLS	$pp \to e\mu,\ e\tau,\ \mu\tau$
	[2] TUMASYAN	23AW	CMS	$pp \to \tau\nu$
	[3] TUMASYAN	23BC	CMS	$pp \to \gamma j$
	[4] TUMASYAN	23H	CMS	$pp \to e\mu,\ e\tau,\ \mu\tau$
	[5] AAD	20T	ATLS	$pp \to jj$
	[6] AABOUD	18BA	ATLS	$pp \to \gamma j$
	[7] SIRUNYAN	18AT	CMS	$pp \to e\mu$
	[8] SIRUNYAN	18DD	CMS	$pp \to$ dijet, ang. distrib.
	[9] SIRUNYAN	17CP	CMS	$pp \to jj$
	[10] KHACHATRY...	16BE	CMS	$pp \to e\mu$
	[11] KHACHATRY...	15V	CMS	$pp \to jj$
	[12] AAD	14AL	ATLS	$pp \to \ell j$
	[13] AAD	14V	ATLS	$pp \to ee,\ \mu\mu$
	[14] CHATRCHYAN	13A	CMS	$pp \to jj$

[1] AAD 23CB use 139 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays with different-flavor high-mass dilepton final states. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in ADD (6 extra dimensions) and RS1 models. Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, mass thresholds below 5.9 (3.8), 5.2 (3.0), and 5.1 (3.0) TeV are excluded in the $e\mu$, $e\tau$ and $\mu\tau$ channels for the ADD (RS1) models, respectively. These limits supersede those in AABOUD 18CM. NODE=S071BHQ;LINKAGE=Q

[2] TUMASYAN 23AW use 138 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays in the tau lepton plus missing transverse momentum final state. Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, threshold masses below 6.6 TeV are excluded in the ADD model with four extra dimensions (see their Figure 8). NODE=S071BHQ;LINKAGE=R

[3] TUMASYAN 23BC use 138 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays to final states with a photon and a jet. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in ADD (6 extra dimensions) and RS1 models. Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, mass thresholds below 7.5 TeV and 5.2 TeV are excluded for the ADD and RS1 models, respectively (see their Figure 9). NODE=S071BHQ;LINKAGE=S

[4] TUMASYAN 23H use 138 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays with different-flavor high-mass dilepton final states. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in the ADD model (with 4 extra dimensions). Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, mass thresholds below 5.6, 5.2, and 5.0 TeV are excluded in the $e\mu$, $e\tau$ and $\mu\tau$ channels, respectively. NODE=S071BHQ;LINKAGE=P

[5] AAD 20T use 139 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays to final states with dijets. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in an ADD (6 extra dimensions) model. Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, mass thresholds below 9.4 TeV are excluded. This limit supersedes AABOUD 17AK. NODE=S071BHQ;LINKAGE=O

[6] AABOUD 18BA use 36.7 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays to final states with a photon and a jet. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in ADD (6 extra dimensions) and RS1 models. Assuming the black hole mass threshold is equal to the Planck scale, mass thresholds below 7.1 TeV and 4.4 TeV are excluded for the ADD and RS1 models, respectively. These limits supersede those in AAD 16AI. NODE=S071BHQ;LINKAGE=A

[7] SIRUNYAN 18AT use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays to $e\mu$ final states. In Figure 4, lower mass limits of 5.3, 5.5 and 5.6 TeV are placed in a model with 4, 5 and 6 extra dimensions, respectively, and a lower mass limit of 3.6 TeV is found for a single warped dimension. NODE=S071BHQ;LINKAGE=B

[8] SIRUNYAN 18DD use 35.9 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black hole decays in dijet angular distributions. A lower mass limit of 5.9 (8.2) TeV is placed in the RS (ADD) model with one (six) extra dimension(s). NODE=S071BHQ;LINKAGE=M

[9] SIRUNYAN 17CP use 2.3 fb^{-1} of data from pp collisions at $\sqrt{s} = 13$ TeV to search for quantum black holes decaying to dijet final states. No excess of events above the expected level of standard model background was observed. Limits on the quantum black hole mass threshold are set as a function of the higher-dimensional Planck scale, under the assumption that the mass threshold must exceed the above Planck scale. Depending on the model, mass thresholds in the range up to 5.1–9.0 TeV are excluded. NODE=S071BHQ;LINKAGE=L

[10] KHACHATRYAN 16BE use 19.7 fb^{-1} of data from pp collisions at $\sqrt{s} = 8$ TeV to search for quantum black holes undergoing lepton flavor violating decay to the $e\mu$ final state. No excess of events above the expected level of standard model background was observed. Exclusion limits at 95% CL are set on mass thresholds for black hole production in the ADD (2–6 flat extra dimensions), RS1 (1 warped extra dimension), and a model NODE=S071BHQ;LINKAGE=K

with a Planck scale at the TeV scale from a renormalization of the gravitational constant (no extra dimensions). Limits on the black hole mass threshold are set assuming that it is equal to the higher-dimensional Planck scale. Mass thresholds for quantum black holes in the range up to 3.15–3.63 TeV are excluded in the ADD model. In the RS1 model, mass thresholds below 2.81 TeV are excluded in the PDG convention for the Schwarzschild radius. In the model with no extra dimensions, mass thresholds below 1.99 TeV are excluded.

[11] KHACHATRYAN 15V use 19.7 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 8 TeV to search for quantum black holes decaying to dijet final states. No excess of events above the expected level of standard model background was observed. Exclusion limits at 95% CL are set on mass thresholds for black hole production in the ADD (2–6 flat extra dimensions) and RS1 (1 warped extra dimension) model. Limits on the black hole mass threshold are set as a function of the higher-dimensional Planck scale, under the assumption that the mass threshold must exceed the above Planck scale. Depending on the model, mass thresholds in the range up to 5.0–6.3 TeV are excluded. This paper supersedes CHATRCHYAN 13AD. NODE=S071BHQ;LINKAGE=J

[12] AAD 14AL use 20.3 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 8 TeV to search for quantum black hole decays to final states with high-invariant-mass lepton + jet. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in an ADD (6 extra dimensions) model. Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, mass thresholds below 5.3 TeV are excluded. NODE=S071BHQ;LINKAGE=E

[13] AAD 14V use 20.3 (20.5) fb^{-1} of data in the dielectron (dimuon) channels from pp collisions at $\sqrt{s}$ = 8 TeV to search for quantum black hole decays involving high-mass dilepton resonances. No excess of events above the expected level of Standard Model background was observed. Exclusion limits at 95% C.L. are set on mass thresholds for black hole production in ADD (6 extra dimensions) and RS1 models. Assuming the black hole mass threshold is equal to the higher-dimensional Planck scale, mass thresholds below 3.65 TeV and 2.24 TeV are excluded for the ADD and RS1 models, respectively. NODE=S071BHQ;LINKAGE=F

[14] CHATRCHYAN 13A use 5 fb^{-1} of data from pp collisions at $\sqrt{s}$ = 7 TeV to search for quantum black holes decaying to dijet final states. No excess of events above the expected level of standard model background was observed. Exclusion limits at 95% CL are set on mass thresholds for black hole production in the ADD (2–6 flat extra dimensions) and RS (1 warped extra dimension) model. Limits on the black hole mass threshold are set as a function of the higher-dimensional Planck scale, under assumption that the mass threshold must exceed the above Planck scale. Depending on the model, mass thresholds in the range up to 4.0–5.3 TeV are excluded. NODE=S071BHQ;LINKAGE=H

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NODE=S071

AAD	23CB	JHEP 2310 082	G. Aad *et al.*	(ATLAS Collab.)	REFID=62540
FLORES	23	IJMP A38 2350002	M.M. Flores *et al.*	(WITS, WARS, NIP-UPD+)	REFID=62385
TUMASYAN	23AP	PL B844 137813	A. Tumasyan *et al.*	(CMS Collab.)	REFID=62406
TUMASYAN	23AW	JHEP 2309 051	A. Tumasyan *et al.*	(CMS Collab.)	REFID=62524
TUMASYAN	23BC	JHEP 2312 189	A. Tumasyan *et al.*	(CMS Collab.)	REFID=62577
TUMASYAN	23H	JHEP 2305 227	A. Tumasyan *et al.*	(CMS Collab.)	REFID=62114
AAD	22F	PR D105 092002	G. Aad *et al.*	(ATLAS Collab.)	REFID=61741
TUMASYAN	22C	PL B832 137263	A. Tumasyan *et al.*	(CMS Collab.)	REFID=61709
TUMASYAN	22D	PR D105 032008	A. Tumasyan *et al.*	(CMS Collab.)	REFID=61720
TUMASYAN	22J	PR D106 012004	A. Tumasyan *et al.*	(CMS Collab.)	REFID=61758
TUMASYAN	22R	JHEP 2204 087	A. Tumasyan *et al.*	(CMS Collab.)	REFID=61792
TUMASYAN	22U	JHEP 2205 005	A. Tumasyan *et al.*	(CMS Collab.)	REFID=61795
AAD	21AF	EPJ C81 332	G. Aad *et al.*	(ATLAS Collab.)	REFID=61333
AAD	21AY	PL B822 136651	G. Aad *et al.*	(ATLAS Collab.)	REFID=61547
AAD	21F	PR D103 112006	G. Aad *et al.*	(ATLAS Collab.)	REFID=61247
AVNISH	21	PR D103 115011	Avnish *et al.*		REFID=61250;ERROR=1
BLAKEMORE	21	PR D104 L061101	C.P. Blakemore *et al.*	(STAN)	REFID=61392
HEACOCK	21	SCI 373 1239	B. Heacock *et al.*	(NIST, RIKEN, NAGO+)	REFID=61546
SIRUNYAN	21A	EPJ C81 13	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=60775
Also		EPJ C81 333 (errat.)	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=61322
SIRUNYAN	21N	JHEP 2107 208	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=61317
TUMASYAN	21D	JHEP 2111 153	A. Tumasyan *et al.*	(CMS Collab.)	REFID=61559
AAD	20AT	EPJ C80 1165	G. Aad *et al.*	(ATLAS Collab.)	REFID=60774
AAD	20C	PL B800 135103	G. Aad *et al.*	(ATLAS Collab.)	REFID=60147
AAD	20T	JHEP 2003 145	G. Aad *et al.*	(ATLAS Collab.)	REFID=60484
LEE	20	PRL 124 101101	J.G. Lee *et al.*	(WASH)	REFID=60536
SIRUNYAN	20AI	JHEP 2005 033	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=60492
SIRUNYAN	20F	PR D102 032003	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=60346
SIRUNYAN	20Q	EPJ C80 237	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=60424
TAN	20A	PRL 124 051301	W.-H. Tan *et al.*		REFID=60531
AABOUD	19A	JHEP 1901 030	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59370
AABOUD	19AS	PR D99 092004	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59874
AABOUD	19O	JHEP 1904 092	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59672
AAD	19D	JHEP 1909 091	G. Aad *et al.*	(ATLAS Collab.)	REFID=60081
Also		JHEP 2006 042 (errat.)	G. Aad *et al.*	(ATLAS Collab.)	REFID=60499
SIRUNYAN	19	PL B788 7	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59412
SIRUNYAN	19AC	JHEP 1904 114	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59676
SIRUNYAN	19AL	EPJ C79 208	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59711
SIRUNYAN	19BE	PRL 122 121803	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59768
SIRUNYAN	19CF	JHEP 1910 125	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=60085
AABOUD	18AK	JHEP 1803 042	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59092
AABOUD	18AL	JHEP 1803 009	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59093
AABOUD	18AV	JHEP 1807 089	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59130
AABOUD	18BA	EPJ C78 102	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59157
AABOUD	18BF	EPJ C78 293	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59172
AABOUD	18BI	EPJ C78 565	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59191
AABOUD	18CE	JHEP 1812 039	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59369
AABOUD	18CJ	PR D98 052008	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59441

AABOUD	18CM	PR D98 092008	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59474
AABOUD	18CQ	PRL 121 191801	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59529
AABOUD	18CW	JHEP 1811 040	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=59567
AABOUD	18F	PL B777 91	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=58626
AABOUD	18I	JHEP 1801 126	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=58792
BERGE	18	PRL 120 141101	J. Berge *et al.*	(MICROSCOPE Collab.)	REFID=58817
FAYET	18	PR D97 055039	P. Fayet	(EPOL)	REFID=58909
FAYET	18A	PR D99 055043	P. Fayet	(ENSP, EPOL)	REFID=59312
HADDOCK	18	PR D97 062002	C. Haddock *et al.*	(NAGO, KEK, OSAK+)	REFID=58911
SIRUNYAN	18AF	PL B781 244	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59016
SIRUNYAN	18AS	JHEP 1803 003	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59094
SIRUNYAN	18AT	JHEP 1804 073	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59095
SIRUNYAN	18AX	JHEP 1805 088	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59103
SIRUNYAN	18BB	JHEP 1806 120	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59112
SIRUNYAN	18BK	JHEP 1807 075	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59131
SIRUNYAN	18BO	JHEP 1808 130	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59142
SIRUNYAN	18BV	EPJ C78 291	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59173
SIRUNYAN	18CW	JHEP 1808 152	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59249
SIRUNYAN	18DA	JHEP 1811 042	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59306
SIRUNYAN	18DD	EPJ C78 789	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59318
SIRUNYAN	18DJ	JHEP 1809 101	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59341
SIRUNYAN	18DU	PR D98 092001	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59469
SIRUNYAN	18F	JHEP 1801 054	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=58787
SIRUNYAN	18I	PRL 120 201801	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=58825
SIRUNYAN	18S	PR D97 092005	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=58938
AABOUD	17AK	PR D96 052004	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=58217
AABOUD	17AP	PL B775 105	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=58296
KLIMCHITSK...	17A	PR D95 123013	G.L. Klimchitskaya, V.M. Mostepanenko		REFID=58006
SIRUNYAN	17AQ	JHEP 1710 073	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=58337
SIRUNYAN	17CP	PL B774 279	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=59307
SIRUNYAN	17F	JHEP 1707 013	A.M. Sirunyan *et al.*	(CMS Collab.)	REFID=57820
AABOUD	16F	JHEP 1606 059	M. Aaboud *et al.*	(ATLAS Collab.)	REFID=57335
AAD	16AI	JHEP 1603 041	G. Aad *et al.*	(ATLAS Collab.)	REFID=57322
AAD	16N	JHEP 1603 026	G. Aad *et al.*	(ATLAS Collab.)	REFID=57168
AAD	16O	PL B760 520	G. Aad *et al.*	(ATLAS Collab.)	REFID=57169
AAD	16R	PL B755 285	G. Aad *et al.*	(ATLAS Collab.)	REFID=57172
KHACHATRY...	16BE	EPJ C76 317	V. Khachatryan *et al.*	(CMS Collab.)	REFID=57373
KHACHATRY...	16BQ	PR D94 052012	V. Khachatryan *et al.*	(CMS Collab.)	REFID=57433
KHACHATRY...	16M	PRL 117 051802	V. Khachatryan *et al.*	(CMS Collab.)	REFID=57196
KHACHATRY...	16N	PL B755 102	V. Khachatryan *et al.*	(CMS Collab.)	REFID=57207
AAD	15AU	EPJ C75 69	G. Aad *et al.*	(ATLAS Collab.)	REFID=56662
AAD	15AZ	EPJ C75 209	G. Aad *et al.*	(ATLAS Collab.)	REFID=56667
Also		EPJ C75 370 (errat.)	G. Aad *et al.*	(ATLAS Collab.)	REFID=57174
AAD	15CP	JHEP 1512 055	G. Aad *et al.*	(ATLAS Collab.)	REFID=57021
AAD	15CS	PR D91 012008	G. Aad *et al.*	(ATLAS Collab.)	REFID=57143
Also		PR D92 059903 (errat.)	G. Aad *et al.*	(ATLAS Collab.)	REFID=56964
ACCOMANDO	15	MPL A30 1540010	E. Accomando	(SHMP)	REFID=56870
KHACHATRY...	15AE	JHEP 1504 025	V. Khachatryan *et al.*	(CMS Collab.)	REFID=56632
KHACHATRY...	15R	PL B749 560	V. Khachatryan *et al.*	(CMS Collab.)	REFID=56531
KHACHATRY...	15T	PR D91 092005	V. Khachatryan *et al.*	(CMS Collab.)	REFID=56533
KHACHATRY...	15V	PR D91 052009	V. Khachatryan *et al.*	(CMS Collab.)	REFID=56535
AAD	14AL	PRL 112 091804	G. Aad *et al.*	(ATLAS Collab.)	REFID=56056
AAD	14BE	EPJ C74 3134	G. Aad *et al.*	(ATLAS Collab.)	REFID=56304
AAD	14V	PR D90 052005	G. Aad *et al.*	(ATLAS Collab.)	REFID=55884
AAD	13A	PL B718 860	G. Aad *et al.*	(ATLAS Collab.)	REFID=54789
AAD	13AO	PR D87 112006	G. Aad *et al.*	(ATLAS Collab.)	REFID=55156
AAD	13AQ	PR D88 012004	G. Aad *et al.*	(ATLAS Collab.)	REFID=55168
AAD	13AS	NJP 15 043007	G. Aad *et al.*	(ATLAS Collab.)	REFID=55270
AAD	13AW	PR D88 072001	G. Aad *et al.*	(ATLAS Collab.)	REFID=55454
AAD	13C	PRL 110 011802	G. Aad *et al.*	(ATLAS Collab.)	REFID=54791
AAD	13D	JHEP 1301 029	G. Aad *et al.*	(ATLAS Collab.)	REFID=54792
AAD	13E	PR D87 015010	G. Aad *et al.*	(ATLAS Collab.)	REFID=54828
CHATRCHYAN	13A	JHEP 1301 013	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=54770
CHATRCHYAN	13AD	JHEP 1307 178	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=55045
CHATRCHYAN	13AQ	PR D87 072005	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=55140
CHATRCHYAN	13BM	PRL 111 211804	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=55492
Also		PRL 112 119903 (errat.)	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=55841
CHATRCHYAN	13W	JHEP 1303 111	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=55028
CHEN	13A	CP C37 063102	J-B. Chen *et al.*	(DALI)	REFID=55360
EDELHAUSER	13	JHEP 1308 091	L. Edelhauser, T. Flacke, M. Kramer	(AACH, KAIST)	REFID=55261
XU	13	JP G40 035107	J. Xu *et al.*		REFID=55264
AAD	12AD	PL B712 331	G. Aad *et al.*	(ATLAS Collab.)	REFID=54177
AAD	12BV	JHEP 1209 041	G. Aad *et al.*	(ATLAS Collab.)	REFID=54583
AAD	12CC	JHEP 1211 138	G. Aad *et al.*	(ATLAS Collab.)	REFID=54653
AAD	12CP	PL B718 411	G. Aad *et al.*	(ATLAS Collab.)	REFID=54725
AAD	12X	PL B710 519	G. Aad *et al.*	(ATLAS Collab.)	REFID=54160
AAD	12Y	PL B710 538	G. Aad *et al.*	(ATLAS Collab.)	REFID=54161
AALTONEN	12V	PR D85 012008	T. Aaltonen *et al.*	(CDF Collab.)	REFID=54361
ABAZOV	12M	PRL 108 131802	V.M. Abazov *et al.*	(D0 Collab.)	REFID=54209
AJELLO	12	JCAP 1202 012	M. Ajello *et al.*	(Fermi-LAT Collab.)	REFID=54444
BAAK	12	EPJ C72 2003	M. Baak *et al.*	(Gfitter Group)	REFID=54138
CHATRCHYAN	12R	PRL 108 111801	S. Chatrchyan *et al.*	(CMS Collab.)	REFID=54207
FLACKE	12	PR D85 126007	T. Flacke, C. Pasold	(WURZ)	REFID=54500
NISHIWAKI	12	PL B707 506	K. Nishiwaki *et al.*	(KOBE, OSAK)	REFID=54151
AAD	11F	PRL 106 121803	G. Aad *et al.*	(ATLAS Collab.)	REFID=16559
AAD	11X	EPJ C71 1744	G. Aad *et al.*	(ATLAS Collab.)	REFID=53926
AALTONEN	11G	PR D83 112008	T. Aaltonen *et al.*	(CDF Collab.)	REFID=16433
AALTONEN	11R	PRL 107 051801	T. Aaltonen *et al.*	(CDF Collab.)	REFID=16455
AALTONEN	11U	PR D83 011102	T. Aaltonen *et al.*	(CDF Collab.)	REFID=53703
AARON	11C	PL B705 52	F. D. Aaron *et al.*	(H1 Collab.)	REFID=53845
ABAZOV	11H	PRL 107 011801	V.M. Abazov *et al.*	(D0 Collab.)	REFID=16473
BEZERRA	11	PR D83 075004	V.B. Bezerra *et al.*		REFID=16525
SUSHKOV	11	PRL 107 171101	A.O. Sushkov *et al.*		REFID=53828

AALTONEN	10N	PRL 104 241801	T. Aaltonen *et al.*	(CDF Collab.)	REFID=53313
ABAZOV	10F	PRL 104 241802	V.M. Abazov *et al.*	(D0 Collab.)	REFID=53318
ABAZOV	10P	PRL 105 221802	V.M. Abazov *et al.*	(D0 Collab.)	REFID=53443
BEZERRA	10	PR D81 055003	V.B. Bezerra *et al.*		REFID=53391
ABAZOV	09AE	PRL 103 191803	V.M. Abazov *et al.*	(D0 Collab.)	REFID=53064
ABAZOV	09D	PRL 102 051601	V.M. Abazov *et al.*	(D0 Collab.)	REFID=52649
MASUDA	09	PRL 102 171101	M. Masuda, M. Sasaki	(ICRR)	REFID=52844
AALTONEN	08AC	PRL 101 181602	T. Aaltonen *et al.*	(CDF Collab.)	REFID=52563
AALTONEN	08S	PR D78 012008	T. Aaltonen *et al.*	(CDF Collab.)	REFID=52373
ABAZOV	08J	PRL 100 091802	V.M. Abazov *et al.*	(D0 Collab.)	REFID=52388
ABAZOV	08S	PRL 101 011601	V.M. Abazov *et al.*	(D0 Collab.)	REFID=52397
DAS	08	PR D78 063011	P.K. Das, V.H.S. Kumar, P.K. Suresh		REFID=52530
GERACI	08	PR D78 022002	A.A. Geraci *et al.*	(STAN)	REFID=52417
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AALTONEN	07G	PRL 99 171801	T. Aaltonen *et al.*	(CDF Collab.)	REFID=52012
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DECCA	07A	EPJ C51 963	R.S. Decca *et al.*		REFID=51880
HAISCH	07	PR D76 034014	U. Haisch, A. Weiler		REFID=51860
KAPNER	07	PRL 98 021101	D.J. Kapner *et al.*		REFID=51607
SCHAEL	07A	EPJ C49 411	S. Schael *et al.*	(ALEPH Collab.)	REFID=51743
TU	07	PRL 98 201101	L.-C. Tu *et al.*		REFID=51808
ABDALLAH	06C	EPJ C45 589	J. Abdallah *et al.*	(DELPHI Collab.)	REFID=51224
ABULENCIA,A	06	PRL 97 171802	A. Abulencia *et al.*	(CDF Collab.)	REFID=51434
GERDES	06	PR D73 112008	D. Gerdes *et al.*		REFID=51286
GOGOLADZE	06	PR D74 093012	I. Gogoladze, C. Macesanu		REFID=51544
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SMULLIN	05	PR D72 122001	S.J. Smullin *et al.*		REFID=51007
ACHARD	04E	PL B587 16	P. Achard *et al.*	(L3 Collab.)	REFID=49893
ACOSTA	04C	PRL 92 121802	D. Acosta *et al.*	(CDF Collab.)	REFID=49935
BARBIERI	04	NP B703 127	R. Barbieri *et al.*		REFID=55653
CASSE	04	PRL 92 111102	M. Casse *et al.*		REFID=50055
CHEKANOV	04B	PL B591 23	S. Chekanov *et al.*	(ZEUS Collab.)	REFID=49912
HOYLE	04	PR D70 042004	C.D. Hoyle *et al.*	(WASH)	REFID=51084
ABAZOV	03	PRL 90 251802	V.M. Abazov *et al.*	(D0 Collab.)	REFID=49443
ABBIENDI	03D	EPJ C26 331	G. Abbiendi *et al.*	(OPAL Collab.)	REFID=49290
ACHARD	03D	PL B572 133	P. Achard *et al.*	(L3 Collab.)	REFID=49556
ADLOFF	03	PL B568 35	C. Adloff *et al.*	(H1 Collab.)	REFID=49522
CHIAVERINI	03	PRL 90 151101	J. Chiaverini *et al.*		REFID=49357
GIUDICE	03	NP B663 377	G.F. Giudice, A. Strumia		REFID=49427
HANNESTAD	03	PR D67 125008	S. Hannestad, G.G. Raffelt		REFID=49536
Also		PR D69 029901(errat.)	S. Hannestad, G.G. Raffelt		REFID=50033
HEISTER	03C	EPJ C28 1	A. Heister *et al.*	(ALEPH Collab.)	REFID=49379
LONG	03	NAT 421 922	J.C. Long *et al.*		REFID=49699
ACHARD	02	PL B524 65	P. Achard *et al.*	(L3 Collab.)	REFID=48524
ACHARD	02D	PL B531 28	P. Achard *et al.*	(L3 Collab.)	REFID=48663
HANNESTAD	02	PRL 88 071301	S. Hannestad, G. Raffelt		REFID=48698
ABBOTT	01	PRL 86 1156	B. Abbott *et al.*	(D0 Collab.)	REFID=48041
FAIRBAIRN	01	PL B508 335	M. Fairbairn		REFID=48264
HANHART	01	PL B509 1	C. Hanhart *et al.*		REFID=48160
HOYLE	01	PRL 86 1418	C.D. Hoyle *et al.*		REFID=48082
ABBIENDI	00R	EPJ C13 553	G. Abbiendi *et al.*	(OPAL Collab.)	REFID=47644
ABREU	00A	PL B491 67	P. Abreu *et al.*	(DELPHI Collab.)	REFID=47825
ABREU	00S	PL B485 45	P. Abreu *et al.*	(DELPHI Collab.)	REFID=47716
ABREU	00Z	EPJ C17 53	P. Abreu *et al.*	(DELPHI Collab.)	REFID=47790
CASSISI	00	PL B481 323	S. Cassisi *et al.*		REFID=47907
CHANG	00B	PRL 85 3765	L.N. Chang *et al.*		REFID=47823
CHEUNG	00	PR D61 015005	K. Cheung		REFID=47885
CORNET	00	PR D61 037701	F. Cornet, M. Relano, J. Rico		REFID=47888
GRAESSER	00	PR D61 074019	M.L. Graesser		REFID=47889
HAN	00	PR D62 125018	T. Han, D. Marfatia, R.-J. Zhang		REFID=47896
MATHEWS	00	JHEP 0007 008	P. Mathews, S. Raychaudhuri, K. Sridhar		REFID=49019
MELE	00	PR D61 117901	S. Mele, E. Sanchez		REFID=47892
RIZZO	00	PR D61 016007	T.G. Rizzo, J.D. Wells		REFID=47887
ABBIENDI	99P	PL B465 303	G. Abbiendi *et al.*	(OPAL Collab.)	REFID=47290
ACCIARRI	99M	PL B464 135	M. Acciarri *et al.*	(L3 Collab.)	REFID=47229
ACCIARRI	99R	PL B470 268	M. Acciarri *et al.*	(L3 Collab.)	REFID=47322
ACCIARRI	99S	PL B470 281	M. Acciarri *et al.*	(L3 Collab.)	REFID=47323
BOURILKOV	99	JHEP 9908 006	D. Bourilkov		REFID=47332
HOSKINS	85	PR D32 3084	J.K. Hoskins *et al.*		REFID=49700