

$K^\pm$

$I(J^P) = \frac{1}{2}(0^-)$

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## THE CHARGED KAON MASS

Revised 1994 by T.G. Trippe, (LBNL).

The average of the six charged kaon mass measurements which we use in the Particle Listings is

$$m_{K^\pm} = 493.677 \pm 0.013 \text{ MeV (S = 2.4)} , \quad (1)$$

where the error has been increased by the scale factor S. The large scale factor indicates a serious disagreement between different input data. The average before scaling the error is

$$m_{K^\pm} = 493.677 \pm 0.005 \text{ MeV} ,$$

$$\chi^2 = 22.9 \text{ for 5 D.F., Prob.} = 0.04\% , \quad (2)$$

where the high  $\chi^2$  and correspondingly low  $\chi^2$  probability further quantify the disagreement.

The main disagreement is between the two most recent and precise results,

$$m_{K^\pm} = 493.696 \pm 0.007 \text{ MeV} \quad \text{DENISOV 91}$$

$$m_{K^\pm} = 493.636 \pm 0.011 \text{ MeV (S = 1.5)} \quad \text{GALL 88}$$

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$$\text{Average} = 493.679 \pm 0.006 \text{ MeV}$$

$$\chi^2 = 21.2 \text{ for 1 D.F., Prob.} = 0.0004\% , \quad (3)$$

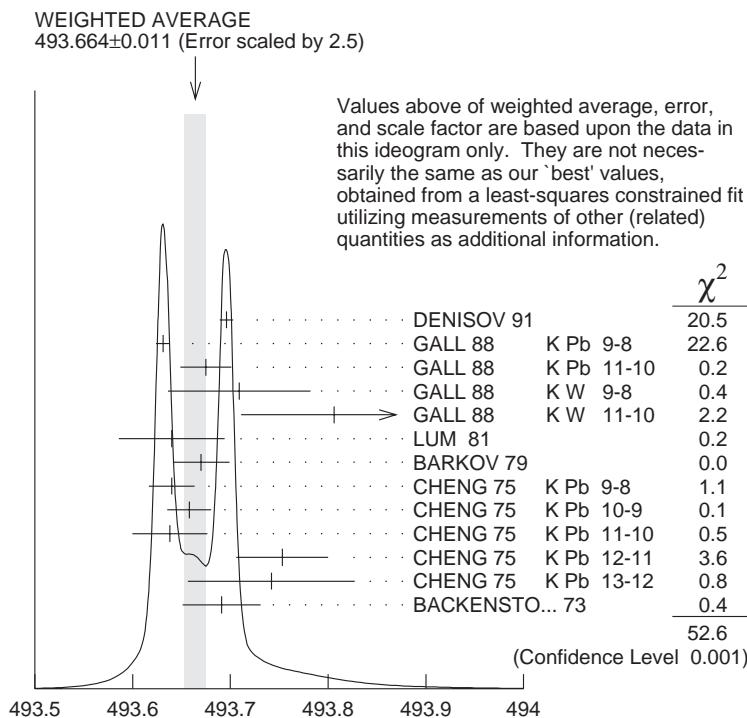
both of which are measurements of x-ray energies from kaonic atoms. Comparing the average in Eq. (3) with the overall average in Eq. (2), it is clear that DENISOV 91 and GALL 88 dominate the overall average, and that their disagreement is responsible for most of the high  $\chi^2$ .

The GALL 88 measurement was made using four different kaonic atom transitions,  $K^-$  Pb ( $9 \rightarrow 8$ ),  $K^-$  Pb ( $11 \rightarrow 10$ ),  $K^-$  W ( $9 \rightarrow 8$ ), and  $K^-$  W ( $11 \rightarrow 10$ ). The  $m_{K^\pm}$  values they obtain from each of these transitions is shown in the Particle Listings and in Fig. 1. Their  $K^-$  Pb ( $9 \rightarrow 8$ )  $m_{K^\pm}$  is below and somewhat inconsistent with their other three transitions. The average of their four measurements is

$$m_{K^\pm} = 493.636 \pm 0.007 ,$$

$$\chi^2 = 7.0 \text{ for 3 D.F., Prob. } = 7.2\% . \quad (4)$$

This is a low but acceptable  $\chi^2$  probability so, to be conservative, GALL 88 scaled up the error on their average by S=1.5 to obtain their published error  $\pm 0.011$  shown in Eq. (3) above and used in the Particle Listings average.



$m_{K^\pm}$  (MeV)

**Figure 1:** Ideogram of  $m_{K^\pm}$  mass measurements. GALL 88 and CHENG 75 measurements are shown separately for each transition they measured.

The ideogram in Fig. 1 shows that the DENISOV 91 measurement and the GALL 88  $K^-$  Pb (9 → 8) measurement yield two well-separated peaks. One might suspect the GALL 88  $K^-$  Pb (9 → 8) measurement since it is responsible both for the internal inconsistency in the GALL 88 measurements and the disagreement with DENISOV 91.

To see if the disagreement could result from a systematic problem with the  $K^-$  Pb ( $9 \rightarrow 8$ ) transition, we have separated the CHENG 75 data, which also used  $K^-$  Pb, into its separate transitions. Figure 1 shows that the CHENG 75 and GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) values are consistent, suggesting the possibility of a common effect such as contaminant nuclear  $\gamma$  rays near the  $K^-$  Pb ( $9 \rightarrow 8$ ) transition energy, although the CHENG 75 errors are too large to make a strong conclusion. The average of all 13 measurements has a  $\chi^2$  of 52.6 as shown in Fig. 1 and the first line of Table 1, yielding an unacceptable  $\chi^2$  probability of 0.00005%. The second line of Table 1 excludes both the GALL 88 and CHENG 75 measurements of the  $K^-$  Pb ( $9 \rightarrow 8$ ) transition and yields a  $\chi^2$  probability of 43%. The third [fourth] line of Table 1 excludes only the GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) [DENISOV 91] measurement and yields a  $\chi^2$  probability of 20% [8.6%]. Table 1 shows that removing both measurements of the  $K^-$  Pb ( $9 \rightarrow 8$ ) transition produces the most consistent set of data, but that excluding only the GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) transition or DENISOV 91 also produces acceptable probabilities.

**Table 1:**  $m_{K^\pm}$  averages for some combinations of Fig. 1 data.

$m_{K^\pm}$ (MeV)	$\chi^2$	D.F.	Prob. (%)	Measurements used
$493.664 \pm 0.004$	52.6	12	0.00005	all 13 measurements
$493.690 \pm 0.006$	10.1	10	43	no $K^-$ Pb( $9 \rightarrow 8$ )
$493.687 \pm 0.006$	14.6	11	20	no GALL 88 $K^-$ Pb( $9 \rightarrow 8$ )
$493.642 \pm 0.006$	17.8	11	8.6	no DENISOV 91

Yu.M. Ivanov, representing DENISOV 91, has estimated corrections needed for the older experiments because of improved  $^{192}\text{Ir}$  and  $^{198}\text{Au}$  calibration  $\gamma$ -ray energies. He estimates that CHENG 75 and BACKENSTOSS 73  $m_{K^\pm}$  values could be raised by about 15 keV and 22 keV, respectively. With these estimated corrections, Table 1 becomes Table 2. The last line of Table 2 shows that if such corrections are assumed, then GALL 88  $K^- \text{Pb} (9 \rightarrow 8)$  is inconsistent with the rest of the data even when DENISOV 91 is excluded. Yu.M. Ivanov warns that these are rough estimates. Accordingly, we do not use Table 2 to reject the GALL 88  $K^- \text{Pb} (9 \rightarrow 8)$  transition, but we note that a future reanalysis of the CHENG 75 data could be useful because it might provide supporting evidence for such a rejection.

**Table 2:**  $m_{K^\pm}$  averages for some combinations of Fig. 1 data after raising CHENG 75 and BACKENSTOSS 73 values by 0.015 and 0.022 MeV respectively.

$m_{K^\pm}$ (MeV)	$\chi^2$	D.F.	Prob. (%)	Measurements used
$493.666 \pm 0.004$	53.9	12	0.00003	all 13 measurements
$493.693 \pm 0.006$	9.0	10	53	no $K^- \text{Pb}(9 \rightarrow 8)$
$493.690 \pm 0.006$	11.5	11	40	no GALL 88 $K^- \text{Pb}(9 \rightarrow 8)$
$493.645 \pm 0.006$	23.0	11	1.8	no DENISOV 91

The GALL 88 measurement uses a Ge semiconductor spectrometer which has a resolution of about 1 keV, so they run the risk of some contaminant nuclear  $\gamma$  rays. Studies of  $\gamma$  rays following stopped  $\pi^-$  and  $\Sigma^-$  absorption in nucleii (unpublished) do not show any evidence for contaminants according to GALL 88 spokesperson, B.L. Roberts. The DENISOV 91 measurement uses a crystal diffraction spectrometer with a

resolution of 6.3 eV for radiation at 22.1 keV to measure the 4f-3d transition in  $K^-$   $^{12}\text{C}$ . The high resolution and the light nucleus reduce the probability for overlap by contaminant  $\gamma$  rays, compared with the measurement of GALL 88. The DENISOV 91 measurement is supported by their high-precision measurement of the 4d-2p transition energy in  $\pi^-$   $^{12}\text{C}$ , which is good agreement with the calculated energy.

While we suspect that the GALL 88  $K^-$  Pb ( $9 \rightarrow 8$ ) measurements could be the problem, we are unable to find clear grounds for rejecting it. Therefore, we retain their measurement in the average and accept the large scale factor until further information can be obtained from new measurements and/or from reanalysis of GALL 88 and CHENG 75 data.

We thank B.L. Roberts (Boston Univ.) and Yu.M. Ivanov (Petersburg Nuclear Physics Inst.) for their extensive help in understanding this problem.

## $K^\pm$ MASS

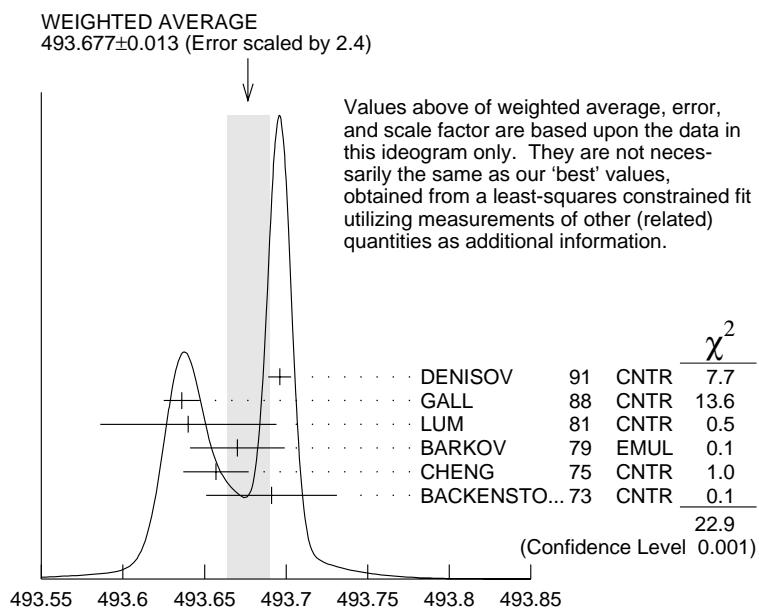
VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>493.677±0.016 OUR FIT</b>	Error includes scale factor of 2.8.			
<b>493.677±0.013 OUR AVERAGE</b>	Error includes scale factor of 2.4. See the ideogram below.			
493.696±0.007	<sup>1</sup> DENISOV	91	CNTR	— Kaonic atoms
493.636±0.011	<sup>2</sup> GALL	88	CNTR	— Kaonic atoms
493.640±0.054	LUM	81	CNTR	— Kaonic atoms
493.670±0.029	BARKOV	79	EMUL	$\pm e^+ e^- \rightarrow K^+ K^-$
493.657±0.020	<sup>2</sup> CHENG	75	CNTR	— Kaonic atoms
493.691±0.040	BACKENSTO...73	CNTR	—	Kaonic atoms
• • • We do not use the following data for averages, fits, limits, etc. • • •				
493.631±0.007	GALL	88	CNTR	$K^-$ Pb ( $9 \rightarrow 8$ )
493.675±0.026	GALL	88	CNTR	$K^-$ Pb ( $11 \rightarrow 10$ )
493.709±0.073	GALL	88	CNTR	$K^-$ W ( $9 \rightarrow 8$ )
493.806±0.095	GALL	88	CNTR	$K^-$ W ( $11 \rightarrow 10$ )
493.640±0.022±0.008	<sup>3</sup> CHENG	75	CNTR	$K^-$ Pb ( $9 \rightarrow 8$ )
493.658±0.019±0.012	<sup>3</sup> CHENG	75	CNTR	$K^-$ Pb ( $10 \rightarrow 9$ )
493.638±0.035±0.016	<sup>3</sup> CHENG	75	CNTR	$K^-$ Pb ( $11 \rightarrow 10$ )

493.753±0.042±0.021	<sup>3</sup> CHENG	75	CNTR	—	$K^-$ Pb (12→11)
493.742±0.081±0.027	<sup>3</sup> CHENG	75	CNTR	—	$K^-$ Pb (13→12)
493.662±0.19	KUNSELMAN	74	CNTR	—	Kaonic atoms
493.78 ±0.17	GREINER	65	EMUL	+	
493.7 ±0.3	BARKAS	63	EMUL	—	
493.9 ±0.2	COHEN	57	RVUE	+	

<sup>1</sup> Error increased from 0.0059 based on the error analysis in IVANOV 92.

<sup>2</sup> This value is the authors' combination of all of the separate transitions listed for this paper.

<sup>3</sup> The CHENG 75 values for separate transitions were calculated from their Table 7 transition energies. The first error includes a 20% systematic error in the noncircular contaminant shift. The second error is due to a ±5 eV uncertainty in the theoretical transition energies.



$m_{K^\pm}$  (MeV)

$m_{K^+} - m_{K^-}$

Test of CPT.

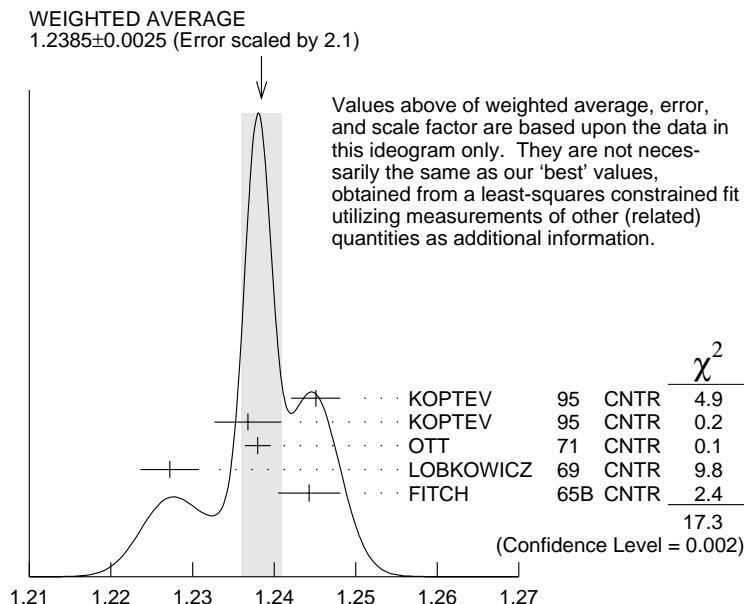
VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG
<b>-0.032±0.090</b>	1.5M	<sup>4</sup> FORD	72	ASPK ±

<sup>4</sup> FORD 72 uses  $m_{\pi^+} - m_{\pi^-} = +28 \pm 70$  keV.

**$K^\pm$  MEAN LIFE**

VALUE ( $10^{-8}$ s)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.2386±0.0024 OUR FIT</b>		Error includes scale factor of 2.0.			
<b>1.2385±0.0025 OUR AVERAGE</b>		Error includes scale factor of 2.1. See the ideogram below.			
1.2451±0.0030	250k	KOPTEV	95	CNTR	$K$ at rest, U target
1.2368±0.0041	150k	KOPTEV	95	CNTR	$K$ at rest, Cu target
1.2380±0.0016	3M	OTT	71	CNTR +	$K$ at rest
1.2272±0.0036		LOBKOWICZ	69	CNTR +	$K$ in flight
1.2443±0.0038		FITCH	65B	CNTR +	$K$ at rest
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
1.2415±0.0024	400k	<sup>5</sup> KOPTEV	95	CNTR	$K$ at rest
1.221 ±0.011		FORD	67	CNTR	±
1.231 ±0.011		BOYARSKI	62	CNTR	+
1.25 $\begin{array}{l} +0.22 \\ -0.17 \end{array}$		BARKAS	61	EMUL	
1.27 $\begin{array}{l} +0.36 \\ -0.23 \end{array}$	51	BHOWMIK	61	EMUL	
1.31 ±0.08	293	NORDIN	61	HBC	—
1.24 ±0.07		NORDIN	61	RVUE	—
1.38 ±0.24	33	FREDEN	60B	EMUL	
1.21 ±0.06		BURROWES	59	CNTR	
1.60 ±0.3	52	EISENBERG	58	EMUL	
0.95 $\begin{array}{l} +0.36 \\ -0.25 \end{array}$		ILOFF	56	EMUL	

<sup>5</sup> KOPTEV 95 report this weighted average of their U-target and Cu-target results, where they have weighted by  $1/\sigma$  rather than  $1/\sigma^2$ .



$K^\pm$  mean life ( $10^{-8}$  s)

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$$(\tau_{K^+} - \tau_{K^-}) / \tau_{\text{average}}$$

This quantity is a measure of *CPT* invariance in weak interactions.

VALUE (%)	DOCUMENT ID	TECN
<b>0.11 ±0.09 OUR AVERAGE</b>	Error includes scale factor of 1.2.	
0.090±0.078	LOBKOWICZ	69 CNTR
0.47 ± 0.30	FORD	67 CNTR

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## RARE KAON DECAYS

(Revised April 2000 by L. Littenberg, BNL and G. Valencia, Iowa State University)

**A. Introduction:** There are several useful reviews on rare kaon decays and related topics [1–11]. The current activity in rare kaon decays can be divided roughly into four categories:

1. Searches for explicit violations of the Standard Model
2. Measurements of Standard Model parameters
3. Searches for *CP* violation
4. Studies of strong interactions at low energy.

The paradigm of Category 1 is the lepton flavor violating decay  $K_L \rightarrow \mu e$ . Category 2 includes processes such as  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ , which is sensitive to  $|V_{td}|$ . Much of the interest in Category 3 is focussed on the decays  $K_L \rightarrow \pi^0 \ell \bar{\ell}$ , where  $\ell \equiv e, \mu, \nu$ . Category 4 includes reactions like  $K^+ \rightarrow \pi^+ \ell^+ \ell^-$  which constitute a testing ground for the ideas of chiral perturbation theory. Other reactions of this type are  $K_L \rightarrow \pi^0 \gamma \gamma$ , which also scales a  $CP$ -conserving background to  $CP$  violation in  $K_L \rightarrow \pi^0 \ell^+ \ell^-$  and  $K_L \rightarrow \gamma \ell^+ \ell^-$ , which could possibly shed light on long distance contributions to  $K_L \rightarrow \mu^+ \mu^-$ .

**B. Explicit violations of the Standard Model:** Most of the activity here is in searches for lepton flavor violation (LFV). This is motivated by the fact that many extensions of the minimal Standard Model violate lepton flavor and by the potential to access very high energy scales. For example, the tree-level exchange of a LFV vector boson of mass  $M_X$  that couples to left-handed fermions with electroweak strength and without mixing angles yields  $B(K_L \rightarrow \mu e) = 4.7 \times 10^{-12} (148 \text{ TeV}/M_X)^4$  [5]. This simple dimensional analysis may be used to read from Table 1 that the reaction  $K_L \rightarrow \mu e$  is already probing scales of over 100 TeV. Table 1 summarizes the present experimental situation vis a vis LFV, along with the expected near-future progress. The decays  $K_L \rightarrow \mu^\pm e^\mp$  and  $K^+ \rightarrow \pi^+ e^\mp \mu^\pm$  (or  $K_L \rightarrow \pi^0 e^\mp \mu^\pm$ ) provide complementary information on potential family number violating interactions since the former is sensitive to parity-odd couplings and the latter is sensitive to parity-even couplings. Related searches in  $\mu$  and  $\tau$  process are discussed in our section “Tests of Conservation Laws”.

**Table 1:** Searches for lepton flavor violation in  $K$  decay

Mode	90% CL upper limit	Exp't	Yr./Ref.	(Near-) future aim
$K^+ \rightarrow \pi^+ e\mu$	$4.8 \times 10^{-11}^*$	BNL-865	99/12	$9 \times 10^{-12}$ (BNL-865)
$K_L \rightarrow \mu e$	$4.7 \times 10^{-12}$	BNL-871	98/13	
$K_L \rightarrow \pi^0 e\mu$	$3.2 \times 10^{-9}$	FNAL-799	94/14	$5 \times 10^{-11}$ (KTeV)

\*preliminary

Another forbidden decay currently being pursued is  $K^+ \rightarrow \pi^+ X^0$ , where  $X^0$  is a very light, noninteracting particle (*e.g.* hyperphoton, axion, familon, *etc.*). The 90% CL upper limit on this process was recently improved to  $1.1 \times 10^{-10}$  [15]. Data already collected by BNL-787 are expected to yield a further factor  $\sim 2$  in sensitivity to this process.

**C. Measurements of Standard Model parameters:** Until 1997, searches for  $K^+ \rightarrow \pi^+ \nu\bar{\nu}$  were motivated by the possibility of observing non-SM physics because the sensitivity attained was far short of the SM prediction for this decay [16] and long-distance contributions were known to be negligible [2]. However, BNL-787 has attained the sensitivity at which the observation of an event can no longer be unambiguously attributed to non-SM physics. In 1997 BNL-787 observed a single candidate event and has recently released the results of further running in which no further events were seen, yielding a branching ratio of  $(1.5^{+3.4}_{-1.2}) \times 10^{-10}$  [15]. Further data already collected are expected to increase the sensitivity by approximately a factor 2, and there are plans for an upgrade to the experiment to collect roughly an order of magnitude more sensitivity [17]. This reaction is now interesting from the point of view of constraining

SM parameters. The branching ratio can be written in terms of the very well-measured rate of  $K_{e3}$  as [2]:

$$\begin{aligned} \text{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) &= \frac{\alpha^2 \text{B}(K^+ \rightarrow \pi^0 e^+ \nu)}{V_{us}^2 2\pi^2 \sin^4 \theta_W} \\ &\times \sum_{l=e,\mu,\tau} |V_{cs}^* V_{cd} X_{NL}^\ell + V_{ts}^* V_{td} X(m_t)|^2 \quad (1) \end{aligned}$$

to eliminate the *a priori* unknown hadronic matrix element. Isospin breaking corrections to the ratio of matrix elements reduce this rate by 10% [18]. In Eq. (1) the Inami-Lim function  $X(m_t)$  is of order 1 [19], and  $X_{NL}^\ell$  is several hundred times smaller. This form exhibits the strong dependence of this branching ratio on  $|V_{td}|$ . QCD corrections, which are contained in  $X_{NL}^\ell$ , are relatively small and now known [10] to  $\leq 10\%$ . Evaluating the constants in Eq. (1) with  $m_t = 175$  GeV, one can cast this result in terms of the CKM parameters  $A$ ,  $\rho$  and  $\eta$  (see our Section on “The Cabibbo-Kobayashi-Maskawa mixing matrix”) [10]

$$\text{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) \approx 1.0 \times 10^{-10} A^4 [\eta^2 + (\rho_o - \rho)^2] \quad (2)$$

where  $\rho_o \equiv 1 + (\frac{2}{3}X_{NL}^e + \frac{1}{3}X_{NL}^\tau)/(A^2 V_{us}^4 X(m_t)) \approx 1.4$ . Thus,  $\text{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$  determines a circle in the  $\rho$ ,  $\eta$  plane with center  $(\rho_o, 0)$  and radius  $\approx \frac{1}{A^2} \sqrt{\frac{\text{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})}{1.0 \times 10^{-10}}}$ .

The decay  $K_L \rightarrow \mu^+ \mu^-$  also has a short distance contribution sensitive to the CKM parameter  $\rho$ . For  $m_t = 175$  GeV it is given by [10]:

$$\text{B}_{\text{SD}}(K_L \rightarrow \mu^+ \mu^-) \approx 1.7 \times 10^{-9} A^4 (\rho'_o - \rho)^2 \quad (3)$$

where  $\rho'_o$  depends on the charm quark mass and is around 1.2. This decay, however, is dominated by a long-distance

contribution from a two-photon intermediate state. The absorptive (imaginary) part of the long-distance component is calculated in terms of the measured rate for  $K_L \rightarrow \gamma\gamma$  to be  $B_{\text{abs}}(K_L \rightarrow \mu^+\mu^-) = (7.07 \pm 0.18) \times 10^{-9}$ ; and it almost completely saturates the observed rate  $B(K_L \rightarrow \mu^+\mu^-) = (7.18 \pm 0.17) \times 10^{-9}$  [20]. The difference between the observed rate and the absorptive component can be attributed to the (coherent) sum of the short-distance amplitude and the real part of the long-distance amplitude. In order to use this mode to constrain  $\rho$  it is, therefore, necessary to know the real part of the long-distance contribution. Unlike the absorptive part, the real part of the long-distance contribution cannot be derived from the measured rate for  $K_L \rightarrow \gamma\gamma$ . At present, it is not possible to compute this long-distance component reliably and, therefore, it is not possible to constrain  $\rho$  from this mode in a model independent way [21]. Several models exist to estimate this long-distance component [22,23] that are sufficient to place rough bounds on new physics from the measured rate for  $K_L \rightarrow \mu^+\mu^-$  [24]. The decay  $K_L \rightarrow e^+e^-$  is completely dominated by long distance physics and is easier to estimate. The result,  $B(K_L \rightarrow e^+e^-) \sim 9 \times 10^{-12}$  [21,23], is in good agreement with the recent measurement [25]. It is expected that studies of the reactions  $K_L \rightarrow \ell^+\ell^-\gamma$ , and  $K_L \rightarrow \ell^+\ell^-\ell'^+\ell'^-$  for  $\ell, \ell' = e$  or  $\mu$ , currently under active study by the KTeV and NA48 experiments, will improve our understanding of the long distance effects in  $K_L \rightarrow \mu^+\mu^-$  (the current data is parameterized in terms of  $\alpha_K^*$ , discussed on page 25 of the  $K_L^0$  Particle Properties Listing in our 1999 WWW update).

**D. Searches for direct CP violation:** The mode  $K_L \rightarrow \pi^0\nu\bar{\nu}$  is dominantly  $CP$ -violating and free of hadronic uncertainties [2,26]. The Standard Model predicts a branching ratio

$(3.0 \pm 1.3) \times 10^{-11}$ ; for  $m_t = 175$  GeV it is given approximately by [10]:

$$B(K_L \rightarrow \pi^0 \nu \bar{\nu}) \approx 4.1 \times 10^{-10} A^4 \eta^2 . \quad (4)$$

The current upper bound is  $B(K_L \rightarrow \pi^0 \nu \bar{\nu}) \leq 5.9 \times 10^{-7}$  [27] and KTeV (FNAL799II) is expected to place a bound of order  $10^{-8}$  [28]. The 90% CL bound on  $K^+ \rightarrow \pi^+ \nu \bar{\nu}$  provides a nearly model independent bound  $B(K_L \rightarrow \pi^0 \nu \bar{\nu}) < 3 \times 10^{-9}$  [29]. A KEK experiment to reach the  $10^{-10}$ /event level is in preparation [30]. The BNL-926 [31] proposal aims to make a  $\sim 15\%$  measurement of  $B(K_L \rightarrow \pi^0 \nu \bar{\nu})$ . There is also a Fermilab EOI [32] with comparable goals.

There has been much recent theoretical work on possible contributions to  $\epsilon'/\epsilon$  and rare  $K$  decays within a generic supersymmetric extension of the Standard Model with  $R$  parity conservation and minimal particle content [24,33]. These conclude that contributions to rare decays much larger than those of the Standard Model are possible without violating current phenomenological constraints.

The decay  $K_L \rightarrow \pi^0 e^+ e^-$  also has sensitivity to the product  $A^4 \eta^2$ . It has a direct  $CP$ -violating component that for  $m_t = 175$  GeV is given by [10]:

$$B_{\text{dir}}(K_L \rightarrow \pi^0 e^+ e^-) \approx 6.7 \times 10^{-11} A^4 \eta^2 . \quad (5)$$

However, like  $K_L \rightarrow \mu^+ \mu^-$  this mode suffers from large theoretical uncertainties due to long distance strong interaction effects. It has an indirect  $CP$ -violating component given by:

$$B_{\text{ind}}(K_L \rightarrow \pi^0 e^+ e^-) = |\epsilon|^2 \frac{\tau_{K_L}}{\tau_{K_S}} B(K_S \rightarrow \pi^0 e^+ e^-) , \quad (6)$$

that has been estimated to be less than  $10^{-12}$  [34], but that will not be known precisely until a measurement of  $K_S \rightarrow \pi^0 e^+ e^-$

is available [4,35]. There is also a  $CP$ -conserving component dominated by a two-photon intermediate state that cannot be computed reliably at present. This component has an absorptive part that can be, in principle, determined from a detailed analysis of  $K_L \rightarrow \pi^0 \gamma\gamma$ .

To understand the rate and the shape of the distribution  $d\Gamma/dm_{\gamma\gamma}$  in  $K_L \rightarrow \pi^0 \gamma\gamma$  within chiral perturbation theory it is necessary to go beyond leading order. The measured rate and spectrum can be accommodated naturally, for example, by allowing only one of the free parameters that occur,  $a_V$ , to vary [36]. There is new data on this decay from KTeV [37] and a fit to the distribution has given  $a_V = -0.72 \pm 0.05 \pm 0.06$ . This value suggests that the absorptive part of the  $CP$ -conserving contribution to  $K_L \rightarrow \pi^0 e^+ e^-$  could be comparable to the direct  $CP$ -violating component [37,35]. The related process,  $K_L \rightarrow \pi^0 \gamma e^+ e^-$ , is potentially an additional background in some region of phase space [38]. This process has recently been observed with a branching ratio of  $(2.42 \pm 0.38_{stat} \pm 0.11_{sys}) \times 10^{-8}$  [39]. Finally, BNL-845 observed a potential background to  $K_L \rightarrow \pi^0 e^+ e^-$  from the decay  $K_L \rightarrow \gamma\gamma e^+ e^-$  [40]. This has recently been confirmed with a 500-fold larger sample by FNAL-799 [41], which measured additional kinematic quantities. It has been estimated that this background will enter at the level of  $10^{-11}$  [42], comparable to the signal level. Because of this, the observation of  $K_L \rightarrow \pi^0 e^+ e^-$  will depend on background subtraction with good statistics.

The current 90% CL preliminary upper bound for the process  $K_L \rightarrow \pi^0 e^+ e^-$  is  $5.64 \times 10^{-10}$  [41]. For the closely related muonic process, the corresponding upper bound is  $B(K_L \rightarrow \pi^0 \mu^+ \mu^-) \leq 3.8 \times 10^{-10}$  [43]. KTeV expects to reach a sensitivity of roughly  $10^{-11}$  for both reactions [28].

### **E. Other long distance dominated modes:**

The decays  $K^+ \rightarrow \pi^+ \ell^+ \ell^-$  ( $\ell = e$  or  $\mu$ ) are described by leading order chiral perturbation theory in terms of one parameter,  $\omega^+$  [44]. It now appears that this parameterization is not sufficient to account for both the rate and the detailed shape of the spectrum in  $K^+ \rightarrow \pi^+ e^+ e^-$  [45]. An analysis beyond leading order in chiral perturbation theory can accommodate both the rate and the spectrum [46], at the cost of introducing at least one new parameter.

## **References**

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## **$K^+$ DECAY MODES**

$K^-$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1 \mu^+ \nu_\mu$	( $63.51 \pm 0.18$ ) %	S=1.3
$\Gamma_2 e^+ \nu_e$	( $1.55 \pm 0.07$ ) $\times 10^{-5}$	
$\Gamma_3 \pi^+ \pi^0$	( $21.16 \pm 0.14$ ) %	S=1.1
$\Gamma_4 \pi^+ \pi^+ \pi^-$	( $5.59 \pm 0.05$ ) %	S=1.8
$\Gamma_5 \pi^+ \pi^0 \pi^0$	( $1.73 \pm 0.04$ ) %	S=1.2
$\Gamma_6 \pi^0 \mu^+ \nu_\mu$	( $3.18 \pm 0.08$ ) %	S=1.5
Called $K_{\mu 3}^+$ .		
$\Gamma_7 \pi^0 e^+ \nu_e$	( $4.82 \pm 0.06$ ) %	S=1.3
Called $K_{e 3}^+$ .		
$\Gamma_8 \pi^0 \pi^0 e^+ \nu_e$	( $2.1 \pm 0.4$ ) $\times 10^{-5}$	
$\Gamma_9 \pi^+ \pi^- e^+ \nu_e$	( $3.91 \pm 0.17$ ) $\times 10^{-5}$	
$\Gamma_{10} \pi^+ \pi^- \mu^+ \nu_\mu$	( $1.4 \pm 0.9$ ) $\times 10^{-5}$	
$\Gamma_{11} \pi^0 \pi^0 \pi^0 e^+ \nu_e$	< 3.5 $\times 10^{-6}$	CL=90%
$\Gamma_{12} \mu^+ \nu_\mu \nu \bar{\nu}$	< 6.0 $\times 10^{-6}$	CL=90%
$\Gamma_{13} e^+ \nu_e \nu \bar{\nu}$	< 6 $\times 10^{-5}$	CL=90%
$\Gamma_{14} \mu^+ \nu_\mu e^+ e^-$	( $1.3 \pm 0.4$ ) $\times 10^{-7}$	
$\Gamma_{15} e^+ \nu_e e^+ e^-$	( $3.0 \pm 3.0$ ) $\times 10^{-8}$	
$\Gamma_{16} e^+ \nu_e \mu^+ \mu^-$	< 5 $\times 10^{-7}$	CL=90%
$\Gamma_{17} \mu^+ \nu_\mu \mu^+ \mu^-$	< 4.1 $\times 10^{-7}$	CL=90%
$\Gamma_{18} \mu^+ \nu_\mu \gamma$	[a,b] ( $5.50 \pm 0.28$ ) $\times 10^{-3}$	
$\Gamma_{19} \pi^+ \pi^0 \gamma$	[a,b] ( $2.75 \pm 0.15$ ) $\times 10^{-4}$	
$\Gamma_{20} \pi^+ \pi^0 \gamma$ (DE)	[b,c] ( $1.8 \pm 0.4$ ) $\times 10^{-5}$	
$\Gamma_{21} \pi^+ \pi^+ \pi^- \gamma$	[a,b] ( $1.04 \pm 0.31$ ) $\times 10^{-4}$	
$\Gamma_{22} \pi^+ \pi^0 \pi^0 \gamma$	[a,b] ( $7.5 \pm 5.5$ ) $\times 10^{-6}$	
$\Gamma_{23} \pi^0 \mu^+ \nu_\mu \gamma$	[a,b] < 6.1 $\times 10^{-5}$	CL=90%
$\Gamma_{24} \pi^0 e^+ \nu_e \gamma$	[a,b] ( $2.62 \pm 0.20$ ) $\times 10^{-4}$	
$\Gamma_{25} \pi^0 e^+ \nu_e \gamma$ (SD)	[d] < 5.3 $\times 10^{-5}$	CL=90%
$\Gamma_{26} \pi^0 \pi^0 e^+ \nu_e \gamma$	< 5 $\times 10^{-6}$	CL=90%
$\Gamma_{27} \pi^+ \gamma \gamma$	[b] ( $1.10 \pm 0.32$ ) $\times 10^{-6}$	
$\Gamma_{28} \pi^+ 3\gamma$	[b] < 1.0 $\times 10^{-4}$	CL=90%

**Lepton Family number (*LF*), Lepton number (*L*),  $\Delta S = \Delta Q$  (*SQ*)  
violating modes, or  $\Delta S = 1$  weak neutral current (*S1*) modes**

$\Gamma_{29}$	$\pi^+ \pi^+ e^- \bar{\nu}_e$	<i>SQ</i>	$< 1.2 \times 10^{-8}$	CL=90%
$\Gamma_{30}$	$\pi^+ \pi^+ \mu^- \bar{\nu}_\mu$	<i>SQ</i>	$< 3.0 \times 10^{-6}$	CL=95%
$\Gamma_{31}$	$\pi^+ e^+ e^-$	<i>S1</i>	$(2.88 \pm 0.13) \times 10^{-7}$	
$\Gamma_{32}$	$\pi^+ \mu^+ \mu^-$	<i>S1</i>	$(7.6 \pm 2.1) \times 10^{-8}$	$S=3.4$
$\Gamma_{33}$	$\pi^+ \nu \bar{\nu}$	<i>S1</i>	$(1.5^{+3.4}_{-1.2}) \times 10^{-10}$	
$\Gamma_{34}$	$\mu^- \nu e^+ e^+$	<i>LF</i>	$< 2.0 \times 10^{-8}$	CL=90%
$\Gamma_{35}$	$\mu^+ \nu_e$	<i>LF</i>	$[e] < 4 \times 10^{-3}$	CL=90%
$\Gamma_{36}$	$\pi^+ \mu^+ e^-$	<i>LF</i>	$< 2.1 \times 10^{-10}$	CL=90%
$\Gamma_{37}$	$\pi^+ \mu^- e^+$	<i>LF</i>	$< 7 \times 10^{-9}$	CL=90%
$\Gamma_{38}$	$\pi^- \mu^+ e^+$	<i>L</i>	$< 7 \times 10^{-9}$	CL=90%
$\Gamma_{39}$	$\pi^- e^+ e^+$	<i>L</i>	$< 1.0 \times 10^{-8}$	CL=90%
$\Gamma_{40}$	$\pi^- \mu^+ \mu^+$	<i>L</i>	$[e] < 1.5 \times 10^{-4}$	CL=90%
$\Gamma_{41}$	$\mu^+ \bar{\nu}_e$	<i>L</i>	$[e] < 3.3 \times 10^{-3}$	CL=90%
$\Gamma_{42}$	$\pi^0 e^+ \bar{\nu}_e$	<i>L</i>	$< 3 \times 10^{-3}$	CL=90%
$\Gamma_{43}$	$\pi^+ \gamma$	[f]		

[a] Most of this radiative mode, the low-momentum  $\gamma$  part, is also included in the parent mode listed without  $\gamma$ 's.

[b] See the Particle Listings below for the energy limits used in this measurement.

[c] Direct-emission branching fraction.

[d] Structure-dependent part.

[e] Derived from an analysis of neutrino-oscillation experiments.

[f] Violates angular-momentum conservation.

## CONSTRAINED FIT INFORMATION

An overall fit to the mean life, 2 decay rate, and 20 branching ratios uses 60 measurements and one constraint to determine 8 parameters. The overall fit has a  $\chi^2 = 78.1$  for 53 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \cdot \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	-58						
$x_4$	-41	-12					
$x_5$	-27	-4	21				
$x_6$	-48	-17	14	2			
$x_7$	-50	-16	34	6	39		
$x_8$	-3	-1	2	0	2	6	
$\Gamma$	7	2	-18	-4	-2	-6	0
	$x_1$	$x_3$	$x_4$	$x_5$	$x_6$	$x_7$	$x_8$

	Mode	Rate ( $10^8 \text{ s}^{-1}$ )	Scale factor
$\Gamma_1$	$\mu^+ \nu_\mu$	$0.5128 \pm 0.0018$	1.5
$\Gamma_3$	$\pi^+ \pi^0$	$0.1708 \pm 0.0012$	1.1
$\Gamma_4$	$\pi^+ \pi^+ \pi^-$	$0.0452 \pm 0.0004$	1.8
$\Gamma_5$	$\pi^+ \pi^0 \pi^0$	$0.01399 \pm 0.00032$	1.2
$\Gamma_6$	$\pi^0 \mu^+ \nu_\mu$ Called $K_{\mu 3}^+$	$0.0257 \pm 0.0006$	1.5
$\Gamma_7$	$\pi^0 e^+ \nu_e$ Called $K_{e3}^+$	$0.0389 \pm 0.0005$	1.3
$\Gamma_8$	$\pi^0 \pi^0 e^+ \nu_e$	$(1.69 \quad +0.34 \quad -0.29) \times 10^{-5}$	

## $K^\pm$ DECAY RATES

$\Gamma(\mu^+ \nu_\mu)$	$\Gamma_1$
$\frac{\text{VALUE} (10^6 \text{ s}^{-1})}{\text{OUR FIT}}$	$\frac{\text{DOCUMENT ID}}{\text{FORD}}$
$51.28 \pm 0.18$	Error includes scale factor of 1.5.
$51.2 \pm 0.8$	$67 \text{ CNTR} \pm$

$\Gamma(\pi^+ \pi^+ \pi^-)$	$\Gamma_4$
$\frac{\text{VALUE} (10^6 \text{ s}^{-1})}{\text{OUR FIT}}$	$\frac{\text{EVTS}}{6}$
$4.52 \pm 0.04$	Error includes scale factor of 1.8.
$4.511 \pm 0.024$	$70 \text{ ASPK}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

4.529±0.032	3.2M	<sup>6</sup> FORD	70	ASPK
4.496±0.030		<sup>6</sup> FORD	67	CNTR ±

<sup>6</sup> First FORD 70 value is second FORD 70 combined with FORD 67.

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### ( $\Gamma(K^+) - \Gamma(K^-)$ ) / $\Gamma(K)$

#### $K^\pm \rightarrow \mu^\pm \nu_\mu$ RATE DIFFERENCE/AVERAGE

Test of  $CPT$  conservation.

VALUE (%)	DOCUMENT ID	TECN
-0.54±0.41	FORD	67 CNTR

#### $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$ RATE DIFFERENCE/AVERAGE

Test of  $CP$  conservation.

VALUE (%)	EVTS	DOCUMENT ID	TECN	CHG
<b>0.07±0.12 OUR AVERAGE</b>				

0.08±0.12		<sup>7</sup> FORD	70	ASPK
-0.50±0.90		FLETCHER	67	OSPK
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-0.02±0.16		<sup>8</sup> SMITH	73	ASPK ±
0.10±0.14	3.2M	<sup>7</sup> FORD	70	ASPK
-0.04±0.21		<sup>7</sup> FORD	67	CNTR

<sup>7</sup> First FORD 70 value is second FORD 70 combined with FORD 67.

<sup>8</sup> SMITH 73 value of  $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$  rate difference is derived from SMITH 73 value of  $K^\pm \rightarrow \pi^\pm 2\pi^0$  rate difference.

#### $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$ RATE DIFFERENCE/AVERAGE

Test of  $CP$  conservation.

VALUE (%)	EVTS	DOCUMENT ID	TECN	CHG
<b>0.0 ±0.6 OUR AVERAGE</b>				
0.08±0.58		SMITH	73	ASPK ±
-1.1 ±1.8	1802	HERZO	69	OSPK

#### $K^\pm \rightarrow \pi^\pm \pi^0$ RATE DIFFERENCE/AVERAGE

Test of  $CPT$  conservation.

VALUE (%)	DOCUMENT ID	TECN
<b>0.8±1.2</b>	HERZO	69 OSPK

#### $K^\pm \rightarrow \pi^\pm \pi^0 \gamma$ RATE DIFFERENCE/AVERAGE

Test of  $CP$  conservation.

VALUE (%)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.9± 3.3 OUR AVERAGE</b>					
0.8± 5.8	2461	SMITH	76	WIRE ±	$E_\pi$ 55–90 MeV
1.0± 4.0	4000	ABRAMS	73B	ASPK ±	$E_\pi$ 51–100 MeV
0.0±24.0	24	EDWARDS	72	OSPK	$E_\pi$ 58–90 MeV

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**$K^+$  BRANCHING RATIOS** $\Gamma(\mu^+ \nu_\mu)/\Gamma_{\text{total}}$  $\Gamma_1/\Gamma$ 

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b><math>63.51 \pm 0.18</math> OUR FIT</b>		Error includes scale factor of 1.3.			
<b><math>63.24 \pm 0.44</math></b>	62k	CHIANG	72	OSPK +	$1.84 \text{ GeV}/c K^+$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
56.9 $\pm 2.6$		<sup>9</sup> ALEXANDER	57	EMUL +	
58.5 $\pm 3.0$		<sup>9</sup> BIRGE	56	EMUL +	

<sup>9</sup> Old experiments not included in averaging. $\Gamma(\mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$  $\Gamma_1/\Gamma_4$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>
<b><math>11.35 \pm 0.12</math> OUR FIT</b>		Error includes scale factor of 1.8.		
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$10.38 \pm 0.82$	427	<sup>10</sup> YOUNG	65	EMUL +

<sup>10</sup> Deleted from overall fit because YOUNG 65 constrains his results to add up to 1. Only YOUNG 65 measured ( $\mu\nu$ ) directly.

 $\Gamma(e^+ \nu_e)/\Gamma_{\text{total}}$  $\Gamma_2/\Gamma$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
$2.1^{+1.8}_{-1.3}$	4	BOWEN	67B	OSPK +	
<160.0	95	BORREANI	64	HBC +	

 $\Gamma(e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$  $\Gamma_2/\Gamma_1$ 

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>
<b><math>2.45 \pm 0.11</math> OUR AVERAGE</b>				
2.51 $\pm 0.15$	404	HEINTZE	76	SPEC +
2.37 $\pm 0.17$	534	HEARD	75B	SPEC +
2.42 $\pm 0.42$	112	CLARK	72	OSPK +
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$1.8^{+0.8}_{-0.6}$	8	MACEK	69	ASPK +
1.9 $\pm 0.7$	10	BOTTERILL	67	ASPK +

 $\Gamma(\pi^+ \pi^0)/\Gamma_{\text{total}}$  $\Gamma_3/\Gamma$ 

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b><math>21.16 \pm 0.14</math> OUR FIT</b>		Error includes scale factor of 1.1.			
<b><math>21.18 \pm 0.28</math></b>	16k	CHIANG	72	OSPK +	$1.84 \text{ GeV}/c K^+$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$					
21.0 $\pm 0.6$		CALLAHAN	65	HLBC	See $\Gamma(\pi^+ \pi^0)/\Gamma(\pi^+ \pi^+ \pi^-)$
21.6 $\pm 0.6$		TRILLING	65B	RVUE	
23.2 $\pm 2.2$	<sup>11</sup> ALEXANDER	57	EMUL	+	
27.7 $\pm 2.7$	<sup>11</sup> BIRGE	56	EMUL	+	

<sup>11</sup> Earlier experiments not averaged.

$\Gamma(\pi^+\pi^0)/\Gamma(\mu^+\nu_\mu)$  $\Gamma_3/\Gamma_1$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>0.3331±0.0028 OUR FIT</b>		Error includes scale factor of 1.1.			

**0.3316±0.0032 OUR AVERAGE**

0.3329±0.0047±0.0010	45k	USHER	92	SPEC	+	$p\bar{p}$ at rest
0.3355±0.0057	<sup>12</sup>	WEISSENBE...	76	SPEC	+	
0.305 ±0.018	1600	ZELLER	69	ASPK	+	
0.3277±0.0065	4517	AUERBACH	67	OSPK	+	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
0.328 ±0.005	25k	<sup>12</sup> WEISSENBE...	74	STRC	+	

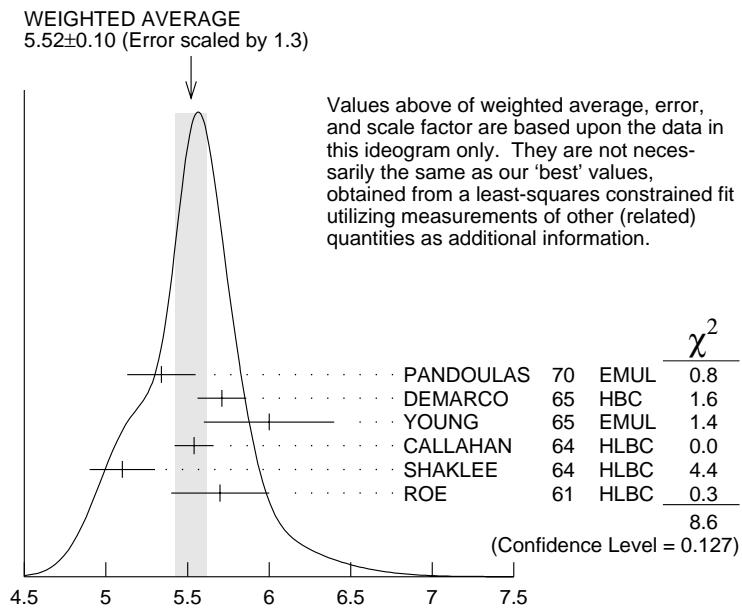
<sup>12</sup> WEISSENBERG 76 revises WEISSENBERG 74.<sup>13</sup> AUERBACH 67 changed from  $0.3253 \pm 0.0065$ . See comment with ratio  $\Gamma(\pi^0\mu^+\nu_\mu)/\Gamma(\mu^+\nu_\mu)$ . $\Gamma(\pi^+\pi^0)/\Gamma(\pi^+\pi^+\pi^-)$  $\Gamma_3/\Gamma_4$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	
<b>3.78±0.04 OUR FIT</b>		Error includes scale factor of 1.5.			
<b>3.84±0.27 OUR AVERAGE</b>		Error includes scale factor of 1.9.			
3.96±0.15	1045	CALLAHAN	66	FBC	+
3.24±0.34	134	YOUNG	65	EMUL	+

 $\Gamma(\pi^+\pi^+\pi^-)/\Gamma_{\text{total}}$  $\Gamma_4/\Gamma$ 

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>5.59±0.05 OUR FIT</b>		Error includes scale factor of 1.8.			
<b>5.52±0.10 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.			
5.34±0.21	693	<sup>14</sup> PANDOULAS	70	EMUL	+
5.71±0.15		DEMARCO	65	HBC	
6.0 ±0.4	44	YOUNG	65	EMUL	+
5.54±0.12	2332	CALLAHAN	64	HLBC	+
5.1 ±0.2	540	SHAKLEE	64	HLBC	+
5.7 ±0.3		ROE	61	HLBC	+
• • • We do not use the following data for averages, fits, limits, etc. • • •					
5.56±0.20	2330	<sup>15</sup> CHIANG	72	OSPK	+
5.2 ±0.3		<sup>16</sup> TAYLOR	59	EMUL	+
6.8 ±0.4		<sup>16</sup> ALEXANDER	57	EMUL	+
5.6 ±0.4		<sup>16</sup> BIRGE	56	EMUL	+

<sup>14</sup> Includes events of TAYLOR 59.<sup>15</sup> Value is not independent of CHIANG 72  $\Gamma(\mu^+\nu_\mu)/\Gamma_{\text{total}}$ ,  $\Gamma(\pi^+\pi^0)/\Gamma_{\text{total}}$ ,  $\Gamma(\pi^+\pi^0\pi^0)/\Gamma_{\text{total}}$ ,  $\Gamma(\pi^0\mu^+\nu_\mu)/\Gamma_{\text{total}}$ , and  $\Gamma(\pi^0e^+\nu_e)/\Gamma_{\text{total}}$ .<sup>16</sup> Earlier experiments not averaged.



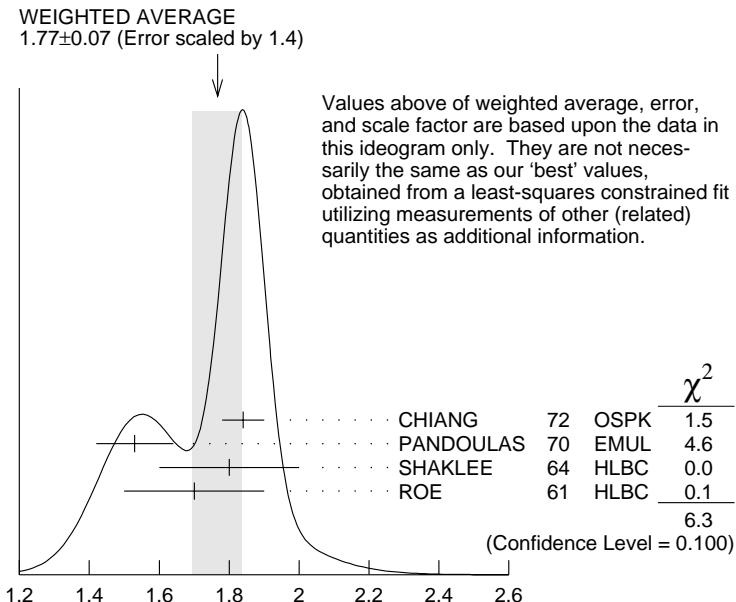
$$\Gamma(\pi^+ \pi^+ \pi^-)/\Gamma_{\text{total}} \text{ (units } 10^{-2})$$

### $\Gamma(\pi^+ \pi^0 \pi^0)/\Gamma_{\text{total}}$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>1.73 \pm 0.04</math> OUR FIT</b>		Error includes scale factor of 1.2.			
<b><math>1.77 \pm 0.07</math> OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.			
1.84 ± 0.06	1307	CHIANG	72	OSPK +	1.84 GeV/c $K^+$
1.53 ± 0.11	198	<sup>17</sup> PANDOULAS	70	EMUL +	
1.8 ± 0.2	108	SHAKLEE	64	HLBC +	
1.7 ± 0.2		ROE	61	HLBC +	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.5 ± 0.2		<sup>18</sup> TAYLOR	59	EMUL +	
2.2 ± 0.4		<sup>18</sup> ALEXANDER	57	EMUL +	
2.1 ± 0.5		<sup>18</sup> BIRGE	56	EMUL +	

<sup>17</sup> Includes events of TAYLOR 59.

<sup>18</sup> Earlier experiments not averaged.



$$\Gamma(\pi^+ \pi^0 \pi^0)/\Gamma_{\text{total}} \text{ (units } 10^{-2})$$

### $\Gamma(\pi^+ \pi^0 \pi^0)/\Gamma(\pi^+ \pi^0)$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_5/\Gamma_3$
<b>0.0819±0.0020 OUR FIT</b>		Error includes scale factor of 1.2.				
<b>0.081 ±0.005</b>	574	<sup>19</sup> LUCAS	73B	HBC	–	Dalitz pairs only
19 LUCAS 73B gives $N(\pi^+ \pi^0 \pi^0) = 574 \pm 5.9\%$ , $N(2\pi) = 3564 \pm 3.1\%$ . We quote $0.5N(\pi^+ \pi^0 \pi^0)/N(2\pi)$ where 0.5 is because only Dalitz pair $\pi^0$ 's were used.						

### $\Gamma(\pi^+ \pi^0 \pi^0)/\Gamma(\pi^+ \pi^+ \pi^-)$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_5/\Gamma_4$
<b>0.310±0.007 OUR FIT</b>		Error includes scale factor of 1.2.				
<b>0.304±0.009 OUR AVERAGE</b>						
0.303±0.009	2027	BISI	65	BC	+	HBC+HLBC
0.393±0.099	17	YOUNG	65	EMUL	+	

### $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma_{\text{total}}$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	$\Gamma_6/\Gamma$
<b>3.18±0.08 OUR FIT</b>		Error includes scale factor of 1.5.				
<b>3.33±0.16</b>	2345	CHIANG	72	OSPK	+	$1.84 \text{ GeV}/c K^+$
• • • We do not use the following data for averages, fits, limits, etc. • • •						
2.8 ±0.4	20	TAYLOR	59	EMUL	+	
5.9 ±1.3	20	ALEXANDER	57	EMUL	+	
2.8 ±1.0	20	BIRGE	56	EMUL	+	

20 Earlier experiments not averaged.

### $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\mu^+ \nu_\mu)$

### $\Gamma_6/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.0501±0.0013 OUR FIT</b>	Error includes scale factor of 1.5.			
<b>0.0488±0.0026 OUR AVERAGE</b>				
0.054 ± 0.009	240	ZELLER	69	ASPK +
0.0480±0.0037	424	21 GARLAND	68	OSPK +
0.0486±0.0040	307	22 AUERBACH	67	OSPK +

21 GARLAND 68 changed from  $0.055 \pm 0.004$  in agreement with  $\mu$ -spectrum calculation of GAILLARD 70 appendix B. L.G.Pondrom, (private communication 73).

22 AUERBACH 67 changed from  $0.0602 \pm 0.0046$  by erratum which brings the  $\mu$ -spectrum calculation into agreement with GAILLARD 70 appendix B.

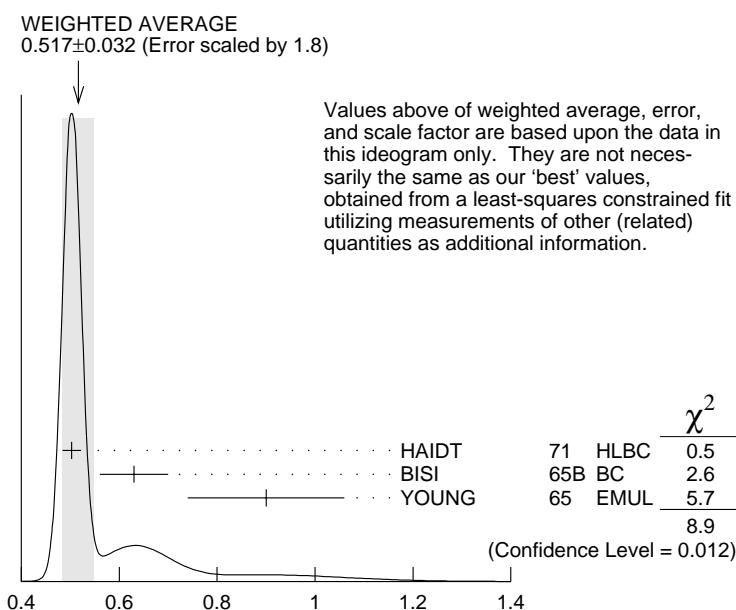
### $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$

### $\Gamma_6/\Gamma_4$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.569±0.014 OUR FIT</b>	Error includes scale factor of 1.5.				
<b>0.517±0.032 OUR AVERAGE</b>	Error includes scale factor of 1.8. See the ideogram below.				
0.503±0.019	1505	23 HAIDT	71	HLBC +	
0.63 ± 0.07	2845	24 BISI	65B	BC +	HBC+HLBC
0.90 ± 0.16	38	YOUNG	65	EMUL +	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.510±0.017	1505	23 EICHTEN	68	HLBC +	

23 HAIDT 71 is a reanalysis of EICHTEN 68.

24 Error enlarged for background problems. See GAILLARD 70.



### $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$

$\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^0 e^+ \nu_e)$   $\Gamma_6/\Gamma_7$ 

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.660±0.015 OUR FIT</b>		Error includes scale factor of 1.5.			
<b>0.680±0.013 OUR AVERAGE</b>					
0.705±0.063	554	25 LUCAS	73B HBC	—	Dalitz pairs only
0.698±0.025	3480	26 CHIANG	72 OSPK	+	1.84 GeV/c $K^+$
0.667±0.017	5601	BOTTERILL	68B ASPK	+	
0.703±0.056	1509	27 CALLAHAN	66B HLBC		
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
0.670±0.014		28 HEINTZE	77 SPEC	+	
0.67 ±0.12		WEISSENBE...	76 SPEC	+	
0.608±0.014	1585	29 BRAUN	75 HLBC	+	
0.596±0.025		30 HAIDT	71 HLBC	+	
0.604±0.022	1398	30 EICHTEN	68 HLBC		

25 LUCAS 73B gives  $N(K_{\mu 3}) = 554 \pm 7.6\%$ ,  $N(K_{e3}) = 786 \pm 3.1\%$ . We divide.

26 CHIANG 72  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^0 e^+ \nu_e)$  is statistically independent of CHIANG 72  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma_{\text{total}}$  and  $\Gamma(\pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$ .

27 From CALLAHAN 66B we use only the  $K_{\mu 3}/K_{e3}$  ratio and do not include in the fit the ratios  $K_{\mu 3}/(\pi\pi^+\pi^0)$  and  $K_{e3}/(\pi\pi^+\pi^0)$ , since they show large disagreements with the rest of the data.

28 HEINTZE 77 value from fit to  $\lambda_0$ . Assumes  $\mu$ - $e$  universality.

29 BRAUN 75 value is from form factor fit. Assumes  $\mu$ - $e$  universality.

30 HAIDT 71 is a reanalysis of EICHTEN 68. Only individual ratios included in fit (see  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$  and  $\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\pi^+ \pi^+ \pi^-)$ ).

 $[\Gamma(\pi^+ \pi^0) + \Gamma(\pi^0 \mu^+ \nu_\mu)]/\Gamma_{\text{total}}$   $(\Gamma_3 + \Gamma_6)/\Gamma$ 

We combine these two modes for experiments measuring them in xenon bubble chamber because of difficulties of separating them there.

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>24.34±0.15 OUR FIT</b>		Error includes scale factor of 1.2.		
<b>24.6 ±1.0 OUR AVERAGE</b>		Error includes scale factor of 1.4.		
25.4 ±0.9	886	SHAKLEE	64 HLBC	+
23.4 ±1.1		ROE	61 HLBC	+

 $\Gamma(\pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$ 

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>4.82±0.06 OUR FIT</b>		Error includes scale factor of 1.3.			
<b>4.85±0.09 OUR AVERAGE</b>					
4.86±0.10	3516	CHIANG	72 OSPK	+	1.84 GeV/c $K^+$
4.7 ±0.3	429	SHAKLEE	64 HLBC	+	
5.0 ±0.5		ROE	61 HLBC	+	
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
5.1 ±1.3		31 ALEXANDER	57 EMUL	+	
3.2 ±1.3		31 BIRGE	56 EMUL	+	

31 Earlier experiments not averaged.

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$   $\Gamma_7/\Gamma_1$

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.0759±0.0011 OUR FIT</b>		Error includes scale factor of 1.4.		
<b>0.0752±0.0024 OUR AVERAGE</b>				
0.069 ± 0.006	350	ZELLER	69 ASPK	+
0.0775±0.0033	960	BOTTERILL	68C ASPK	+
0.069 ± 0.006	561	GARLAND	68 OSPK	+
0.0791±0.0054	295	32 AUERBACH	67 OSPK	+

32 AUERBACH 67 changed from  $0.0797 \pm 0.0054$ . See comment with ratio  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\mu^+ \nu_\mu)$ . The value  $0.0785 \pm 0.0025$  given in AUERBACH 67 is an average of AUERBACH 67  $\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$  and CESTER 66  $\Gamma(\pi^0 e^+ \nu_e)/[\Gamma(\mu^+ \nu_\mu) + \Gamma(\pi^+ \pi^0)]$ .

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\pi^+ \pi^0)$   $\Gamma_7/\Gamma_3$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.2280±0.0035 OUR FIT</b>		Error includes scale factor of 1.3.			
<b>0.221 ± 0.012</b>	786	33 LUCAS	73B HBC	-	Dalitz pairs only

33 LUCAS 73B gives  $N(K_{e3}) = 786 \pm 3.1\%$ ,  $N(2\pi) = 3564 \pm 3.1\%$ . We divide.

$\Gamma(\pi^0 e^+ \nu_e)/\Gamma(\pi^+ \pi^+ \pi^-)$   $\Gamma_7/\Gamma_4$

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.862±0.011 OUR FIT</b>		Error includes scale factor of 1.3.		
<b>0.860±0.014 OUR AVERAGE</b>				
0.867±0.027	2768	BARMIN	87 XEBC	+
0.856±0.040	2827	BRAUN	75 HLBC	+
0.850±0.019	4385	34 HAIDT	71 HLBC	+
0.94 ± 0.09	854	BELLOTTI	67B HLBC	
0.90 ± 0.06	230	BORREANI	64 HBC	+
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.846±0.021	4385	34 EICHTEN	68 HLBC	+
0.90 ± 0.16	37	YOUNG	65 EMUL	+

34 HAIDT 71 is a reanalysis of EICHTEN 68.

$\Gamma(\pi^0 e^+ \nu_e)/[\Gamma(\mu^+ \nu_\mu) + \Gamma(\pi^+ \pi^0)]$   $\Gamma_7/(\Gamma_1+\Gamma_3)$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG
<b>5.70±0.08 OUR FIT</b>		Error includes scale factor of 1.4.		
<b>6.01±0.15 OUR AVERAGE</b>				

5.92±0.65	35 WEISSENBE...	76 SPEC	+
6.16±0.22	5110 ESCHSTRUTH	68 OSPK	+
5.89±0.21	1679 CESTER	66 OSPK	+

35 Value calculated from WEISSENBERG 76 ( $\pi^0 e\nu$ ), ( $\mu\nu$ ), and ( $\pi\pi^0$ ) values to eliminate dependence on our 1974 ( $\pi^2\pi^0$ ) and ( $\pi\pi^+\pi^-$ ) fractions.

$\Gamma(\pi^0 \pi^0 e^+ \nu_e)/\Gamma(\pi^0 e^+ \nu_e)$

$\Gamma_8/\Gamma_7$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	-----	------	-------------	------	-----

**$4.3^{+0.9}_{-0.7}$  OUR FIT**

**$4.1^{+1.0}_{-0.7}$  OUR AVERAGE**

$4.2^{+1.0}_{-0.9}$  25 BOLOTOV 86B CALO –

$3.8^{+5.0}_{-1.2}$  2 LJUNG 73 HLBC +

• • • We do not use the following data for averages, fits, limits, etc. • • •

<37.0 90 0 ROMANO 71 HLBC +

$\Gamma(\pi^0 \pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$

$\Gamma_8/\Gamma$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	------	-------------	------	-----

**$2.1 \pm 0.4$  OUR FIT**

**$2.54 \pm 0.89$**  10 BARMIN 88B HLBC +

$\Gamma(\pi^+ \pi^- e^+ \nu_e)/\Gamma(\pi^+ \pi^+ \pi^-)$

$\Gamma_9/\Gamma_4$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	------	-------------	------	-----

**$6.99 \pm 0.30$  OUR AVERAGE** Error includes scale factor of 1.2.

$7.21 \pm 0.32$  30k ROSENLETT 77 SPEC +

$7.36 \pm 0.68$  500 BOURQUIN 71 ASPK

$7.0 \pm 0.9$  106 SCHWEINB... 71 HLBC +

$5.83 \pm 0.63$  269 ELY 69 HLBC +

• • • We do not use the following data for averages, fits, limits, etc. • • •

6.7 ± 1.5 69 BIRGE 65 FBC +

$\Gamma(\pi^+ \pi^- \mu^+ \nu_\mu)/\Gamma_{\text{total}}$

$\Gamma_{10}/\Gamma$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	------	-------------	------	-----

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.77^{+0.54}_{-0.50}$  1 CLINE 65 FBC +

$\Gamma(\pi^+ \pi^- \mu^+ \nu_\mu)/\Gamma(\pi^+ \pi^+ \pi^-)$

$\Gamma_{10}/\Gamma_4$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	------	-------------	------	-----

**$2.57 \pm 1.55$**  7 BISI 67 DBC +

• • • We do not use the following data for averages, fits, limits, etc. • • •

~ 2.5 1 GREINER 64 EMUL +

$\Gamma(\pi^0 \pi^0 \pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$

$\Gamma_{11}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	-----	------	-------------	------	-----

**<3.5** 90 0 BOLOTOV 88 SPEC –

• • • We do not use the following data for averages, fits, limits, etc. • • •

<9 90 0 BARMIN 92 XEBC +

$\Gamma(\mu^+ \nu_\mu \nu \bar{\nu})/\Gamma_{\text{total}}$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;6.0</b>	90	0	36 PANG	73 CNTR	+

36 PANG 73 assumes  $\mu$  spectrum from  $\nu$ - $\nu$  interaction of BARDIN 70.

$\Gamma_{12}/\Gamma$

$\Gamma(e^+ \nu_e \nu \bar{\nu})/\Gamma(e^+ \nu_e)$

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;3.8</b>	90	0	HEINTZE	79 SPEC	+

$\Gamma_{13}/\Gamma_2$

$\Gamma(\mu^+ \nu_\mu e^+ e^-)/\Gamma(\pi^+ \pi^- e^+ \nu_e)$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>3.3±0.9</b>	14	37 DIAMANT-...	76 SPEC	+	$m_{e^+ e^-} > 140$ MeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

27.  $\pm 8$ . 14 37 DIAMANT-... 76 SPEC + Extrapolated BR

37 DIAMANT-BERGER 76 gives this result times our 1975  $\pi^+ \pi^- e \nu$  BR ratio. The second DIAMANT-BERGER 76 value is the first value extrapolated to 0 to include low mass  $e^+ e^-$  pairs. More recent calculations (BIJNENS 93) of this extrapolation disagree with those of DIAMANT-BERGER 76.

$\Gamma_{14}/\Gamma_9$

$\Gamma(e^+ \nu_e e^+ e^-)/\Gamma(\pi^+ \pi^- e^+ \nu_e)$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.76±0.76</b>	4	38 DIAMANT-...	76 SPEC	+	$m_{e^+ e^-} > 140$ MeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

5.4  $^{+5.4}_{-2.7}$  4 38 DIAMANT-... 76 SPEC + Extrapolated BR

38 DIAMANT-BERGER 76 gives this result times our 1975  $\pi^+ \pi^- e \nu$  BR ratio. The second DIAMANT-BERGER 76 value is the first value extrapolated to 0 to include low mass  $e^+ e^-$  pairs. More recent calculations (BIJNENS 93) of this extrapolation disagree with those of DIAMANT-BERGER 76.

$\Gamma_{15}/\Gamma_9$

$\Gamma(e^+ \nu_e \mu^+ \mu^-)/\Gamma_{\text{total}}$

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt;5 × 10<sup>-7</sup></b>	90	ADLER	98 B787

$\Gamma_{16}/\Gamma$

$\Gamma(\mu^+ \nu_\mu \mu^+ \mu^-)/\Gamma_{\text{total}}$

VALUE (units $10^{-7}$ )	CL%	DOCUMENT ID	TECN	CHG
<b>&lt;4.1</b>	90	ATIYA	89 B787	+

$\Gamma_{17}/\Gamma$

$\Gamma(\mu^+ \nu_\mu \gamma)/\Gamma_{\text{total}}$

VALUE (units $10^{-3}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>5.50±0.28 OUR AVERAGE</b>					
6.6 $\pm 1.5$	39,40 DEMIDOV	90 XEBC			$P(\mu) < 231.5$ MeV/c
6.0 $\pm 0.9$	BARMIN	88 HLBC	+		$P(\mu) < 231.5$ MeV/c
5.4 $\pm 0.3$	41 AKIBA	85 SPEC			$P(\mu) < 231.5$ MeV/c

$\Gamma_{18}/\Gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.5 \pm 0.8$	40,42	DEMIDOV	90	XEBC	$E(\gamma) > 20$ MeV
$3.2 \pm 0.5$	57	43 BARMIN	88	HLBC	$+ E(\gamma) > 20$ MeV
$5.8 \pm 3.5$	12	WEISSENBE...	74	STRC	$+ E(\gamma) > 9$ MeV

<sup>39</sup>  $P(\mu)$  cut given in DEMIDOV 90 paper, 235.1 MeV/c, is a misprint according to authors (private communication).

<sup>40</sup> DEMIDOV 90 quotes only inner bremsstrahlung (IB) part.

<sup>41</sup> Assumes  $\mu$ -e universality and uses constraints from  $K \rightarrow e\nu\gamma$ .

<sup>42</sup> Not independent of above DEMIDOV 90 value. Cuts differ.

<sup>43</sup> Not independent of above BARMIN 88 value. Cuts differ.

### $\Gamma(\pi^+ \pi^0 \gamma)/\Gamma_{\text{total}}$

$\Gamma_{19}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.75±0.15 OUR AVERAGE</b>						
2.71±0.45		140	BOLOTOV	87	WIRE	—
2.87±0.32		2461	SMITH	76	WIRE	$\pm$
2.71±0.19		2100	ABRAMS	72	ASPK	$\pm$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.5^{+1.1}_{-0.6}$	44	LJUNG	73	HLBC	$+$	$T\pi^+ 55-80$ MeV
$2.6^{+1.5}_{-1.1}$	44	LJUNG	73	HLBC	$+$	$T\pi^+ 55-90$ MeV
$6.8^{+3.7}_{-2.1}$	17	44 LJUNG	73	HLBC	$+$	$T\pi^+ 55-102$ MeV
$2.4 \pm 0.8$	24	EDWARDS	72	OSPK		$T\pi^+ 58-90$ MeV
$<1.0$	0	45 MALTSEV	70	HLBC	$+$	$T\pi^+ < 55$ MeV
$<1.9$	90	0	EMMERSON	69	OSPK	$T\pi^+ 55-80$ MeV
$2.2 \pm 0.7$	18	CLINE	64	FBC	$+$	$T\pi^+ 55-80$ MeV

<sup>44</sup> The LJUNG 73 values are not independent.

<sup>45</sup> MALTSEV 70 selects low  $\pi^+$  energy to enhance direct emission contribution.

### $\Gamma(\pi^+ \pi^0 \gamma(\text{DE}))/\Gamma_{\text{total}}$

$\Gamma_{20}/\Gamma$

Direct emission part of  $\Gamma(\pi^+ \pi^0 \gamma)/\Gamma_{\text{total}}$ .

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.8 ±0.4 OUR AVERAGE</b>				
$2.05 \pm 0.46^{+0.39}_{-0.23}$	BOLOTOV	87	WIRE	—
$2.3 \pm 3.2$	SMITH	76	WIRE	$\pm$
$1.56 \pm 0.35 \pm 0.5$	ABRAMS	72	ASPK	$\pm$

### $\Gamma(\pi^+ \pi^+ \pi^- \gamma)/\Gamma_{\text{total}}$

$\Gamma_{21}/\Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.04±0.31 OUR AVERAGE</b>					
1.10±0.48	7	BARMIN	89	XEBC	$E(\gamma) > 5$ MeV
$1.0 \pm 0.4$		STAMER	65	EMUL	$+ E(\gamma) > 11$ MeV

### $\Gamma(\pi^+ \pi^0 \pi^0 \gamma)/\Gamma(\pi^+ \pi^0 \pi^0)$

$\Gamma_{22}/\Gamma_5$

VALUE (units $10^{-4}$ )	DOCUMENT ID	TECN	CHG	COMMENT
$4.3^{+3.2}_{-1.7}$	BOLOTOV	85	SPEC	— $E(\gamma) > 10$ MeV

### $\Gamma(\pi^0 \mu^+ \nu_\mu \gamma)/\Gamma_{\text{total}}$

$\Gamma_{23}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<6.1	90	0	LJUNG	73	HLBC	+ $E(\gamma) > 30$ MeV

### $\Gamma(\pi^0 e^+ \nu_e \gamma)/\Gamma(\pi^0 e^+ \nu_e)$

$\Gamma_{24}/\Gamma_7$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.54 ± 0.04 OUR AVERAGE</b>		Error includes scale factor of 1.1.			

$0.46 \pm 0.08$	82	46 BARMIN	91 XEBC		$E(\gamma) > 10$ MeV, $0.6 < \cos\theta_e \gamma < 0.9$
$0.56 \pm 0.04$	192	47 BOLOTOV	86B CALO	-	$E(\gamma) > 10$ MeV
$0.76 \pm 0.28$	13	48 ROMANO	71 HLBC		$E(\gamma) > 10$ MeV
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
$1.51 \pm 0.25$	82	46 BARMIN	91 XEBC		$E(\gamma) > 10$ MeV, $\cos\theta_e \gamma < 0.98$
$0.48 \pm 0.20$	16	49 LJUNG	73 HLBC	+	$E(\gamma) > 30$ MeV
$0.22^{+0.15}_{-0.10}$		49 LJUNG	73 HLBC	+	$E(\gamma) > 30$ MeV
$0.53 \pm 0.22$		48 ROMANO	71 HLBC	+	$E(\gamma) > 30$ MeV
$1.2 \pm 0.8$		BELLOTTI	67 HLBC	+	$E(\gamma) > 30$ MeV

<sup>46</sup> BARMIN 91 quotes branching ratio  $\Gamma(K \rightarrow e\pi^0\nu\gamma)/\Gamma_{\text{all}}$ . The measured normalization is  $[\Gamma(K \rightarrow e\pi^0\nu) + \Gamma(K \rightarrow \pi^+\pi^-\pi^0)]$ . For comparison with other experiments we used  $\Gamma(K \rightarrow e\pi^0\nu)/\Gamma_{\text{all}} = 0.0482$  to calculate the values quoted here.

<sup>47</sup>  $\cos\theta(e\gamma)$  between 0.6 and 0.9.

<sup>48</sup> Both ROMANO 71 values are for  $\cos\theta(e\gamma)$  between 0.6 and 0.9. Second value is for comparison with second LJUNG 73 value. We use lowest  $E(\gamma)$  cut for Summary Table value. See ROMANO 71 for  $E_\gamma$  dependence.

<sup>49</sup> First LJUNG 73 value is for  $\cos\theta(e\gamma) < 0.9$ , second value is for  $\cos\theta(e\gamma)$  between 0.6 and 0.9 for comparison with ROMANO 71.

### $\Gamma(\pi^0 e^+ \nu_e \gamma(\text{SD}))/\Gamma_{\text{total}}$

$\Gamma_{25}/\Gamma$

Structure-dependent part.

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	CHG
<5.3	90	BOLOTOV	86B CALO	-

### $\Gamma(\pi^0 \pi^0 e^+ \nu_e \gamma)/\Gamma_{\text{total}}$

$\Gamma_{26}/\Gamma$

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<5	90	0	BARMIN	92 XEBC	+	$E_\gamma > 10$ MeV

### $\Gamma(\pi^+ \gamma\gamma)/\Gamma_{\text{total}}$

$\Gamma_{27}/\Gamma$

All values given here assume a phase space pion energy spectrum.

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
11 ± 3 ± 1	31	50 KITCHING	97 B787			

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 10	90	0	ATIYA	90B	B787	T $\pi$	117–127 MeV
< 84	90	0	ASANO	82	CNTR	+	T $\pi$ 117–127 MeV
–420 ± 520		0	ABRAMS	77	SPEC	+	T $\pi$ <92 MeV
< 350	90	0	LJUNG	73	HLBC	+	6–102, 114–127 MeV
< 500	90	0	KLEMS	71	OSPK	+	T $\pi$ <117 MeV
–100 ± 600			CHEN	68	OSPK	+	T $\pi$ 60–90 MeV

50 KITCHING 97 is extrapolated from their model-independent branching fraction ( $6.0 \pm 1.5 \pm 0.7$ )  $\times 10^{-7}$  for  $100 \text{ MeV}/c < P_{\pi^+} < 180 \text{ MeV}/c$  using Chiral Perturbation Theory.

### $\Gamma(\pi^+ 3\gamma)/\Gamma_{\text{total}}$

$\Gamma_{28}/\Gamma$

Values given here assume a phase space pion energy spectrum.

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<1.0	90	ASANO	82	CNTR	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

<3.0	90	KLEMS	71	OSPK	+	$T(\pi) > 117 \text{ MeV}$
------	----	-------	----	------	---	----------------------------

### $\Gamma(\pi^+ \pi^+ e^- \bar{\nu}_e)/\Gamma_{\text{total}}$

$\Gamma_{29}/\Gamma$

Test of  $\Delta S = \Delta Q$  rule.

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
--------------------------	-----	------	-------------	------	-----

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 9.0	95	0	SCHWEINB...	71	HLBC	+
< 6.9	95	0	ELY	69	HLBC	+
<20.	95		BIRGE	65	FBC	+

### $\Gamma(\pi^+ \pi^+ e^- \bar{\nu}_e)/\Gamma(\pi^+ \pi^- e^+ \nu_e)$

$\Gamma_{29}/\Gamma_9$

Test of  $\Delta S = \Delta Q$  rule.

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN
< 3	90	3	51 BLOCH	76 SPEC

• • • We do not use the following data for averages, fits, limits, etc. • • •

<130.	95	0	BOURQUIN	71	ASPK
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51 BLOCH 76 quotes  $3.6 \times 10^{-4}$  at CL = 95%, we convert.

### $\Gamma(\pi^+ \pi^+ \mu^- \bar{\nu}_\mu)/\Gamma_{\text{total}}$

$\Gamma_{30}/\Gamma$

Test of  $\Delta S = \Delta Q$  rule.

VALUE (units $10^{-6}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<3.0	95	0	BIRGE	65	FBC

### $\Gamma(\pi^+ e^+ e^-)/\Gamma_{\text{total}}$

$\Gamma_{31}/\Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by combined first-order weak and electromagnetic interactions.

VALUE (units $10^{-7}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.88 ± 0.13 OUR AVERAGE</b>						
2.94 ± 0.05 ± 0.14		10300	52 APPEL	99	SPEC	+
2.75 ± 0.23 ± 0.13		500	53 ALLIEGRO	92	SPEC	+
2.7 ± 0.5		41	54 BLOCH	75	SPEC	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 17	90	CENCE	74	ASPK	+	Three track evts
< 2.7	90	CENCE	74	ASPK	+	Two track events
<320	90	BEIER	72	OSPK	±	
< 44	90	BISI	67	DBC	+	
< 8.8	90	CLINE	67B	FBC	+	
< 24.5	90	1	CAMERINI	64	FBC	+

<sup>52</sup> APPEL 99 establishes vector nature of this decay and determines form factor  $f(Z) = f_0(1+\delta Z)$ ,  $Z=M_{ee}^2/m_K^2$ ,  $\delta=2.14 \pm 0.13 \pm 0.15$ .

<sup>53</sup> ALLIEGRO 92 assumes a vector interaction with a form factor given by  $\lambda = 0.105 \pm 0.035 \pm 0.015$  and a correlation coefficient of  $-0.82$ .

<sup>54</sup> BLOCH 75 assumes a vector interaction.

### $\Gamma(\pi^+ \mu^+ \mu^-)/\Gamma_{\text{total}}$

### $\Gamma_{32}/\Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-8}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>7.6 ±2.1 OUR AVERAGE</b>			Error includes scale factor of 3.4.		
9.22 ± 0.60 ± 0.49	402	MA	00	B865	+
5.0 ± 0.4 ± 0.9	207	55 ADLER	97C	B787	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 23	90	ATIYA	89	B787	+
< 240	90	BISI	67	DBC	+
< 300	90	CAMERINI	65	FBC	+

<sup>55</sup> ADLER 97C gives systematic error  $0.7 \times 10^{-8}$  and theoretical uncertainty  $0.6 \times 10^{-8}$ , which we combine in quadrature to obtain our second error.

### $\Gamma(\pi^+ \nu \bar{\nu})/\Gamma_{\text{total}}$

### $\Gamma_{33}/\Gamma$

Test for  $\Delta S = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.15<sup>+0.34</sup><sub>-0.12</sub> OUR AVERAGE</b>						
<b>0.15<sup>+0.34</sup><sub>-0.12</sub></b>	1	ADLER	00	B787		

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.42 <sup>+0.97</sup> <sub>-0.35</sub>	1	ADLER	97	B787		
< 2.4	90	ADLER	96	B787		
< 7.5	90	ATIYA	93	B787	+	$T(\pi)$ 115–127 MeV
< 5.2	90	56 ATIYA	93	B787	+	
< 17	90	0 ATIYA	93B	B787	+	$T(\pi)$ 60–100 MeV
< 34	90	ATIYA	90	B787	+	
< 140	90	ASANO	81B	CNTR	+	$T(\pi)$ 116–127 MeV
< 940	90	57 CABLE	73	CNTR	+	$T(\pi)$ 60–105 MeV
< 560	90	57 CABLE	73	CNTR	+	$T(\pi)$ 60–127 MeV
< 57000	90	0 58 LJUNG	73	HLBC	+	
< 1400	90	57 KLEMS	71	OSPK	+	$T(\pi)$ 117–127 MeV

**56** Combining ATIYA 93 and ATIYA 93B results. Superseded by ADLER 96.

**57** KLEMS 71 and CABLE 73 assume  $\pi$  spectrum same as  $K_{e3}$  decay. Second CABLE 73 limit combines CABLE 73 and KLEMS 71 data for vector interaction.

**58** LJUNG 73 assumes vector interaction.

### $\Gamma(\mu^- \nu e^+ e^+)/\Gamma(\pi^+ \pi^- e^+ \nu_e)$

$\Gamma_{34}/\Gamma_9$

Test of lepton family number conservation.

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt;0.5</b>	90	0	59 DIAMANT-...	76 SPEC	+

59 DIAMANT-BERGER 76 quotes this result times our 1975  $\pi^+ \pi^- e \nu$  BR ratio.

### $\Gamma(\mu^+ \nu_e)/\Gamma_{\text{total}}$

$\Gamma_{35}/\Gamma$

Forbidden by lepton family number conservation.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;0.004</b>	90	0	60 LYONS	81 HLBC	0	200 GeV $K^+$ narrow band $\nu$ beam

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.012 90 60 COOPER 82 HLBC Wideband  $\nu$  beam

60 COOPER 82 and LYONS 81 limits on  $\nu_e$  observation are here interpreted as limits on lepton family number violation in the absence of mixing.

### $\Gamma(\pi^+ \mu^+ e^-)/\Gamma_{\text{total}}$

$\Gamma_{36}/\Gamma$

Test of lepton family number conservation.

VALUE (units $10^{-10}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt; 2.1</b>	90	0	LEE	90 SPEC	+	

• • • We do not use the following data for averages, fits, limits, etc. • • •

<11 90 0 CAMPAGNARI 88 SPEC + In LEE 90

<48 90 0 DIAMANT-... 76 SPEC +

### $\Gamma(\pi^+ \mu^- e^+)/\Gamma_{\text{total}}$

$\Gamma_{37}/\Gamma$

Test of lepton family number conservation.

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt; 7</b>	90	0	61 DIAMANT-... 76	SPEC	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

<28 90 61 BEIER 72 OSPK ±

61 Measurement actually applies to the sum of the  $\pi^+ \mu^- e^+$  and  $\pi^- \mu^+ e^+$  modes.

### $\Gamma(\pi^- \mu^+ e^+)/\Gamma_{\text{total}}$

$\Gamma_{38}/\Gamma$

Test of total lepton number conservation.

VALUE (units $10^{-9}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<b>&lt; 7</b>	90	0	62 DIAMANT-... 76	SPEC	+

• • • We do not use the following data for averages, fits, limits, etc. • • •

<28 90 62 BEIER 72 OSPK ±

62 Measurement actually applies to the sum of the  $\pi^+ \mu^- e^+$  and  $\pi^- \mu^+ e^+$  modes.

### $\Gamma(\pi^+ \mu^- e^+)/\Gamma_{\text{total}}$

$\Gamma_{37}/\Gamma$

VALUE (units $10^{-8}$ )	CL%	DOCUMENT ID	TECN	CHG
<b>&lt;1.4</b>	90	BEIER	72 OSPK	±

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.4 90 BEIER 72 OSPK ±

### $\Gamma(\pi^- e^+ e^+)/\Gamma_{\text{total}}$

$\Gamma_{39}/\Gamma$

Test of total lepton number conservation.

VALUE (units $10^{-5}$ )	DOCUMENT ID	TECN	CHG
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$			
<1.5	CHANG	68	HBC

### $\Gamma(\pi^- e^+ e^+)/\Gamma(\pi^+ \pi^- e^+ \nu_e)$

$\Gamma_{39}/\Gamma_9$

Test of total lepton number conservation.

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	CHG
<2.5	90	0	63 DIAMANT-...	76	SPEC +

<sup>63</sup> DIAMANT-BERGER 76 quotes this result times our 1975 BR ratio.

### $\Gamma(\pi^- \mu^+ \mu^+)/\Gamma_{\text{total}}$

$\Gamma_{40}/\Gamma$

Forbidden by total lepton number conservation.

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN
<1.5	90	64 LITTENBERG	92 HBC

<sup>64</sup> LITTENBERG 92 is from retroactive data analysis of CHANG 68 bubble chamber data.

### $\Gamma(\mu^+ \bar{\nu}_e)/\Gamma_{\text{total}}$

$\Gamma_{41}/\Gamma$

Forbidden by total lepton number conservation.

VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<3.3	90	65 COOPER	82 HLBC	Wideband $\nu$ beam

<sup>65</sup> COOPER 82 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton number violation in the absence of mixing.

### $\Gamma(\pi^0 e^+ \bar{\nu}_e)/\Gamma_{\text{total}}$

$\Gamma_{42}/\Gamma$

Forbidden by total lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.003	90	66 COOPER	82 HLBC	Wideband $\nu$ beam

<sup>66</sup> COOPER 82 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton number violation in the absence of mixing.

### $\Gamma(\pi^+ \gamma)/\Gamma_{\text{total}}$

$\Gamma_{43}/\Gamma$

Violates angular momentum conservation. Not listed in Summary Table.

VALUE (units $10^{-6}$ )	CL%	DOCUMENT ID	TECN	CHG
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
<1.4	90	ASANO	82 CNTR	+
<4.0	90	67 KLEMS	71 OSPK	+

<sup>67</sup> Test of model of Selleri, Nuovo Cimento **60A** 291 (1969).

## $K^+$ LONGITUDINAL POLARIZATION OF EMITTED $\mu^+$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<-0.990	90	68 AOKI	94	SPEC	+
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>					
<-0.990	90	IMAZATO	92	SPEC	Repl. by AOKI 94
-0.970 ± 0.047		69 YAMANAKA	86	SPEC	+
-1.0 ± 0.1		69 CUTTS	69	SPRK	+
-0.96 ± 0.12		69 COOMBES	57	CNTR	+

<sup>68</sup> AOKI 94 measures  $\xi P_\mu = -0.9996 \pm 0.0030 \pm 0.0048$ . The above limit is obtained by summing the statistical and systematic errors in quadrature, normalizing to the physically significant region ( $|\xi P_\mu| < 1$ ) and assuming that  $\xi=1$ , its maximum value.

<sup>69</sup> Assumes  $\xi=1$ .

## DALITZ PLOT PARAMETERS FOR $K \rightarrow 3\pi$ DECAYS

Revised 1999 by T.G. Trippe (LBNL).

The Dalitz plot distribution for  $K^\pm \rightarrow \pi^\pm \pi^\pm \pi^\mp$ ,  $K^\pm \rightarrow \pi^0 \pi^0 \pi^\pm$ , and  $K_L^0 \rightarrow \pi^+ \pi^- \pi^0$  can be parameterized by a series expansion such as that introduced by Weinberg [1]. We use the form

$$\begin{aligned} |M|^2 \propto & 1 + g \frac{(s_3 - s_0)}{am_{\pi^+}^2} + h \left[ \frac{s_3 - s_0}{m_{\pi^+}^2} \right]^2 \\ & + j \frac{(s_2 - s_1)}{m_{\pi^+}^2} + k \left[ \frac{s_2 - s_1}{m_{\pi^+}^2} \right]^2 \\ & + f \frac{(s_2 - s_1)}{m_{\pi^+}^2} \frac{(s_3 - s_0)}{m_{\pi^+}^2} + \dots , \end{aligned} \quad (1)$$

where  $m_{\pi^+}^2$  has been introduced to make the coefficients  $g$ ,  $h$ ,  $j$ , and  $k$  dimensionless, and

$$s_i = (P_K - P_i)^2 = (m_K - m_i)^2 - 2m_K T_i, \quad i = 1, 2, 3,$$

$$s_0 = \frac{1}{3} \sum_i s_i = \frac{1}{3} (m_K^2 + m_1^2 + m_2^2 + m_3^2)$$

Here the  $P_i$  are four-vectors,  $m_i$  and  $T_i$  are the mass and kinetic energy of the  $i^{th}$  pion, and the index 3 is used for the odd pion.

The coefficient  $g$  is a measure of the slope in the variable  $s_3$  (or  $T_3$ ) of the Dalitz plot, while  $h$  and  $k$  measure the quadratic dependence on  $s_3$  and  $(s_2 - s_1)$ , respectively. The coefficient  $j$  is related to the asymmetry of the plot and must be zero if  $CP$  invariance holds. Note also that if  $CP$  is good,  $g$ ,  $h$ , and  $k$  must be the same for  $K^+ \rightarrow \pi^+\pi^+\pi^-$  as for  $K^- \rightarrow \pi^-\pi^-\pi^+$ .

Since different experiments use different forms for  $|M|^2$ , in order to compare the experiments we have converted to  $g$ ,  $h$ ,  $j$ , and  $k$  whatever coefficients have been measured. Where such conversions have been done, the measured coefficient  $a_y$ ,  $a_t$ ,  $a_u$ , or  $a_v$  is given in the comment at the right. For definitions of these coefficients, details of this conversion, and discussion of the data, see the April 1982 version of this note [2].

## References

1. S. Weinberg, Phys. Rev. Lett. **4**, 87 (1960).
  2. Particle Data Group, Phys. Lett. **111B**, 69 (1982).
- 

### ENERGY DEPENDENCE OF $K^\pm$ DALITZ PLOT

$$|\text{matrix element}|^2 = 1 + gu + hu^2 + kv^2$$

where  $u = (s_3 - s_0) / m_\pi^2$  and  $v = (s_1 - s_2) / m_\pi^2$

### LINEAR COEFFICIENT $g_{\tau^+}$ FOR $K^+ \rightarrow \pi^+\pi^+\pi^-$

Some experiments use Dalitz variables  $x$  and  $y$ . In the comments we give  $a_y$  = coefficient of  $y$  term. See note above on "Dalitz Plot Parameters for  $K \rightarrow 3\pi$  Decays." For discussion of the conversion of  $a_y$  to  $g$ , see the earlier version of the same note in the Review published in Physics Letters **111B** 70 (1982).

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>-0.2154 \pm 0.0035</math> OUR AVERAGE</b>					Error includes scale factor of 1.4. See the ideogram below.
$-0.2221 \pm 0.0065$	225k	DEVAUX	77	SPEC	$+ a_y = .2814 \pm .0082$
$-0.2157 \pm 0.0028$	750k	FORD	72	ASPK	$+ a_y = .2734 \pm .0035$
$-0.200 \pm 0.009$	39819	<sup>70</sup> HOFFMASTER	72	HLBC	$+ a_y = .2734 \pm .0035$

• • • We do not use the following data for averages, fits, limits, etc. • • •

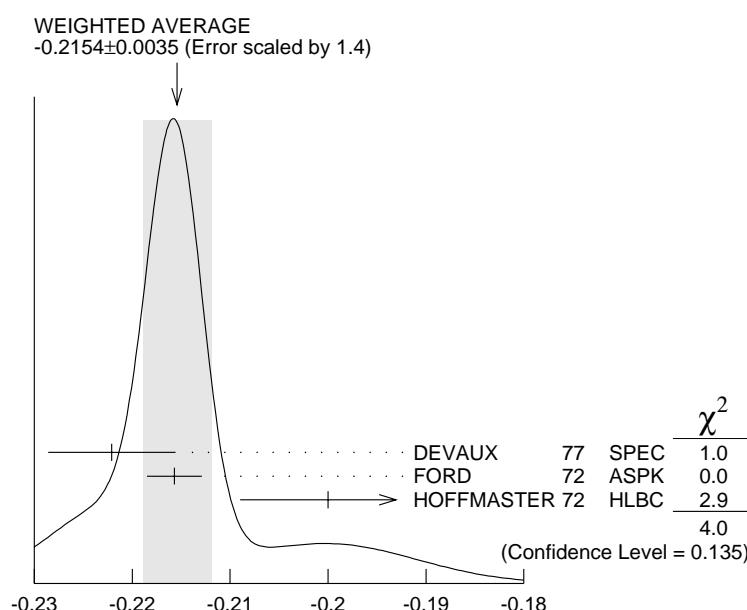
-0.196 ± 0.012	17898	<sup>71</sup> GRAUMAN	70	HLBC	+	$a_y = 0.228 \pm 0.030$
-0.218 ± 0.016	9994	<sup>72</sup> BUTLER	68	HBC	+	$a_y = 0.277 \pm 0.020$
-0.22 ± 0.024	5428	<sup>72,73</sup> ZINCHENKO	67	HBC	+	$a_y = 0.28 \pm 0.03$

<sup>70</sup> HOFFMASTER 72 includes GRAUMAN 70 data.

<sup>71</sup> Emulsion data added — all events included by HOFFMASTER 72.

<sup>72</sup> Experiments with large errors not included in average.

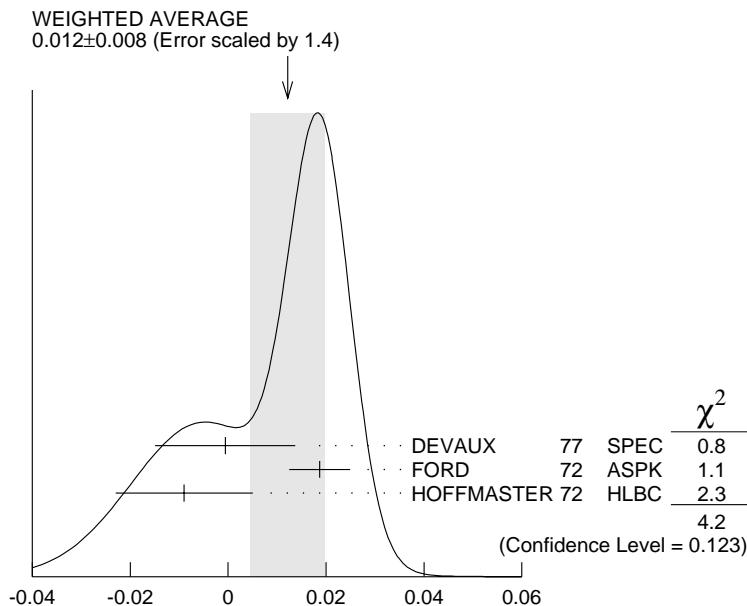
<sup>73</sup> Also includes DBC events.



Linear energy dependence for  $K^+ \rightarrow \pi^+ \pi^+ \pi^-$

### QUADRATIC COEFFICIENT $h$ FOR $K^+ \rightarrow \pi^+ \pi^+ \pi^-$

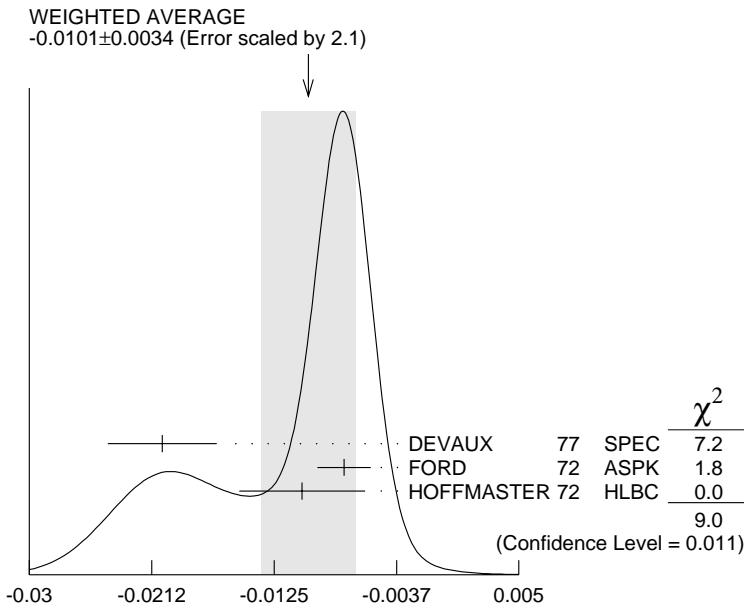
VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.012 ± 0.008 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
-0.0006 ± 0.0143	225k	DEVAUX	77	SPEC
0.0187 ± 0.0062	750k	FORD	72	ASPK
-0.009 ± 0.014	39819	HOFFMASTER 72	HLBC	+



Quadratic coefficient  $h$  for  $K^+ \rightarrow \pi^+ \pi^+ \pi^-$

### QUADRATIC COEFFICIENT $k$ FOR $K^+ \rightarrow \pi^+ \pi^+ \pi^-$

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>-0.0101 ± 0.0034 OUR AVERAGE</b>				Error includes scale factor of 2.1. See the ideogram below.
-0.0205 ± 0.0039	225k	DEVAUX	77	SPEC +
-0.0075 ± 0.0019	750k	FORD	72	ASPK +
-0.0105 ± 0.0045	39819	HOFFMASTER 72	HLBC	+ (from CHG column)



Quadratic coefficient  $k$  for  $K^+ \rightarrow \pi^+ \pi^+ \pi^-$

### LINEAR COEFFICIENT $g_{\tau^-}$ FOR $K^- \rightarrow \pi^- \pi^- \pi^+$

Some experiments use Dalitz variables  $x$  and  $y$ . In the comments we give  $a_y$  = coefficient of  $y$  term. See note above on "Dalitz Plot Parameters for  $K \rightarrow 3\pi$  Decays." For discussion of the conversion of  $a_y$  to  $g$ , see the earlier version of the same note in the Review published in Physics Letters **111B** 70 (1982).

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.217 ± 0.007 OUR AVERAGE</b>					Error includes scale factor of 2.5.
-0.2186 ± 0.0028	750k	FORD	72	ASPK	$a_y = 0.2770 \pm 0.0035$
-0.193 ± 0.010	50919	MAST	69	HBC	$a_y = 0.244 \pm 0.013$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
-0.199 ± 0.008	81k	74 LUCAS	73	HBC	$a_y = 0.252 \pm 0.011$
-0.190 ± 0.023	5778	75,76 MOSCOSO	68	HBC	$a_y = 0.242 \pm 0.029$
-0.220 ± 0.035	1347	77 FERRO-LUZZI	61	HBC	$a_y = 0.28 \pm 0.045$

<sup>74</sup> Quadratic dependence is required by  $K_L^0$  experiments. For comparison we average only those  $K^\pm$  experiments which quote quadratic fit values.

<sup>75</sup> Experiments with large errors not included in average.

<sup>76</sup> Also includes DBC events.

<sup>77</sup> No radiative corrections included.

### QUADRATIC COEFFICIENT $h$ FOR $K^- \rightarrow \pi^- \pi^- \pi^+$

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.010 ± 0.006 OUR AVERAGE</b>				
0.0125 ± 0.0062	750k	FORD	72	ASPK
-0.001 ± 0.012	50919	MAST	69	HBC

## QUADRATIC COEFFICIENT $k$ FOR $K^- \rightarrow \pi^- \pi^- \pi^+$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>
<b>-0.0084 ± 0.0019 OUR AVERAGE</b>				
-0.0083 ± 0.0019	750k	FORD	72	ASPK
-0.014 ± 0.012	50919	MAST	69	HBC

## $(g_{\tau^+} - g_{\tau^-}) / (g_{\tau^+} + g_{\tau^-})$ FOR $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$

A nonzero value for this quantity indicates  $CP$  violation.

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
-0.70 ± 0.53	3.2M	FORD	70

## LINEAR COEFFICIENT $g$ FOR $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

Unless otherwise stated, all experiments include terms quadratic

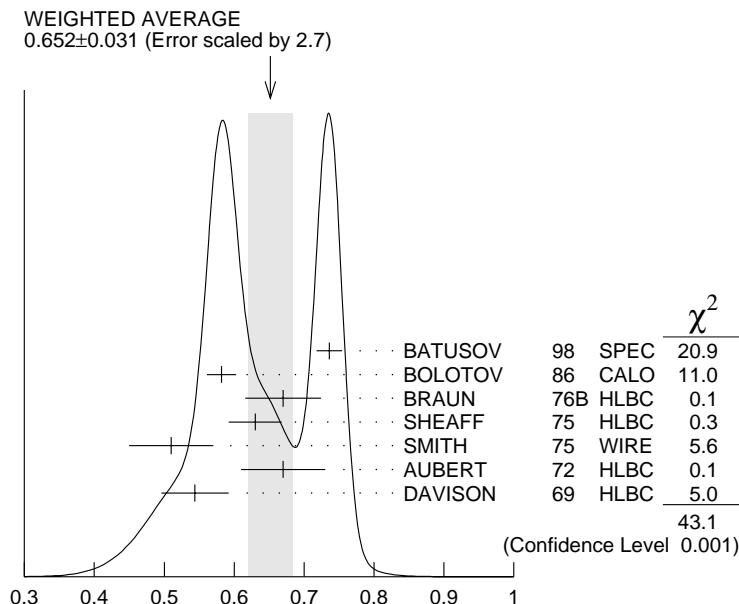
in  $(s_3 - s_0) / m_{\pi^+}^2$ . See note above on "Dalitz Plot Parameters for  $K \rightarrow 3\pi$  Decays."

See BATUSOV 98 for a discussion of the discrepancy between their result and others, especially BOLOTOV 86. At this time we have no way to resolve the discrepancy so we depend on the large scale factor as a warning.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>0.652 ± 0.031 OUR AVERAGE</b>					Error includes scale factor of 2.7. See the ideogram below.
0.736 ± 0.014 ± 0.012	33k	BATUSOV	98	SPEC	+
0.582 ± 0.021	43k	BOLOTOV	86	CALO	-
0.670 ± 0.054	3263	BRAUN	76B	HLBC	+
0.630 ± 0.038	5635	SHEAFF	75	HLBC	+
0.510 ± 0.060	27k	SMITH	75	WIRE	+
0.67 ± 0.06	1365	AUBERT	72	HLBC	+
0.544 ± 0.048	4048	DAVISON	69	HLBC	+
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.806 ± 0.220	4639	78 BERTRAND	76	EMUL	+
0.484 ± 0.084	574	79 LUCAS	73B	HBC	-
0.527 ± 0.102	198	78 PANDOULAS	70	EMUL	+
0.586 ± 0.098	1874	79 BISI	65	HLBC	+
0.48 ± 0.04	1792	79 KALMUS	64	HLBC	+

<sup>78</sup> Experiments with large errors not included in average.

<sup>79</sup> Authors give linear fit only.

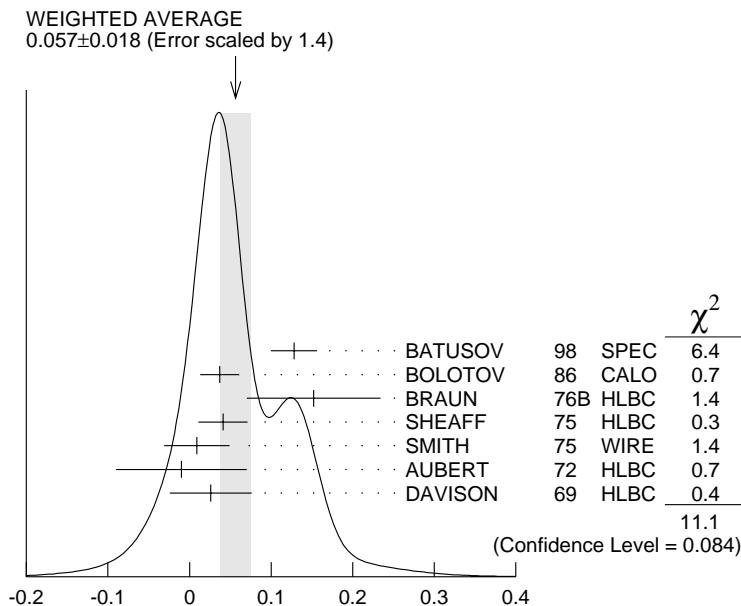


Linear energy dependence for  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

### QUADRATIC COEFFICIENT $h$ FOR $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.057±0.018 OUR AVERAGE</b>					Error includes scale factor of 1.4. See the ideogram below.
0.128±0.015±0.024	33k	BATUSOV	98	SPEC	+
0.037±0.024	43k	BOLOTOV	86	CALO	-
0.152±0.082	3263	BRAUN	76B	HLBC	+
0.041±0.030	5635	SHEAFF	75	HLBC	+
0.009±0.040	27k	SMITH	75	WIRE	+
-0.01 ± 0.08	1365	AUBERT	72	HLBC	+
0.026±0.050	4048	DAVISON	69	HLBC	+
					Also emulsion
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.164±0.121	4639	80 BERTRAND	76	EMUL	+
0.018±0.124	198	80 PANDOULAS	70	EMUL	+

80 Experiments with large errors not included in average.



Quadratic coefficient  $h$  FOR  $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

### QUADRATIC COEFFICIENT $k$ FOR $K^\pm \rightarrow \pi^\pm \pi^0 \pi^0$

VALUE	EVTS	DOCUMENT ID	TECN	CHG
<b>0.0197±0.0045±0.0029</b>	33k	BATUSOV	98	SPEC +

## $K_{\ell 3}^\pm$ AND $K_{\ell 3}^0$ FORM FACTORS

Written by T.G. Trippe (LBNL).

Assuming that only the vector current contributes to  $K \rightarrow \pi \ell \nu$  decays, we write the matrix element as

$$M \propto f_+(t) [(P_K + P_\pi)_\mu \bar{\ell} \gamma_\mu (1 + \gamma_5) \nu] + f_-(t) [m_\ell \bar{\ell} (1 + \gamma_5) \nu], \quad (1)$$

where  $P_K$  and  $P_\pi$  are the four-momenta of the  $K$  and  $\pi$  mesons,  $m_\ell$  is the lepton mass, and  $f_+$  and  $f_-$  are dimensionless form factors which can depend only on  $t = (P_K - P_\pi)^2$ , the square of the four-momentum transfer to the leptons. If time-reversal invariance holds,  $f_+$  and  $f_-$  are relatively real.  $K_{\mu 3}$  experiments measure  $f_+$  and  $f_-$ , while  $K_{e 3}$  experiments are sensitive only

to  $f_+$  because the small electron mass makes the  $f_-$  term negligible.

**(a)  $K_{\mu 3}$  experiments.** Analyses of  $K_{\mu 3}$  data frequently assume a linear dependence of  $f_+$  and  $f_-$  on  $t$ , *i.e.*,

$$f_{\pm}(t) = f_{\pm}(0) [1 + \lambda_{\pm}(t/m_{\pi}^2)] \quad (2)$$

Most  $K_{\mu 3}$  data are adequately described by Eq. (2) for  $f_+$  and a constant  $f_-$  (*i.e.*,  $\lambda_- = 0$ ). There are two equivalent parametrizations commonly used in these analyses:

**(1)  $\lambda_+, \xi(0)$  parametrization.** Analyses of  $K_{\mu 3}$  data often introduce the ratio of the two form factors

$$\xi(t) = f_-(t)/f_+(t) . \quad (3)$$

The  $K_{\mu 3}$  decay distribution is then described by the two parameters  $\lambda_+$  and  $\xi(0)$  (assuming time reversal invariance and  $\lambda_- = 0$ ). These parameters can be determined by three different methods:

**Method A.** By studying the Dalitz plot or the pion spectrum of  $K_{\mu 3}$  decay. The Dalitz plot density is (see, *e.g.*, Chouinet *et al.* [1]):

$$\rho(E_{\pi}, E_{\mu}) \propto f_+^2(t) [A + B\xi(t) + C\xi(t)^2] ,$$

where

$$A = m_K (2E_{\mu}E_{\nu} - m_K E'_{\pi}) + m_{\mu}^2 \left( \frac{1}{4}E'_{\pi} - E_{\nu} \right) ,$$

$$B = m_{\mu}^2 \left( E_{\nu} - \frac{1}{2}E'_{\pi} \right) ,$$

$$C = \frac{1}{4}m_{\mu}^2 E'_{\pi} ,$$

$$E'_{\pi} = E_{\pi}^{\max} - E_{\pi} = (m_K^2 + m_{\pi}^2 - m_{\mu}^2) / 2m_K - E_{\pi} . \quad (4)$$

Here  $E_\pi$ ,  $E_\mu$ , and  $E_\nu$  are, respectively, the pion, muon, and neutrino energies in the kaon center of mass. The density  $\rho$  is fit to the data to determine the values of  $\lambda_+$ ,  $\xi(0)$ , and their correlation.

**Method B.** By measuring the  $K_{\mu 3}/K_{e 3}$  branching ratio and comparing it with the theoretical ratio (see, *e.g.*, Fearing *et al.* [2]) as given in terms of  $\lambda_+$  and  $\xi(0)$ , assuming  $\mu$ - $e$  universality:

$$\begin{aligned} \Gamma(K_{\mu 3}^\pm)/\Gamma(K_{e 3}^\pm) &= 0.6457 + 1.4115\lambda_+ + 0.1264\xi(0) \\ &\quad + 0.0192\xi(0)^2 + 0.0080\lambda_+\xi(0) , \\ \Gamma(K_{\mu 3}^0)/\Gamma(K_{e 3}^0) &= 0.6452 + 1.3162\lambda_+ + 0.1264\xi(0) \\ &\quad + 0.0186\xi(0)^2 + 0.0064\lambda_+\xi(0) . \end{aligned} \quad (5)$$

This cannot determine  $\lambda_+$  and  $\xi(0)$  simultaneously but simply fixes a relationship between them.

**Method C.** By measuring the muon polarization in  $K_{\mu 3}$  decay. In the rest frame of the  $K$ , the  $\mu$  is expected to be polarized in the direction  $\mathbf{A}$  with  $\mathbf{P} = \mathbf{A}/|\mathbf{A}|$ , where  $\mathbf{A}$  is given (Cabibbo and Maksymowicz [3]) by

$$\begin{aligned} \mathbf{A} &= a_1(\xi)\mathbf{p}_\mu \\ &- a_2(\xi) \left[ \frac{\mathbf{p}_\mu}{m_\mu} \left( m_K - E_\pi + \frac{\mathbf{p}_\pi \cdot \mathbf{p}_\mu}{|\mathbf{p}_\mu|^2} (E_\mu - m_\mu) \right) + \mathbf{p}_\pi \right] \\ &+ m_K \text{Im}\xi(t)(\mathbf{p}_\pi \times \mathbf{p}_\mu) . \end{aligned} \quad (6)$$

If time-reversal invariance holds,  $\xi$  is real, and thus there is no polarization perpendicular to the  $K$ -decay plane. Polarization experiments measure the weighted average of  $\xi(t)$  over the  $t$

range of the experiment, where the weighting accounts for the variation with  $t$  of the sensitivity to  $\xi(t)$ .

**(2)  $\lambda_+, \lambda_0$  parametrization.** Most of the more recent  $K_{\mu 3}$  analyses have parameterized in terms of the form factors  $f_+$  and  $f_0$  which are associated with vector and scalar exchange, respectively, to the lepton pair.  $f_0$  is related to  $f_+$  and  $f_-$  by

$$f_0(t) = f_+(t) + [t/(m_K^2 - m_\pi^2)] f_-(t). \quad (7)$$

Here  $f_0(0)$  must equal  $f_+(0)$  unless  $f_-(t)$  diverges at  $t = 0$ . The earlier assumption that  $f_+$  is linear in  $t$  and  $f_-$  is constant leads to  $f_0$  linear in  $t$ :

$$f_0(t) = f_0(0) [1 + \lambda_0(t/m_\pi^2)]. \quad (8)$$

With the assumption that  $f_0(0) = f_+(0)$ , the two parametrizations,  $(\lambda_+, \xi(0))$  and  $(\lambda_+, \lambda_0)$  are equivalent as long as correlation information is retained.  $(\lambda_+, \lambda_0)$  correlations tend to be less strong than  $(\lambda_+, \xi(0))$  correlations.

The experimental results for  $\xi(0)$  and its correlation with  $\lambda_+$  are listed in the  $K^\pm$  and  $K_L^0$  sections of the Particle Listings in section  $\xi_A$ ,  $\xi_B$ , or  $\xi_C$  depending on whether method A, B, or C discussed above was used. The corresponding values of  $\lambda_+$  are also listed.

Because recent experiments tend to use the  $(\lambda_+, \lambda_0)$  parametrization, we include a subsection for  $\lambda_0$  results. Whenever possible we have converted  $\xi(0)$  results into  $\lambda_0$  results and vice versa.

See the 1982 version of this note [4] for additional discussion of the  $K_{\mu 3}^0$  parameters, correlations, and conversion between parametrizations, and also for a comparison of the experimental results.

**(b)  $K_{e3}$  experiments.** Analysis of  $K_{e3}$  data is simpler than that of  $K_{\mu 3}$  because the second term of the matrix element assuming a pure vector current [Eq. (1) above] can be neglected. Here  $f_+$  is usually assumed to be linear in  $t$ , and the linear coefficient  $\lambda_+$  of Eq. (2) is determined.

If we remove the assumption of a pure vector current, then the matrix element for the decay, in addition to the terms in Eq. (1), would contain

$$\begin{aligned} & +2m_K f_S \bar{\ell}(1+\gamma_5)\nu \\ & +(2f_T/m_K)(P_K)_\lambda(P_\pi)_\mu \bar{\ell}\sigma_{\lambda\mu}(1+\gamma_5)\nu , \end{aligned} \quad (9)$$

where  $f_S$  is the scalar form factor, and  $f_T$  is the tensor form factor. In the case of the  $K_{e3}$  decays where the  $f_-$  term can be neglected, experiments have yielded limits on  $|f_S/f_+|$  and  $|f_T/f_+|$ .

## References

1. L.M. Chouquet, J.M. Gaillard, and M.K. Gaillard, Phys. Reports **4C**, 199 (1972).
  2. H.W. Fearing, E. Fischbach, and J. Smith, Phys. Rev. **D2**, 542 (1970).
  3. N. Cabibbo and A. Maksymowicz, Phys. Lett. **9**, 352 (1964).
  4. Particle Data Group, Phys. Lett. **111B**, 73 (1982).
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## $K_{e3}^{\pm}$ FORM FACTORS

In the form factor comments, the following symbols are used.

$f_+$  and  $f_-$  are form factors for the vector matrix element.

$f_S$  and  $f_T$  refer to the scalar and tensor term.

$f_0 = f_+ + f_- t/(m_K^2 - m_\pi^2)$ .

$\lambda_+$ ,  $\lambda_-$ , and  $\lambda_0$  are the linear expansion coefficients of  $f_+$ ,  $f_-$ , and  $f_0$ .

$\lambda_+$  refers to the  $K_{\mu 3}^{\pm}$  value except in the  $K_{e3}^{\pm}$  sections.

$d\xi(0)/d\lambda_+$  is the correlation between  $\xi(0)$  and  $\lambda_+$  in  $K_{\mu 3}^{\pm}$ .

$d\lambda_0/d\lambda_+$  is the correlation between  $\lambda_0$  and  $\lambda_+$  in  $K_{\mu 3}^{\pm}$ .

$t$  = momentum transfer to the  $\pi$  in units of  $m_\pi^2$ .

DP = Dalitz plot analysis.

PI =  $\pi$  spectrum analysis.

MU =  $\mu$  spectrum analysis.

POL =  $\mu$  polarization analysis.

BR =  $K_{\mu 3}^{\pm}/K_{e3}^{\pm}$  branching ratio analysis.

E = positron or electron spectrum analysis.

RC = radiative corrections.

### $\lambda_+$ (LINEAR ENERGY DEPENDENCE OF $f_+$ IN $K_{e3}^{\pm}$ DECAY)

For radiative correction of  $K_{e3}^{\pm}$  Dalitz plot, see GINSBERG 67 and BECHERRAWY 70.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.0276±0.0021 OUR AVERAGE</b>					
0.018 ± 0.007	3k	ARTEMOV	97B SPEC	–	DP
0.0284±0.0027±0.0020	32k	81 AKIMENKO	91 SPEC		PI, no RC
0.029 ± 0.004	62k	82 BOLOTOV	88 SPEC		PI, no RC
0.027 ± 0.008		83 BRAUN	73B HLBC	+	DP, no RC
0.029 ± 0.011	4017	CHIANG	72 OSPK	+	DP, RC negligible
0.027 ± 0.010	2707	STEINER	71 HLBC	+	DP, uses RC
0.045 ± 0.015	1458	BOTTERILL	70 OSPK		PI, uses RC
0.08 ± 0.04	960	BOTTERILL	68C ASPK	+	$e^+$ , uses RC
–0.02 ± 0.08 –0.12	90	EISLER	68 HLBC	+	PI, uses RC
0.045 ± 0.017 –0.018	854	BELLOTTI	67B FBC	+	DP, uses RC
+0.016 ± 0.016	1393	IMLAY	67 OSPK	+	DP, no RC
+0.028 ± 0.013 –0.014	515	KALMUS	67 FBC	+	$e^+$ , PI, no RC
–0.04 ± 0.05	230	BORREANI	64 HBC	+	$e^+$ , no RC
–0.010 ± 0.029	407	JENSEN	64 XEBC	+	PI, no RC
+0.036 ± 0.045	217	BROWN	62B XEBC	+	PI, no RC

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.025 ± 0.007                    84 BRAUN                    74 HLBC                     $K_{\mu 3}/K_{e3}$  vs.  $t$

81 AKIMENKO 91 state that radiative corrections would raise  $\lambda_+$  by 0.0013.

82 BOLOTOV 88 state radiative corrections of GINSBERG 67 would raise  $\lambda_+$  by 0.002.

- <sup>83</sup> BRAUN 73B states that radiative corrections of GINSBERG 67 would lower  $\lambda_+^e$  by 0.002 but that radiative corrections of BECHERRAWY 70 disagrees and would raise  $\lambda_+^e$  by 0.005.  
<sup>84</sup> BRAUN 74 is a combined  $K_{\mu 3}$ - $K_{e3}$  result. It is not independent of BRAUN 73C ( $K_{\mu 3}$ ) and BRAUN 73B ( $K_{e3}$ ) form factor results.

### $\xi_A = f_-/f_+$ (determined from $K_{\mu 3}^\pm$ spectra)

The parameter  $\xi$  is redundant with  $\lambda_0$  below and is not put into the Meson Summary Table.

VALUE	$d\xi(0)/d\lambda_+$	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>-0.31 \pm 0.15</math> OUR EVALUATION</b>			Error includes scale factor of 1.6. Correlation is $d\xi(0)/d\lambda_+ = -14$ . From a fit discussed in note on $K_{\ell 3}$ form factors in 1982 edition, PL <b>111B</b> (April 1982).			
$-0.27 \pm 0.25$	-17	3973	WHITMAN	80	SPEC	+
$-0.8 \pm 0.8$	-20	490	85 ARNOLD	74	HLBC	+
$-0.57 \pm 0.24$	-9	6527	86 MERLAN	74	ASPK	+
$-0.36 \pm 0.40$	-19	1897	87 BRAUN	73C	HLBC	+
$-0.62 \pm 0.28$	-12	4025	88 ANKENBRA...	72	ASPK	+
$+0.45 \pm 0.28$	-15	3480	89 CHIANG	72	OSPK	+
$-1.1 \pm 0.56$	-29	3240	90 HAIDT	71	HLBC	+
$-0.5 \pm 0.8$	-26	2041	91 KIJEWSKI	69	OSPK	+
$+0.72 \pm 0.93$	-17	444	CALLAHAN	66B	FBC	+
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>						
$-0.5 \pm 0.9$	none	78	EISLER	68	HLBC	+
$0.0^{+1.1}_{-0.9}$		2648	92 CALLAHAN	66B	FBC	+
$+0.7 \pm 0.5$		87	GIACOMELLI	64	EMUL	+
$-0.08 \pm 0.7$		93	JENSEN	64	XEBC	+
$+1.8 \pm 0.6$		76	BROWN	62B	XEBC	+
						$\lambda_+ = 0$
						$\mu, \lambda_+ = 0$
						$MU+BR, \lambda_+ = 0$
						$DP+BR$
						$DP+BR, \lambda_+ = 0$

<sup>85</sup> ARNOLD 74 figure 4 was used to obtain  $\xi_A$  and  $d\xi(0)/d\lambda_+$ .

<sup>86</sup> MERLAN 74 figure 5 was used to obtain  $d\xi(0)/d\lambda_+$ .

<sup>87</sup> BRAUN 73C gives  $\xi(t) = -0.34 \pm 0.20$ ,  $d\xi(t)/d\lambda_+ = -14$  for  $\lambda_+ = 0.027$ ,  $t = 6.6$ .  
We calculate above  $\xi(0)$  and  $d\xi(0)/d\lambda_+$  for their  $\lambda_+ = 0.025 \pm 0.017$ .

<sup>88</sup> ANKENBRANDT 72 figure 3 was used to obtain  $d\xi(0)/d\lambda_+$ .

<sup>89</sup> CHIANG 72 figure 10 was used to obtain  $d\xi(0)/d\lambda_+$ . Fit had  $\lambda_- = \lambda_+$  but would not change for  $\lambda_- = 0$ . L.Pondrom, (private communication 74).

<sup>90</sup> HAIDT 71 table 8 (Dalitz plot analysis) gives  $d\xi(0)/d\lambda_+ = (-1.1 + 0.5)/(0.050 - 0.029) = -29$ , error raised from 0.50 to agree with  $d\xi(0) = 0.20$  for fixed  $\lambda_+$ .

<sup>91</sup> KIJEWSKI 69 figure 17 was used to obtain  $d\xi(0)/d\lambda_+$  and errors.

<sup>92</sup> CALLAHAN 66 table 1 ( $\pi$  analysis) gives  $d\xi(0)/d\lambda_+ = (0.72 - 0.05)/(0 - 0.04) = -17$ , error raised from 0.80 to agree with  $d\xi(0) = 0.37$  for fixed  $\lambda_+$ .  $t$  unknown.

<sup>93</sup> JENSEN 64 gives  $\lambda_+^\mu = \lambda_+^e = -0.020 \pm 0.027$ .  $d\xi(0)/d\lambda_+$  unknown. Includes SHAKLEE 64  $\xi_B(K_{\mu 3}/K_{e3})$ .

### $\xi_B = f_-/f_+$ (determined from $K_{\mu 3}^\pm/K_{e3}^\pm$ )

The  $K_{\mu 3}^\pm/K_{e3}^\pm$  branching ratio fixes a relationship between  $\xi(0)$  and  $\lambda_+$ . We quote the author's  $\xi(0)$  and associated  $\lambda_+$  but do not average because the  $\lambda_+$  values differ. The fit result and scale factor given below are not obtained from these  $\xi_B$  values. Instead they are obtained directly from the fitted  $K_{\mu 3}^\pm/K_{e3}^\pm$  ratio  $\Gamma(\pi^0 \mu^+ \nu_\mu)/\Gamma(\pi^0 e^+ \nu_e)$ , with the exception of HEINTZE 77. The parameter  $\xi$  is redundant with  $\lambda_0$  below and is not put into the Meson Summary Table.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.31±0.15 OUR EVALUATION</b>		Error includes scale factor of 1.6. Correlation is $d\xi(0)/d\lambda_+ = -14$ . From a fit discussed in note on $K_{\ell 3}$ form factors in 1982 edition, PL <b>111B</b> (April 1982).			
-0.12±0.12	55k	94 HEINTZE	77 CNTR	+	$\lambda_+ = 0.029$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.0 ± 0.15	5825	CHIANG	72 OSPK	+	$\lambda_+ = 0.03$ , fig.10
-0.81±0.27	1505	95 HAIDT	71 HLBC	+	$\lambda_+ = 0.028$ , fig.8
-0.35±0.22		96 BOTTERILL	70 OSPK	+	$\lambda_+ = 0.045 \pm 0.015$
+0.91±0.82		ZELLER	69 ASPK	+	$\lambda_+ = 0.023$
-0.08±0.15	5601	96 BOTTERILL	68B ASPK	+	$\lambda_+ = 0.023 \pm 0.008$
-0.60±0.20	1398	95 EICHTEN	68 HLBC	+	See note
+1.0 ± 0.6	986	GARLAND	68 OSPK	+	$\lambda_+ = 0$
+0.75±0.50	306	AUERBACH	67 OSPK	+	$\lambda_+ = 0$
+0.4 ± 0.4	636	CALLAHAN	66B FBC	+	$\lambda_+ = 0$
+0.6 ± 0.5		BISI	65B HBC	+	$\lambda_+ = 0$
+0.8 ± 0.6	500	CUTTS	65 OSPK	+	$\lambda_+ = 0$
-0.17 <sup>+0.75</sup> <sub>-0.99</sub>		SHAKLEE	64 XEBC	+	$\lambda_+ = 0$

<sup>94</sup> Calculated by us from  $\lambda_0$  and  $\lambda_+$  given below.

<sup>95</sup> EICHTEN 68 has  $\lambda_+ = 0.023 \pm 0.008$ ,  $t = 4$ , independent of  $\lambda_-$ . Replaced by HAIDT 71.

<sup>96</sup> BOTTERILL 70 is re-evaluation of BOTTERILL 68B with different  $\lambda_+$ .

### $\xi_C = f_-/f_+$ (determined from $\mu$ polarization in $K_{\mu 3}^\pm$ )

The  $\mu$  polarization is a measure of  $\xi(t)$ . No assumptions on  $\lambda_+$  necessary,  $t$  (weighted by sensitivity to  $\xi(t)$ ) should be specified. In  $\lambda_+$ ,  $\xi(0)$  parametrization this is  $\xi(0)$  for  $\lambda_+=0$ .  $d\xi/d\lambda = \xi t$ . For radiative correction to muon polarization in  $K_{\mu 3}^\pm$ , see GINSBERG 71. The parameter  $\xi$  is redundant with  $\lambda_0$  below and is not put into the Meson Summary Table.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.31±0.15 OUR EVALUATION</b>		Error includes scale factor of 1.6. Correlation is $d\xi(0)/d\lambda_+ = -14$ . From a fit discussed in note on $K_{\ell 3}$ form factors in 1982 edition, PL <b>111B</b> (April 1982).			
-0.25±1.20	1585	97 BRAUN	75 HLBC	+	POL, $t=4.2$
-0.95±0.3	3133	98 CUTTS	69 OSPK	+	Total pol. $t=4.0$
-1.0 ± 0.3	6000	99 BETTELS	68 HLBC	+	Total pol. $t=4.9$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.64 \pm 0.27$	40k	100 MERLAN	74 ASPK	+	POL, $d\xi(0)/d\lambda_+$ $= +1.7$
$-1.4 \pm 1.8$	397	101 CALLAHAN	66B FBC	+	Total pol.
$-0.7^{+0.9}_{-3.3}$	2950	101 CALLAHAN	66B FBC	+	Long. pol.
$+1.2^{+2.4}_{-1.8}$	2100	101 BORREANI	65 HLBC	+	Polarization
$-4.0$ to $+1.7$	500	101 CUTTS	65 OSPK	+	Long. pol.

97 BRAUN 75  $d\xi(0)/d\lambda_+ = \xi t = -0.25 \times 4.2 = -1.0$ .

98 CUTTS 69  $t = 4.0$  was calculated from figure 8.  $d\xi(0)/d\lambda_+ = \xi t = -0.95 \times 4 = -3.8$ .

99 BETTELS 68  $d\xi(0)/d\lambda_+ = \xi t = -1.0 \times 4.9 = -4.9$ .

100 MERLAN 74 polarization result (figure 5) not possible. See discussion of polarization experiments in note on " $K_{\ell 3}$  Form Factors" in the 1982 edition of this Review [Physics Letters **111B** (1982)].

101  $t$  value not given.

### Im( $\xi$ ) in $K_{\mu 3}^{\pm}$ DECAY (from transverse $\mu$ pol.)

Test of  $T$  reversal invariance.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>-0.014 \pm 0.014</math> OUR AVERAGE</b>					
$-0.013 \pm 0.016 \pm 0.003$	3.9M	ABE	99S CNTR	+	$p_T K^+$ at rest
$-0.016 \pm 0.025$	20M	CAMPBELL	81 CNTR	+	Pol.
$-0.3^{+0.3}_{-0.4}$	3133	CUTTS	69 OSPK	+	Total pol. fig.7
$-0.1 \pm 0.3$	6000	BETTELS	68 HLBC	+	Total pol.
$0.0 \pm 1.0$	2648	CALLAHAN	66B FBC	+	MU
$+1.6 \pm 1.3$	397	CALLAHAN	66B FBC	+	Total pol.
$0.5^{+1.4}_{-0.5}$	2950	CALLAHAN	66B FBC	+	Long. pol.

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.010 \pm 0.019$	32M	102 BLATT	83 CNTR	Polarization
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102 Combined result of MORSE 80 ( $K_{\mu 3}^0$ ) and CAMPBELL 81 ( $K_{\mu 3}^+$ ).

### $\lambda_+$ (LINEAR ENERGY DEPENDENCE OF $f_+$ IN $K_{\mu 3}^{\pm}$ DECAY)

See also the corresponding entries and footnotes in sections  $\xi_A$ ,  $\xi_C$ , and  $\lambda_0$ . For radiative correction of  $K_{\mu 3}^{\pm}$  Dalitz plot, see GINSBERG 70 and BECHERRAWY 70.

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>0.031 \pm 0.008</math> OUR EVALUATION</b>					

Error includes scale factor of 1.6. From a fit discussed in note on  $K_{\ell 3}$  form factors in 1982 edition, PL **111B** (April 1982).

Average is meaningless.

$0.014 \pm 0.024$	3k	ARTEMOV	97B SPEC	-	DP
$+0.050 \pm 0.013$	3973	WHITMAN	80 SPEC	+	DP
$0.025 \pm 0.030$	490	ARNOLD	74 HLBC	+	DP
$0.027 \pm 0.019$	6527	MERLAN	74 ASPK	+	DP
$0.025 \pm 0.017$	1897	BRAUN	73C HLBC	+	DP
$0.024 \pm 0.019$	4025	103 ANKENBRA...	72 ASPK	+	PI
$-0.006 \pm 0.015$	3480	CHIANG	72 OSPK	+	DP
$0.050 \pm 0.018$	3240	HAIDT	71 HLBC	+	DP
$0.009 \pm 0.026$	2041	KIJEWSKI	69 OSPK	+	PI
$0.0 \pm 0.05$	444	CALLAHAN	66B FBC	+	PI

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.029±0.024        3000    <sup>104</sup> ARTEMOV    97    SPEC   –    DP  
<sup>103</sup> ANKENBRANDT 72  $\lambda_+$  from figure 3 to match  $d\xi(0)/d\lambda_+$ . Text gives 0.024 ± 0.022.  
<sup>104</sup> Superseded by ARTEMOV 97B.

## $\lambda_0$ (LINEAR ENERGY DEPENDENCE OF $f_0$ IN $K_{\mu 3}^\pm$ DECAY)

Wherever possible, we have converted the above values of  $\xi(0)$  into values of  $\lambda_0$  using the associated  $\lambda_+^\mu$  and  $d\xi/d\lambda_+$ .

VALUE	$d\lambda_0/d\lambda_+$	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.006±0.007 OUR EVALUATION</b>			Error includes scale factor of 1.6. Correlation is $d\lambda_0/d\lambda_+ = -0.16$ . From a fit discussed in note on $K_{\ell 3}$ form factors in 1982 edition, PL <b>111B</b> (April 1982).			

Average is meaningless.

+0.058±0.020	0.0	3k	105 ARTEMOV	97B SPEC	–	DP
+0.029±0.011	-0.37	3973	WHITMAN	80 SPEC	+	DP
+0.019±0.010	+0.03	55k	106 HEINTZE	77 SPEC	+	BR
+0.008±0.097	+0.92	1585	107 BRAUN	75 HLBC	+	POL
-0.040±0.040	-0.62	490	ARNOLD	74 HLBC	+	DP
-0.019±0.015	+0.27	6527	108 MERLAN	74 ASPK	+	DP
-0.008±0.020	-0.53	1897	109 BRAUN	73C HLBC	+	DP
-0.026±0.013	+0.03	4025	110 ANKENBRA...	72 ASPK	+	PI
+0.030±0.014	-0.21	3480	110 CHIANG	72 OSPK	+	DP
-0.039±0.029	-1.34	3240	110 HAIDT	71 HLBC	+	DP
-0.056±0.024	+0.69	3133	107 CUTTS	69 OSPK	+	POL
-0.031±0.045	-1.10	2041	110 KIJEWSKI	69 OSPK	+	PI
-0.063±0.024	+0.60	6000	107 BETTELS	68 HLBC	+	POL
+0.058±0.036	-0.37	444	110 CALLAHAN	66B FBC	+	PI

• • • We do not use the following data for averages, fits, limits, etc. • • •

+0.062±0.024	0.0	3000	111 ARTEMOV	97 SPEC	–	DP
-0.017±0.011			112 BRAUN	74 HLBC	+	$K_{\mu 3}/K_{e3}$ vs. $t$

105 ARTEMOV 97B does not give  $d\lambda_0/d\lambda_+$  so we take it to be zero.

106 HEINTZE 77 uses  $\lambda_+ = 0.029 \pm 0.003$ .  $d\lambda_0/d\lambda_+$  estimated by us.

107  $\lambda_0$  value is for  $\lambda_+ = 0.03$  calculated by us from  $\xi(0)$  and  $d\xi(0)/d\lambda_+$ .

108 MERLAN 74  $\lambda_0$  and  $d\lambda_0/d\lambda_+$  were calculated by us from  $\xi_A$ ,  $\lambda_+^\mu$ , and  $d\xi(0)/d\lambda_+$ . Their figure 6 gives  $\lambda_0 = -0.025 \pm 0.012$  and no  $d\lambda_0/d\lambda_+$ .

109 This value and error are taken from BRAUN 75 but correspond to the BRAUN 73C  $\lambda_+^\mu$  result.  $d\lambda_0/d\lambda_+$  is from BRAUN 73C  $d\xi(0)/d\lambda_+$  in  $\xi_A$  above.

110  $\lambda_0$  calculated by us from  $\xi(0)$ ,  $\lambda_+^\mu$ , and  $d\xi(0)/d\lambda_+$ .

111 ARTEMOV 97 does not give  $d\lambda_0/d\lambda_+$  so we take it to be zero. Superseded by ARTEMOV 97B.

112 BRAUN 74 is a combined  $K_{\mu 3}$ - $K_{e3}$  result. It is not independent of BRAUN 73C ( $K_{\mu 3}$ ) and BRAUN 73B ( $K_{e3}$ ) form factor results.

## **$|f_S/f_+|$ FOR $K_{e3}^\pm$ DECAY**

Ratio of scalar to  $f_+$  couplings.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.084±0.023 OUR AVERAGE</b>	Error includes scale factor of 1.2.					
0.070±0.016±0.016		32k	AKIMENKO	91	SPEC	$\lambda_+, f_S, f_T,$ $\phi$ fit
0.00 ±0.10		2827	BRAUN	75	HLBC	+
0.14 +0.03 -0.04		2707	STEINER	71	HLBC	+
						$\lambda_+, f_S, f_T,$ $\phi$ fit

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.13	90	4017	CHIANG	72	OSPK	+
<0.23	90		BOTTERILL	68C	ASPK	
<0.18	90		BELLOTTI	67B	HLBC	
<0.30	95		KALMUS	67	HLBC	+

## **$|f_T/f_+|$ FOR $K_{e3}^\pm$ DECAY**

Ratio of tensor to  $f_+$  couplings.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.38±0.11 OUR AVERAGE</b>	Error includes scale factor of 1.1.					
0.53 +0.09 -0.10 ±0.10		32k	AKIMENKO	91	SPEC	$\lambda_+, f_S, f_T,$ $\phi$ fit
0.07±0.37		2827	BRAUN	75	HLBC	+
0.24 +0.16 -0.14		2707	STEINER	71	HLBC	+
						$\lambda_+, f_S, f_T,$ $\phi$ fit

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.75	90	4017	CHIANG	72	OSPK	+
<0.58	90		BOTTERILL	68C	ASPK	
<0.58	90		BELLOTTI	67B	HLBC	
<1.1	95		KALMUS	67	HLBC	+

## **$f_T/f_+$ FOR $K_{\mu 3}^\pm$ DECAY**

Ratio of tensor to  $f_+$  couplings.

VALUE	EVTS	DOCUMENT ID	TECN
<b>0.02±0.12</b>	1585	BRAUN	75

## **DECAY FORM FACTORS FOR $K^\pm \rightarrow \pi^+ \pi^- e^\pm \nu_e$**

Given in ROSSELET 77, BEIER 73, and BASILE 71C.

## **DECAY FORM FACTOR FOR $K^\pm \rightarrow \pi^0 \pi^0 e^\pm \nu$**

Given in BOLOTOV 86B and BARMIN 88B.

## $K^\pm \rightarrow \ell^\pm \nu \gamma$ FORM FACTORS

For definitions of the axial-vector  $F_A$  and vector  $F_V$  form factor, see the "Note on  $\pi^\pm \rightarrow \ell^\pm \nu \gamma$  and  $K^\pm \rightarrow \ell^\pm \nu \gamma$  Form Factors" in the  $\pi^\pm$  section. In the kaon literature, often different definitions  $a_K = F_A/m_K$  and  $v_K = F_V/m_K$  are used.

### $F_A + F_V$ , SUM OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR $K \rightarrow e \nu e \gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.148±0.010 OUR AVERAGE</b>				
0.147±0.011	51	113 HEINTZE	79 SPEC	$K \rightarrow e \nu \gamma$
0.150 <sup>+0.018</sup> <sub>-0.023</sub>	56	114 HEARD	75 SPEC	$K \rightarrow e \nu \gamma$

113 HEINTZE 79 quotes absolute value of  $|F_A + F_V| \sin\theta_C$ . We use  $\sin\theta_C = V_{us} = 0.2205$ .  
 114 HEARD 75 quotes absolute value of  $|F_A + F_V| \sin\theta_C$ . We use  $\sin\theta_C = V_{us} = 0.2205$ .

### $F_A + F_V$ , SUM OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR $K \rightarrow \mu \nu \mu \gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
< 0.23	90	115 AKIBA	85 SPEC	$K \rightarrow \mu \nu \gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
-1.2 to 1.1	90	DEMIDOV	90 XEBC	$K \rightarrow \mu \nu \gamma$

115 AKIBA 85 quotes absolute value.

### $F_A - F_V$ , DIFFERENCE OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR $K \rightarrow e \nu e \gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<0.49	90	116 HEINTZE	79 SPEC	$K \rightarrow e \nu \gamma$

116 HEINTZE 79 quotes  $|F_A - F_V| < \sqrt{11} |F_A + F_V|$ .

### $F_A - F_V$ , DIFFERENCE OF AXIAL-VECTOR AND VECTOR FORM FACTOR FOR $K \rightarrow \mu \nu \mu \gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>-2.2 to 0.3 OUR EVALUATION</b>				
-2.2 to 0.6	90	DEMIDOV	90 XEBC	$K \rightarrow \mu \nu \gamma$
-2.5 to 0.3	90	AKIBA	85 SPEC	$K \rightarrow \mu \nu \gamma$

## $K^\pm$ REFERENCES

ADLER	00	PRL 84 3768	S. Adler <i>et al.</i>	(BNL 787 Collab.)
MA	00	PRL 84 2580	H. Ma <i>et al.</i>	(BNL 865 Collab.)
ABE	99S	PRL 83 4253	M. Abe <i>et al.</i>	(KEK-E246 Collab.)
APPEL	99	PRL 83 4482	R. Appel <i>et al.</i>	(BNL 865 Collab.)
ADLER	98	PR D58 012003	S. Adler <i>et al.</i>	(BNL 787 Collab.)
BATUSOV	98	NP B516 3	V.Y. Batusov <i>et al.</i>	
ADLER	97	PRL 79 2204	S. Adler <i>et al.</i>	(BNL 787 Collab.)
ADLER	97C	PRL 79 4756	S. Adler <i>et al.</i>	(BNL 787 Collab.)
ARTEMOV	97	PAN 60 218	V.M. Artemov <i>et al.</i>	(JINR)
		Translated from YAF 60 277.		
ARTEMOV	97B	PAN 60 2023	V.M. Artemov <i>et al.</i>	
		Translated from YAF 60 2205.		

KITCHING	97	PRL 79 4079	P. Kitching <i>et al.</i>	(BNL 787 Collab.)
ADLER	96	PRL 76 1421	S. Adler <i>et al.</i>	(BNL 787 Collab.)
KOPTEV	95	JETPL 61 877	V.P. Koptev <i>et al.</i>	(PNPI)
		Translated from ZETFP 61 865.		
AOKI	94	PR D50 69	M. Aoki <i>et al.</i>	(INUS, KEK, TOKMS)
ATIYA	93	PRL 70 2521	M.S. Atiya <i>et al.</i>	(BNL 787 Collab.)
Also	93C	PRL 71 305 (erratum)	M.S. Atiya <i>et al.</i>	(BNL 787 Collab.)
ATIYA	93B	PR D48 R1	M.S. Atiya <i>et al.</i>	(BNL 787 Collab.)
BIJNENS	93	NP B396 81	J. Bijnens, G. Ecker, J. Gasser	(CERN, BERN)
ALLIEGRO	92	PRL 68 278	C. Alliegro <i>et al.</i>	(BNL, FNAL, PSI+)
BARMIN	92	SJNP 55 547	V.V. Barmin <i>et al.</i>	(ITEP)
		Translated from YAF 55 976.		
IMAZATO	92	PRL 69 877	J. Imazato <i>et al.</i>	(KEK, INUS, TOKY+)
IVANOV	92	THESIS	Ivanov	(PNPI)
LITTENBERG	92	PRL 68 443	L.S. Littenberg, R.E. Shrock	(BNL, STON)
USHER	92	PR D45 3961	T. Usher <i>et al.</i>	(UCI)
AKIMENKO	91	PL B259 225	S.A. Akimenko <i>et al.</i>	(SERP, JINR, TBIL+)
BARMIN	91	SJNP 53 606	V.V. Barmin <i>et al.</i>	(ITEP)
		Translated from YAF 53 981.		
DENISOV	91	JETPL 54 558	A.S. Denisov <i>et al.</i>	(PNPI)
		Translated from ZETFP 54 557.		
Also	92	THESIS	Ivanov	(PNPI)
ATIYA	90	PRL 64 21	M.S. Atiya <i>et al.</i>	(BNL 787 Collab.)
ATIYA	90B	PRL 65 1188	M.S. Atiya <i>et al.</i>	(BNL 787 Collab.)
DEMIDOV	90	SJNP 52 1006	V.S. Demidov <i>et al.</i>	(ITEP)
		Translated from YAF 52 1595.		
LEE	90	PRL 64 165	A.M. Lee <i>et al.</i>	(BNL, FNAL, VILL, WASH+)
ATIYA	89	PRL 63 2177	M.S. Atiya <i>et al.</i>	(BNL 787 Collab.)
BARMIN	89	SJNP 50 421	V.V. Barmin <i>et al.</i>	(ITEP)
		Translated from YAF 50 679.		
BARMIN	88	SJNP 47 643	V.V. Barmin <i>et al.</i>	(ITEP)
		Translated from YAF 47 1011.		
BARMIN	88B	SJNP 48 1032	V.V. Barmin <i>et al.</i>	(ITEP)
		Translated from YAF 48 1719.		
BOLOTOV	88	JETPL 47 7	V.N. Bolotov <i>et al.</i>	(ASCI)
		Translated from ZETFP 47 8.		
CAMPAGNARI	88	PRL 61 2062	C. Campagnari <i>et al.</i>	(BNL, FNAL, PSI+)
GALL	88	PRL 60 186	K.P. Gall <i>et al.</i>	(BOST, MIT, WILL, CIT+)
BARMIN	87	SJNP 45 62	V.V. Barmin <i>et al.</i>	(ITEP)
		Translated from YAF 45 97.		
BOLOTOV	87	SJNP 45 1023	V.N. Bolotov <i>et al.</i>	(INRM)
		Translated from YAF 45 1652.		
BOLOTOV	86	SJNP 44 73	V.N. Bolotov <i>et al.</i>	(INRM)
		Translated from YAF 44 117.		
BOLOTOV	86B	SJNP 44 68	V.N. Bolotov <i>et al.</i>	(INRM)
		Translated from YAF 44 108.		
YAMANAKA	86	PR D34 85	T. Yamanaka <i>et al.</i>	(KEK, TOKY)
Also	84	PRL 52 329	R.S. Hayano <i>et al.</i>	(TOKY, KEK)
AKIBA	85	PR D32 2911	Y. Akiba <i>et al.</i>	(TOKY, TINT, TSUK, KEK)
BOLOTOV	85	JETPL 42 481	V.N. Bolotov <i>et al.</i>	(INRM)
		Translated from ZETFP 42 390.		
BLATT	83	PR D27 1056	S.R. Blatt <i>et al.</i>	(YALE, BNL)
ASANO	82	PL 113B 195	Y. Asano <i>et al.</i>	(KEK, TOKY, INUS, OSAK)
COOPER	82	PL 112B 97	A.M. Cooper <i>et al.</i>	(RL)
PDG	82	PL 111B	M. Roos <i>et al.</i>	(HELS, CIT, CERN)
PDG	82B	PL 111B 70	M. Roos <i>et al.</i>	(HELS, CIT, CERN)
ASANO	81B	PL 107B 159	Y. Asano <i>et al.</i>	(KEK, TOKY, INUS, OSAK)
CAMPBELL	81	PRL 47 1032	M.K. Campbell <i>et al.</i>	(YALE, BNL)
Also	83	PR D27 1056	S.R. Blatt <i>et al.</i>	(YALE, BNL)
LUM	81	PR D23 2522	G.K. Lum <i>et al.</i>	(LBL, NBS+)
LYONS	81	ZPHY C10 215	L. Lyons, C. Albajar, G. Myatt	(OXF)
MORSE	80	PR D21 1750	W.M. Morse <i>et al.</i>	(BNL, YALE)
WHITMAN	80	PR D21 652	R. Whitman <i>et al.</i>	(ILLC, BNL, ILL)
BARKOV	79	NP B148 53	L.M. Barkov <i>et al.</i>	(NOVO, KIAE)
HEINTZE	79	NP B149 365	J. Heintze <i>et al.</i>	(HEIDP, CERN)
ABRAMS	77	PR D15 22	R.J. Abrams <i>et al.</i>	(BNL)
DEVAUX	77	NP B126 11	B. Devaux <i>et al.</i>	(SACL, GEVA)
HEINTZE	77	PL 70B 482	J. Heintze <i>et al.</i>	(HEIDP, CERN)
ROSSELET	77	PR D15 574	L. Rosselet <i>et al.</i>	(GEVA, SACL)
BERTRAND	76	NP B114 387	D. Bertrand <i>et al.</i>	(BRUX, KIDR, DUUC+)
BLOCH	76	PL 60B 393	P. Bloch <i>et al.</i>	(GEVA, SACL)
BRAUN	76B	LNC 17 521	H.M. Braun <i>et al.</i>	(AACH3, BARI, BELG+)
DIAMANT-...	76	PL 62B 485	A.M. Diamant-Berger <i>et al.</i>	(SACL, GEVA)
HEINTZE	76	PL 60B 302	J. Heintze <i>et al.</i>	(HEIDP)

SMITH	76	NP B109 173	K.M. Smith <i>et al.</i>	(GLAS, LIVP, OXF+)
WEISSENBE...	76	NP B115 55	A.O. Weissenberg <i>et al.</i>	(ITEP, LEBD)
BLOCH	75	PL 56B 201	P. Bloch <i>et al.</i>	(SACL, GEVA)
BRAUN	75	NP B89 210	H.M. Braun <i>et al.</i>	(AACH3, BARI, BRUX+)
CHENG	75	NP A254 381	S.C. Cheng <i>et al.</i>	(COLU, YALE)
HEARD	75	PL 55B 324	K.S. Heard <i>et al.</i>	(CERN, HEIDH)
HEARD	75B	PL 55B 327	K.S. Heard <i>et al.</i>	(CERN, HEIDH)
SHEAFF	75	PR D12 2570	M. Sheaff	(WISC)
SMITH	75	NP B91 45	K.M. Smith <i>et al.</i>	(GLAS, LIVP, OXF+)
ARNOLD	74	PR D9 1221	C.L. Arnold, B.P. Roe, D. Sinclair	(MICH)
BRAUN	74	PL 51B 393	H.M. Braun <i>et al.</i>	(AACH3, BARI, BRUX+)
CENCE	74	PR D10 776	R.J. Cence <i>et al.</i>	(HAWA, LBL, WISC)
Also	73	Thethesis unpub.	D.B. Clarke	(WISC)
KUNSELMAN	74	PR C9 2469	R. Kunselman	(WYOM)
MERLAN	74	PR D9 107	S. Merlan <i>et al.</i>	(YALE, BNL, LASL)
WEISSENBE...	74	PL 48B 474	A.O. Weissenberg <i>et al.</i>	(ITEP, LEBD)
ABRAMS	73B	PRL 30 500	R.J. Abrams <i>et al.</i>	(BNL)
BACKENSTO...	73	PL 43B 431	G. Backenstoss <i>et al.</i>	(CERN, KARLK, KARLE+)
BEIER	73	PRL 30 399	E.W. Beier <i>et al.</i>	(PENN)
BRAUN	73B	PL 47B 185	H.M. Braun, M. Cornelssen	(AACH3, BARI, BRUX+)
Also	75	NP B89 210	H.M. Braun <i>et al.</i>	(AACH3, BARI, BRUX+)
BRAUN	73C	PL 47B 182	H.M. Braun, M. Cornelssen	(AACH3, BARI, BRUX+)
Also	75	NP B89 210	H.M. Braun <i>et al.</i>	(AACH3, BARI, BRUX+)
CABLE	73	PR D8 3807	G.D. Cable <i>et al.</i>	(EFI, LBL)
LJUNG	73	PR D8 1307	D. Ljung, D. Cline	(WISC)
Also	72	PRL 28 523	D. Ljung	(WISC)
Also	72	PRL 28 1287	D. Cline, D. Ljung	(WISC)
Also	69	PRL 23 326	U. Camerini <i>et al.</i>	(WISC)
LUCAS	73	PR D8 719	P.W. Lucas, H.D. Taft, W.J. Willis	(YALE)
LUCAS	73B	PR D8 727	P.W. Lucas, H.D. Taft, W.J. Willis	(YALE)
PANG	73	PR D8 1989	C.Y. Pang <i>et al.</i>	(EFI, ARIZ, LBL)
Also	72	PL 40B 699	G.D. Cable <i>et al.</i>	(EFI, LBL)
SMITH	73	NP B60 411	K.M. Smith <i>et al.</i>	(GLAS, LIVP, OXF+)
ABRAMS	72	PRL 29 1118	R.J. Abrams <i>et al.</i>	(BNL)
ANKENBRA...	72	PRL 28 1472	C.M. Ankenbrandt <i>et al.</i>	(BNL, LASL, FNAL+)
AUBERT	72	NC 12A 509	B. Aubert <i>et al.</i>	(ORSAY, BRUX, EPOL)
BEIER	72	PRL 29 678	E.W. Beier <i>et al.</i>	(PENN)
CHIANG	72	PR D6 1254	I.H. Chiang <i>et al.</i>	(ROCH, WISC)
CLARK	72	PRL 29 1274	A.R. Clark <i>et al.</i>	(LBL)
EDWARDS	72	PR D5 2720	R.T. Edwards <i>et al.</i>	(ILL)
FORD	72	PL 38B 335	W.T. Ford <i>et al.</i>	(PRIN)
HOFFMASTER	72	NP B36 1	S. Hoffmaster <i>et al.</i>	(STEV, SETO, LEHI)
BASILE	71C	PL 36B 619	P. Basile <i>et al.</i>	(SACL, GEVA)
BOURQUIN	71	PL 36B 615	M.H. Bourquin <i>et al.</i>	(GEVA, SACL)
GINSBERG	71	PR D4 2893	E.S. Ginsberg	(MIT)
HAIDT	71	PR D3 10	D. Haidt	(AACH, BARI, CERN, EPOL, NIJM+)
Also	69	PL 29B 691	D. Haidt <i>et al.</i>	(AACH, BARI, CERN, EPOL+)
KLEMS	71	PR D4 66	J.H. Klems, R.H. Hildebrand, R. Stiening	(CHIC+)
Also	70	PRL 24 1086	J.H. Klems, R.H. Hildebrand, R. Stiening	(LRL+)
Also	70B	PRL 25 473	J.H. Klems, R.H. Hildebrand, R. Stiening	(LRL+)
OTT	71	PR D3 52	R.J. Ott, T.W. Pritchard	(LOQM)
ROMANO	71	PL 36B 525	F. Romano <i>et al.</i>	(BARI, CERN, ORSAY)
SCHWEINB...	71	PL 36B 246	W. Schweinberger	(AACH, BELG, CERN, NIJM+)
STEINER	71	PL 36B 521	H.J. Steiner	(AACH, BARI, CERN, EPOL, ORSAY+)
BARDIN	70	PL 32B 121	D.Y. Bardin, S.N. Bilenky, B.M. Pontecorvo	(JINR)
BECHERRAWY	70	PR D1 1452	T. Becherrawy	(ROCH)
BOTTERILL	70	PL 31B 325	D.R. Botterill <i>et al.</i>	(OXF)
FORD	70	PRL 25 1370	W.T. Ford <i>et al.</i>	(PRIN)
GAILLARD	70	CERN 70-14	J.M. Gaillard, L.M. Chouinet	(CERN, ORSAY)
GINSBERG	70	PR D1 229	E.S. Ginsberg	(HAIF)
GRAUMAN	70	PR D1 1277	J. Grauman <i>et al.</i>	(STEV, SETO, LEHI)
Also	69	PRL 23 737	J.U. Grauman <i>et al.</i>	(STEV, SETO, LEHI)
MALTSEV	70	SJNP 10 678	E.I. Maltsev <i>et al.</i>	(JINR)

Translated from YAF 10 1195.

PANDOULAS	70	PR D2 1205	D. Pandoulas <i>et al.</i>	(STEV, SETO)
CUTTS	69	PR 184 1380	D. Cutts <i>et al.</i>	(LRL, MIT)
Also	68	PRL 20 955	D. Cutts <i>et al.</i>	(LRL, MIT)
DAVISON	69	PR 180 1333	D.C. Davison <i>et al.</i>	(UCR)
ELY	69	PR 180 1319	R.P.J. Ely <i>et al.</i>	(LOUC, WISC, LRL)
EMMERSOHN	69	PRL 23 393	J.M.L. Emmerson, T.W. Quirk	(OXF)
HERZO	69	PR 186 1403	D. Herzo <i>et al.</i>	(ILL)
KIJEWSKI	69	Thesis UCRL 18433	P.K. Kijewski	(LBL)
LOBKOWICZ	69	PR 185 1676	F. Lobkowicz <i>et al.</i>	(ROCH, BNL)
Also	66	PRL 17 548	F. Lobkowicz <i>et al.</i>	(ROCH, BNL)
MACEK	69	PRL 22 32	R.J. Macek <i>et al.</i>	(PENN, TEMP)
MAST	69	PR 183 1200	T.S. Mast <i>et al.</i>	(LRL)
SELLERI	69	NC 60A 291	F. Selleri	
ZELLER	69	PR 182 1420	M.E. Zeller <i>et al.</i>	(UCLA, LRL)
BETTELS	68	NC 56A 1106	J. Bettels	(AACH, BARI, BERG, CERN, EPOL+)
Also	71	PR D3 10	D. Haidt	(AACH, BARI, CERN, EPOL, NIJM+)
BOTTERILL	68B	PRL 21 766	D.R. Botterill <i>et al.</i>	(OXF)
BOTTERILL	68C	PR 174 1661	D.R. Botterill <i>et al.</i>	(OXF)
BUTLER	68	UCRL 18420	W.D. Butler <i>et al.</i>	(LRL)
CHANG	68	PRL 20 510	C.Y. Chang <i>et al.</i>	(UMD, RUTG)
CHEN	68	PRL 20 73	M. Chen <i>et al.</i>	(LRL, MIT)
EICHENTH	68	PL 27B 586	T. Eichten	(AACH, BARI, CERN, EPOL, ORSAY+)
EISLER	68	PR 169 1090	F.R. Eisler <i>et al.</i>	(RUTG)
ESCHSTRUTH	68	PR 165 1487	P.T. Eschstruth <i>et al.</i>	(PRIN, PENN)
GARLAND	68	PR 167 1225	R. Garland <i>et al.</i>	(COLU, RUTG, WISC)
MOSCOSO	68	Thesis	L. Moscoso	(ORSAY)
AUERBACH	67	PR 155 1505	L.B. Auerbach <i>et al.</i>	(PENN, PRIN)
Also	74	PR D9 3216	L.B. Auerbach	
Erratum.				
BELLOTTI	67	Heidelberg Conf.	E. Bellotti, A. Pullia	(MILA)
BELLOTTI	67B	NC 52A 1287	E. Bellotti, E. Fiorini, A. Pullia	(MILA)
Also	66B	PL 20 690	E. Bellotti <i>et al.</i>	(MILA)
BISI	67	PL 25B 572	V. Bisi <i>et al.</i>	(TORI)
BOTTERILL	67	PRL 19 982	D.R. Botterill <i>et al.</i>	(OXF)
Also	68	PR 171 1402	D.R. Botterill <i>et al.</i>	(OXF)
BOWEN	67B	PR 154 1314	D.R. Bowen <i>et al.</i>	(PPA)
CLINE	67B	Hercog Novi Tbl. 4	D. Cline	
Proc. International School on Elementary Particle Physics.				
FLETCHER	67	PRL 19 98	C.R. Fletcher <i>et al.</i>	(ILL)
FORD	67	PRL 18 1214	W.T. Ford <i>et al.</i>	(PRIN)
GINSBERG	67	PR 162 1570	E.S. Ginsberg	(MASB)
IMLAY	67	PR 160 1203	R.L. Imlay <i>et al.</i>	(PRIN)
KALMUS	67	PR 159 1187	G.E. Kalmus, A. Kernan	(LRL)
ZINCHENKO	67	Thesis Rutgers	A.I. Zinchenko	(RUTG)
CALLAHAN	66	NC 44A 90	A.C. Callahan	(WISC)
CALLAHAN	66B	PR 150 1153	A.C. Callahan <i>et al.</i>	(WISC, LRL, UCR+)
CESTER	66	PL 21 343	R. Cester <i>et al.</i>	(PPA)
See footnote 1 in AUERBACH 67.				
Also	67	PR 155 1505	L.B. Auerbach <i>et al.</i>	(PENN, PRIN)
BIRGE	65	PR 139B 1600	R.W. Birge <i>et al.</i>	(LRL, WISC)
BISI	65	NC 35 768	V. Bisi <i>et al.</i>	(TORI)
BISI	65B	PR 139B 1068	V. Bisi <i>et al.</i>	(TORI)
BORREANI	65	PR 140B 1686	G. Borreani <i>et al.</i>	(BARI, TORI)
CALLAHAN	65	PRL 15 129	A. Callahan, D. Cline	(WISC)
CAMERINI	65	NC 37 1795	U. Camerini <i>et al.</i>	(WISC, LRL)
CLINE	65	PL 15 293	D. Cline, W.F. Fry	(WISC)
CUTTS	65	PR 138B 969	D. Cutts, T. Elliott, R. Stiening	(LRL)
DEMARCO	65	PR 140B 1430	A. de Marco, C. Grossi, G. Rinaudo	(TORI, CERN)
FITCH	65B	PR 140B 1088	V.L. Fitch, C.A. Quarles, H.C. Wilkins	(PRIN+)
GREINER	65	ARNS 15 67	D. Greiner	(LRL)
STAMER	65	PR 138B 440	P. Stamer <i>et al.</i>	(STEV)
TRILLING	65B	UCRL 16473	G.N. Trilling	(LRL)
Updated from 1965 Argonne Conference, page 5.				
YOUNG	65	Thesis UCRL 16362	P.S. Young	(LRL)
Also	67	PR 156 1464	P.S. Young, W.Z. Osborne, W.H. Barkas	(LRL)
BORREANI	64	PL 12 123	G. Borreani, G. Rinaudo, A.E. Werbrouck	(TORI)
CALLAHAN	64	PR 136B 1463	A. Callahan, R. March, R. Stark	(WISC)
CAMERINI	64	PRL 13 318	U. Camerini <i>et al.</i>	(WISC, LRL)
CLINE	64	PRL 13 101	D. Cline, W.F. Fry	(WISC)

GIACOMELLI	64	NC 34 1134	G. Giacomelli <i>et al.</i>	(BGNA, MUNI)
GREINER	64	PRL 13 284	D.E. Greiner, W.Z. Osborne, W.H. Barkas	(LRL)
JENSEN	64	PR 136B 1431	G.L. Jensen <i>et al.</i>	(MICH)
KALMUS	64	PRL 13 99	G.E. Kalmus <i>et al.</i>	(LRL, WISC)
SHAKLEE	64	PR 136B 1423	F.S. Shaklee <i>et al.</i>	(MICH)
BARKAS	63	PR 11 26	W.H. Barkas, J.N. Dyer, H.H. Heckman	(LRL)
BOYARSKI	62	PR 128 2398	A.M. Boyarski <i>et al.</i>	(MIT)
BROWN	62B	PRL 8 450	J.L. Brown <i>et al.</i>	(LRL, MICH)
BARKAS	61	PR 124 1209	W.H. Barkas <i>et al.</i>	(LRL)
BHOWMIK	61	NC 20 857	B. Bhowmik, P.C. Jain, P.C. Mathur	(DELH)
FERRO-LUZZI	61	NC 22 1087	M. Ferro-Luzzi <i>et al.</i>	(LRL)
NORDIN	61	PR 123 2166	P. Nordin	(LRL)
ROE	61	PRL 7 346	B.P. Roe <i>et al.</i>	(MICH, LRL)
FREDEN	60B	PR 118 564	S.C. Freden, F.C. Gilbert, R.S. White	(LRL)
BURROWES	59	PRL 2 117	H.C. Burrowes <i>et al.</i>	(MIT)
TAYLOR	59	PR 114 359	S. Taylor <i>et al.</i>	(COLU)
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Also	65	PL 14 72	N. Cabibbo, Maksymowicz	(CERN)
BIRGE	63	PRL 11 35	R.W. Birge <i>et al.</i>	(LRL, WISC, BARI)
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