

Reference = ACHASOV 16A; PR D93 092001
 Verifier code = DRUZHININ

Normally we send all verifications for one experiment to one person, usually the spokesperson or data-analysis coordinator, who then distributes them to the appropriate people. Please tell us if we should send the verifications for your experiment to someone else.

PLEASE READ NOW

**PLEASE
REPLY
WITHIN
ONE WEEK**

Vladimir P. Druzhinin

EMAIL: v.p.druzhinin@inp.nsk.su

March 20, 2017

Dear Colleague,

- (1) Please check the results of your experiment carefully. They are marked.
- (2) Please reply within one week.
- (3) Please reply even if everything is correct.
- (4) IMPORTANT!! Please tell WHICH papers you are verifying. We have lots of requests out.
- (5) Feel free to make comments on our treatment of any of the results (not just yours) you see.

Thank you for helping us make the Review accurate and useful.

Sincerely,

Simon Eidelman
 BINP, Budker Inst. of Nuclear Physics
 Prospekt Lavrent'eva 11
 RU-630090 Novosibirsk
 Russian Federation

EMAIL: simon.eidelman@cern.ch

LIGHT UNFLAVORED MESONS ($S = C = B = 0$)

For $I = 1$ (π, b, ρ, a): $u\bar{d}, (u\bar{u} - d\bar{d})/\sqrt{2}, d\bar{u}$;
for $I = 0$ ($\eta, \eta', h, h', \omega, \phi, f, f'$): $c_1(u\bar{u} + d\bar{d}) + c_2(s\bar{s})$

$\rho(770)$

$I^G(J^{PC}) = 1^+(1^{--})$

$\rho(770) \Gamma(e^+e^-)\Gamma(i)/\Gamma^2(\text{total})$

| $\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$ | $\Gamma_{12}/\Gamma \times \Gamma_8/\Gamma$ |
|---|---|
|---|---|

| VALUE (units 10^{-8}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

2.22 ± 0.26 OUR AVERAGE

YOUR DATA $1.98 \pm 0.22 \pm 0.10$ Error includes scale factor of 1.3. See the ideogram below.
 $2.90 \pm 0.60 \pm 0.18$ 18k $1^{\text{ACHASOV}} 16A \text{ SND } 0.60\text{-}1.38 e^+e^- \rightarrow \pi^0\gamma$
 $2.37 \pm 0.53 \pm 0.33$ 36k $2^{\text{ACHASOV}} 03 \text{ SND } 0.60\text{-}0.97 e^+e^- \rightarrow \pi^0\gamma$
 $3.61 \pm 0.74 \pm 0.49$ 10k $3^{\text{DOLINSKY}} 89 \text{ ND } e^+e^- \rightarrow \pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •

1.875 ± 0.026 $4^{\text{BENAYOUN}} 10 \text{ RVUE } 0.4\text{-}1.05 e^+e^-$

YOUR NOTE

¹ From the VMD model with the rho(770), omega(782), phi(1020) resonances, and an additional resonance describing the total contribution of the rho(1450) and omega(1420) states. Supersedes ACHASOV 03.

² Using $\sigma_\phi \rightarrow \pi^0\gamma$ from ACHASOV 00 and $m_\rho = 775.97$ MeV in the model with the energy-independent phase of rho-omega interference equal to $(-10.2 \pm 7.0)^\circ$.

³ Recalculated by us from the cross section in the peak.

⁴ A simultaneous fit of $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$ data.

$\rho(770) \text{ BRANCHING RATIOS}$

| $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$ | Γ_8/Γ |
|---|-------------------|
|---|-------------------|

| VALUE (units 10^{-4}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

• • • We do not use the following data for averages, fits, limits, etc. • • •

YOUR DATA

4.20 ± 0.52 $1^{\text{ACHASOV}} 16A \text{ SND } 0.60\text{-}1.38 e^+e^- \rightarrow \pi^0\gamma$
 $6.21^{+1.28}_{-1.18} \pm 0.39$ 18k $2,3^{\text{AKHMETSHIN}} 05 \text{ CMD2 } 0.60\text{-}1.38 e^+e^- \rightarrow \pi^0\gamma$
 $5.22 \pm 1.17 \pm 0.75$ 36k $3,4^{\text{ACHASOV}} 03 \text{ SND } 0.60\text{-}0.97 e^+e^- \rightarrow \pi^0\gamma$
 6.8 ± 1.7 $5^{\text{BENAYOUN}} 96 \text{ RVUE } 0.54\text{-}1.04 e^+e^- \rightarrow \pi^0\gamma$
 7.9 ± 2.0 $3^{\text{DOLINSKY}} 89 \text{ ND } e^+e^- \rightarrow \pi^0\gamma$

YOUR NOTE

¹ Using $B(\rho \rightarrow e^+e^-)$ from PDG 15. Supersedes ACHASOV 03.

² Using $B(\rho \rightarrow e^+e^-) = (4.67 \pm 0.09) \times 10^{-5}$.

³ Not independent of the corresponding $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$.

⁴ Using $B(\rho \rightarrow e^+e^-) = (4.54 \pm 0.10) \times 10^{-5}$.

⁵ Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

$\rho(770) \text{ REFERENCES}$

| | | | | | |
|------------|------------|-----|-------------------------|-------------------------------|-----------------------------|
| YOUR PAPER | ACHASOV | 16A | PR D93 092001 | M.N. Achasov <i>et al.</i> | (SND Collab.) |
| | PDG | 15 | RPP 2015 at pdg.lbl.gov | | (PDG Collab.) |
| | BENAYOUN | 10 | EPJ C65 211 | M. Benayoun <i>et al.</i> | |
| | AKHMETSHIN | 05 | PL B605 26 | R.R. Akhmetshin <i>et al.</i> | (Novosibirsk CMD-2 Collab.) |
| | ACHASOV | 03 | PL B559 171 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) |
| | ACHASOV | 00 | EPJ C12 25 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) |
| | BENAYOUN | 96 | ZPHY C72 221 | M. Benayoun <i>et al.</i> | (IPNP, NOVO) |
| | DOLINSKY | 91 | PRPL 202 99 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| | DOLINSKY | 89 | ZPHY C42 511 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| | DRUZHININ | 84 | PL 144B 136 | V.P. Druzhinin <i>et al.</i> | (NOVO) |

$\omega(782)$

$I^G(J^{PC}) = 0^-(1^{--})$

NODE=MXXX005

NODE=MXXX005

NODE=M009

NODE=M009233

NODE=M009G2

NODE=M009G2

NODE=M009G2;LINKAGE=B

NODE=M009G;LINKAGE=SH

NODE=M009G2;LINKAGE=LP

NODE=M009G2;LINKAGE=BE

NODE=M009235

NODE=M009R9

NODE=M009R9

NODE=M009R9;LINKAGE=C

NODE=M009R9;LINKAGE=AK

NODE=M009R9;LINKAGE=BZ

NODE=M009R9;LINKAGE=AS

NODE=M009R9;LINKAGE=A

NODE=M009

REFID=57513

REFID=56977;ERROR=2

REFID=53212

REFID=50330

REFID=49187

REFID=47417

REFID=45753

REFID=41369

REFID=41003

REFID=20561

NODE=M001

$\omega(782) \Gamma(e^+ e^-) \Gamma(i)/\Gamma^2(\text{total})$

| $\Gamma(e^+ e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}$ | | $\Gamma_9/\Gamma \times \Gamma_2/\Gamma$ | | | |
|---|--|--|-------------|--|--|
| <u>VALUE (units 10^{-6})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> | |
| 6.37 ± 0.09 OUR AVERAGE | | | | | |
| YOUR DATA | 6.336 ± 0.056 ± 0.089 | ¹ ACHASOV 16A | SND | 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 6.47 ± 0.14 ± 0.39 | 18k AKHMETSHIN 05 | CMD2 | 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 6.50 ± 0.11 ± 0.20 | 36k ACHASOV 03 | SND | 0.60–0.97 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 6.34 ± 0.21 ± 0.21 | 10k DOLINSKY 89 | ND | $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| | 6.80 ± 0.13 | ⁴ BENAYOUN 10 | RVUE | 0.4–1.05 $e^+ e^-$ | |
| YOUR NOTE | 1 From the VMD model with the interfering $\rho(770)$, $\omega(782)$, $\phi(1020)$, and an additional resonance describing the total contribution of the $\rho(1450)$ and $\omega(1420)$ states. Supersedes ACHASOV 03. | | | | |
| | 2 Using $\sigma_{\phi \rightarrow \pi^0 \gamma}$ from ACHASOV 00 and $m_\omega = 782.57$ MeV in the model with the energy-independent phase of ρ - ω interference equal to $(-10.2 \pm 7.0)^\circ$. | | | | |
| | 3 Recalculated by us from the cross section in the peak. | | | | |
| | 4 A simultaneous fit of $e^+ e^- \rightarrow \pi^+ \pi^-$, $\pi^+ \pi^- \pi^0$, $\pi^0 \gamma$, $\eta \gamma$ data. | | | | |

$\omega(782)$ BRANCHING RATIOS

| $\Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}$ | | Γ_2/Γ | | | |
|---|---|----------------------------------|-------------|--|--|
| <u>VALUE (units 10^{-2})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | |
| YOUR DATA | 8.88 ± 0.18 | ¹ ACHASOV 16A | SND | 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 8.09 ± 0.14 | ² AMBROSINO 08G | KLOE | $e^+ e^- \rightarrow \pi^+ \pi^- 2\pi^0$, $2\pi^0 \gamma$ | |
| | 9.06 ± 0.20 ± 0.57 | 18k ^{3,4} AKHMETSHIN 05 | CMD2 | 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 9.34 ± 0.15 ± 0.31 | 36k ACHASOV 03 | SND | 0.60–0.97 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 8.65 ± 0.16 ± 0.42 | 1.2M ACHASOV 03D | RVUE | 0.44–2.00 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$ | |
| | 8.39 ± 0.24 | 9k ⁷ BENAYOUN 96 | RVUE | $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 8.88 ± 0.62 | 10k DOLINSKY 89 | ND | $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| YOUR NOTE | 1 Using $B(\omega \rightarrow e^+ e^-)$ from PDG 15. Supersedes ACHASOV 03. | | | | |
| | 2 Not independent of $\Gamma(\pi^0 \gamma)/\Gamma(\pi^+ \pi^- \pi^0)$ from AMBROSINO 08G. | | | | |
| | 3 Using $B(\omega \rightarrow e^+ e^-) = (7.14 \pm 0.13) \times 10^{-5}$. | | | | |
| | 4 Not independent of the corresponding $\Gamma(e^+ e^-) \times \Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}^2$. | | | | |
| | 5 Using ACHASOV 03, ACHASOV 03D and $B(\omega \rightarrow \pi^+ \pi^-) = (1.70 \pm 0.28)\%$. | | | | |
| | 6 Not independent of the corresponding $\Gamma(e^+ e^-) \times \Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}^2$. | | | | |
| | 7 Reanalysis of DRUZHININ 84, DOLINSKY 89, DOLINSKY 91 taking into account the triangle anomaly contributions. | | | | |

$\omega(782)$ REFERENCES

| | | | | |
|------------|---------------|-------------------------|-------------------------------|-----------------------------|
| YOUR PAPER | ACHASOV 16A | PR D93 092001 | M.N. Achasov <i>et al.</i> | (SND Collab.) |
| | PDG 15 | RPP 2015 at pdg.lbl.gov | | (PDG Collab.) |
| | BENAYOUN 10 | EPJ C65 211 | M. Benayoun <i>et al.</i> | |
| | AMBROSINO 08G | PL B669 223 | F. Ambrosino <i>et al.</i> | (KLOE Collab.) |
| | AKHMETSHIN 05 | PL B605 26 | R.R. Akhmetshin <i>et al.</i> | (Novosibirsk CMD-2 Collab.) |
| | ACHASOV 03 | PL B559 171 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) |
| | ACHASOV 03D | PR D68 052006 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) |
| | ACHASOV 00 | EPJ C12 25 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) |
| | BENAYOUN 96 | ZPHY C72 221 | M. Benayoun <i>et al.</i> | (IPNP, NOVO) |
| | DOLINSKY 91 | PRPL 202 99 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| | DOLINSKY 89 | ZPHY C42 511 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| | DRUZHININ 84 | PL 144B 136 | V.P. Druzhinin <i>et al.</i> | (NOVO) |

$\phi(1020)$

$I^G(J^{PC}) = 0^-(1^{--})$

| $\phi(1020) \Gamma(i) \Gamma(e^+ e^-)/\Gamma^2(\text{total})$ | | | | | |
|---|--------------------|--|-------------|--|--|
| $\Gamma(\pi^0 \gamma)/\Gamma_{\text{total}} \times \Gamma(e^+ e^-)/\Gamma_{\text{total}}$ | | $\Gamma_7/\Gamma \times \Gamma_9/\Gamma$ | | | |
| <u>VALUE (units 10^{-7})</u> | <u>EVTS</u> | <u>DOCUMENT ID</u> | <u>TECN</u> | <u>COMMENT</u> | |
| 3.87 ± 0.15 OUR AVERAGE | | | | | |
| YOUR DATA | 4.04 ± 0.09 ± 0.19 | ¹ ACHASOV 16A | SND | 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$ | |
| | 3.75 ± 0.11 ± 0.29 | 18k AKHMETSHIN 05 | CMD2 | 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$ | |

NODE=M001225

NODE=M001G4

NODE=M001G4

NODE=M001G4;LINKAGE=A

NODE=M001G;LINKAGE=SH

NODE=M001G4;LINKAGE=LP

NODE=M001G4;LINKAGE=BE

NODE=M001220

NODE=M001R28

NODE=M001R28

NODE=M001R28;LINKAGE=A

NODE=M001R28;LINKAGE=AM

NODE=M001R;LINKAGE=AH

NODE=M001R;LINKAGE=VL

NODE=M001R28;LINKAGE=VF

NODE=M001R28;LINKAGE=ZL

NODE=M001R28;LINKAGE=A1

NODE=M001

REFID=57513

REFID=56977;ERROR=3

REFID=53212

REFID=52573

REFID=50330

REFID=49187

REFID=49577

REFID=47417

REFID=45753

REFID=41369

REFID=41003

REFID=20561

NODE=M004

NODE=M004224

NODE=M004G3

NODE=M004G3

$3.67 \pm 0.10^{+0.27}_{-0.25}$ ² ACHASOV 00 SND $e^+ e^- \rightarrow \pi^0 \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.29 ± 0.11 ³ BENAYOUN 10 RVUE 0.4–1.05 $e^+ e^-$

YOUR NOTE

¹ From the VMD model with the interfering $\rho(770)$, $\omega(782)$, $\phi(1020)$ resonances, and an additional resonance describing the total contribution of the $\rho(1450)$ and $\omega(1420)$ states. Supersedes ACHASOV 00.

³ A simultaneous fit of $e^+ e^- \rightarrow \pi^+ \pi^-$, $\pi^+ \pi^- \pi^0$, $\pi^0 \gamma$, $\eta \gamma$ data.

² From the $\pi^0 \rightarrow 2\gamma$ decay and using $B(\pi^0 \rightarrow 2\gamma) = (98.798 \pm 0.032) \times 10^{-2}$.

ERROR=4

NODE=M004G3;LINKAGE=B

NODE=M004G7;LINKAGE=BE

NODE=M004G3;LINKAGE=A

$\phi(1020)$ BRANCHING RATIOS

$\Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}$

| VALUE (units 10^{-3}) | EVTS | DOCUMENT ID | TECN | COMMENT | Γ_7/Γ |
|--------------------------|------|-------------|------|---------|-------------------|
|--------------------------|------|-------------|------|---------|-------------------|

1.31 ± 0.13 OUR AVERAGE

1.30 ± 0.13 DRUZHININ 84 ND $e^+ e^- \rightarrow 3\gamma$

1.4 ± 0.5 32 COSME 76 OSPK $e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

YOUR DATA

1.367 ± 0.072 ¹ ACHASOV 16A SND 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$

1.258 ± 0.037 ± 0.077 18k ^{2,3} AKHMETSHIN 05 CMD2 0.60–1.38 $e^+ e^- \rightarrow \pi^0 \gamma$

1.226 ± 0.036 ^{+0.096} _{-0.089} ⁴ ACHASOV 00 SND $e^+ e^- \rightarrow \pi^0 \gamma$

1.26 ± 0.17 ⁵ BENAYOUN 96 RVUE 0.54–1.04 $e^+ e^- \rightarrow \pi^0 \gamma$

NODE=M004R17

NODE=M004R17

OCCUR=3

ERROR=5

YOUR NOTE

¹ Using $B(\phi \rightarrow e^+ e^-)$ from PDG 15. Supersedes ACHASOV 00.

⁵ Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

² Using $B(\phi \rightarrow e^+ e^-) = (2.98 \pm 0.04) \times 10^{-4}$.

³ Not independent of the corresponding $\Gamma(e^+ e^-) \times \Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}^2$.

⁴ From the $\pi^0 \rightarrow 2\gamma$ decay and using $B(\phi \rightarrow e^+ e^-) = (2.99 \pm 0.08) \times 10^{-4}$.

NODE=M004R17;LINKAGE=D

NODE=M004R;LINKAGE=TS

NODE=M004R17;LINKAGE=AH

NODE=M004R17;LINKAGE=AK

NODE=M004R;LINKAGE=3G

$\phi(1020)$ REFERENCES

YOUR PAPER

| | | | | |
|------------|-----|-------------------------|-------------------------------|-----------------------------|
| ACHASOV | 16A | PR D93 092001 | M.N. Achasov <i>et al.</i> | (SND Collab.) |
| PDG | 15 | RPP 2015 at pdg.lbl.gov | | (PDG Collab.) |
| BENAYOUN | 10 | EPJ C65 211 | M. Benayoun <i>et al.</i> | |
| AKHMETSHIN | 05 | PL B605 26 | R.R. Akhmetshin <i>et al.</i> | (Novosibirsk CMD-2 Collab.) |
| ACHASOV | 00 | EPJ C12 25 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) |
| BENAYOUN | 96 | ZPHY C72 221 | M. Benayoun <i>et al.</i> | (IPNP, NOVO) |
| DOLINSKY | 91 | PRPL 202 99 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| DOLINSKY | 89 | ZPHY C42 511 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| DRUZHININ | 84 | PL 144B 136 | V.P. Druzhinin <i>et al.</i> | (NOVO) |
| COSME | 76 | PL 63B 352 | G. Cosme <i>et al.</i> | (ORsay) |

NODE=M004

REFID=57513

REFID=56977;ERROR=6

REFID=53212

REFID=50330

REFID=47417

REFID=45753

REFID=41369

REFID=41003

REFID=20561

REFID=20529

Reference = ACHASOV 16B; PR D94 092002
 Verifier code = DRUZHININ

Normally we send all verifications for one experiment to one person, usually the spokesperson or data-analysis coordinator, who then distributes them to the appropriate people. Please tell us if we should send the verifications for your experiment to someone else.

PLEASE READ NOW

**PLEASE
REPLY
WITHIN
ONE WEEK**

Vladimir P. Druzhinin

EMAIL: v.p.druzhinin@inp.nsk.su

March 20, 2017

Dear Colleague,

- (1) Please check the results of your experiment carefully. They are marked.
- (2) Please reply within one week.
- (3) Please reply even if everything is correct.
- (4) IMPORTANT!! Please tell WHICH papers you are verifying. We have lots of requests out.
- (5) Feel free to make comments on our treatment of any of the results (not just yours) you see.

Thank you for helping us make the Review accurate and useful.

Sincerely,

Simon Eidelman
 BINP, Budker Inst. of Nuclear Physics
 Prospekt Lavrent'eva 11
 RU-630090 Novosibirsk
 Russian Federation

EMAIL: simon.eidelman@cern.ch

LIGHT UNFLAVORED MESONS ($S = C = B = 0$)

For $I = 1$ (π, b, ρ, a): $u\bar{d}, (u\bar{u} - d\bar{d})/\sqrt{2}, d\bar{u}$;
 for $I = 0$ ($\eta, \eta', h, h', \omega, \phi, f, f'$): $c_1(u\bar{u} + d\bar{d}) + c_2(s\bar{s})$

 $\omega(1420)$
 $J^P(C) = 0^-(1^{--})$
 $\omega(1420) \Gamma(i)\Gamma(e^+e^-)/\Gamma^2(\text{total})$

$$\Gamma(\omega\eta)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}} \quad \Gamma_3/\Gamma \times \Gamma_5/\Gamma$$

| VALUE (units 10^{-8}) | EVTS | DOCUMENT ID | TECN | COMMENT |
|--------------------------|------|-------------|------|---------|
|--------------------------|------|-------------|------|---------|

• • • We do not use the following data for averages, fits, limits, etc. • • •

YOUR DATA

| | | | | |
|---------------------|-----|-----------------------|---------|---|
| $1.6^{+0.9}_{-0.7}$ | 898 | ²⁴ ACHASOV | 16B SND | $1.34-2.00 e^+e^- \rightarrow \omega\eta$ |
|---------------------|-----|-----------------------|---------|---|

YOUR NOTE

²⁴ From a fit with contributions from $\omega(1420)$, $\omega(1650)$, and $\phi(1680)$. The mass and the width of $\omega(1420)$ are fixed to the 2014 edition (PDG 14) of this review.

NODE=MXXX005

NODE=MXXX005

NODE=M125

NODE=M125230

NODE=M125G6

NODE=M125G6

 $\omega(1420) \text{ REFERENCES}$

| | | | | | |
|------------|----------------|-----------|--------------------------------|--|--------------------------------|
| YOUR PAPER | ACHASOV PDG | 16B 14 | PR D94 092002 CP C38 070001 | M.N. Achasov <i>et al.</i> K. Olive <i>et al.</i> | (SND Collab.) (PDG Collab.) |
|------------|----------------|-----------|--------------------------------|--|--------------------------------|

 $\omega(1650)$
 $J^P(C) = 0^-(1^{--})$
 $\omega(1650) \text{ MASS}$

$$\text{VALUE (MeV)} \quad \text{EVTS} \quad \text{DOCUMENT ID} \quad \text{TECN} \quad \text{COMMENT}$$

1670 ± 30 OUR ESTIMATE

• • • We do not use the following data for averages, fits, limits, etc. • • •

YOUR DATA

| | | | | |
|------------------------|-------|-------------------------|-----------|---|
| 1660 \pm 10 | 898 | ¹ ACHASOV | 16B SND | $1.34-2.00 e^+e^- \rightarrow \omega\eta$ |
| 1680 \pm 10 | 13.1k | ² AULCHENKO | 15A SND | $1.05-1.80 e^+e^- \rightarrow \pi^+\pi^-\pi^0$ |
| 1667 \pm 13 \pm 6 | | AUBERT | 07AU BABR | $10.6 e^+e^- \rightarrow \omega\pi^+\pi^-\gamma$ |
| 1645 \pm 8 | 13 | AUBERT | 06D BABR | $10.6 e^+e^- \rightarrow \omega\eta\gamma$ |
| 1660 \pm 10 \pm 2 | | AUBERT,B | 04N BABR | $10.6 e^+e^- \rightarrow \pi^+\pi^-\pi^0\gamma$ |
| 1770 \pm 50 \pm 60 | 1.2M | ³ ACHASOV | 03D RVUE | $0.44-2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$ |
| 1619 \pm 5 | | ⁴ HENNER | 02 RVUE | $1.2-2.0 e^+e^- \rightarrow \rho\pi, \omega\pi\pi$ |
| 1700 \pm 20 | | EUGENIO | 01 SPEC | $18 \pi^-p \rightarrow \omega\eta n$ |
| 1705 \pm 26 | 612 | ⁵ AKHMETSHIN | 00D CMD2 | $e^+e^- \rightarrow \omega\pi^+\pi^-$ |
| 1820^{+190}_{-150} | | ⁶ ACHASOV | 98H RVUE | $e^+e^- \rightarrow \pi^+\pi^-\pi^0$ |
| 1840^{+100}_{-70} | | ⁷ ACHASOV | 98H RVUE | $e^+e^- \rightarrow \omega\pi^+\pi^-$ |
| 1780^{+170}_{-300} | | ⁸ ACHASOV | 98H RVUE | $e^+e^- \rightarrow K^+K^-$ |
| ~ 2100 | | ⁹ ACHASOV | 98H RVUE | $e^+e^- \rightarrow K_S^0 K^\pm\pi^\mp$ |
| 1606 \pm 9 | | ¹⁰ CLEGG | 94 RVUE | |
| 1662 \pm 13 | 750 | ¹¹ ANTONELLI | 92 DM2 | $1.34-2.4 e^+e^- \rightarrow \rho\pi, \omega\pi\pi$ |
| 1670 \pm 20 | | ATKINSON | 83B OMEG | $20-70 \gamma p \rightarrow 3\pi X$ |
| 1657 \pm 13 | | CORDIER | 81 DM1 | $e^+e^- \rightarrow \omega 2\pi$ |
| 1679 \pm 34 | 21 | ESPOSITO | 80 FRAM | $e^+e^- \rightarrow 3\pi$ |
| 1652 \pm 17 | | COSME | 79 OSPK | $e^+e^- \rightarrow 3\pi$ |

NODE=M126M

 NODE=M126M
 → UNCHECKED ←

OCCUR=2

OCCUR=2

OCCUR=3

OCCUR=4

OCCUR=5

OCCUR=4

YOUR NOTE

- ¹ From a fit with contributions from $\omega(1420)$, $\omega(1650)$, and $\phi(1680)$.
² From a fit with contributions from $\omega(782)$, $\phi(1020)$, $\omega(1420)$, and $\omega(1650)$.
³ From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the $\pi^+ \pi^- \pi^0$ and ANTONELLI 92 on the $\omega \pi^+ \pi^-$ final states. Supersedes ACHASOV 99E and ACHASOV 02E.
⁴ Using results of CORDIER 81 and preliminary data of DOLINSKY 91 and ANTONELLI 92.
⁵ Using the data of AKHMETSHIN 00D and ANTONELLI 92. The $\rho\pi$ dominance for the energy dependence of the $\omega(1420)$ and $\omega(1650)$ width assumed.
⁶ Using data from BARKOV 87, DOLINSKY 91, and ANTONELLI 92.
⁷ Using the data from ANTONELLI 92.
⁸ Using the data from IVANOV 81 and BISELLO 88B.
⁹ Using the data from BISELLO 91C.
¹⁰ From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.
¹¹ From the combined fit of the $\rho\pi$ and $\omega\pi\pi$ final states.

NODE=M126M;LINKAGE=E
 NODE=M126M;LINKAGE=A
 NODE=M126M;LINKAGE=VH

NODE=M126M;LINKAGE=AB

NODE=M126M;LINKAGE=KI

NODE=M126M;LINKAGE=L1
 NODE=M126M;LINKAGE=L2
 NODE=M126M;LINKAGE=L3
 NODE=M126M;LINKAGE=L4
 NODE=M126M;LINKAGE=AD

NODE=M126M;LINKAGE=AE

NODE=M126W

NODE=M126W

→ UNCHECKED ←

| VALUE (MeV) | EVTS | DOCUMENT ID | TECN | COMMENT |
|---|-------|-------------------|-----------|--|
| 315± 35 OUR ESTIMATE | | | | |
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | |
| 110± 20 | 898 | 12 ACHASOV | 16B SND | 1.34–2.00 $e^+ e^- \rightarrow \omega\eta$ |
| 310± 30 | 13.1k | 13 AULCHENKO | 15A SND | 1.05–1.80 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$ |
| 222± 25± 20 | | AUBERT | 07AU BABR | 10.6 $e^+ e^- \rightarrow \omega\pi^+ \pi^- \gamma$ |
| 114± 14 | 13 | AUBERT | 06D BABR | 10.6 $e^+ e^- \rightarrow \omega\eta\gamma$ |
| 230± 30± 20 | | AUBERT,B | 04N BABR | 10.6 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \gamma$ |
| 490 ⁺²⁰⁰ ₋₁₅₀ ± 130 | 1.2M | 14 ACHASOV | 03D RVUE | 0.44–2.00 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$ |
| 250± 14 | | 15 HENNER | 02 RVUE | 1.2–2.0 $e^+ e^- \rightarrow \rho\pi, \omega\pi\pi$ |
| 250± 50 | | EUGENIO | 01 SPEC | 18 $\pi^- p \rightarrow \omega\eta n$ |
| 370± 25 | 612 | 16 AKHMETSHIN 00D | CMD2 | $e^+ e^- \rightarrow \omega\pi^+ \pi^-$ |
| 113± 20 | | 17 CLEGG | 94 RVUE | |
| 280± 24 | 750 | 18 ANTONELLI | 92 DM2 | 1.34–2.4 $e^+ e^- \rightarrow \rho\pi, \omega\pi\pi$ |
| 160± 20 | | ATKINSON | 83B OMEG | 20–70 $\gamma p \rightarrow 3\pi X$ |
| 136± 46 | | CORDIER | 81 DM1 | $e^+ e^- \rightarrow \omega 2\pi$ |
| 99± 49 | 21 | ESPOSITO | 80 FRAM | $e^+ e^- \rightarrow 3\pi$ |
| 42± 17 | | COSME | 79 OSPK | $e^+ e^- \rightarrow 3\pi$ |

YOUR DATA

- 12 From a fit with contributions from $\omega(1420)$, $\omega(1650)$, and $\phi(1680)$.
 13 From a fit with contributions from $\omega(782)$, $\phi(1020)$, $\omega(1420)$, and $\omega(1650)$.
 14 From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the $\pi^+ \pi^- \pi^0$ and ANTONELLI 92 on the $\omega \pi^+ \pi^-$ final states. Supersedes ACHASOV 99E and ACHASOV 02E.
 15 Using results of CORDIER 81 and preliminary data of DOLINSKY 91 and ANTONELLI 92.
 16 Using the data of AKHMETSHIN 00D and ANTONELLI 92. The $\rho\pi$ dominance for the energy dependence of the $\omega(1420)$ and $\omega(1650)$ width assumed.
 17 From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.
 18 From the combined fit of the $\rho\pi$ and $\omega\pi\pi$ final states.

OCCUR=2

OCCUR=5

OCCUR=4

YOUR NOTE

- 12 From a fit with contributions from $\omega(1420)$, $\omega(1650)$, and $\phi(1680)$.
 13 From a fit with contributions from $\omega(782)$, $\phi(1020)$, $\omega(1420)$, and $\omega(1650)$.
 14 From the combined fit of ANTONELLI 92, ACHASOV 01E, ACHASOV 02E, and ACHASOV 03D data on the $\pi^+ \pi^- \pi^0$ and ANTONELLI 92 on the $\omega \pi^+ \pi^-$ final states. Supersedes ACHASOV 99E and ACHASOV 02E.
 15 Using results of CORDIER 81 and preliminary data of DOLINSKY 91 and ANTONELLI 92.
 16 Using the data of AKHMETSHIN 00D and ANTONELLI 92. The $\rho\pi$ dominance for the energy dependence of the $\omega(1420)$ and $\omega(1650)$ width assumed.
 17 From a fit to two Breit-Wigner functions and using the data of DOLINSKY 91 and ANTONELLI 92.
 18 From the combined fit of the $\rho\pi$ and $\omega\pi\pi$ final states.

NODE=M126W;LINKAGE=E
 NODE=M126W;LINKAGE=A
 NODE=M126W;LINKAGE=VH

NODE=M126W;LINKAGE=AB

NODE=M126W;LINKAGE=KI

NODE=M126W;LINKAGE=AD

NODE=M126W;LINKAGE=AE

NODE=M126230

NODE=M126G5

NODE=M126G5

| Γ(ωη)/Γ _{total} × Γ(e ⁺ e ⁻)/Γ _{total} | Γ ₃ /Γ × Γ ₄ /Γ |
|---|---|
| VALUE (units 10⁻⁶) CL% | |
| 0.44±0.05 | 898 19 ACHASOV 16B SND 1.34–2.00 $e^+ e^- \rightarrow \omega\eta$ |
| 0.57±0.06 | 13 AUBERT 06D BABR 10.6 $e^+ e^- \rightarrow \omega\eta\gamma$ |
| <6 | 90 28 AKHMETSHIN 03B CMD2 $e^+ e^- \rightarrow \eta\pi^0\gamma$ |

YOUR NOTE

- 19 From a fit with contributions from $\omega(1420)$, $\omega(1650)$, and $\phi(1680)$.
 28 $\omega(1650)$ mass and width fixed at 1700 MeV and 250 MeV, respectively.

ERROR=7

NODE=M126G5;LINKAGE=A

NODE=M126G5;LINKAGE=KH

$\omega(1650)$ REFERENCES

NODE=M126

| | | | | | | |
|------------|----------------------|------------|--|--|--------------------------------|----------------------------|
| YOUR PAPER | ACHASOV AULCHENKO | 16B 15A | PR D94 092002 JETP 121 27 Translated from ZETF 148 34. | M.N. Achasov <i>et al.</i> V.M. Aulchenko <i>et al.</i> | (SND Collab.) (SND Collab.) | REFID=57537 REFID=56843 |
| | AUBERT | 07AU | PR D76 092005 | B. Aubert <i>et al.</i> | (BABAR Collab.) | REFID=52049 |
| | AUBERT | 06D | PR D73 052003 | B. Aubert <i>et al.</i> | (BABAR Collab.) | REFID=51047 |
| | AUBERT,B | 04N | PR D70 072004 | B. Aubert <i>et al.</i> | (BABAR Collab.) | REFID=50184 |
| | ACHASOV | 03D | PR D68 052006 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) | REFID=49577 |
| | AKHMETSHIN | 03B | PL B562 173 | R.R. Akhmetshin <i>et al.</i> | (Novosibirsk CMD-2 Collab.) | REFID=49406 |
| | ACHASOV | 02E | PR D66 032001 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) | REFID=48815 |
| | HENNER | 02 | EPJ C26 3 | V.K. Henner <i>et al.</i> | (Novosibirsk SND Collab.) | REFID=49177 |
| | ACHASOV | 01E | PR D63 072002 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) | REFID=48311 |
| | EUGENIO | 01 | PL B497 190 | P. Eugenio <i>et al.</i> | (Novosibirsk CMD-2 Collab.) | REFID=48010 |
| | AKHMETSHIN | 00D | PL B489 125 | R.R. Akhmetshin <i>et al.</i> | (Novosibirsk CMD-2 Collab.) | REFID=47935 |
| | ACHASOV | 99E | PL B462 365 | M.N. Achasov <i>et al.</i> | (Novosibirsk SND Collab.) | REFID=47391 |
| | ACHASOV | 98H | PR D57 4334 | N.N. Achasov, A.A. Kozhevnikov | | REFID=46323 |
| | CLEGG | 94 | ZPHY C62 455 | A.B. Clegg, A. Donnachie | (LANC, MCHS) | REFID=44081 |
| | ANTONELLI | 92 | ZPHY C56 15 | A. Antonelli <i>et al.</i> | (DM2 Collab.) | REFID=43168 |
| | BISELLLO | 91C | ZPHY C52 227 | D. Bisello <i>et al.</i> | (DM2 Collab.) | REFID=41867 |
| | DOLINSKY | 91 | PRPL 202 99 | S.I. Dolinsky <i>et al.</i> | (NOVO) | REFID=41369 |
| | BISELLLO | 88B | ZPHY C39 13 | D. Bisello <i>et al.</i> | (PADO, CLER, FRAS+) | REFID=40581 |
| | BARKOV | 87 | JETPL 46 164 | L.M. Barkov <i>et al.</i> | (NOVO) | REFID=40280 |
| | | | Translated from ZETFP 46 132. | | | |
| | ATKINSON | 83B | PL 127B 132 | M. Atkinson <i>et al.</i> | (BONN, CERN, GLAS+) | REFID=21502 |
| | CORDIER | 81 | PL 106B 155 | A. Cordier <i>et al.</i> | (ORSAY) | REFID=21586 |
| | IVANOV | 81 | PL 107B 297 | P.M. Ivanov <i>et al.</i> | (NOVO) | REFID=20553 |
| | ESPOSITO | 80 | LNC 28 195 | B. Esposito <i>et al.</i> | (FRAS, NAPL, PADO+) | REFID=21584 |
| | COSME | 79 | NP B152 215 | G. Cosme <i>et al.</i> | (IPN) | REFID=21475 |

Reference = ACHASOV 16D; PR D94 112001
 Verifier code = DRUZHININ

Normally we send all verifications for one experiment to one person, usually the spokesperson or data-analysis coordinator, who then distributes them to the appropriate people. Please tell us if we should send the verifications for your experiment to someone else.

PLEASE READ NOW

**PLEASE
REPLY
WITHIN
ONE WEEK**

Vladimir P. Druzhinin

EMAIL: v.p.druzhinin@inp.nsk.su

March 20, 2017

Dear Colleague,

- (1) Please check the results of your experiment carefully. They are marked.
- (2) Please reply within one week.
- (3) Please reply even if everything is correct.
- (4) IMPORTANT!! Please tell WHICH papers you are verifying. We have lots of requests out.
- (5) Feel free to make comments on our treatment of any of the results (not just yours) you see.

Thank you for helping us make the Review accurate and useful.

Sincerely,

Simon Eidelman
 BINP, Budker Inst. of Nuclear Physics
 Prospekt Lavrent'eva 11
 RU-630090 Novosibirsk
 Russian Federation

EMAIL: simon.eidelman@cern.ch

LIGHT UNFLAVORED MESONS ($S = C = B = 0$)

For $I = 1$ (π, b, ρ, a): $u\bar{d}, (u\bar{u} - d\bar{d})/\sqrt{2}, d\bar{u}$;
 for $I = 0$ ($\eta, \eta', h, h', \omega, \phi, f, f'$): $c_1(u\bar{u} + d\bar{d}) + c_2(s\bar{s})$

$\rho(1450)$

$I^G(J^{PC}) = 1^+(1^-^-)$

See our mini-review under the $\rho(1700)$.

$\rho(1450)$ MASS

$\omega\pi$ MODE

| | VALUE (MeV) | EVTS | DOCUMENT ID | TECN | COMMENT | |
|--|------------------------------|-------|------------------|----------|--|--|
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | | |
| YOUR DATA | 1510 \pm 7 | 10.2k | 1 ACHASOV | 16D SND | 1.05–2.00 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$ | |
| | 1544 \pm 22 $^{+11}_{-46}$ | 821 | 2 MATVIENKO | 15 BELL | $\bar{B}^0 \rightarrow D^* \pi^+ \omega \pi^-$ | |
| | 1491 \pm 19 | 7815 | 3 ACHASOV | 13 SND | 1.05–2.00 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$ | |
| | 1582 \pm 17 \pm 25 | 2382 | 4 AKHMETSHIN 03B | CMD2 | $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$ | |
| | 1349 \pm 25 $^{+10}_{-5}$ | 341 | 5 ALEXANDER | 01B CLE2 | $B \rightarrow D^(*) \omega \pi^-$ | |
| | 1523 \pm 10 | | 6 EDWARDS | 00A CLE2 | $\tau^- \rightarrow \omega \pi^- \nu_\tau$ | |
| | 1463 \pm 25 | | 7 CLEGG | 94 RVUE | | |
| | 1250 | | 8 ASTON | 80C OMEG | 20–70 $\gamma p \rightarrow \omega \pi^0 p$ | |
| | 1290 \pm 40 | | 8 BARBER | 80C SPEC | 3–5 $\gamma p \rightarrow \omega \pi^0 p$ | |

- YOUR NOTE
- 1 From a phenomenological model based on vector meson dominance with interfering $\rho(770)$, $\rho(1450)$, and $\rho(1700)$. The $\rho(1700)$ mass and width are fixed at 1720 MeV and 250 MeV, respectively. Systematic uncertainties not estimated. Supersedes ACHASOV 13.
 - 2 Using Breit-Wigner parameterization of the $\rho(1450)$ and assuming equal probabilities of the $\rho(1450) \rightarrow \pi\pi$ and $\rho(1450) \rightarrow \omega\pi$ decays.
 - 3 From a phenomenological model based on vector meson dominance with the interfering $\rho(1450)$ and $\rho(1700)$ and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.
 - 4 Using the data of AKHMETSHIN 03B and BISELLO 91B assuming the $\omega\pi^0$ and $\pi^+\pi^-$ mass dependence of the total width. $\rho(1700)$ mass and width fixed at 1700 MeV and 240 MeV, respectively.
 - 5 Using Breit-Wigner parameterization of the $\rho(1450)$ and assuming the $\omega\pi^-$ mass dependence for the total width.
 - 6 Mass-independent width parameterization. $\rho(1700)$ mass and width fixed at 1700 MeV and 235 MeV respectively.
 - 7 Using data from BISELLO 91B, DOLINSKY 86 and ALBRECHT 87L.
 - 8 Not separated from $b_1(1235)$, not pure $J^P = 1^-$ effect.

$\rho(1450)$ WIDTH

$\omega\pi$ MODE

| | VALUE (MeV) | EVTS | DOCUMENT ID | TECN | COMMENT | |
|--|---------------------------------|-------|------------------|----------|--|--|
| • • • We do not use the following data for averages, fits, limits, etc. • • • | | | | | | |
| YOUR DATA | 440 \pm 40 | 10.2k | 1 ACHASOV | 16D SND | 1.05–2.00 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$ | |
| | 303 \pm 31 $^{+69}_{-52 - 7}$ | 821 | 2 MATVIENKO | 15 BELL | $\bar{B}^0 \rightarrow D^* \pi^+ \omega \pi^-$ | |
| | 429 \pm 42 \pm 10 | 2382 | 3 AKHMETSHIN 03B | CMD2 | $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$ | |
| | 547 \pm 86 $^{+46}_{-45}$ | 341 | 4 ALEXANDER | 01B CLE2 | $B \rightarrow D^(*) \omega \pi^-$ | |
| | 400 \pm 35 | | 5 EDWARDS | 00A CLE2 | $\tau^- \rightarrow \omega \pi^- \nu_\tau$ | |
| | 311 \pm 62 | | 6 CLEGG | 94 RVUE | | |
| | 300 | | 7 ASTON | 80C OMEG | 20–70 $\gamma p \rightarrow \omega \pi^0 p$ | |
| | 320 \pm 100 | | 7 BARBER | 80C SPEC | 3–5 $\gamma p \rightarrow \omega \pi^0 p$ | |

NODE=MXXX005

NODE=MXXX005

NODE=M105

NODE=M105

NODE=M105205

NODE=M105M3
NODE=M105M3

OCCUR=2

NODE=M105M3;LINKAGE=D

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NODE=M105M3;LINKAGE=AC

NODE=M105M3;LINKAGE=HK

NODE=M105M3;LINKAGE=3Z

NODE=M105M;LINKAGE=E1

NODE=M105M3;LINKAGE=B

NODE=M105M3;LINKAGE=A

NODE=M105210

NODE=M105W3
NODE=M105W3

OCCUR=3

YOUR NOTE

- ¹ From a phenomenological model based on vector meson dominance with interfering $\rho(770)$, $\rho(1450)$, and $\rho(1700)$. The $\rho(1700)$ mass and width are fixed at 1720 MeV and 250 MeV, respectively. Systematic uncertainties not estimated. Supersedes ACHASOV 13.
- ² Using Breit-Wigner parameterization of the $\rho(1450)$ and assuming equal probabilities of the $\rho(1450) \rightarrow \pi\pi$ and $\rho(1450) \rightarrow \omega\pi$ decays.
- ³ Using the data of AKHMETSHIN 03B and BISELLO 91B assuming the $\omega\pi^0$ and $\pi^+\pi^-$ mass dependence of the total width. $\rho(1700)$ mass and width fixed at 1700 MeV and 240 MeV, respectively.
- ⁴ Using Breit-Wigner parameterization of the $\rho(1450)$ and assuming the $\omega\pi^-$ mass dependence for the total width.
- ⁵ Mass-independent width parameterization. $\rho(1700)$ mass and width fixed at 1700 MeV and 235 MeV respectively.
- ⁶ Using data from BISELLO 91B, DOLINSKY 86 and ALBRECHT 87L.
- ⁷ Not separated from $b_1(1235)$, not pure $J^P = 1^-$ effect.

 $\rho(1450) \Gamma(i)/\Gamma(\text{total}) \times \Gamma(e^+e^-)/\Gamma(\text{total})$

| $\Gamma(\omega\pi)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ | | $\Gamma_3/\Gamma \times \Gamma_9/\Gamma$ | | |
|---|------|--|------|---------|
| VALUE (units 10^{-6}) | EVTS | DOCUMENT ID | TECN | COMMENT |

• • • We do not use the following data for averages, fits, limits, etc. • • •

YOUR DATA

| | | | | |
|---------------|-------|----------------------|---------|---|
| 2.1 \pm 0.4 | 10.2k | ¹ ACHASOV | 16D SND | 1.05–2.00 $e^+e^- \rightarrow \pi^0\pi^0\gamma$ |
| 5.3 \pm 0.4 | 7815 | ² ACHASOV | 13 SND | 1.05–2.00 $e^+e^- \rightarrow \pi^0\pi^0\gamma$ |

YOUR NOTE

- ¹ From a phenomenological model based on vector meson dominance with interfering $\rho(770)$, $\rho(1450)$, and $\rho(1700)$. The $\rho(1700)$ mass and width are fixed at 1720 MeV and 250 MeV, respectively. Systematic uncertainties not estimated. Supersedes ACHASOV 13.
- ² From a phenomenological model based on vector meson dominance with the interfering $\rho(1450)$ and $\rho(1700)$ and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.

 $\rho(1450)$ REFERENCES

YOUR PAPER

| | | | | |
|------------|-----|---------------|-------------------------------|---------------------------------|
| ACHASOV | 16D | PR D94 112001 | M.N. Achasov <i>et al.</i> | (SND Collab.) |
| MATVIENKO | 15 | PR D92 012013 | D. Matvienko <i>et al.</i> | (BELLE Collab.) |
| ACHASOV | 13 | PR D88 054013 | M.N. Achasov <i>et al.</i> | (SND Collab.) |
| AKHMETSHIN | 03B | PL B562 173 | R.R. Akhmetshin <i>et al.</i> | (Novosibirsk CMD-2 Collab.) |
| ALEXANDER | 01B | PR D64 092001 | J.P. Alexander <i>et al.</i> | (CLEO Collab.) |
| EDWARDS | 00A | PR D61 072003 | K.W. Edwards <i>et al.</i> | (CLEO Collab.) |
| CLEGG | 94 | ZPHY C62 455 | A.B. Clegg, A. Donnachie | (LANC, MCHS) |
| BISELLO | 91B | NPBPS B21 111 | D. Bisello | (DM2 Collab.) |
| ALBRECHT | 87L | PL B185 223 | H. Albrecht <i>et al.</i> | (ARGUS Collab.) |
| DOLINSKY | 86 | PL B174 453 | S.I. Dolinsky <i>et al.</i> | (NOVO) |
| ASTON | 80C | PL 92B 211 | D. Aston | (BONN, CERN, EPOL, GLAS, LANC+) |
| BARBER | 80C | ZPHY C4 169 | D.P. Barber <i>et al.</i> | (DARE, LANC, SHEF) |

 $\rho(1700)$

$$\Gamma^G(J^PC) = 1^+(1^{--})$$

 $\rho(1700) \Gamma(i)/\Gamma(\text{total}) \times \Gamma(e^+e^-)/\Gamma(\text{total})$

| $\Gamma(\pi^0\omega)/\Gamma_{\text{total}} \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ | | $\Gamma_{18}/\Gamma \times \Gamma_{17}/\Gamma$ | | |
|---|------|--|------|---------|
| VALUE (units 10^{-6}) | EVTS | DOCUMENT ID | TECN | COMMENT |

• • • We do not use the following data for averages, fits, limits, etc. • • •

YOUR DATA

| | | | | |
|-----------------|-------|----------------------|---------|---|
| 0.09 \pm 0.05 | 10.2k | ¹ ACHASOV | 16D SND | 1.05–2.00 $e^+e^- \rightarrow \pi^0\pi^0\gamma$ |
| 1.7 \pm 0.4 | 7815 | ² ACHASOV | 13 SND | 1.05–2.00 $e^+e^- \rightarrow \pi^0\pi^0\gamma$ |

YOUR NOTE

- ¹ From a phenomenological model based on vector meson dominance with interfering $\rho(700)$, $\rho(1450)$, and $\rho(1700)$. The $\rho(1700)$ mass and width are fixed at 1720 MeV and 250 MeV, respectively. Systematic uncertainty not estimated. Supersedes ACHASOV 13.
- ² From a phenomenological model based on vector meson dominance with the interfering $\rho(1450)$ and $\rho(1700)$ and their widths fixed at 400 and 250 MeV, respectively. Systematic uncertainty not estimated.

 $\rho(1700)$ REFERENCES

YOUR PAPER

| | | | | |
|---------|-----|---------------|----------------------------|---------------|
| ACHASOV | 16D | PR D94 112001 | M.N. Achasov <i>et al.</i> | (SND Collab.) |
| ACHASOV | 13 | PR D88 054013 | M.N. Achasov <i>et al.</i> | (SND Collab.) |

NODE=M105W3;LINKAGE=D

NODE=M105W3;LINKAGE=C

NODE=M105W3;LINKAGE=HK

NODE=M105W3;LINKAGE=3Z

NODE=M105W;LINKAGE=E1

NODE=M105W3;LINKAGE=B

NODE=M105W3;LINKAGE=A

NODE=M105230

NODE=M105R05

NODE=M105R05

OCCUR=3

NODE=M105R05;LINKAGE=A

NODE=M105R05;LINKAGE=AC

NODE=M105

REFID=57618

REFID=56601

REFID=55584

REFID=49406

REFID=48391

REFID=47465

REFID=44081

REFID=41752

REFID=40418

REFID=20246

REFID=20652

REFID=20653

NODE=M065

NODE=M065240

NODE=M065R01

NODE=M065R01

OCCUR=4

NODE=M065R01;LINKAGE=B

NODE=M065R01;LINKAGE=AC

NODE=M065

REFID=57618

REFID=55584