

n $I(J^P) = \frac{1}{2}(\frac{1}{2}^+)$ Status: * * * *

We have omitted some results that have been superseded by later experiments. See our earlier editions.

***n* MASS**

The mass is known much more precisely in u (atomic mass units) than in MeV; see the footnotes. The conversion from u to MeV, $1\text{ u} = 931.494013 \pm 0.000037\text{ MeV}/c^2$ (MOHR 99, the 1998 CODATA value), involves the relatively poorly known electronic charge.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
939.56533 ±0.00004 OUR AVERAGE			
939.56530±0.000038	¹ MOHR	99	RVUE 1998 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
939.56531±0.000037	² KESSLER	99	SPEC $np \rightarrow d\gamma$
939.56565 ±0.00028	^{3,4} DIFILIPPO	94	TRAP Penning trap
939.56563 ±0.00028	⁵ COHEN	87	RVUE 1986 CODATA value
939.56564 ±0.00028	^{4,6} GREENE	86	SPEC $np \rightarrow d\gamma$
939.5731 ±0.0027	⁴ COHEN	73	RVUE 1973 CODATA value

¹ The mass is known much more precisely in u: $m = 1.00866491578 \pm 0.00000000055\text{ u}$.

² We use the 1998 CODATA u-to-MeV conversion factor (see the heading above) to get this mass in MeV from the much more precisely measured KESSLER 99 value of $1.00866491637 \pm 0.00000000082\text{ u}$.

³ The mass is known much more precisely in u: $m = 1.0086649235 \pm 0.0000000023\text{ u}$.
We use the 1986 CODATA conversion factor to get the mass in MeV.

⁴ These determinations are not independent of the $m_n - m_p$ measurements below.

⁵ The mass is known much more precisely in u: $m = 1.008664904 \pm 0.000000014\text{ u}$.

⁶ The mass is known much more precisely in u: $m = 1.008664919 \pm 0.000000014\text{ u}$.

***ñ* MASS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
939.485±0.051	59	⁷ CRESTI	86	HBC $\bar{p}p \rightarrow \bar{n}n$

⁷ This is a corrected result (see the erratum). The error is statistical. The maximum systematic error is 0.029 MeV.

$$(m_n - m_{\bar{n}})/m_n$$

A test of *CPT* invariance. Calculated from the *n* and *ñ* masses, above.

VALUE	DOCUMENT ID
(9±5) × 10⁻⁵ OUR EVALUATION	

$m_n - m_p$

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
1.2933318±0.0000005 OUR AVERAGE			
1.2933318±0.0000005	⁸ MOHR	99	RVUE 1998 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
1.293318 ± 0.000009	⁹ COHEN	87	RVUE 1986 CODATA value
1.293328±0.0000072	GREENE	86	SPEC $np \rightarrow d\gamma$
1.293429 ± 0.000036	COHEN	73	RVUE 1973 CODATA value
⁸ Calculated by us from the MOHR 99 ratio $m_n/m_p = 1.00137841887 \pm 0.00000000058$.			
In u, $m_n - m_p = (1.3884489 \pm 0.0000006) \times 10^{-3}$ u.			
⁹ Calculated by us from the COHEN 87 ratio $m_n/m_p = 1.001378404 \pm 0.000000009$. In u, $m_n - m_p = 0.001388434 \pm 0.000000009$ u.			

 n MEAN LIFE

We now compile only direct measurements of the lifetime, not those inferred from decay correlation measurements. (Limits on lifetimes for *bound* neutrons are given in the section “ p PARTIAL MEAN LIVES.”)

For a review, see EROZOLIMSKII 89 and papers that follow it in an issue of NIM devoted to the “Proceedings of the International Workshop on Fundamental Physics with Slow Neutrons” (Grenoble 1989). For later reviews and/or commentary, see FREEDMAN 90, SCHRECKENBACH 92, and PENDLEBURY 93.

VALUE (s)	DOCUMENT ID	TECN	COMMENT
886.7± 1.9 OUR AVERAGE Error includes scale factor of 1.2.			
889.2± 3.0± 3.8	BYRNE	96	CNTR Penning trap
882.6± 2.7	¹⁰ MAMPE	93	CNTR Gravitational trap
888.4± 3.1± 1.1	NESVIZHEV...	92	CNTR Gravitational trap
878 ± 27 ± 14	KOSSAKOW...	89	TPC Pulsed beam
887.6± 3.0	MAMPE	89	CNTR Gravitational trap
877 ± 10	PAUL	89	CNTR Storage ring
876 ± 10 ± 19	LAST	88	SPEC Pulsed beam
891 ± 9	SPIVAK	88	CNTR Beam
903 ± 13	KOSVINTSEV	86	CNTR Gravitational trap
918 ± 14	CHRISTENSEN	72	CNTR
• • • We do not use the following data for averages, fits, limits, etc. • • •			
888.4± 2.9	ALFIMENKOV	90	CNTR See NESVIZHEVSKII 92
893.6± 3.8± 3.7	BYRNE	90	CNTR See BYRNE 96
937 ± 18	¹¹ BYRNE	80	CNTR
875 ± 95	KOSVINTSEV	80	CNTR
881 ± 8	BONDARENKO...	78	CNTR See SPIVAK 88

¹⁰ IGNATOVICH 95 calls into question some of the corrections and averaging procedures used by MAMPE 93. The response, BONDARENKO 96, denies the validity of the criticisms.

¹¹ This measurement has been withdrawn (J. Byrne, private communication, 1990).

***n* MAGNETIC MOMENT**

See the "Note on Baryon Magnetic Moments" in the Λ Listings.

VALUE (μ_N)	DOCUMENT ID	TECN	COMMENT
-1.9130427 ± 0.0000005 OUR AVERAGE			
-1.91304272 ± 0.00000045	MOHR	99	RVUE 1998 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-1.91304275 ± 0.00000045	COHEN	87	RVUE 1986 CODATA value
-1.91304277 ± 0.00000048	12 GREENE	82	MRS
12 GREENE 82 measures the moment to be $(1.04187564 \pm 0.00000026) \times 10^{-3}$ Bohr magnetons. The value above is obtained by multiplying this by $m_p/m_e = 1836.152701 \pm 0.000037$ (the 1986 CODATA value from COHEN 87).			

***n* ELECTRIC DIPOLE MOMENT d_n**

A nonzero value is forbidden by both T invariance and P invariance.

A number of early results have been omitted. See RAMSEY 90 and GOLUB 94 for reviews.

VALUE (10^{-25} ecm)	CL%	DOCUMENT ID	TECN	COMMENT
< 0.63 (CL = 90%)				
< 0.63	90	13 HARRIS	99	MRS $d = (-0.1 \pm 0.36) \times 10^{-25}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 0.97	90	ALTAREV	96	MRS $(+0.26 \pm 0.40 \pm 0.16) \times 10^{-25}$
< 1.1	95	ALTAREV	92	MRS See ALTAREV 96
< 1.2	95	SMITH	90	MRS See HARRIS 99
< 2.6	95	ALTAREV	86	MRS $d = (-1.4 \pm 0.6) \times 10^{-25}$
0.3 ± 4.8		PENDLEBURY	84	MRS Ultracold neutrons
< 6	90	ALTAREV	81	MRS $d = (2.1 \pm 2.4) \times 10^{-25}$
< 16	90	ALTAREV	79	MRS $d = (4.0 \pm 7.5) \times 10^{-25}$

13 This HARRIS 99 result includes the result of SMITH 90. However, the averaging of the results of these two experiments has been criticized by LAMOREAUX 00.

***n* ELECTRIC POLARIZABILITY α_n**

Following is the electric polarizability α_n defined in terms of the induced electric dipole moment by $\mathbf{D} = 4\pi\epsilon_0\alpha_n\mathbf{E}$. For a review, see SCHMIED-MAYER 89.

VALUE (10^{-3} fm 3)	DOCUMENT ID	TECN	COMMENT
0.98 ± 0.19 OUR AVERAGE	Error includes scale factor of 1.1.		
0.0 ± 0.5	14 KOESTER	95	CNTR n Pb, n Bi transmission
1.20 ± 0.15 ± 0.20	SCHMIEDM...	91	CNTR n Pb transmission
1.07 ± 0.33 -1.07	ROSE	90B	CNTR $\gamma d \rightarrow \gamma np$
0.8 ± 1.0	KOESTER	88	CNTR n Pb, n Bi transmission
1.2 ± 1.0	SCHMIEDM...	88	CNTR n Pb, n C transmission
• • • We do not use the following data for averages, fits, limits, etc. • • •			

$1.17^{+0.43}_{-1.17}$

ROSE

90 CNTR See ROSE 90B

¹⁴ KOESTER 95 uses natural Pb and the isotopes 208, 207, and 206. See this paper for a discussion of methods used by various groups to extract α_n from data.

n CHARGE

See also " $|q_p + q_e|/e$ " in the proton Listings.

VALUE ($10^{-21} e$)	DOCUMENT ID	TECN	COMMENT
- 0.4 ± 1.1	15 BAUMANN 88		Cold <i>n</i> deflection
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-15 ± 22	16 GAEHLER 82	CNTR	Reactor neutrons
¹⁵ The BAUMANN 88 error ±1.1 gives the 68% CL limits about the the value -0.4.			
¹⁶ The GAEHLER 82 error ±22 gives the 90% CL limits about the the value -15.			

LIMIT ON $n\bar{n}$ OSCILLATIONS

Mean Time for $n\bar{n}$ Transition in Vacuum

A test of $\Delta B=2$ baryon number nonconservation. MOHAPATRA 80 and MOHAPATRA 89 discuss the theoretical motivations for looking for $n\bar{n}$ oscillations. DOVER 83 and DOVER 85 give phenomenological analyses. The best limits come from looking for the decay of neutrons bound in nuclei. However, these analyses require model-dependent corrections for nuclear effects. See KABIR 83, DOVER 89, and ALBERICO 91 for discussions. Direct searches for $n \rightarrow \bar{n}$ transitions using reactor neutrons are cleaner but give somewhat poorer limits. We include limits for both free and bound neutrons in the Summary Table.

VALUE (s)	CL %	DOCUMENT ID	TECN	COMMENT
>8.6 × 10⁷ (CL = 90%)				
>8.6 × 10 ⁷	90	BALDO...	94	CNTR Reactor (free) neutrons
>1.2 × 10 ⁸	90	BERGER	90	FREJ <i>n</i> bound in iron
>1.2 × 10 ⁸	90	TAKITA	86	CNTR Kamiokande
• • • We do not use the following data for averages, fits, limits, etc. • • •				
>1 × 10 ⁷	90	BALDO...	90	CNTR See BALDO-CEOLIN 94
>4.9 × 10 ⁵	90	BRESSI	90	CNTR Reactor neutrons
>4.7 × 10 ⁵	90	BRESSI	89	CNTR See BRESSI 90
>1 × 10 ⁶	90	FIDECARO	85	CNTR Reactor neutrons
>8.8 × 10 ⁷	90	PARK	85B	CNTR
>3 × 10 ⁷		BATTISTONI	84	NUSX
>2.7 × 10 ⁷ –1.1 × 10 ⁸		JONES	84	CNTR
>2 × 10 ⁷		CHERRY	83	CNTR

***n* DECAY MODES**

Mode	Fraction (Γ_i/Γ)	Confidence level
$\Gamma_1 \quad p e^- \bar{\nu}_e$	100 %	
$\Gamma_2 \quad \text{hydrogen-atom} \bar{\nu}_e$		
Charge conservation (Q) violating mode		
$\Gamma_3 \quad p \nu_e \bar{\nu}_e$	$Q \quad < \quad 8 \times 10^{-27}$	68%

***n* BRANCHING RATIOS**

$$\Gamma(\text{hydrogen-atom} \bar{\nu}_e)/\Gamma_{\text{total}} \quad \Gamma_2/\Gamma$$

VALUE	CL%	DOCUMENT ID	TECN
-------	-----	-------------	------

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<3 \times 10^{-2}$ 95 17 GREEN 90 RVUE

17 GREEN 90 infers that $\tau(\text{hydrogen-atom} \bar{\nu}_e) > 3 \times 10^4$ s by comparing neutron lifetime measurements made in storage experiments with those made in β -decay experiments. However, the result depends sensitively on the lifetime measurements, and does not of course take into account more recent measurements of same.

$$\Gamma(p \nu_e \bar{\nu}_e)/\Gamma_{\text{total}} \quad \Gamma_3/\Gamma$$

Forbidden by charge conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
-------	-----	-------------	------	---------

$<8 \times 10^{-27}$ 68 18 NORMAN 96 RVUE ${}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge}$ neutrals

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<9.7 \times 10^{-18}$ 90 ROY 83 CNTR ${}^{113}\text{Cd} \rightarrow {}^{113m}\text{In}$ neut.

$<7.9 \times 10^{-21}$ Vайдя 83 CNTR ${}^{87}\text{Rb} \rightarrow {}^{87m}\text{Sr}$ neut.

$<9 \times 10^{-24}$ 90 BARABANOV 80 CNTR ${}^{71}\text{Ga} \rightarrow {}^{71}\text{GeX}$

$<3 \times 10^{-19}$ NORMAN 79 CNTR ${}^{87}\text{Rb} \rightarrow {}^{87m}\text{Sr}$ neut.

18 NORMAN 96 gets this limit by attributing SAGE and GALLEX counting rates to the charge-nonconserving transition ${}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge} + \text{neutrals}$ rather than to solar-neutrino reactions.

BARYON DECAY PARAMETERS

Written 1996 by E.D. Commins (University of California, Berkeley).

Baryon semileptonic decays

The typical spin-1/2 baryon semileptonic decay is described by a matrix element, the hadronic part of which may be written as:

$$\overline{B}_f [f_1(q^2) \gamma_\lambda + i f_2(q^2) \sigma_{\lambda\mu} q^\mu + g_1(q^2) \gamma_\lambda \gamma_5 + g_3(q^2) \gamma_5 q_\lambda] B_i . \quad (1)$$

Here B_i and \bar{B}_f are spinors describing the initial and final baryons, and $q = p_i - p_f$, while the terms in f_1 , f_2 , g_1 , and g_3 account for vector, induced tensor (“weak magnetism”), axial vector, and induced pseudoscalar contributions [1]. Second-class current contributions are ignored here. In the limit of zero momentum transfer, f_1 reduces to the vector coupling constant g_V , and g_1 reduces to the axial-vector coupling constant g_A . The latter coefficients are related by Cabibbo’s theory [2], generalized to six quarks (and three mixing angles) by Kobayashi and Maskawa [3]. The g_3 term is negligible for transitions in which an e^\pm is emitted, and gives a very small correction, which can be estimated by PCAC [4], for μ^\pm modes. Recoil effects include weak magnetism, and are taken into account adequately by considering terms of first order in

$$\delta = \frac{m_i - m_f}{m_i + m_f} , \quad (2)$$

where m_i and m_f are the masses of the initial and final baryons.

The experimental quantities of interest are the total decay rate, the lepton-neutrino angular correlation, the asymmetry coefficients in the decay of a polarized initial baryon, and the polarization of the decay baryon in its own rest frame for an unpolarized initial baryon. Formulae for these quantities are derived by standard means [5] and are analogous to formulae for nuclear beta decay [6]. We use the notation of Ref. 6 in the Listings for neutron beta decay. For comparison with experiments at higher q^2 , it is necessary to modify the form factors at $q^2 = 0$ by a “dipole” q^2 dependence, and for high-precision comparisons to apply appropriate radiative corrections [7].

The ratio g_A/g_V may be written as

$$g_A/g_V = |g_A/g_V| e^{i\phi_{AV}} . \quad (3)$$

The presence of a “triple correlation” term in the transition probability, proportional to $\text{Im}(g_A/g_V)$ and of the form

$$\boldsymbol{\sigma}_i \cdot (\mathbf{p}_\ell \times \mathbf{p}_\nu) \quad (4)$$

for initial baryon polarization or

$$\boldsymbol{\sigma}_f \cdot (\mathbf{p}_\ell \times \mathbf{p}_\nu) \quad (5)$$

for final baryon polarization, would indicate failure of time-reversal invariance. The phase angle ϕ has been measured precisely only in neutron decay (and in ^{19}Ne nuclear beta decay), and the results are consistent with T invariance.

Hyperon nonleptonic decays

The amplitude for a spin-1/2 hyperon decaying into a spin-1/2 baryon and a spin-0 meson may be written in the form

$$M = G_F m_\pi^2 \cdot \overline{B}_f (A - B\gamma_5) B_i , \quad (6)$$

where A and B are constants [1]. The transition rate is proportional to

$$R = 1 + \gamma \hat{\boldsymbol{\omega}}_f \cdot \hat{\boldsymbol{\omega}}_i + (1 - \gamma)(\hat{\boldsymbol{\omega}}_f \cdot \hat{\mathbf{n}})(\hat{\boldsymbol{\omega}}_i \cdot \hat{\mathbf{n}}) \\ + \alpha(\hat{\boldsymbol{\omega}}_f \cdot \hat{\mathbf{n}} + \hat{\boldsymbol{\omega}}_i \cdot \hat{\mathbf{n}}) + \beta \hat{\mathbf{n}} \cdot (\hat{\boldsymbol{\omega}}_f \times \hat{\boldsymbol{\omega}}_i) , \quad (7)$$

where $\hat{\mathbf{n}}$ is a unit vector in the direction of the final baryon momentum, and $\hat{\boldsymbol{\omega}}_i$ and $\hat{\boldsymbol{\omega}}_f$ are unit vectors in the directions of the initial and final baryon spins. (The sign of the last term in the above equation was incorrect in our 1988 and 1990 editions.) The parameters α , β , and γ are defined as

$$\alpha = 2 \text{Re}(s^* p) / (|s|^2 + |p|^2) , \\ \beta = 2 \text{Im}(s^* p) / (|s|^2 + |p|^2) , \\ \gamma = (|s|^2 - |p|^2) / (|s|^2 + |p|^2) , \quad (8)$$

where $s = A$ and $p = |\mathbf{p}_f| B/(E_f + m_f)$; here E_f and \mathbf{p}_f are the energy and momentum of the final baryon. The parameters α , β , and γ satisfy

$$\alpha^2 + \beta^2 + \gamma^2 = 1 . \quad (9)$$

If the hyperon polarization is \mathbf{P}_Y , the polarization \mathbf{P}_B of the decay baryons is

$$\mathbf{P}_B = \frac{(\alpha + \mathbf{P}_Y \cdot \hat{\mathbf{n}})\hat{\mathbf{n}} + \beta(\mathbf{P}_Y \times \hat{\mathbf{n}}) + \gamma\hat{\mathbf{n}} \times (\mathbf{P}_Y \times \hat{\mathbf{n}})}{1 + \alpha\mathbf{P}_Y \cdot \hat{\mathbf{n}}} . \quad (10)$$

Here \mathbf{P}_B is defined in the rest system of the baryon, obtained by a Lorentz transformation along $\hat{\mathbf{n}}$ from the hyperon rest frame, in which $\hat{\mathbf{n}}$ and \mathbf{P}_Y are defined.

An additional useful parameter ϕ is defined by

$$\beta = (1 - \alpha^2)^{1/2} \sin\phi . \quad (11)$$

In the Listings, we compile α and ϕ for each decay, since these quantities are most closely related to experiment and are essentially uncorrelated. When necessary, we have changed the signs of reported values to agree with our sign conventions. In the Baryon Summary Table, we give α , ϕ , and Δ (defined below) with errors, and also give the value of γ without error.

Time-reversal invariance requires, in the absence of final-state interactions, that s and p be relatively real, and therefore that $\beta = 0$. However, for the decays discussed here, the final-state interaction is strong. Thus

$$s = |s| e^{i\delta_s} \text{ and } p = |p| e^{i\delta_p} , \quad (12)$$

where δ_s and δ_p are the pion-baryon s - and p -wave strong interaction phase shifts. We then have

$$\beta = \frac{-2|s||p|}{|s|^2 + |p|^2} \sin(\delta_s - \delta_p) . \quad (13)$$

One also defines $\Delta = -\tan^{-1}(\beta/\alpha)$. If T invariance holds, $\Delta = \delta_s - \delta_p$. For $\Lambda \rightarrow p\pi^-$ decay, the value of Δ may be compared with the s - and p -wave phase shifts in low-energy π^-p scattering, and the results are consistent with T invariance.

Radiative hyperon decays

For the radiative decay of a polarized spin-1/2 hyperon, $B_i \rightarrow B_f \gamma$, the angular distribution of the direction \hat{p} of the final spin-1/2 baryon in the hyperon rest frame is

$$\frac{d\Gamma_\gamma}{d\Omega} = \frac{\Gamma_\gamma}{4\pi} (1 + \alpha_\gamma \hat{p} \cdot \mathbf{P}_i) , \quad (14)$$

where \mathbf{P}_i is the hyperon polarization and the asymmetry parameter α_γ is

$$\alpha_\gamma = \frac{2\text{Re} [g'_1(0)f_M^*(0)]}{|g'_1(0)|^2 + |f_M(0)|^2} . \quad (15)$$

Here $f_M = \frac{(m_i - m_f)}{(m_i + m_f)} [(m_i + m_f)f'_2 - f'_1]$, where $f'_1(q^2)$, $f'_2(q^2)$, and $g'_1(q^2)$ are the $\Delta Q = 0$ analogs of the $|\Delta Q| = 1$ form factors defined above.

References

1. E.D. Commins and P.H. Bucksbaum, *Weak Interactions of Leptons and Quarks* (Cambridge University Press, Cambridge, England, 1983).
2. N. Cabibbo, Phys. Rev. Lett. **10**, 531 (1963).
3. M. Kobayashi and T. Maskawa, Prog. Theor. Phys. **49**, 652 (1973).
4. M.L. Goldberger and S.B. Treiman, Phys. Rev. **111**, 354 (1958).
5. P.H. Frampton and W.K. Tung, Phys. Rev. **D3**, 1114 (1971).
6. J.D. Jackson, S.B. Treiman, and H.W. Wyld, Jr., Phys. Rev. **106**, 517 (1957), and Nucl. Phys. **4**, 206 (1957).
7. Y. Yokoo, S. Suzuki, and M. Morita, Prog. Theor. Phys. **50**, 1894 (1973).

$n \rightarrow p e^- \nu$ DECAY PARAMETERS

See the above "Note on Baryon Decay Parameters." For discussions of recent results, see the references cited at the beginning of the section on the neutron mean life. For discussions of the values of the weak coupling constants g_A and g_V obtained using the neutron lifetime and asymmetry parameter A , comparisons with other methods of obtaining these constants, and implications for particle physics and for astrophysics, see DUBBERS 91 and WOOLCOCK 91. For tests of the $V-A$ theory of neutron decay, see EROZOLIMSKII 91B and MOSTOVOI 96.

g_A / g_V

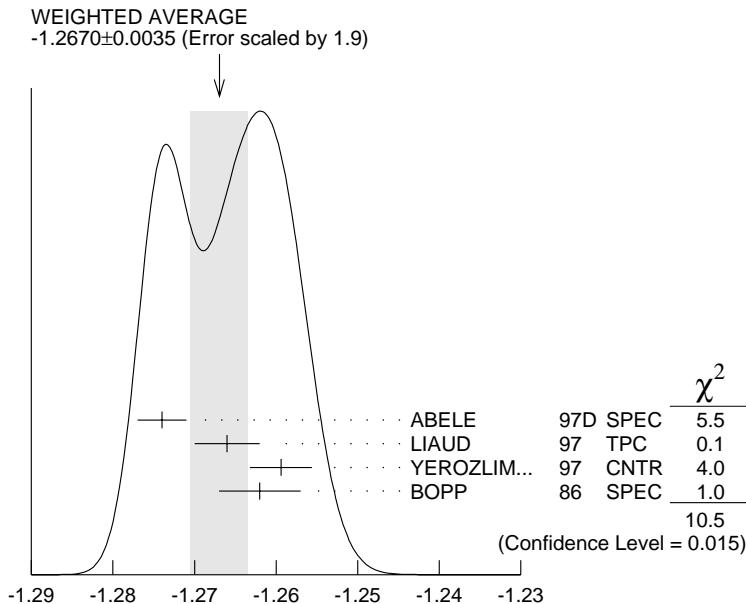
VALUE	DOCUMENT ID	TECN	COMMENT
-1.2670 ± 0.0035 OUR AVERAGE			Error includes scale factor of 1.9. See the ideogram below.
-1.274 ± 0.003	ABELE	97D SPEC	cold n , polarized
-1.266 ± 0.004	LIAUD	97 TPC	e mom- n spin corr.
-1.2594 ± 0.0038	¹⁹ YEROZLIM...	97 CNTR	e mom- n spin corr.
-1.262 ± 0.005	BOPP	86 SPEC	e mom- n spin corr.
• • • We do not use the following data for averages, fits, limits, etc. • • •			
-1.266 ± 0.004	SCHRECK...	95 TPC	See LIAUD 97
-1.2544 ± 0.0036	EROZOLIM...	91 CNTR	See YEROZOLIM-SKY 97
-1.226 ± 0.042	MOSTOVOY	83 RVUE	
-1.261 ± 0.012	²⁰ EROZOLIM...	79 CNTR	e mom- n spin corr.
-1.259 ± 0.017	²⁰ STRATOWA	78 CNTR	proton recoil spectrum
-1.263 ± 0.015	EROZOLIM...	77 CNTR	See EROZOLIMSKII 79
-1.250 ± 0.036	²⁰ DOBROZE...	75 CNTR	See STRATOWA 78
-1.258 ± 0.015	²¹ KROHN	75 CNTR	e mom- n spin corr.
-1.263 ± 0.016	²² KROPF	74 RVUE	n decay alone
-1.250 ± 0.009	²² KROPF	74 RVUE	n decay + nuclear ft

¹⁹ YEROZOLIMSKY 97 makes a correction to the EROZOLIMSKII 91 value.

²⁰ These experiments measure the absolute value of g_A/g_V only.

²¹ KROHN 75 includes events of CHRISTENSEN 70.

²² KROPF 74 reviews all data through 1972.



$$g_A / g_V$$

β ASYMMETRY PARAMETER A

This is the neutron-spin electron-momentum correlation coefficient. Unless otherwise noted, the values are corrected for radiative effects and weak magnetism.

VALUE	DOCUMENT ID	TECN	COMMENT
-0.1162 ± 0.0013 OUR AVERAGE			Error includes scale factor of 1.8. See the ideogram below.
-0.1189 ± 0.0012	ABELE	97D SPEC	cold n , polarized
$-0.1160 \pm 0.0009 \pm 0.0012$	LIAUD	97 TPC	e mom- n spin corr.
-0.1135 ± 0.0014	²³ YEROZLIM...	97 CNTR	e mom- n spin corr.
-0.1146 ± 0.0019	BOPP	86 SPEC	
• • • We do not use the following data for averages, fits, limits, etc. • • •			
$-0.1160 \pm 0.0009 \pm 0.0011$	SCHRECK...	95 TPC	See LIAUD 97
-0.1116 ± 0.0014	EROZOLIM...	91 CNTR	See YEROZOLIM-SKY 97
-0.114 ± 0.005	²⁴ EROZOLIM...	79 CNTR	
-0.113 ± 0.006	²⁴ KROHN	75 CNTR	

²³ YEROZOLIMSKY 97 makes a correction to the EROZOLIMSKII 91 value.

²⁴ These results are not corrected for radiative effects and weak magnetism, but the corrections are small compared to the errors.

STEINBERG	76	PR D13 2469	R.I. Steinberg <i>et al.</i>	(YALE, ISNG)
DOBROZE...	75	PR D11 510	R. Dobrozemsky <i>et al.</i>	(SEIB)
KROHN	75	PL 55B 175	V.E. Krohn, G.R. Ringo	(ANL)
EROZOLIM...	74	JETPL 20 345	B.G. Erozolimsky <i>et al.</i>	
			Translated from ZETFP 20 745.	
KROPF	74	ZPHY 267 129	H. Kropf, E. Paul	(LINZ)
Also	70	NP A154 160	H. Paul	(VIEN)
STEINBERG	74	PRL 33 41	R.I. Steinberg <i>et al.</i>	(YALE, ISNG)
COHEN	73	JPCRD 2 663	E.R. Cohen, B.N. Taylor	(RISC, NBS)
CHRISTENSEN	72	PR D5 1628	C.J. Christensen <i>et al.</i>	(RISO)
CHRISTENSEN	70	PR C1 1693	C.J. Christensen, V.E. Krohn, G.R. Ringo	(ANL)
EROZOLIM...	70C	PL 33B 351	B.G. Erozolimsky <i>et al.</i>	(KIAE)
GRIGOREV	68	SJNP 6 239	V.K. Grigoriev <i>et al.</i>	(ITEP)
			Translated from YAF 6 329.	