

$\pi(1300)$ $I^G(J^{PC}) = 1^-(0^{-+})$ **$\pi(1300)$ MASS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
1300±100 OUR ESTIMATE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1345± 8±10	18k	1 SCHEGELSKY 06	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
1200± 40	90k	SALVINI 04	OBLX	$\bar{p}p \rightarrow 2\pi^+2\pi^-$
1343± 15±24		CHUNG 02	B852	$18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$
1375± 40		ABELE 01	CBAR	$0.0\bar{p}d \rightarrow \pi^-4\pi^0p$
1275± 15		BERTIN 97D	OBLX	$0.05\bar{p}p \rightarrow 2\pi^+2\pi^-$
~ 1114		ABELE 96	CBAR	$0.0\bar{p}p \rightarrow 5\pi^0$
1190± 30		ZIELINSKI 84	SPEC	$200\pi^+Z \rightarrow Z3\pi$
1240± 30		BELLINI 82	SPEC	$40\pi^-A \rightarrow A3\pi$
1273± 50	2 AARON	81	RVUE	
1342± 20		BONESINI 81	OMEG	$12\pi^-p \rightarrow p3\pi$
~ 1400		DAUM 81B	SPEC	$63,94\pi^-p$

¹ From analysis of L3 data at 183–209 GeV.² Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from DAUM 80 and DANKOWYCH 81. **$\pi(1300)$ WIDTH**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
200 to 600 OUR ESTIMATE				
• • • We do not use the following data for averages, fits, limits, etc. • • •				
260± 20±30	18k	3 SCHEGELSKY 06	RVUE	$\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
470±120	90k	SALVINI 04	OBLX	$\bar{p}p \rightarrow 2\pi^+2\pi^-$
449± 39±47		CHUNG 02	B852	$18.3\pi^-p \rightarrow \pi^+\pi^-\pi^-p$
268± 50		ABELE 01	CBAR	$0.0\bar{p}d \rightarrow \pi^-4\pi^0p$
218±100		BERTIN 97D	OBLX	$0.05\bar{p}p \rightarrow 2\pi^+2\pi^-$
~ 340		ABELE 96	CBAR	$0.0\bar{p}p \rightarrow 5\pi^0$
440± 80		ZIELINSKI 84	SPEC	$200\pi^+Z \rightarrow Z3\pi$
360±120		BELLINI 82	SPEC	$40\pi^-A \rightarrow A3\pi$
580±100	4 AARON	81	RVUE	
220± 70		BONESINI 81	OMEG	$12\pi^-p \rightarrow p3\pi$
~ 600		DAUM 81B	SPEC	$63,94\pi^-p$

³ From analysis of L3 data at 183–209 GeV.⁴ Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from DAUM 80 and DANKOWYCH 81.

$\pi(1300)$ DECAY MODES

Mode	Fraction (Γ_i/Γ)
$\Gamma_1 \rho\pi$	seen
$\Gamma_2 \pi(\pi\pi)_{S\text{-wave}}$	seen
$\Gamma_3 \gamma\gamma$	

$\pi(1300) \Gamma(i)\Gamma(\gamma\gamma)/\Gamma(\text{total})$

$\Gamma(\rho\pi) \times \Gamma(\gamma\gamma)/\Gamma_{\text{total}}$	$\Gamma_1\Gamma_3/\Gamma$
<u>VALUE (keV)</u>	<u>CL%</u>
<0.085	90
• • • We do not use the following data for averages, fits, limits, etc. • • •	ACCIARRI 97T L3 $e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$
<0.8	95
<0.54	90
5 From analysis of L3 data at 183–209 GeV.	SCHEGELSKY 06 RVUE $\gamma\gamma \rightarrow \pi^+\pi^-\pi^0$
ALBRECHT 97B ARG $e^+e^- \rightarrow e^+e^-\pi^+\pi^-\pi^0$	

⁵ From analysis of L3 data at 183–209 GeV.

$\pi(1300)$ BRANCHING RATIOS

$\Gamma(\pi(\pi\pi)_{S\text{-wave}})/\Gamma(\rho\pi)$	Γ_2/Γ_1
<u>VALUE</u>	<u>CL%</u> <u>EVTS</u>
2.2 ± 0.4	90k SALVINI 04 OBLX $\bar{p}p \rightarrow 2\pi^+2\pi^-$
seen	CHUNG 02 B852 18.3 $\pi^-p \rightarrow \pi^+2\pi^-p$
<0.15	90 ABELE 01 CBAR 0.0 $\bar{p}d \rightarrow \pi^-4\pi^0p$
2.12	6 AARON 81 RVUE

⁶ Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from DAUM 80 and DANKOWYCH 81.

$\pi(1300)$ REFERENCES

SCHEGELSKY 06	EPJ A27 199	V.A. Schegelsky <i>et al.</i>	
SALVINI 04	EPJ C35 21	P. Salvini <i>et al.</i>	(OBELIX Collab.)
CHUNG 02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
ABELE 01	EPJ C19 667	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ACCIARRI 97T	PL B413 147	M. Acciarri <i>et al.</i>	(L3 Collab.)
ALBRECHT 97B	ZPHY C74 469	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BERTIN 97D	PL B414 220	A. Bertin <i>et al.</i>	(OBELIX Collab.)
ABELE 96	PL B380 453	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ZIELINSKI 84	PR D30 1855	M. Zielinski <i>et al.</i>	(ROCH, MINN, FNAL)
BELLINI 82	PRL 48 1697	G. Bellini <i>et al.</i>	(MILA, BGNA, JINR)
AARON 81	PR D24 1207	R.A. Aaron, R.S. Longacre	(NEAS, BNL)
BONESINI 81	PL 103B 75	M. Bonesini <i>et al.</i>	(MILA, L1VP, DARE+)
DANKOWY... 81	PRL 46 580	J.A. Dankowych <i>et al.</i>	(TNTO, BNL, CARL+)
DAUM 81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
DAUM 80	PL 89B 281	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
BOWLER 75	NP B97 227	M.G. Bowler <i>et al.</i>	(OXFTP, DARE)