

$a_1(1260)$ $I^G(J^{PC}) = 1^-(1^{++})$

See also our review under the $a_1(1260)$ in PDG 06, Journal of Physics, G **33** 1 (2006).

 $a_1(1260)$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
1230 ± 40 OUR ESTIMATE				
$1255 \pm 6^{+7}_{-17}$	420k	ALEKSEEV	10 COMP	$190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1243 $\pm 12 \pm 20$		¹ AUBERT	07AU BABR	$10.6 e^+ e^- \rightarrow \rho^0 \rho^\pm \pi^\mp \gamma$
1230–1270	6360	² LINK	07A FOCS	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
1203 ± 3		³ GOMEZ-DUM..04	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
1330 ± 24	90k	SALVINI	04 OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
1331 $\pm 10 \pm 3$	37k	⁴ ASNER	00 CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
1255 $\pm 7 \pm 6$	5904	⁵ ABREU	98G DLPH	$e^+ e^-$
1207 $\pm 5 \pm 8$	5904	⁶ ABREU	98G DLPH	$e^+ e^-$
1196 $\pm 4 \pm 5$	5904	^{7,8} ABREU	98G DLPH	$e^+ e^-$
1240 ± 10		BARBERIS	98B	$450 pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
1262 $\pm 9 \pm 7$		^{5,9} ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88\text{--}94, \tau \rightarrow 3\pi\nu$
1210 $\pm 7 \pm 2$		^{6,9} ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88\text{--}94, \tau \rightarrow 3\pi\nu$
$1211 \pm 7^{+50}_{-0}$		⁶ ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1121 ± 8		¹⁰ ANDO	92 SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1242 ± 37		¹¹ IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260 ± 14		¹² IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1250 ± 9		¹³ IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1208 ± 15		ARMSTRONG	90 OMEG	$300.0 pp \rightarrow pp \pi^+ \pi^- \pi^0$
1220 ± 15		¹⁴ ISGUR	89 RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260 ± 25		¹⁵ BOWLER	88 RVUE	
1166 $\pm 18 \pm 11$		BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1164 $\pm 41 \pm 23$		BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
1250 ± 40		¹⁴ TORNQVIST	87 RVUE	
1046 ± 11		ALBRECHT	86B ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1056 $\pm 20 \pm 15$		RUCKSTUHL	86 DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1194 $\pm 14 \pm 10$		SCHMIDKE	86 MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1255 ± 23		BELLINI	85 SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1240 ± 80		¹⁶ DANKOWY...	81 SPEC	$8.45 \pi^- p \rightarrow n 3\pi$
1280 ± 30		¹⁶ DAUM	81B CNTR	$63.94 \pi^- p \rightarrow p 3\pi$
1041 ± 13		¹⁷ GAVILLET	77 HBC	$4.2 K^- p \rightarrow \Sigma 3\pi$

¹ The $\rho^\pm \pi^\mp$ state can be also due to the $\pi(1300)$.² Using the Breit-Wigner parameterization; strong correlation between mass and width.³ Using the data of BARATE 98R.⁴ From a fit to the 3π mass spectrum including the $K\bar{K}^*(892)$ threshold.⁵ Uses the model of KUHN 90.

⁶ Uses the model of ISGUR 89.⁷ Includes the effect of a possible a'_1 state.⁸ Uses the model of FEINDT 90.⁹ Supersedes AKERS 95P.¹⁰ Average and spread of values using 2 variants of the model of BOWLER 75.¹¹ Reanalysis of RUCKSTUHL 86.¹² Reanalysis of SCHMIDKE 86.¹³ Reanalysis of ALBRECHT 86B.¹⁴ From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.¹⁵ From a combined reanalysis of ALBRECHT 86B and DAUM 81B.¹⁶ Uses the model of BOWLER 75.¹⁷ Produced in K^- backward scattering.

$a_1(1260)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
250 to 600 OUR ESTIMATE				
$367 \pm 9^{+28}_{-25}$	420k	ALEKSEEV	10 COMP	$190 \pi^- Pb \rightarrow \pi^- \pi^- \pi^+ Pb'$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
410 ± 31 ± 30	18 AUBERT	07AU BABR	$10.6 e^+ e^- \rightarrow \rho^0 \rho^\pm \pi^\mp \gamma$	
520–680	6360	19 LINK	07A FOCS	$D^0 \rightarrow \pi^- \pi^+ \pi^- \pi^+$
480 ± 20	20 GOMEZ-DUM..04	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$	
580 ± 41	90k	SALVINI	04 OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$
460 ± 85	205	21 DRUTSKOY	02 BELL	$B \rightarrow D^{(*)} K^- K^{*0}$
814 ± 36 ± 13	37k	22 ASNER	00 CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
450 ± 50	22k	23 AKHMETSHIN 99E	CMD2	$1.05-1.38 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \pi^0$
570 ± 10	24 BONDAR	99 RVUE	$e^+ e^- \rightarrow 4\pi, \tau \rightarrow 3\pi \nu_\tau$	
587 ± 27 ± 21	5904	25 ABREU	98G DLPH	$e^+ e^-$
478 ± 3 ± 15	5904	26 ABREU	98G DLPH	$e^+ e^-$
425 ± 14 ± 8	5904	27,28 ABREU	98G DLPH	$e^+ e^-$
400 ± 35	BARBERIS	98B		$450 pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
621 ± 32 ± 58	25,29 ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$	
457 ± 15 ± 17	ACKERSTAFF	97R OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$	
446 ± 21^{+140}_{-0}	26 ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$	
239 ± 11	ANDO	92 SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$	
266 ± 13 ± 4	30 ANDO	92 SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$	
465^{+228}_{-143}	31 IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$	
298^{+40}_{-34}	32 IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$	
488 ± 32	33 IVANOV	91 RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$	
430 ± 50	ARMSTRONG	90 OMEG	$300.0 pp \rightarrow pp \pi^+ \pi^- \pi^0$	
420 ± 40	34 ISGUR	89 RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$	
396 ± 43	35 BOWLER	88 RVUE		
405 ± 75 ± 25	BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$	
419 ± 108 ± 57	BAND	87 MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$	

521 \pm 27	ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
476 $^{+132}_{-120} \pm$ 54	RUCKSTUHL	86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
462 \pm 56 \pm 30	SCHMIDKE	86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
292 \pm 40	BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
380 \pm 100	³⁶ DANKOWY...	81	SPEC	$8.45 \pi^- p \rightarrow n 3\pi$
300 \pm 50	³⁶ DAUM	81B	CNTR	$63,94 \pi^- p \rightarrow p 3\pi$
230 \pm 50	³⁷ GAVILLET	77	HBC	$4.2 K^- p \rightarrow \Sigma 3\pi$

¹⁸The $\rho^\pm \pi^\mp$ state can be also due to the $\pi(1300)$.

¹⁹Using the Breit-Wigner parameterization; strong correlation between mass and width.

²⁰Using the data of BARATE 98R.

²¹From a fit of the $K^- K^{*0}$ distribution assuming $m_{a_1} = 1230$ MeV and purely resonant production of the $K^- K^{*0}$ system.

²²From a fit to the 3π mass spectrum including the $K\bar{K}^*(892)$ threshold.

²³Using the $a_1(1260)$ mass of 1230 MeV.

²⁴From AKHMETSHIN 99E and ASNER 00 data using the $a_1(1260)$ mass of 1230 MeV.

²⁵Uses the model of KUHN 90.

²⁶Uses the model of ISGUR 89.

²⁷Includes the effect of a possible a'_1 state.

²⁸Uses the model of FEINDT 90.

²⁹Supersedes AKERS 95P.

³⁰Average and spread of values using 2 variants of the model of BOWLER 75.

³¹Reanalysis of RUCKSTUHL 86.

³²Reanalysis of SCHMIDKE 86.

³³Reanalysis of ALBRECHT 86B.

³⁴From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.

³⁵From a combined reanalysis of ALBRECHT 86B and DAUM 81B.

³⁶Uses the model of BOWLER 75.

³⁷Produced in K^- backward scattering.

$a_1(1260)$ DECAY MODES

Mode	Fraction (Γ_i/Γ)
$\Gamma_1 \pi^+ \pi^- \pi^0$	
$\Gamma_2 \pi^0 \pi^0 \pi^0$	
$\Gamma_3 (\rho\pi)_S$ -wave	seen
$\Gamma_4 (\rho\pi)_D$ -wave	seen
$\Gamma_5 (\rho(1450)\pi)_S$ -wave	seen
$\Gamma_6 (\rho(1450)\pi)_D$ -wave	seen
$\Gamma_7 \sigma\pi$	seen
$\Gamma_8 f_0(980)\pi$	not seen
$\Gamma_9 f_0(1370)\pi$	seen
$\Gamma_{10} f_2(1270)\pi$	seen
$\Gamma_{11} K\bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{12} \pi\gamma$	seen

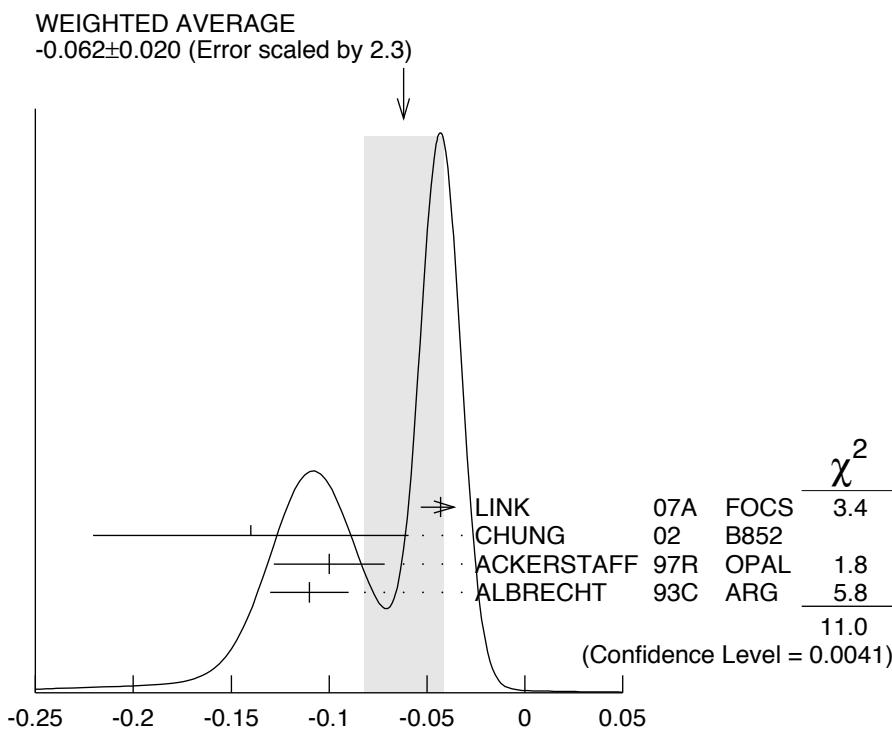
$a_1(1260)$ PARTIAL WIDTHS

$\Gamma(\pi\gamma)$				Γ_{12}
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
640±246	ZIELINSKI	84C	SPEC	200 $\pi^+ Z \rightarrow Z 3\pi^-$

D-wave/S-wave AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

VALUE	DOCUMENT ID	TECN	COMMENT
-0.062±0.020 OUR AVERAGE	Error includes scale factor of 2.3. See the ideogram below.		
-0.043±0.009±0.005	LINK	07A	FOCS $D^0 \rightarrow \pi^-\pi^+\pi^-\pi^+$
-0.14 ± 0.04 ± 0.07	38 CHUNG	02	B852 $18.3 \pi^- p \rightarrow \pi^+\pi^-\pi^- p$
-0.10 ± 0.02 ± 0.02	39,40 ACKERSTAFF	97R	OPAL $E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
-0.11 ± 0.02	39 ALBRECHT	93C	ARG $\tau^+ \rightarrow \pi^+\pi^+\pi^-\nu$

38 Deck-type background not subtracted.
 39 Uses the model of ISGUR 89.
 40 Supersedes AKERS 95P.



D-wave/S-wave AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

$a_1(1260)$ BRANCHING RATIOS

$\Gamma((\rho\pi)_S\text{-wave})/\Gamma_{\text{total}}$				Γ_3/Γ
VALUE (units 10^{-2})	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
60.19	37k	41 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+\tau^-, \tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$

$\Gamma((\rho\pi)_D\text{-wave})/\Gamma_{\text{total}}$ Γ_4/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$1.30 \pm 0.60 \pm 0.22$	37k	41 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma((\rho(1450)\pi)_S\text{-wave})/\Gamma_{\text{total}}$ Γ_5/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$0.56 \pm 0.84 \pm 0.32$	37k	41,42 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma((\rho(1450)\pi)_D\text{-wave})/\Gamma_{\text{total}}$ Γ_6/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$2.04 \pm 1.20 \pm 0.28$	37k	41,42 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma(\sigma\pi)/\Gamma_{\text{total}}$ Γ_7/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
seen		CHUNG	02	B852 $18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
$18.76 \pm 4.29 \pm 1.48$	37k	41,43 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma(f_0(980)\pi)/\Gamma_{\text{total}}$ Γ_8/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
not seen	37k	ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma(f_0(1370)\pi)/\Gamma_{\text{total}}$ Γ_9/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$7.40 \pm 2.71 \pm 1.26$	37k	41,44 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma(f_2(1270)\pi)/\Gamma_{\text{total}}$ Γ_{10}/Γ

<u>VALUE (units 10^{-2})</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$1.19 \pm 0.49 \pm 0.17$	37k	41,45 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma(K\bar{K}^*(892)+c.c.)/\Gamma_{\text{total}}$

<u>VALUE</u> (units 10^{-2})	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	Γ_{11}/Γ
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.2 \pm 0.5	2255	46 COAN	04 CLEO	$\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau$	
8 to 15	205	47 DRUTSKOY	02 BELL	$B \rightarrow D^{(*)} K^- K^{*0}$	
$3.3 \pm 0.5 \pm 0.1$	37k	48 ASNER	00 CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-$, $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$	
2.6 \pm 0.3		49 BARATE	99R ALEP	$\tau \rightarrow K\bar{K}\pi\nu_\tau$	

$\Gamma(\sigma\pi)/\Gamma((\rho\pi)_{S-\text{wave}})$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	Γ_7/Γ_3
• • • We do not use the following data for averages, fits, limits, etc. • • •					
0.06 \pm 0.05	90k	SALVINI	04 OBLX	$\bar{p}p \rightarrow 2\pi^+ 2\pi^-$	
~ 0.3	28k	AKHMETSHIN 99E	CMD2	$1.05-1.38 e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^-$	
0.003 \pm 0.003		50 LONGACRE	82 RVUE		

$\Gamma(\pi^0\pi^0\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>COMMENT</u>	Γ_2/Γ_1
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.008	90	51 BARBERIS	01 450 $p p \rightarrow p_f 3\pi^0 p_s$	
41	From a fit to the Dalitz plot.			
42	Assuming for $\rho(1450)$ mass and width of 1370 and 386 MeV respectively.			
43	Assuming for σ mass and width of 860 and 880 MeV respectively.			
44	Assuming for $f_0(1370)$ mass and width of 1186 and 350 MeV respectively.			
45	Assuming for $f_2(1270)$ mass and width of 1275 and 185 MeV respectively.			
46	Using structure functions from KUHN 92 and DECKER 93A and $B(\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau) = (0.155 \pm 0.006 \pm 0.009)\%$ from BRIERE 03.			
47	From a comparison to ALAM 94 assuming purely resonant production of the $K^- K^{*0}$ system.			
48	From a fit to the 3π mass spectrum including the $K\bar{K}^*(892)$ threshold.			
49	Assuming $a_1(1260)$ dominance and taking $B(\tau \rightarrow a_1(1260)\nu_\tau)$ from BUSKULIC 96.			
50	Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from GAVILLET 77, DAUM 80, and DANKOWYCH 81.			
51	Inconsistent with observations of $\sigma\pi$, $f_0(1370)\pi$, and $f_2(1270)\pi$ decay modes.			

$a_1(1260)$ REFERENCES

ALEKSEEV	10	PRL 104 241803	M.G. Alekseev <i>et al.</i>	(COMPASS Collab.)
AUBERT	07AU	PR D76 092005	B. Aubert <i>et al.</i>	(BABAR Collab.)
LINK	07A	PR D75 052003	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
PDG	06	JPG 33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
COAN	04	PRL 92 232001	T.E. Coan <i>et al.</i>	(CLEO Collab.)
GOMEZ-DUM...	04	PR D69 073002	D. Gomez Dumm, A. Pich, J. Portoles	
SALVINI	04	EPJ C35 21	P. Salvini <i>et al.</i>	(OBELIX Collab.)
BRIERE	03	PRL 90 181802	R. A. Briere <i>et al.</i>	(CLEO Collab.)
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
DRUTSKOY	02	PL B542 171	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
BARBERIS	01	PL B507 14	D. Barberis <i>et al.</i>	
ASNER	00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
AKHMETSHIN	99E	PL B466 392	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BARATE	99R	EPJ C11 599	R. Barate <i>et al.</i>	(ALEPH Collab.)
BONDAR	99	PL B466 403	A.E. Bondar <i>et al.</i>	(Novosibirsk CMD-2 Collab.)

ABREU	98G	PL B426 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BARATE	98R	EPJ C4 409	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARBERIS	98B	PL B422 399	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ACKERSTAFF	97R	ZPHY C75 593	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BUSKULIC	96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
ALAM	94	PR D50 43	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DECKER	93A	ZPHY C58 445	R. Decker <i>et al.</i>	
ANDO	92	PL B291 496	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
KUHN	92	ZPHY C56 661	J.H. Kuhn, E. Mirkes	
IVANOV	91	ZPHY C49 563	Y.P. Ivanov, A.A. Osipov, M.K. Volkov	(JINR)
ARMSTRONG	90	ZPHY C48 213	T.A. Armstrong, M. Benayoun, W. Beusch	(WA76 Coll.)
FEINDT	90	ZPHY C48 681	M. Feindt	(HAMB)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ISGUR	89	PR D39 1357	N. Isgur, C. Morningstar, C. Reader	(TNT)
BOWLER	88	PL B209 99	M.G. Bowler	(OXF)
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
TORNQVIST	87	ZPHY C36 695	N.A. Tornqvist	(HELS)
ALBRECHT	86B	ZPHY C33 7	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
RUCKSTUHL	86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)
SCHMIDKE	86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)
BELLINI	85	SJNP 41 781	D. Bellini <i>et al.</i>	
ZIELINSKI	84C	Translated from YAF 41 1223.	M. Zielinski <i>et al.</i>	(ROCH, MINN, FNAL)
LONGACRE	82	PR D26 82	R.S. Longacre	(BNL)
DANKOWY...	81	PRL 46 580	J.A. Dankowych <i>et al.</i>	(TNT, BNL, CARL+)
DAUM	81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
DAUM	80	PL 89B 281	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+) JP
GAVILLET	77	PL 69B 119	P. Gavillet <i>et al.</i>	(AMST, CERN, NIJM+) JP
BOWLER	75	NP B97 227	M.G. Bowler <i>et al.</i>	(OXFTP, DARE)