Supersymmetric Particle Searches

The exclusion of particle masses within a mass range (m_1, m_2) will be denoted with the notation "none $m_1 - m_2$ " in the VALUE column of the following Listings. The latest unpublished results are described in the "Supersymmetry: Experiment" review.

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 - $-\widetilde{\chi}_1^0$ -p elastic cross section Spin-dependent interactions Spin-independent interactions
 - Other bounds on $\widetilde{\chi}_1^0$ from astrophysics and cosmology
 - Unstable $\widetilde{\chi}_1^0$ (Lightest Neutralino) Mass Limit

 $\widetilde{\chi}^0_2$, $\widetilde{\chi}^0_3$, $\widetilde{\chi}^0_4$ (Neutralinos) Mass Limits

 $\tilde{\chi}_1^{\pm}$, $\tilde{\chi}_2^{\pm}$ (Charginos) Mass Limits

Long-lived $\tilde{\chi}^{\pm}$ (Chargino) Mass Limits

 $\widetilde{\nu}$ (Sneutrino) Mass Limit

Charged Sleptons

- $-\widetilde{e}$ (Selectron) Mass Limit
- $-\widetilde{\mu}$ (Smuon) Mass Limit
- $-\widetilde{ au}$ (Stau) Mass Limit
- Degenerate Charged Sleptons
- $-\ell$ (Slepton) Mass Limit

 \tilde{q} (Squark) Mass Limit

Long-lived \tilde{q} (Squark) Mass Limit

b (Sbottom) Mass Limit

 \tilde{t} (Stop) Mass Limit

Heavy \tilde{g} (Gluino) Mass Limit

Long-lived \tilde{g} (Gluino) Mass Limit

Light G (Gravitino) Mass Limits from Collider Experiments

Supersymmetry Miscellaneous Results

Most of the results shown below, unless stated otherwise, are based on the Minimal Supersymmetric Standard Model (MSSM), as described in the Note on Supersymmetry. Unless otherwise indicated, this includes the assumption of common gaugino and scalar masses at the scale of Grand Unification (GUT), and use of the resulting relations in the spectrum and

decay branching ratios. Unless otherwise indicated, it is also assumed that R-parity (R) is conserved and that:

- 1) The $\tilde{\chi}_1^0$ is the lighest supersymmetric particle (LSP)
- 2) $m_{\widetilde{f}_L} = m_{\widetilde{f}_R}$, where $\widetilde{f}_{L,R}$ refer to the scalar partners of left-and right-handed fermions.

Limits involving different assumptions are identified in the Comments or in the Footnotes. We summarize here the notations used in this Chapter to characterize some of the most common deviations from the MSSM (for further details, see the Note on Supersymmetry).

Theories with R-parity violation (R) are characterized by a superpotential of the form: $\lambda_{ijk}L_iL_je_k^c + \lambda'_{ijk}L_iQ_jd_k^c + \lambda''_{ijk}u_i^cd_j^cd_k^c$, where i,j,k are generation indices. The presence of any of these couplings is often identified in the following by the symbols $LL\overline{E}$, $LQ\overline{D}$, and $\overline{U}D\overline{D}$. Mass limits in the presence of R will often refer to "direct" and "indirect" decays. Direct refers to R decays of the particle in consideration. Indirect refers to cases where R appears in the decays of the LSP.

In several models, most notably in theories with so-called Gauge Mediated Supersymmetry Breaking (GMSB), the gravitino (\widetilde{G}) is the LSP. It is usually much lighter than any other massive particle in the spectrum, and $m_{\widetilde{G}}$ is then neglected in all decay processes involving gravitinos. In these scenarios, particles other than the neutralino are sometimes considered as the next-to-lighest supersymmetric particle (NLSP), and are assumed to decay to their even-R partner plus \widetilde{G} . If the lifetime is short enough for the decay to take place within the detector, \widetilde{G} is assumed to be undetected and to give rise to missing energy (E) or missing transverse energy (E) signatures.

When needed, specific assumptions on the eigenstate content of $\tilde{\chi}^0$ and $\tilde{\chi}^{\pm}$ states are indicated, using the notation $\tilde{\gamma}$

(photino), \widetilde{H} (higgsino), \widetilde{W} (wino), and \widetilde{Z} (zino) to signal that the limit of pure states was used. The terms gaugino is also used, to generically indicate wino-like charginos and zino-like neutralinos.

In the listings we have made use of the following abbreviations for simplified models employed by the experimental collaborations in supersymmetry searches published in the past year.

Simplified Models Table

Tglu1A: gluino pair production with $\tilde{g} \to q\bar{q}\tilde{\chi}_1^0$

Tglu1B: gluino pair production with $\tilde{g} \to qq'\tilde{\chi}_1^{\pm}$, $\tilde{\chi}_1^{\pm} \to W^{\pm}\tilde{\chi}_1^0$ **Tglu1C:** gluino pair production with a 2/3 probability of having a $\tilde{g} \to q q' \tilde{\chi}_1^{\pm}, \ \tilde{\chi}_1^{\pm} \to W^{\pm} \tilde{\chi}_1^0$ decay and a 1/3 probability of having a $\tilde{g} \to qq\tilde{\chi}_2^0$, $\tilde{\chi}_2^0 \to Z^{\pm}\tilde{\chi}_1^0$ decay.

Tglu1D: gluino pair production with one gluino decaying to $q\bar{q}'\tilde{\chi}_1^{\pm}$ with $\tilde{\chi}_1^{\pm} \to W^{\pm} + \tilde{G}$, and the other gluino decaying to $q\bar{q}\tilde{\chi}_1^0$ with

Tglu1E: gluino pair production with $\tilde{g} \to qq'\tilde{\chi}_1^{\pm}$, $\tilde{\chi}_1^{\pm} \to W^{\pm}\tilde{\chi}_2^0$ and $\tilde{\chi}_2^0 \to Z^{\pm}\tilde{\chi}_1^0$ where $m_{\tilde{\chi}_1^{\pm}} = (m_{\tilde{g}} + m_{\tilde{\chi}_1^0})/2$, $m_{\tilde{\chi}_2^0} = (m_{\tilde{\chi}_1^{\pm}} + m_{\tilde{\chi}_1^0})/2$ $m_{\tilde{\chi}_1^0})/2.$

Tglu1F: gluino pair production with $\tilde{g} \to qq'\tilde{\chi}_1^{\pm}$ or $\tilde{g} \to qq'\tilde{\chi}_2^0$ with equal branching ratios, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate scalar tau lepton or sneutrino to $\tau\nu\tilde{\chi}_1^0$ and where $\tilde{\chi}_2^0$ decays through an intermediate scalar tau lepton or sneutrino to $au^+ au^-\tilde{\chi}_1^0$ or $\nuar{
u}\tilde{\chi}_1^0$; the mass hierarchy is such that $m_{\chi^\pm_+}\sim$ $m_{\tilde{\chi}_2^0} = (m_{\tilde{g}} + m_{\chi_1^0})/2$ and $m_{\tilde{\tau},\tilde{\nu}} = (m_{\tilde{\chi}_1^{\pm}} + m_{\tilde{\chi}_1^0})/2$.

Tglu2A: gluino pair production with $\tilde{g} \to b\bar{b}\tilde{\chi}_1^0$ **Tglu3A:** gluino pair production with $\tilde{g} \to t\bar{t}\tilde{\chi}_1^0$

Tglu3B: gluino pair production with $\tilde{g} \to t\tilde{t}$ where \tilde{t} decays exclusively

Tglu3C: gluino pair production with $\tilde{g} \to t\bar{t}$ where \tilde{t} decays exclusively

Tglu4A: gluino pair production with one gluino decaying to $q\bar{q}'\tilde{\chi}_1^{\pm}$ with $\tilde{\chi}_1^{\pm} \to W^{\pm} + \tilde{G}$, and the other gluino decaying to $q\bar{q}\tilde{\chi}_1^0$ with

Tglu4B: gluino pair production with gluinos decaying to $q\bar{q}\tilde{\chi}_1^0$ and $\tilde{\chi}_1^0 \to \gamma + \tilde{G}$.

- **Tglu4C:** gluino pair production with gluinos decaying to $\tilde{g} \to q\bar{q}\tilde{\chi}_1^0$ and $\tilde{\chi}_1^0 \to Z + \tilde{G}$.
- **Tstop1:** stop pair production with $\tilde{t} \to t \tilde{\chi}_1^0$
- **Tstop2:** stop pair production with $\tilde{t} \to b\tilde{\chi}_1^{\pm}$ with $\tilde{\chi}_1^{\pm} \to W^{\pm}\tilde{\chi}_1^0$ **Tstop3:** stop pair production with the subsequent four-body decay $\tilde{t} \rightarrow b f f' \tilde{\chi}_1^0$ where f represents a lepton or a quark
- **Tstop4:** stop pair production with $\tilde{t} \to c\tilde{\chi}_1^0$
- **Tstop5:** stop pair production with $\tilde{t} \to b\bar{\nu}\tilde{\tau}$ with $\tilde{\tau} \to \tau\tilde{G}$
- **Tstop1RPV:** stop pair production with $\tilde{t} \to \bar{b}\bar{s}$ via RPV coupling λ_{323}''
 - **Tsbot1:** sbottom pair production with $\tilde{b} \to b\tilde{\chi}_1^0$
 - **Tsbot2:** sbottom pair production with $\tilde{b} \to t \chi_1^-, \chi_1^- \to W^- \tilde{\chi}_1^0$
 - **Tsbot3:** sbottom pair production with $\tilde{b} \to b\tilde{\chi}_2^0$, where one of the $\tilde{\chi}_2^0 \to Z^{(*)}\tilde{\chi}_1^0 \to f\bar{f}\tilde{\chi}_1^0$ and the other $\tilde{\chi}_2^0 \to \ell\ell^+ \to \ell^+\ell^-\tilde{\chi}_1^0$ **Tsqk1:** squark pair production with $\tilde{q} \to q\tilde{\chi}_1^0$
- **Tchi1chi1A:** electroweak pair and associated production of nearly mass-degenerate charginos $\tilde{\chi}_1^{\pm}$ and neutralinos $\tilde{\chi}_1^0$, where $\tilde{\chi}_1^{\pm}$ decays to $\tilde{\chi}_1^0$ plus soft radiation, and where one of the $\tilde{\chi}_1^0$ decays to $\gamma + \tilde{G}$ while the other one decays to $Z/H + \tilde{G}$ (with equal probability).
- **Tchi1chi1B:** electroweak pair production of charginos $\tilde{\chi}_1^{\pm}$, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate slepton or sneutrino to $l\nu\tilde{\chi}_1^0$ and where the slepton or sneutrino mass is 5%, 25%, 50%, 75% and 95% of the $\tilde{\chi}_1^{\pm}$ mass.
- **Tchi1chi1C:** electroweak pair production of charginos $\tilde{\chi}_1^{\pm}$, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate slepton or sneutrino to $l\nu\tilde{\chi}_1^0$ and where $m_{\tilde{\ell},\tilde{\nu}} = (m_{\tilde{\chi}_1^{\pm}} + m_{\tilde{\chi}_1^0})/2$.

 - Tchi1n1A: electroweak associated production of mass-degenerate charginos $\tilde{\chi}_1^{\pm}$ and neutralinos $\tilde{\chi}_1^0$, where $\tilde{\chi}_1^{\pm}$ decays exclusively to $W^{\pm} + \tilde{G}$ and $\tilde{\chi}_1^0$ decays exclusively to $\gamma + \tilde{G}$.

 Tchi1n2A: electroweak associated production of mass-degenerate charginos $\tilde{\chi}_1^{\pm}$ and neutralinos $\tilde{\chi}_2^0$, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate slepton or sneutrino to $l\nu\tilde{\chi}_1^0$ and where $\tilde{\chi}_2^0$ decays through an intermediate slepton or sneutrino to $l^+l^-\tilde{\chi}_1^0$ or $u\bar{u}\tilde{\chi}_1^0$ $\nu\bar{\nu}\tilde{\chi}_1^0$.
 - **Tchi1n2B:** electroweak associated production of mass-degenerate charginos $\tilde{\chi}_1^{\pm}$ and neutralinos $\tilde{\chi}_2^0$, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate slepton or sneutrino to $l\nu\tilde{\chi}_1^0$ and where $chiz_2$ decays through an intermediate slepton or sneutrino to $l^+l^-\tilde{\chi}_1^0$ or $\nu\bar{\nu}\tilde{\chi}_1^0$ and where the slepton or sneutrino mass is 5%, 25%, 50%, 75% and 95% of the $\tilde{\chi}_1^{\pm}$ mass.
 - **Tchi1n2C:** electroweak associated production of mass-degenerate charginos $\tilde{\chi}_1^{\pm}$ and neutralinos $\tilde{\chi}_2^0$, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate slepton or sneutrino to $l\nu\tilde{\chi}_1^0$ and where $chiz_2$

decays through an intermediate slepton or sneutrino to $l^+l^-\tilde{\chi}_1^0$ or $\nu\bar{\nu}\tilde{\chi}_1^0$ and where $m_{\tilde{\ell},\tilde{\nu}}=(m_{\tilde{\chi}_1^{\pm}}+m_{\tilde{\chi}_1^0})/2$.

- **Tchi1n2D:** electroweak associated production of mass-degenerate charginos $\tilde{\chi}_1^{\pm}$ and neutralinos $\tilde{\chi}_2^0$, where $\tilde{\chi}_1^{\pm}$ decays through an intermediate scalar tau lepton or sneutrino to $\tau \nu \tilde{\chi}_1^0$ and where $chiz_2$ decays through an intermediate scalar tau lepton or sneutrino to $\tau^+\tau^-\tilde{\chi}_1^0$ or $\nu\bar{\nu}\tilde{\chi}_1^0$ and where $m_{\tilde{\tau},\tilde{\nu}}=(m_{\tilde{\chi}_1^{\pm}}+m_{\tilde{\chi}_1^0})/2$.
 - **Tn2n3A:** electroweak associated production of mass-degenerate neutralinos $\tilde{\chi}_2^0$ and $\tilde{\chi}_3^0$, where $\tilde{\chi}_2^0$ and $\tilde{\chi}_3^0$ decay through intermediate sleptons to $l^+l^-\tilde{\chi}_1^0$ and where the slepton mass is 5%, 25%, 50%, 75% and 95% of the $\tilde{\chi}_2^0$ mass.
 - **Tn2n3B:** electroweak associated production of mass-degenerate neutralinos $\tilde{\chi}_2^0$ and $\tilde{\chi}_3^0$, where $\tilde{\chi}_2^0$ and $\tilde{\chi}_3^0$ decay through intermediate sleptons to $l^+l^-\tilde{\chi}_1^0$ and where $m_{\tilde{\ell}} = (m_{\tilde{\chi}_2^0} + m_{\tilde{\chi}_1^0})/2$.
 - **Tglu1D:** gluino pair production with $\tilde{g} \to qq\tilde{\chi}_2^{\tilde{0}}$, and $\tilde{\chi}_2^0$ decaying through an intermediate slepton or sneutrino to $l^+l^-\tilde{\chi}_1^0$ or $\nu\bar{\nu}\tilde{\chi}_1^0$ where $m_{\tilde{\chi}_2^0} = (m_{\tilde{g}} + m_{\tilde{\chi}_1^0})/2$ and $m_{\tilde{\ell},\tilde{\nu}} = (m_{\tilde{\chi}_2^0} + m_{\tilde{\chi}_1^0})/2$.

$\widetilde{\chi}_1^0$ (Lightest Neutralino) MASS LIMIT

 $\widetilde{\chi}_1^0$ is often assumed to be the lightest supersymmetric particle (LSP). See also the $\widetilde{\chi}_2^0$, $\widetilde{\chi}_3^0$, $\widetilde{\chi}_4^0$ section below.

We have divided the $\widetilde{\chi}^0_1$ listings below into five sections:

- 1) Accelerator limits for stable $\widetilde{\chi}^0_1$,
- 2) Bounds on $\widetilde{\chi}_1^0$ from dark matter searches,
- 3) $\widetilde{\chi}_1^0-p$ elastic cross section (spin-dependent, spin-independent interactions),
- 4) Other bounds on $\widetilde{\chi}_1^0$ from astrophysics and cosmology, and
- 5) Unstable $\tilde{\chi}_1^0$ (Lightest Neutralino) mass limit.

- Accelerator limits for stable $\widetilde{\chi}_1^0$ -

Unless otherwise stated, results in this section assume spectra, production rates, decay modes, and branching ratios as evaluated in the MSSM, with gaugino and sfermion mass unification at the GUT scale. These papers generally study production of $\widetilde{\chi}_i^0\,\widetilde{\chi}_j^0$ ($i\geq 1,\, j\geq 2$), $\widetilde{\chi}_1^+\,\widetilde{\chi}_1^-$, and (in the case of hadronic collisions) $\widetilde{\chi}_1^+\,\widetilde{\chi}_2^0$ pairs. The mass limits on $\widetilde{\chi}_1^0$ are either direct, or follow indirectly from the constraints set by the non-observation of $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_2^0$ states on the gaugino and higgsino MSSM parameters M_2 and μ . In some cases, information is used from the nonobservation of slepton decays.

Obsolete limits obtained from e^+e^- collisions up to \sqrt{s} =184 GeV have been removed from this compilation and can be found in the 2000 Edition (The European Physical Journal **C15** 1 (2000)) of this Review.

$$\Delta m = m_{\widetilde{\chi}_2^0} - m_{\widetilde{\chi}_1^0}.$$

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
		$^{ m 1}$ DREINER	09	THEO	
>40	95	² ABBIENDI	04H	OPAL	all $tan\beta$, $\Delta m > 5$ GeV,
					$m_0 > 500 \text{ GeV}, A_0 = 0$
>42.4	95	³ HEISTER	04	ALEP	all $ aneta$, all Δm , all m_0
>39.2	95	⁴ ABDALLAH	03M	DLPH	all tan β , $m_{\widetilde{\nu}} >$ 500 Ge \widetilde{V}
>46	95	⁵ ABDALLAH	03M	DLPH	all tan β , all Δm , all m_0
>32.5	95	⁶ ACCIARRI	00 D	L3	$\tan \beta > 0.7$, $\Delta m > 3$ GeV, all m_0

• • • We do not use the following data for averages, fits, limits, etc. • • •

 1 DREINER 09 show that in the general MSSM with non-universal gaugino masses there exists no model-independent laboratory bound on the mass of the lightest neutralino. An essentially massless χ^0_1 is allowed by the experimental and observational data, imposing some constraints on other MSSM parameters, including $M_2,\ \mu$ and the slepton and squark masses.

² ABBIENDI 04H search for charginos and neutralinos in events with acoplanar leptons+jets and multi-jet final states in the 192–209 GeV data, combined with the results on leptonic final states from ABBIENDI 04. The results hold for a scan over the parameter space covering the region 0 < M_2 <5000 GeV, $-1000 < \mu < 1000$ GeV and $\tan \beta$ from 1 to 40. This limit supersedes ABBIENDI 00H.

³ HEISTER 04 data collected up to 209 GeV. Updates earlier analysis of selectrons from HEISTER 02E, includes a new analysis of charginos and neutralinos decaying into stau and uses results on charginos with initial state radiation from HEISTER 02J. The limit is based on the direct search for charginos and neutralinos, the constraints from the slepton search and the Higgs mass limits from HEISTER 02 using a top mass of 175 GeV, interpreted in a framework with universal gaugino and sfermion masses. Assuming the mixing in the stau sector to be negligible, the limit improves to 43.1 GeV. Under the assumption of MSUGRA with unification of the Higgs and sfermion masses, the limit improves to 50 GeV, and reaches 53 GeV for $A_0=0$. These limits include and update the results of BARATE 01.

⁴ ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV. A limit on the mass of $\widetilde{\chi}_1^0$ is derived from direct searches for neutralinos combined with the chargino search. Neutralinos are searched in the production of $\widetilde{\chi}_1^0\widetilde{\chi}_2^0$, $\widetilde{\chi}_1^0\widetilde{\chi}_3^0$, as well as $\widetilde{\chi}_2^0\widetilde{\chi}_3^0$ and $\widetilde{\chi}_2^0\widetilde{\chi}_4^0$ giving rise to cascade decays, and $\widetilde{\chi}_1^0\widetilde{\chi}_2^0$ and $\widetilde{\chi}_1^0\widetilde{\chi}_2^0$, followed by the decay $\widetilde{\chi}_2^0 \to \widetilde{\tau}\tau$. The results hold for the parameter space defined by values of $M_2 < 1$ TeV, $|\mu| \le 2$ TeV with the $\widetilde{\chi}_1^0$ as LSP. The limit is obtained for $\tan\beta = 1$ and large m_0 , where $\widetilde{\chi}_2^0\widetilde{\chi}_4^0$ and chargino pair production are important. If the constraint from Higgs searches is also imposed, the limit improves to 49.0 GeV in the $m_h^{\rm max}$ scenario with $m_t=174.3$ GeV. These limits update the results of ABREU 00J.

 5 ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV. An indirect limit on the mass of $\widetilde{\chi}^0_1$ is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays and $\widetilde{\tau}\tau$ final states), for charginos (for all Δm_+) and for sleptons, stop and sbottom. The results hold for the full parameter space defined by values of $M_2<1$ TeV, $|\mu|\leq 2$ TeV with the $\widetilde{\chi}^0_1$ as LSP. Constraints from the Higgs search in the $m_h^{\rm max}$ scenario assuming m_t =174.3 GeV are included. The limit is obtained for $\tan\beta\geq 5$ when stau mixing leads to mass degeneracy between $\widetilde{\tau}_1$

and $\widetilde{\chi}_1^0$ and the limit is based on $\widetilde{\chi}_2^0$ production followed by its decay to $\widetilde{\tau}_1 \tau$. In the pathological scenario where m_0 and $|\mu|$ are large, so that the $\widetilde{\chi}_2^0$ production cross section is negligible, and where there is mixing in the stau sector but not in stop nor sbottom, the limit is based on charginos with soft decay products and an ISR photon. The limit then degrades to 39 GeV. See Figs. 40–42 for the dependence of the limit on $\tan\beta$ and $m_{\widetilde{\nu}}$. These limits update the results of ABREU 00W.

⁶ ACCIARRI 00D data collected at \sqrt{s} =189 GeV. The results hold over the full parameter space defined by 0.7 \leq tan β \leq 60, 0 \leq M_2 \leq 2 TeV, m_0 \leq 500 GeV, $|\mu|$ \leq 2 TeV. The minimum mass limit is reached for tan β =1 and large m_0 . The results of slepton searches from ACCIARRI 99W are used to help set constraints in the region of small m_0 . The limit improves to 48 GeV for m_0 \gtrsim 200 GeV and tan β \gtrsim 10. See their Figs. 6–8 for the tan β and m_0 dependence of the limits. Updates ACCIARRI 98F.

 7 AAD 14K sets limits on the χ -nucleon spin-dependent and spin-independent cross sections out to $m_\chi=10$ TeV.

- Bounds on $\widetilde{\chi}^0_1$ from dark matter searches

These papers generally exclude regions in the M_2 - μ parameter plane assuming that $\widetilde{\chi}^0_1$ is the dominant form of dark matter in the galactic halo. These limits are based on the lack of detection in laboratory experiments, telescopes, or by the absence of a signal in underground neutrino detectors. The latter signal is expected if $\widetilde{\chi}^0_1$ accumulates in the Sun or the Earth and annihilates into high-energy $\nu^{\rm T}$ s.

VALUE DOCUMENT ID TECN

• • • We do not use the following data for averages, fits, limits, etc. • •

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<sup>1</sup> AARTSEN
                        16c ICCB
 <sup>2</sup> AARTSEN
                       16D ICCB
 <sup>3</sup> ABDALLAH
                              HESS
 <sup>4</sup> ABDALLAH
                       16A HESS
 <sup>5</sup> ADRIAN-MAR..16
                              ANTR
 <sup>6</sup> AHNEN
 <sup>7</sup> AVRORIN
                              BAIK
                       16
 <sup>8</sup> CIRELLI
                       16
                              THEO
 <sup>8</sup> LEITE
                              THEO
 <sup>9</sup> AARTSEN
                       15E ICCB
<sup>10</sup> ABRAMOWSKI15
                              HESS
<sup>11</sup> ACKERMANN 15
                              FLAT
<sup>12</sup> ACKERMANN 15A FLAT
<sup>13</sup> ACKERMANN 15B
                             FLAT
<sup>14</sup> ADRIAN-MAR..15
                              ANTR
<sup>15</sup> BUCKLEY
                              THEO
<sup>16</sup> CHOI
                              SKAM
<sup>17</sup> ALEKSIC
                              MGIC
<sup>18</sup> AVRORIN
                       14
                              BAIK
<sup>19</sup> AARTSEN
                       13
                              ICCB
<sup>20</sup> AARTSEN
                       13C ICCB
<sup>21</sup> ABRAMOWSKI13
                              HESS
<sup>22</sup> ADRIAN-MAR..13
                              ANTR
<sup>23</sup> BERGSTROM 13
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24	BOLIEV	13	BAKS
23	JIN	13	ASTR
23	KOPP	13	COSM
25	ABBASI	12	ICCB
26	ABRAMOWSK	l11	HESS
27	ABDO	10	FLAT
28	ACKERMANN	10	FLAT
29	ABBASI	09 B	ICCB
30	ACHTERBERG	06	AMND
31	ACKERMANN	06	AMND
32	DEBOER	06	RVUE
33	DESAI	04	SKAM
33	AMBROSIO	99	MCRO
34	LOSECCO	95	RVUE
35	MORI	93	KAMI
36	BOTTINO	92	COSM
37	BOTTINO	91	RVUE
38	GELMINI	91	COSM
39	KAMIONKOW.	91	RVUE
40	MORI	91 B	KAMI
41	OLIVE	88	COSM

none 4-15 GeV

- 1 AARTSEN 16 C is based on data collected during 317 effective days with the IceCube 79-string detector including the DeepCore sub-array. They looked for interactions of ν 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the spin dependent neutralino-proton cross section for neutralino masses in the range 10–10000 GeV. This updates AARTSEN 13.
- 2 AARTSEN 16D is based on 329 live days of running with the DeepCore subdetector of the IceCube detector. They set a limit of $10^{-23}~{\rm cm}^3{\rm s}^{-1}$ on the annihilation cross section to $\nu\overline{\nu}$. This updates AARTSEN 15C.
- ³ ABDALLAH 16 places constraints on the dark matter annihilation cross section for annihilations in the Galactic center for masses between 200 GeV to 70 TeV. This updates ABRAMOWSKI 15.
- ⁴ ABDALLAH 16A place upper limits on the annihilation cross section with final states in the energy range of 0.1 to 2 TeV. This complements ABRAMOWSKI 13.
- 5 ADRIAN-MARTINEZ 16 is based on data from the ANTARES neutrino telescope. They looked for interactions of ν 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon neutrino flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 50 to 5,000 GeV. This updates ADRIAN-MARTINEZ 13.
- ⁶ AHNEN 16 combines 158 hours of Segue 1 observations with MAGIC with 6 year observations of 15 dwarf satellite galaxies by Fermi-LAT to set limits on annihilation cross sections for dark matter masses between 10 GeV and 100 TeV.
- ⁷ AVRORIN 16 is based on 2.76 years with Lake Baikal neutrino telescope. They derive 90% upper limits on the annihilation cross section from dark matter annihilations in the Galactic center.
- ⁸ CIRELLI 16 and LEITE 16 derive bounds on the annihilation cross section from radio observations.
- 9 AARTSEN 15E is based on 319.7 live days of running with the IceCube 79-string detector. They set a limit of $4\times 10^{-24}~\text{cm}^3\text{s}^{-1}$ on the annihilation cross section to $\nu\overline{\nu}$ for dark matter with masses between 30–10000 GeV annihilating in the Galactic center assuming an NFW profile.
- ¹⁰ ABRAMOWSKI 15 places constraints on the dark matter annihilation cross section for annihilations in the Galactic center for masses between 300 GeV to 10 TeV.

- ACKERMANN 15 is based on 5.8 years of data with Fermi-LAT and search for monochromatic gamma-rays in the energy range of 0.2–500 GeV from dark matter annihilations. This updates ACKERMANN 13A.
- ¹² ACKERMANN 15A is based on 50 months of data with Fermi-LAT and search for dark matter annihilation signals in the isotropic gamma-ray background as well as galactic subhalos in the energy range of a few GeV to a few tens of TeV.
- 13 ACKERMANN 15 B is based on 6 years of data with Fermi-LAT observations of Milky Way dwarf spheroidal galaxies. Set limits on the annihilation cross section from $m_{\chi}=2$ GeV to 10 TeV. This updates ACKERMANN 14.
- 14 ADRIAN-MARTINEZ 15 is based on data from the ANTARES neutrino telescope. They looked for interactions of ν_{μ} 's from neutralino annihilations in the galactic center over a background of atmospheric neutrinos and set 90% CL limits on the muon neutrino flux. They also set limits on the annihilation cross section for wimp masses of 25–10000 GeV.
- ¹⁵ BUCKLEY 15 is based on 5 years of Fermi-LAT data searching for dark matter annihilation signals from Large Magellanic Cloud.
- 16 CHOI 15 is based on 3903 days of SuperKamiokande data searching for neutrinos produced from dark matter annihilations in the sun. They place constraints on the dark matter-nucleon scattering cross section for dark matter masses between 4–200 GeV.
- 17 ALEKSIC 14 is based on almost 160 hours of observations of Segue 1 satellite dwarf galaxy using the MAGIC telescopes between 2011 and 2013. Sets limits on the annihilation cross section out to $m_{\chi}=10$ TeV.
- ¹⁸ AVRORIN 14 is based on almost 2.76 years with Lake Baikal neutrino telescope. They derive 90% upper limits on the fluxes of muons and muon neutrinos from dark matter annihilations in the Sun.
- AARTSEN 13 is based on data collected during 317 effective days with the IceCube 79-string detector including the DeepCore sub-array. They looked for interactions of ν_{μ} 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 20–5000 GeV.
- AARTSEN 13C is based on data collected during 339.8 effective days with the IceCube 59-string detector. They looked for interactions of ν_{μ} 's from neutralino annihilations in nearby galaxies and galaxy clusters. They obtain limits on the neutralino annihilation cross section for neutralino masses in the range 30–100,000 GeV.
- 21 ABRAMOWSKI 13 place upper limits on the annihilation cross section with $\gamma\gamma$ final states in the energy range of 0.5–25 TeV.
- 22 ADRIAN-MARTINEZ 13 is based on data from the ANTARES neutrino telescope. They looked for interactions of ν_{μ} 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 50–10, 000 GeV.
- ²³ BERGSTROM 13, JIN 13, and KOPP 13 derive limits on the mass and annihilation cross section using AMS-02 data. JIN 13 also sets a limit on the lifetime of the dark matter particle.
- 24 BOLIEV 13 is based on data collected during 24.12 years of live time with the Bakson Underground Scintillator Telescope. They looked for interactions of ν_{μ} 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent and spin independent neutralino-proton cross section for neutralino masses in the range 10–1000 GeV.
- 25 ABBASI 12 is based on data collected during 812 effective days with AMANDA II and 149 days of the IceCube 40-string detector combined with the data of ABBASI 09B. They looked for interactions of ν_{μ} 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. No excess is observed. They also obtain limits on the spin dependent neutralino-proton cross section for neutralino masses in the range 50–5000 GeV.

- 26 ABRAMOWSKI 11 place upper limits on the annihilation cross section with $\gamma\gamma$ final states.
- ²⁷ ABDO 10 place upper limits on the annihilation cross section with $\gamma\gamma$ or $\mu^+\mu^-$ final states.
- ²⁸ ACKERMANN 10 place upper limits on the annihilation cross section with $b\overline{b}$ or $\mu^+\mu^-$ final states.
- ABBASI 09B is based on data collected during 104.3 effective days with the IceCube 22-string detector. They looked for interactions of ν_{μ} 's from neutralino annihilations in the Sun over a background of atmospheric neutrinos and set 90% CL limits on the muon flux. They also obtain limits on the spin dependent neutralino–proton cross section for neutralino masses in the range 250–5000 GeV.
- 30 ACHTERBERG 06 is based on data collected during 421.9 effective days with the AMANDA detector. They looked for interactions of $\nu_{\mu} s$ from the centre of the Earth over a background of atmospheric neutrinos and set 90 % CL limits on the muon flux. Their limit is compared with the muon flux expected from neutralino annihilations into $W^+\,W^-$ and $b\,\overline{b}$ at the centre of the Earth for MSSM parameters compatible with the relic dark matter density, see their Fig. 7.
- 31 ACKERMANN 06 is based on data collected during 143.7 days with the AMANDA-II detector. They looked for interactions of ν_{μ} s from the Sun over a background of atmospheric neutrinos and set 90 % CL limits on the muon flux. Their limit is compared with the muon flux expected from neutralino annihilations into $W^+\,W^-$ in the Sun for SUSY model parameters compatible with the relic dark matter density, see their Fig. 3.
- 32 DEBOER 06 interpret an excess of diffuse Galactic gamma rays observed with the EGRET satellite as originating from π^0 decays from the annihilation of neutralinos into quark jets. They analyze the corresponding parameter space in a supergravity inspired MSSM model with radiative electroweak symmetry breaking, see their Fig. 3 for the preferred region in the $(m_0, m_{1/2})$ plane of a scenario with large $\tan\beta$.
- ³³ AMBROSIO 99 and DESAI 04 set new neutrino flux limits which can be used to limit the parameter space in supersymmetric models based on neutralino annihilation in the Sun and the Earth.
- 34 LOSECCO 95 reanalyzed the IMB data and places lower limit on $m_{\widetilde{\chi}^0_1}$ of 18 GeV if the LSP is a photino and 10 GeV if the LSP is a higgsino based on LSP annihilation in the sun producing high-energy neutrinos and the limits on neutrino fluxes from the IMB detector.
- ³⁵ MORI 93 excludes some region in M_2 - μ parameter space depending on $\tan\beta$ and lightest scalar Higgs mass for neutralino dark matter $m_{\widetilde{\chi}0} > m_W$, using limits on upgoing muons produced by energetic neutrinos from neutralino annihilation in the Sun and the Earth.
- 36 BOTTINO 92 excludes some region M_2 - μ parameter space assuming that the lightest neutralino is the dark matter, using upgoing muons at Kamiokande, direct searches by Ge detectors, and by LEP experiments. The analysis includes top radiative corrections on Higgs parameters and employs two different hypotheses for nucleon-Higgs coupling. Effects of rescaling in the local neutralino density according to the neutralino relic abundance are taken into account.
- 37 BOTTINO 91 excluded a region in $M_2-\mu$ plane using upgoing muon data from Kamioka experiment, assuming that the dark matter surrounding us is composed of neutralinos and that the Higgs boson is not too heavy.
- 38 GELMINI 91 exclude a region in $M_2-\mu$ plane using dark matter searches.
- 39 KAMIONKOWSKI 91 excludes a region in the M_2 - μ plane using IMB limit on upgoing muons originated by energetic neutrinos from neutralino annihilation in the sun, assuming that the dark matter is composed of neutralinos and that $m_{H_1^0} \lesssim 50$ GeV. See Fig. 8 in the paper.
- ⁴⁰ MORI 91B exclude a part of the region in the M_2 - μ plane with $m_{\widetilde{\chi}_1^0} \lesssim$ 80 GeV using a limit on upgoing muons originated by energetic neutrinos from neutralino annihilation

in the earth, assuming that the dark matter surrounding us is composed of neutralinos and that $m_{H_1^0}\lesssim$ 80 GeV.

⁴¹ OLIVE 88 result assumes that photinos make up the dark matter in the galactic halo. Limit is based on annihilations in the sun and is due to an absence of high energy neutrinos detected in underground experiments. The limit is model dependent.

$\widetilde{\chi}_1^0$ -ho elastic cross section $\overline{}$

Experimental results on the $\widetilde{\chi}_1^0$ -p elastic cross section are evaluated at $m_{\widetilde{\chi}_1^0}$ =100 GeV. The experimental results on the cross section are often mass dependent. Therefore, the mass and cross section results are also given where the limit is strongest, when appropriate. Results are quoted separately for spin-dependent interactions (based on an effective 4-Fermi Lagrangian of the form $\overline{\chi}\gamma^\mu\gamma^5\chi\overline{q}\gamma_\mu\gamma^5q$) and spin-independent interactions ($\overline{\chi}\chi\overline{q}\,q$). For calculational details see GRIEST 88B, ELLIS 88D, BAR-BIERI 89C, DREES 93B, ARNOWITT 96, BERGSTROM 96, and BAER 97 in addition to the theory papers listed in the Tables. For a description of the theoretical assumptions and experimental techniques underlying most of the listed papers, see the review on "Dark matter" in this "Review of Particle Physics," and references therein. Most of the following papers use galactic halo and nuclear interaction assumptions from (LEWIN 96).

Spin-dependent interactions

VAL	UE (pb)		CL%		DOCUMENT ID		TECN	COMMENT
• •	• We	do not use the	followin	g d	ata for averages	, fits,	limits, e	tc. • • •
<	7	$\times 10^{-3}$	90	1	AKERIB	16A	LUX	Xe
<	5	\times 10 ⁻⁴	90		AMOLE	16	PICO	CF ₃ I
<	1	\times 10 ⁻³	90		AMOLE	16A	PICO	C_3F_8
<	6.8	$\times 10^{-3}$	90		APRILE	16 B	X100	Xe
<	6.3	\times 10 ⁻³	90		FELIZARDO	14	SMPL	C ₂ CIF ₅
<	0.01		90		AKIMOV	12	ZEP3	Xe
<	0.07	_	90		ARCHAMBAU.	.12	PICA	F
<	7	\times 10 ⁻³			BEHNKE	12	COUP	
<	-	_	90		DAW	12	DRFT	CS ₂ ; CF ₄
<	8.5	\times 10 ⁻³			FELIZARDO	12	SMPL	C ₂ CIF ₅
<			90		KIM	12	KIMS	Csl
$5 \times$	10^{-10}	0 to 10^{-5}	95	12	BUCHMUEL	11 B	THEO	
<			90	13	ANGLE		XE10	Xe
<	0.055			14	BEDNYAKOV		HDMS	
<	0.33		90	15	BEHNKE	80	COUP	3
<	5			10	AKERIB	06	CDMS	
<				1/	SHIMIZU		CNTR	-
	0.4			10	ALNER	05		Nal Spin Dep.
<				19	BARNABE-HE.		PICA	C
$2 \times$	10-1	$^{ m l}$ to $1 imes 10^{-4}$		20	ELLIS	04		$\mu > 0$
<	8.0			21	AHMED	03		Nal Spin Dep.
< '	40			22	TAKEDA	03		NaF Spin Dep.
< 1	10			23	ANGLOHER	02	CRES	Saphire

8×10^{-7} to 2×10^{-5}	²⁴ ELLIS	01 C	THEO	$ an\!eta\!\le 10$
< 3.8	²⁵ BERNABEI	00 D	DAMA	Xe
< 0.8	SPOONER	00	UKDM	Nal
< 4.8	²⁶ BELLI	99C	DAMA	F
<100	²⁷ OOTANI	99	BOLO	LiF
< 0.6	BERNABEI	98C	DAMA	Xe
< 5	²⁶ BERNABEI	97	DAMA	F

 $^{^{1}}$ The strongest limit is 2.9×10^{-3} pb at $m_{\chi}=33$ GeV. The limit for scattering on neutrons is 2×10^{-4} pb at 100 GeV and is 9.4×10^{-5} pb at 33 GeV. ² The strongest limit is 5×10^{-4} pb at $m_\chi=80$ GeV.

 $^{\mbox{GeV}.}$ 8 The strongest limit is 6 \times 10 $^{-3}$ at $m_{\chi}=$ 60 GeV.

- $^{10}\,\mathrm{The}$ strongest limit is 5.7×10^{-3} at $m_\chi=35$ GeV.
- 11 This result updates LEE 07A. The strongest limit is at $m_\chi=80$ GeV.
- 12 Predictions for the spin-dependent elastic cross section based on a frequentist approach to electroweak observables in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 13 The strongest limit is 0.6 pb and occurs at $m_\chi =$ 30 GeV. The limit for scattering on neutrons is 0.01 pb at $m_{\gamma} = 100$ GeV, and the strongest limit is 0.0045 pb at $m_{\gamma} =$
- 14 Limit applies to neutron elastic cross section.
- $^{15}\,\mathrm{The}$ strongest upper limit is 0.25 pb and occurs at $m_\chi \simeq$ 40 GeV.
- 16 The strongest upper limit is 4 pb and occurs at $m_\chi\simeq 60$ GeV. The limit on the neutron spin-dependent elastic cross section is 0.07 pb. This latter limit is improved in AHMED 09, where a limit of 0.02 pb is obtained at $m_\chi=100$ GeV. The strongest limit in AHMED 09 is 0.018 pb and occurs at $m_\chi =$ 60 GeV.

 $^{^3}$ The strongest limit is 6.5×10^{-4} pb at $m_\chi=30$ GeV. This updates AMOLE 15.

 $^{^4}$ The strongest limit is 5.2 \times 10 $^{-3}$ pb at 50 GeV. The limit for scattering on neutrons is 2.8×10^{-4} pb at 100 GeV and the strongest limit is 2.0 \times 10 $^{-4}$ pb at 50 GeV. This updates APRILE 13.

 $^{^5}$ The strongest limit is 0.0043 pb and occurs at $m_\chi=$ 35 GeV. FELIZARDO 14 also presents limits for the scattering on neutrons. At $m_\chi^\Lambda=100$ GeV, the upper limit is 0.13 pb and the strongest limit is 0.066 pb at $m_\chi=35$ GeV.

 $^{^6}$ This result updates LEBEDENKO 09A. The strongest limit is $8 imes 10^{-3}$ pb at $m_{_Y} = 50$ GeV. Limit applies to the neutralino neutron elastic cross section.

 $^{^7\,{\}rm This}$ result updates ARCHAMBAULT 09. The strongest limit is 0.032 pb at $m_\chi=20$

 $^{^9\,\}mathrm{The}$ strongest limit is 1.8 pb and occurs at $m_\chi=100$ GeV.

 $^{^{17}}$ The strongest upper limit is 1.2 pb and occurs at $m_\chi~\simeq~$ 40 GeV. The limit on the neutron spin-dependent cross section is 35 pb.

 $^{^{18}}$ The strongest upper limit is 0.35 pb and occurs at $m_\chi~\simeq~60$ GeV.

 $^{^{19}\,\}mathrm{The}$ strongest upper limit is 1.2 pb and occurs $m_\chi~\simeq~30$ GeV.

 $^{^{20}}$ ELLIS 04 calculates the χp elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry, but without universal scalar masses. In the case of universal squark and slepton masses, but non-universal Higgs masses, the limit becomes 2×10^{-4} , see ELLIS 03E.

 $^{^{21}\,\}mathrm{The}$ strongest upper limit is 0.75 pb and occurs at $m_\chi\approx 70$ GeV.

 $^{^{22}}$ The strongest upper limit is 30 pb and occurs at $m_{\chi}~\approx~20$ GeV.

 $^{^{23}\,\}text{The}$ strongest upper limit is 8 pb and occurs at $m_\chi \simeq 30$ GeV.

Spin-independent interactions

VALUE (pb)	CL%	DOCUMENT ID		TECN	COMMENT
• • • We do not use the following	owing o	lata for averages, fits	, limi	ts, etc.	• • •
$< 1.8 \times 10^{-10}$	90	¹ AKERIB	17	LUX	Xe
$< 3 \times 10^{-8}$	90	AMOLE	16	PICO	CF ₃ I
$< 1.5 \times 10^{-9}$	90	² APRILE	16 B	X100	Xe
$<$ 4 \times 10 ⁻⁹	90	³ TAN	16	PANX	Xe
$< 4 \times 10^{-10}$	90	⁴ TAN	16 B	PANX	Xe
$< 6.1 \times 10^{-8}$	90	AGNES	15	DSID	Ar
$< 2.2 \times 10^{-8}$	90	⁵ AGNESE	15 B	CDMS	Ge
$< 1.5 \times 10^{-9}$	90	⁶ AKERIB	14	LUX	Xe
10^{-11} – 10^{-7}	95	⁷ BUCHMUEL		THEO	
$< 4.6 \times 10^{-6}$	90	⁸ FELIZARDO	14	SMPL	C ₂ CIF ₅
10^{-11} -10^{-8}	95	⁹ ROSZKOWSKI	14	THEO	
$< 2.2 \times 10^{-6}$	90	10 AGNESE	13	CDMS	Si
$< 5 \times 10^{-8}$	90	11 AKIMOV	12	ZEP3	Xe
1.6×10^{-6} ; 3.7×10^{-5}		12 ANGLOHER	12	CRES	CaWO ₄
$3 \times 10^{-12} \text{ to } 3 \times 10^{-9}$	95	13 BECHTLE	12	THEO	
$< 1.6 \times 10^{-7}$		14 BEHNKE	12	COUP	
$< 6.5 \times 10^{-6}$		¹⁵ FELIZARDO	12		C ₂ CIF ₅
$< 2.3 \times 10^{-7}$	90	¹⁶ KIM	12	KIMS	Csl
$< 3.3 \times 10^{-8}$	90	17 AHMED	11A		Ge
$< 4.4 \times 10^{-8}$	90	18 ARMENGAUD		EDE2	Ge
$< 7 \times 10^{-7}$	90	¹⁹ ANGLOHER	09	CRES	CaWO ₄
$< 1 \times 10^{-7}$	90	²⁰ ANGLE	80	XE10	Xe
$< 1 \times 10^{-6}$	90	BENETTI 21 ALNED	08	WARP	
$< 7.5 \times 10^{-7}$	90	21 ALNER	07A	ZEP2	Xe
$< 2 \times 10^{-7} $ $< 90 \times 10^{-7}$		²² AKERIB	06A	CDMS	
–		ALNER ²³ ALNER	05 05 A	NAIA ZEPL	Nal Spin Indep.
$<12 \times 10^{-7}$ $<20 \times 10^{-7}$		²⁴ ANGLOHER	05A 05	CRES	CaWO ₄
$< 20 \times 10^{-7}$		SANGLARD	05 05	EDEL	Ge
$< 4 \times 10^{-7}$		25 AKERIB	03	CDMS	
2×10^{-11} to 1.5×10^{-7}	95	26 BALTZ	04	THEO	Ge
2×10^{-11} to 8×10^{-6}	95	27,28 ELLIS	04	THEO	<i>u</i> > 0
$< 5 \times 10^{-8}$		²⁹ PIERCE	04A	THEO	$\mu > 0$
$< 2 \times 10^{-5}$		30 AHMED	03	NAIA	Nal Spin Indep.
$< 3 \times 10^{-6}$		31 AKERIB	03	CDMS	
2×10^{-13} to 2×10^{-7}		32 BAER	03A	THEO	3 0
$< 1.4 \times 10^{-5}$		33 KLAPDOR-K		HDMS	Ge
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 $^{^{24}}$ ELLIS 01C calculates the χ -p elastic scattering cross section in the framework of $N\!\!=\!1$ supergravity models with radiative breaking of the electroweak gauge symmetry. In models with nonuniversal Higgs masses, the upper limit to the cross section is 6×10^{-4} .

 $^{^{25}}$ The strongest upper limit is 3 pb and occurs at $m_\chi \simeq$ 60 GeV. The limits are for inelastic scattering $X^0+{}^{129}{\rm Xe} \to X^0+{}^{129}{\rm Xe}^*$ (39.58 keV). ²⁶ The strongest upper limit is 4.4 pb and occurs at $m_\chi \simeq$ 60 GeV.

 $^{^{27}}$ The strongest upper limit is about 35 pb and occurs at $m_\chi \simeq 15$ GeV.

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<sup>34</sup> ABRAMS
< 6 \times 10^{-6}
                                                                    02
                                                                          CDMS Ge
< 1.4 \times 10^{-6}
                                             <sup>35</sup> BENOIT
                                                                          EDEL Ge
1 \times 10^{-12} to 7 \times 10^{-6}
                                             <sup>27</sup> KIM
                                                                    02B THEO
                                             <sup>36</sup> MORALES
 < 3 \times 10^{-5}
                                                                    02B CSME Ge
< 1 \times 10^{-5}
                                             <sup>37</sup> MORALES
                                                                    02C IGEX
< 1 \times 10^{-6}
                                                 BALTZ
                                                                          THEO
                                             <sup>38</sup> BAUDIS
< 3 \times 10^{-5}
                                                                          HDMS Ge
< 4.5 \times 10^{-6}
                                                 BENOIT
                                                                          EDEL
                                             <sup>39</sup> BOTTINO
< 7 \times 10^{-6}
                                                                          THEO
                                             <sup>40</sup> CORSETTI
< 1 \times 10^{-8}
                                                                    01
                                                                          THEO tan \beta \leq 25
5\times10^{-10} to 1.5\times10^{-8}
                                             <sup>41</sup> ELLIS
                                                                    01C THEO tan \beta \leq 10
<~4~~\times10^{-6}
                                             <sup>40</sup> GOMEZ
                                                                          THEO
2 \times 10^{-10} to 1 \times 10^{-7}
                                             <sup>40</sup> LAHANAS
                                                                          THEO
< 3 \times 10^{-6}
                                                 ABUSAIDI
                                                                    00
                                                                          CDMS Ge, Si
                                             42 ACCOMANDO 00
< 6 \times 10^{-7}
                                                                          THEO
                                             <sup>43</sup> BERNABEI
                                                                          DAMA Nal
2.5 \times 10^{-9} to 3.5 \times 10^{-8}
                                             44 FENG
                                                                          THEO tan\beta=10
< 1.5 \times 10^{-5}
                                                 MORALES
                                                                    00
                                                                          IGEX Ge
 < 4 \times 10^{-5}
                                                                          UKDM Nal
                                                 SPOONER
                                                                    00
< 7 \times 10^{-6}
                                                                          HDMO <sup>76</sup>Ge
                                                                    99
                                                 BAUDIS
 < 7 \times 10^{-6}
                                                 BERNABEI
                                                                    98C DAMA Xe
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 $^{^1}$ The strongest limit is $1 imes 10^{-10}$ at 40 GeV. This updates AKERIB 16.

The strongest limit is 1×10^{-9} pb at 50 GeV. This updates APRILE 12. The strongest limit is 3×10^{-9} pb at $m_{\chi} = 45$ GeV. This updates XIAO 15.

⁴ The strongest limit is 2.5×10^{-10} pb at $m_\chi = 40$ GeV.

 $^{^{5}}$ AGNESE 15B result updates AHMED 10 and AHMED 09. The strongest limit is 1.8 imes 10^{-8} pb and occurs at $m_{\chi} = 60$ GeV.

⁶ The strongest upper limit is 7.6×10^{-10} at $m_V = 33$ GeV.

 $^{^7}$ Predictions for the spin-independent elastic cross section based on a frequentist approach to electroweak observables in the framework of ${\it N}=1$ supergravity models with radiative breaking of the electroweak gauge symmetry using the 20 fb $^{-1}$ 8 TeV and the 5 fb $^{-1}$ 7 TeV LHC data and the LUX data.

 $^{^8}$ The strongest limit is 3.6×10^{-6} pb and occurs at $m_\chi=35$ GeV.

 $^{^{9}}$ Predictions for the spin-independent elastic cross section based on a Bayesian approach to electroweak observables in the framework of ${\it N}=1$ supergravity models with radiative breaking of the electroweak gauge symmetry using the 20 fb $^{-1}$ LHC data and LUX.

 $^{^{10}\,\}mathsf{AGNESE}$ 13 presents 90% CL limits on the elastic cross section for masses in the range 7–100 GeV using the Si based detector. The strongest upper limit is 1.8×10^{-6} pb at $m_\chi=50$ GeV. This limit is improved to 7×10^{-7} pb in AGNESE 13A.

¹¹ This result updates LEBEDENKO 09. The strongest limit is 3.9×10^{-8} pb at $m_V =$ 52 GeV.

¹² ANGLOHER 12 presents results of 730 kg days from the CRESST-II dark matter detector. They find two maxima in the likelihood function corresponding to best fit WIMP masses of 25.3 and 11.6 GeV with elastic cross sections of 1.6×10^{-6} and 3.7×10^{-5} pb respectively, see their Table 4. The statistical significance is more than 4σ .

 $^{^{13}}$ Predictions for the spin-independent elastic cross section based on a frequentist approach to electroweak observables in the framework of ${\it N}=1$ supergravity models with radiative breaking of the electroweak gauge symmetry using the 5 fb $^{-1}$ LHC data and XENON100. ¹⁴ The strongest limit is 1.4×10^{-7} at $m_\chi=60$ GeV.

 $^{^{15}\,\}text{The strongest limit is } 4.7\times 10^{-6}\,\,\text{at }m_\chi^{}=35\,\,\text{GeV}.$

- 16 This result updates LEE 07A. The strongest limit is 2.1×10^{-7} at $m_\chi = 70$ GeV.
- $^{
 m 17}$ AHMED 11A gives combined results from CDMS and EDELWEISS. The strongest limit is at $m_{\chi}=$ 90 GeV.
- 18 ARMENGAUD 11 updates result of ARMENGAUD 10. Strongest limit at $m_{_Y}=85~{\rm GeV}.$
- 19 The strongest upper limit is 4.8×10^{-7} pb and occurs at $m_{_Y} = 50$ GeV.
- 20 The strongest upper limit is 5.1×10^{-8} pb and occurs at $m_{\gamma} \simeq 30$ GeV. The values quoted here are based on the analysis performed in ANGLE 08 with the update from SORENSEN 09.
- ²¹ The strongest upper limit is 6.6×10^{-7} pb and occurs at $m_\chi \simeq 65$ GeV.
- 22 AKERIB 06A updates the results of AKERIB 05. The strongest upper limit is 1.6 imes $10^{-7}~\mathrm{pb}$ and occurs at $m_\chi~\approx~60~\mathrm{GeV}.$
- ²³ The strongest upper limit is also close to 1.0×10^{-6} pb and occurs at $m_{\chi} \simeq 70$ GeV. BENOIT 06 claim that the discrimination power of ZEPLIN-I measurement (ALNER 05A) is not reliable enough to obtain a limit better than 1×10^{-3} pb. However, SMITH 06do not agree with the criticisms of BENOIT 06.
- ²⁴ The strongest upper limit is also close to 1.4×10^{-6} pb and occurs at $m_{\chi} \simeq 70$ GeV.
- 25 AKERIB 04 is incompatible with BERNABEI 00 most likely value, under the assumption of standard WIMP-halo interactions. The strongest upper limit is 4×10^{-7} pb and occurs at $m_\chi \simeq 60$ GeV.
- 26 Predictions for the spin-independent elastic cross section in the framework of ${\it N}=1$ supergravity models with radiative breaking of the electroweak gauge symmetry.
- 27 KIM 02 and ELLIS 04 calculate the χp elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry, but without universal scalar masses.
- 28 In the case of universal squark and slepton masses, but non-universal Higgs masses, the limit becomes 2×10^{-6} (2×10^{-11} when constraint from the BNL g-2 experiment are included), see ELLIS 03E. ELLIS 05 display the sensitivity of the elastic scattering cross section to the π -Nucleon Σ term.
- PIERCE 04A calculates the χp elastic scattering cross section in the framework of models
- with very heavy scalar masses. See Fig. 2 of the paper. 30 The strongest upper limit is 1.8×10^{-5} pb and occurs at $m_\chi\approx 80$ GeV.
- $^{
 m 31}$ Under the assumption of standard WIMP-halo interactions, Akerib 03 is incompatible with BERNABEI 00 most likely value at the 99.98% CL. See Fig. 4.
- 32 BAER 03A calculates the χp elastic scattering cross section in several models including the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 33 The strongest upper limit is 7×10^{-6} pb and occurs at $m_\chi\simeq 30$ GeV.
- 34 ABRAMS 02 is incompatible with the DAMA most likely value at the 99.9% CL. The strongest upper limit is 3×10^{-6} pb and occurs at $m_{\chi} \simeq 30$ GeV.
- 35 BENOIT 02 excludes the central result of DAMA at the 99.8%CL. 36 The strongest upper limit is 2 \times 10 $^{-5}$ pb and occurs at $m_\chi \simeq$ 40 GeV.
- 37 The strongest upper limit is 7×10^{-6} pb and occurs at $m_\chi^{\sim} \simeq$ 46 GeV.
- 38 The strongest upper limit is 1.8×10^{-5} pb and occurs at $m_{_Y} \simeq$ 32 GeV
- 39 BOTTINO 01 calculates the χ -p elastic scattering cross section in the framework of the following supersymmetric models: N=1 supergravity with the radiative breaking of the electroweak gauge symmetry, N=1 supergravity with nonuniversal scalar masses and an effective MSSM model at the electroweak scale.
- 40 Calculates the χ -p elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.

- ⁴¹ ELLIS 01C calculates the χ -p elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. ELLIS 02B find a range 2×10^{-8} – 1.5×10^{-7} at $\tan\beta$ =50. In models with nonuniversal Higgs masses, the upper limit to the cross section is 4×10^{-7} .
- 42 ACCOMANDO 00 calculate the $\chi\text{-}p$ elastic scattering cross section in the framework of minimal $N\!\!=\!\!1$ supergravity models with radiative breaking of the electroweak gauge symmetry. The limit is relaxed by at least an order of magnitude when models with nonuniversal scalar masses are considered. A subset of the authors in ARNOWITT 02 updated the limit to $<9\times10^{-8}$ (tan β <55).
- 43 BERNABEI 00 search for annual modulation of the WIMP signal. The data favor the hypothesis of annual modulation at 4σ and are consistent, for a particular model framework quoted there, with $m_{\chi^0}{=}44^{+12}_{-9}$ GeV and a spin-independent χ^0 -proton cross section of (5.4 \pm 1.0) \times 10 $^{-6}$ pb. See also BERNABEI 01 and BERNABEI 00c.
- ⁴⁴ FENG 00 calculate the χ -p elastic scattering cross section in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry with a particular emphasis on focus point models. At $\tan\beta$ =50, the range is 8×10^{-8} - 4×10^{-7} .

Other bounds on $\tilde{\chi}_1^0$ from astrophysics and cosmology

Most of these papers generally exclude regions in the M_2 – μ parameter plane by requiring that the $\widetilde{\chi}^0_1$ contribution to the overall cosmological density is less than some maximal value to avoid overclosure of the Universe. Those not based on the cosmological density are indicated. Many of these papers also include LEP and/or other bounds.

VALUE		DOCUMENT ID		TECN	COMMENT
>46 GeV	1	ELLIS	00	RVUE	
• • • We do not use the			/erage	es, fits, li	imits, etc. • • •
	2	BUCHMUEL	14	COSM	
		BUCHMUEL			
		ROSZKOWSKI		COSM	
		CABRERA	13	COSM	
		ELLIS	13B		
		STREGE	13	COSM	
		AKULA	12	COSM	
		ARBEY	12A	COSM	
		BAER	12	COSM	
		BALAZS	12	COSM	
	8	BECHTLE	12	COSM	
		BESKIDT	12	COSM	
> 18 GeV	10	BOTTINO	12	COSM	
	2	BUCHMUEL	12	COSM	
	2	CAO	12A	COSM	
		ELLIS	12 B	COSM	
	11	FENG	12 B	COSM	
	2	KADASTIK	12	COSM	
		STREGE	12	COSM	
	12	BUCHMUEL	11	COSM	
		ROSZKOWSKI	11	COSM	
		ELLIS	10	COSM	
		BUCHMUEL	09	COSM	
	16	DREINER	09	THEO	

	¹⁷ BUCHMUEL	08	COSM	
	¹³ ELLIS	08	COSM	
	¹⁸ CALIBBI	07	COSM	
	¹⁹ ELLIS	07	COSM	
	²⁰ ALLANACH	06	COSM	
	²¹ DE-AUSTRI	06	COSM	
	¹³ BAER	05	COSM	
	²² BALTZ	04	COSM	
> 6 GeV	^{10,23} BELANGER	04	THEO	
	²⁴ ELLIS	04 B	COSM	
	²⁵ PIERCE	04A	COSM	
	²⁶ BAER	03	COSM	
> 6 GeV	¹⁰ BOTTINO	03	COSM	
, , , , , , , , , , , , , , , , , , , ,	²⁶ CHATTOPAD.		COSM	
	²⁷ ELLIS	03	COSM	
	¹³ ELLIS	03 B	COSM	
	²⁶ ELLIS	03C	COSM	
	²⁶ LAHANAS	03	COSM	
	²⁸ LAHANAS	02	COSM	
	²⁹ BARGER	01C	COSM	
	30 ELLIS	01B	COSM	
	²⁷ BOEHM	00B	COSM	
	31 FENG	00	COSM	
< 600 GeV	32 ELLIS	98B	COSM	
< 000 GCV	33 EDSJO	97		Co-annihilation
	³⁴ BAER	96	COSM	Co-ammination
	¹³ BEREZINSKY		COSM	
	35 FALK	95		CP-violating phases
	36 DREES	93		Minimal supergravity
	37 FALK	93		Sfermion mixing
	36 KELLEY	93		Minimal supergravity
	38 MIZUTA	93		Co-annihilation
	39 LOPEZ	92		Minimal supergravity,
		92	COSIVI	$m_0 = A = 0$
	⁴⁰ MCDONALD	92	COSM	
	41 GRIEST	91	COSM	
	⁴² NOJIRI	91	COSM	Minimal supergravity
	⁴³ OLIVE	91	COSM	
	⁴⁴ ROSZKOWSK	91	COSM	
	⁴⁵ GRIEST	90	COSM	
	⁴³ OLIVE	89	COSM	
none 100 eV – 15 GeV	SREDNICKI	88		$\widetilde{\gamma}$; $m_{\widetilde{f}}{=}100~{ m GeV}$
none 100 eV-5 GeV	ELLIS	84		$\widetilde{\gamma}$; for $m_{\widetilde{f}} = 100 \text{ GeV}$
	GOLDBERG	83	COSM	$\widetilde{\gamma}$
	⁴⁶ KRAUSS	83	COSM	
	VYSOTSKII	83	COSM	$\widetilde{\gamma}$

¹ ELLIS 00 updates ELLIS 98. Uses LEP e^+e^- data at \sqrt{s} =202 and 204 GeV to improve bound on neutralino mass to 51 GeV when scalar mass universality is assumed and 46 GeV when Higgs mass universality is relaxed. Limits on tanβ improve to > 2.7 (μ > 0), > 2.2 (μ < 0) when scalar mass universality is assumed and > 1.9 (both signs of μ) when Higgs mass universality is relaxed.

- 2 Implications of the LHC result on the Higgs mass and on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 3 BUCHMUELLER 14A places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches using the 20 fb $^{-1}$ 8 TeV and the 5 fb $^{-1}$ 7 TeV LHC and the LUX data.
- 4 ROSZKOWSKI 14 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using Bayesian statistics and indirect experimental searches using the 20 fb $^{-1}$ LHC and the LUX data.
- ⁵ CABRERA 13 and STREGE 13 place constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry with and without non-universal Higgs masses using the 5.8 fb⁻¹, $\sqrt{s}=7$ TeV ATLAS supersymmetry searches and XENON100 results.
- 6 ELLIS 13B place constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry with and without Higgs mass universality. Models with universality below the GUT scale are also considered.
- ⁷ BALAZS 12 and STREGE 12 place constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using the 1 fb⁻¹ LHC supersymmetry searches, the 5 fb⁻¹ Higgs mass constraints, both with $\sqrt{s}=7$ TeV, and XENON100 results.
- ⁸ BECHTLE 12 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches, using the 5 fb⁻¹ LHC and XENON100 data.
- ⁹ BESKIDT 12 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches, the 5 fb⁻¹ LHC and the XENON100 data.
- ¹⁰ BELANGER 04 and BOTTINO 12 (see also BOTTINO 03, BOTTINO 03A and BOTTINO 04) do not assume gaugino or scalar mass unification.
- 11 FENG 12B places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry and large sfermion masses using the 1 fb $^{-1}$ LHC supersymmetry searches, the 5 fb $^{-1}$ LHC Higgs mass constraints both with $\sqrt{s}=7$ TeV, and XENON100 results.
- 12 BUCHMUELLER 11 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches and including supersymmetry breaking relations between A and B parameters.
- 13 Places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry but non-Universal Higgs masses.
- 14 ELLIS 10 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry with universality above the GUT scale.
- 15 BUCHMUELLER 09 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches.
- 16 DREINER 09 show that in the general MSSM with non-universal gaugino masses there exists no model-independent laboratory bound on the mass of the lightest neutralino. An essentially massless χ_1^0 is allowed by the experimental and observational data, imposing some constraints on other MSSM parameters, including $\mathit{M}_2,~\mu$ and the slepton and squark masses.
- 17 BUCHMUELLER 08 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry using indirect experimental searches.

- 18 CALIBBI 07 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry with universality above the GUT scale including the effects of right-handed neutrinos.
- 19 ELLIS 07 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry with universality below the GUT scale.
- 20 ALLANACH 06 places constraints on the SUSY parameter space in the framework of N = 1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 21 DE-AUSTRI 06 places constraints on the SUSY parameter space in the framework of N = 1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 22 BALTZ 04 places constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- ²³ Limit assumes a pseudo scalar mass < 200 GeV. For larger pseudo scalar masses, $m_{\chi} > 18(29)$ GeV for $\tan\beta = 50(10)$. Bounds from WMAP, $(g-2)_{\mu}$, $b \rightarrow s\gamma$, LEP.
- ²⁴ ELLIS 04B places constraints on the SUSY parameter space in the framework of *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry including supersymmetry breaking relations between A and B parameters. See also ELLIS 03D.
- 25 PIERCE 04A places constraints on the SUSY parameter space in the framework of models with very heavy scalar masses.
- 26 BAER 03, CHATTOPADHYAY 03, ELLIS 03C and LAHANAS 03 place constraints on the SUSY parameter space in the framework of N=1 supergravity models with radiative breaking of the electroweak gauge symmetry based on WMAP results for the cold dark matter density.
- ²⁷BOEHM 00B and ELLIS 03 place constraints on the SUSY parameter space in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. Includes the effect of χ - \tilde{t} co-annihilations.
- 28 LAHANAS 02 places constraints on the SUSY parameter space in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. Focuses on the role of pseudo-scalar Higgs exchange.
- ²⁹ BARGER 01C use the cosmic relic density inferred from recent CMB measurements to constrain the parameter space in the framework of minimal *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 30 ELLIS 01B places constraints on the SUSY parameter space in the framework of minimal N=1 supergravity models with radiative breaking of the electroweak gauge symmetry. Focuses on models with large $\tan \beta$.
- ³¹ FENG 00 explores cosmologically allowed regions of MSSM parameter space with multi-TeV masses.
- 32 ELLIS 98B assumes a universal scalar mass and radiative supersymmetry breaking with universal gaugino masses. The upper limit to the LSP mass is increased due to the inclusion of $\chi-\widetilde{\tau}_R$ coannihilations.
- 33 EDSJO 97 included all coannihilation processes between neutralinos and charginos for any neutralino mass and composition.
- 34 Notes the location of the neutralino Z resonance and h resonance annihilation corridors in minimal supergravity models with radiative electroweak breaking.
- 35 Mass of the bino (=LSP) is limited to $m_{\widetilde{B}} \lesssim 350$ GeV for $m_t = 174$ GeV.
- ³⁶ DREES 93, KELLEY 93 compute the cosmic relic density of the LSP in the framework of minimal *N*=1 supergravity models with radiative breaking of the electroweak gauge symmetry.
- 37 FALK 93 relax the upper limit to the LSP mass by considering sfermion mixing in the MSSM
- 38 MIZUTA 93 include coannihilations to compute the relic density of Higgsino dark matter.
- $^{
 m 39}\,{\rm LOPEZ}$ 92 calculate the relic LSP density in a minimal SUSY GUT model.
- ⁴⁰ MCDONALD 92 calculate the relic LSP density in the MSSM including exact tree-level annihilation cross sections for all two-body final states.

- Unstable $\widetilde{\chi}_1^0$ (Lightest Neutralino) MASS LIMIT :

Unless otherwise stated, results in this section assume spectra and production rates as evaluated in the MSSM. Unless otherwise stated, the goldstino or gravitino mass $m_{\widetilde{G}}$ is assumed to be negligible relative to all other masses. In the following, \widetilde{G} is assumed to be undetected and to give rise to a missing energy $(\not\!E)$ signature.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>380	95	$^{ m 1}$ KHACHATRY.	14L CMS	$\widetilde{\chi}_1^0 ightarrow ~Z\widetilde{G}$ simplified models,GMSB
• • • We do n	ot use	the following data	for averages,	, fits, limits, etc. • • •
		² KHACHATRY.	16 BX	$\geq 3\ell^{\pm}$, RPV, λ or λ' couplings, wino- or higgsino-like neutralinos
		³ AAD	14BH ATLS	$2\gamma + E_T$, GMSB, SPS8
		⁴ AAD	13AP ATLS	$2\gamma+\cancel{\cancel{E}_T}$, GMSB, SPS8
none 220-380	95	⁵ AAD	13Q ATLS	$\gamma + b + ot\!$
		⁶ AAD	13R ATLS	•
		⁷ AALTONEN	13ı CDF	$\widetilde{\chi}_{1}^{ar{0}} ightarrow \ \gamma \widetilde{G}$, $ ot\!\!\!/ _{T}$, GMSB
>220	95	⁸ CHATRCHYAN	N 13AH CMS	$\widetilde{\chi}_{f 1}^{f 0} ightarrow \ \gamma \widetilde{\it G}$, GMSB, SPS8, $c au$ $<$
		⁹ AAD	12CP ATLS	$500~ ext{mm}$ $2\gamma + \cancel{\cancel{E}}_T$, GMSB
		¹⁰ AAD	12CT ATLS	\geq 4 ℓ^{\pm} , R
		¹¹ AAD	12R ATLS	$\widetilde{\chi}_1^0 ightarrow \ \mu jj, \not\!\! R, \lambda_{211}' \neq 0$
		¹² ABAZOV	12AD D 0	$\widetilde{\chi}_1^0 \widetilde{\chi}_1^0 \rightarrow \gamma Z \widetilde{G} \widetilde{G}$, GMSB
		¹³ CHATRCHYAN		$2\gamma+ ot\!\!\!E_T$, GMSB
		¹⁴ CHATRCHYAN	N 11B CMS	$W^0 \rightarrow \gamma G, W^{\pm} \rightarrow \ell^{\pm} G, \text{GMSB}$
>149	95	¹⁵ AALTONEN	10 CDF	$p\overline{p} \rightarrow \widetilde{\chi}\widetilde{\chi}, \widetilde{\chi} = \widetilde{\chi}_2^0, \widetilde{\chi}_1^{\pm}, \widetilde{\chi}_1^0 \rightarrow$
>175	95	¹⁶ ABAZOV	10P D0	$\gamma \widetilde{G}$, GMSB $\widetilde{\chi}_1^0 o \gamma \widetilde{G}$, GMSB

⁴¹ GRIEST 91 improve relic density calculations to account for coannihilations, pole effects, and threshold effects.

⁴² NOJIRI 91 uses minimal supergravity mass relations between squarks and sleptons to narrow cosmologically allowed parameter space.

⁴³ Mass of the bino (=LSP) is limited to $m_{\widetilde{B}} \lesssim 350$ GeV for $m_t \leq 200$ GeV. Mass of the higgsino (=LSP) is limited to $m_{\widetilde{H}} \lesssim 1$ TeV for $m_t \leq 200$ GeV.

 $^{^{44}}$ ROSZKOWSKI 91 calculates LSP relic density in mixed gaugino/higgsino region.

 $^{^{45}\,\}rm Mass$ of the bino (=LSP) is limited to $m_{\widetilde{B}}\lesssim 550$ GeV. Mass of the higgsino (=LSP) is limited to $m_{\widetilde{H}}\lesssim 3.2$ TeV.

 $^{^{46}}$ KRAUSS 83 finds $m_{\widetilde{\gamma}}$ not 30 eV to 2.5 GeV. KRAUSS 83 takes into account the gravitino decay. Find that limits depend strongly on reheated temperature. For example a new allowed region $m_{\widetilde{\gamma}}=$ 4–20 MeV exists if $m_{\rm gravitino}$ <40 TeV. See figure 2.

>125 95 17 ABAZOV 08F D0
$$p\overline{p} \rightarrow \widetilde{\chi}\widetilde{\chi}, \ \widetilde{\chi} = \widetilde{\chi}_{2}^{0}, \ \widetilde{\chi}_{1}^{\pm}, \ \widetilde{\chi}_{1}^{0} \rightarrow \widetilde{G}, \ GMSB$$
18 ABULENCIA 07H CDF R , $LL\overline{E}$
> 96.8 95 19 ABBIENDI 06B OPAL $e^{+}e^{-} \rightarrow \widetilde{B}\widetilde{B}, \ (\widetilde{B} \rightarrow \widetilde{G}\gamma)$
20 ABDALLAH 05B DLPH $e^{+}e^{-} \rightarrow \widetilde{G}\widetilde{\chi}_{1}^{0}, \ (\widetilde{\chi}_{1}^{0} \rightarrow \widetilde{G}\gamma)$
> 96 95 21 ABDALLAH 05B DLPH $e^{+}e^{-} \rightarrow \widetilde{B}\widetilde{B}, \ (\widetilde{B} \rightarrow \widetilde{G}\gamma)$

- 1 KHACHATRYAN 14 L searched in $^{19.5}$ fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for evidence of direct pair production of neutralinos with Higgs or Z-bosons in the decay chain, leading to $HH,\ HZ$ and ZZ final states with missing transverse energy. The decays of $^{16-20}$. a Higgs boson to a b-quark pair, to a photon pair, and to final states with leptons are considered in conjunction with hadronic and leptonic decay modes of the Z and W bosons. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of GMSB simplified models where the decays $\widetilde{\chi}_1^0 \to H\widetilde{G}$ or $\widetilde{\chi}_1^0 \to Z\widetilde{G}$ take place either 100% or 50% of the time, see Figs. 16–20.
- 2 KHACHATRYAN 16BX searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing 3 or more leptons coming from the electroweak production of wino- or higgsino-like neutralinos, assuming non-zero R-parity-violating leptonic couplings λ_{122} , λ_{123} , and λ_{233} or semileptonic couplings λ'_{131} , λ'_{233} , λ'_{331} , and λ'_{333} . No excess over the expected background is observed and limits are derived on the neutralino mass, see Figs. 24 and 25.
- ³ AAD 14BH searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing non-pointing photons in a diphoton plus missing transverse energy final state. No excess is observed above the background expected from Standard Model processes. The results are used to set 95% C.L. exclusion limits in the contact of gauge-mediated supersymmetric breaking models, with the lightest neutralino being the next-to-lightest supersymmetric particle and decaying with a lifetime in the range from 0.25 ns to about 100 ns into a photon and a gravitino. For limits on the NLSP lifetime versus Λ plane, for the SPS8 model, see their Fig. 7.
- ⁴ AAD 13AP searched in 4.8 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing non-pointing photons in a diphoton plus missing transverse energy final state. No excess is observed above the background expected from Standard Model processes. The results are used to set 95% C.L. exclusion limits in the context of gauge-mediated supersymmetric breaking models, with the lightest neutralino being the next-to-lightest supersymmetric particle and decaying with a lifetime in excess of 0.25 ns into a photon and a gravitino. For limits in the NLSP lifetime versus Λ plane, for the SPS8 model, see their Fig. 8.
- ⁵ AAD 13Q searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing a high- p_T isolated photon, at least one jet identified as originating from a bottom quark, and high missing transverse momentum. Such signatures may originate from supersymmetric models with gauge-mediated supersymmetry breaking in events in which one of a pair of higgsino-like neutralinos decays into a photon and a gravitino while the other decays into a Higgs boson and a gravitino. No significant excess above the expected background was found and limits were set on the neutralino mass in a generalized GMSB model (GGM) with a higgsino-like neutralino NLSP, see their Fig. 4. Intermediate neutralino masses between 220 and 380 GeV are excluded at 95% C.L, regardless of the squark and gluino masses, purely on the basis of the expected weak production.
- ⁶ AAD 13R looked in 4.4 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing new, heavy particles that decay at a significant distance from their production point into a final state containing a high-momentum muon and charged hadrons. No excess over the expected background is observed and limits are placed on the production cross-section of neutralinos via squarks for various $m_{\widetilde{q}}$, $m_{\widetilde{\chi}_1^0}$ in an R-parity violating scenario with
- $\lambda'_{211} \neq$ 0, as a function of the neutralino lifetime, see their Fig. 6.
- ⁷ AALTONEN 131 searched in 6.3 fb⁻¹ of $p\bar{p}$ collisions at $\sqrt{s}=1.96$ TeV for events containing E_T and a delayed photon that arrives late in the detector relative to the

time expected from prompt production. No evidence of delayed photon production is observed.

- 8 CHATRCHYAN 13AH searched in 4.9 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events containing E_T and a delayed photon that arrives late in the detector relative to the time expected from prompt production. No significant excess above the expected background was found and limits were set on the pair production of $\widetilde{\chi}_1^0$ depending on the neutralino proper decay length, see Fig. 8. Supersedes CHATRCHYAN 12BK.
- 9 AAD 12CP searched in 4.8 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with two photons and large E_T due to $\widetilde{\chi}^0_1\to\gamma\,\widetilde{G}$ decays in a GMSB framework. No significant excess above the expected background was found and limits were set on the neutralino mass in a generalized GMSB model (GGM) with a bino-like neutralino NLSP, see Figs. 6 and 7. The other sparticle masses were decoupled, $\tan\beta=2$ and $c\tau_{NLSP}<0.1$ mm. Also, in the framework of the SPS8 model, limits are presented in Fig. 8.
- 10 AAD 12CT searched in 4.7 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events containing four or more leptons (electrons or muons) and either moderate values of missing transverse momentum or large effective mass. No significant excess is found in the data. Limits are presented in a simplified model of R-parity violating supersymmetry in which charginos are pair-produced and then decay into a W-boson and a $\tilde{\chi}_1^0$, which in turn decays through an RPV coupling into two charged leptons ($e^\pm\,e^\mp$ or $\mu^\pm\,\mu^\mp$) and a neutrino. In this model, limits are set on the neutralino mass as a function of the chargino mass, see Fig. 3a. Limits are also set in an R-parity violating mSUGRA model, see Fig. 3b.
- 11 AAD 12 R looked in 33 pb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events containing new, heavy particles that decay at a significant distance from their production point into a final state containing a high-momentum muon and charged hadrons. No excess over the expected background is observed and limits are placed on the production cross-section of neutralinos via squarks for various $(m_{\widetilde{q}},\ m_{\widetilde{\chi}^0_1})$ in an R-parity violating scenario with

 $\lambda_{211}^{\prime} \neq 0$, as a function of the neutralino lifetime, see their Fig. 8. Superseded by AAD 13R.

- ABAZOV 12AD looked in 6.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=1.96$ TeV for events with a photon, a Z-boson, and large $\not E_T$ in the final state. This topology corresponds to a GMSB model where pairs of neutralino NLSPs are either pair produced promptly or from decays of other supersymmetric particles and then decay to either $Z \ G$ or $\gamma \ G$. No significant excess over the SM expectation is observed and a limit at 95% C.L. on the cross section is derived as a function of the effective SUSY breaking scale Λ , see Fig. 3. Assuming $N_{mes}=2$, $M_{mes}=3$ Λ , $\tan\beta=3$, $\mu=0.75$ M_1 , and $C_{grav}=1$, the model is excluded at 95% C.L. for values of $\Lambda<87$ TeV.
- model is excluded at 95% C.L. for values of $\Lambda <$ 87 TeV. 13 CHATRCHYAN 12BK searched in 2.23 fb $^{-1}$ of pp collisions at $\sqrt{s}=$ 7 TeV for events with two photons and large $\not\!\!E_T$ due to $\widetilde{\chi}_1^0 \to \gamma \, \widetilde{\mathcal{G}}$ decays in a GMSB framework. No significant excess above the expected background was found and limits were set on the pair production of $\widetilde{\chi}_1^0$ depending on the neutralino lifetime, see Fig. 6.
- 14 CHATRCHYAN 11 B looked in 35 pb $^{-1}$ of pp collisions at $\sqrt{s}{=}7$ TeV for events with an isolated lepton (e or μ), a photon and E_T which may arise in a generalized gauge mediated model from the decay of Wino-like NLSPs. No evidence for an excess over the expected background is observed. Limits are derived in the plane of squark/gluino mass versus Wino mass (see Fig. 4). Mass degeneracy of the produced squarks and gluinos is assumed.
- 15 AALTONEN 10 searched in 2.6 fb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for diphoton events with large $\not\!\!E_T$. They may originate from the production of $\widetilde{\chi}^\pm$ in pairs or associated to a $\widetilde{\chi}^0_2$, decaying into $\widetilde{\chi}^0_1$ which itself decays in GMSB to $\gamma\widetilde{G}$. There is no excess of events beyond expectation. An upper limit on the cross section is calculated in the GMSB model as a function of the $\widetilde{\chi}^0_1$ mass and lifetime, see their Fig. 2. A limit is derived on the $\widetilde{\chi}^0_1$ mass of 149 GeV for $\tau_{\widetilde{\chi}^0_1} \ll 1$ ns, which improves the results of previous searches.

- 16 ABAZOV 10P looked in 6.3 fb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with at least two isolated γs and large E_T . These could be the signature of $\widetilde{\chi}_2^0$ and $\widetilde{\chi}_1^\pm$ production, decaying to $\widetilde{\chi}_1^0$ and finally $\widetilde{\chi}_1^0 \to \gamma \, \widetilde{G}$ in a GMSB framework. No significant excess over the SM expectation is observed, and a limit at 95% C.L. on the cross section is derived for $N_{mes}=1$, $\tan\beta=15$ and $\mu>0$, see their Fig. 2. This allows them to set a limit on the effective SUSY breaking scale $\Lambda>124$ TeV, from which the excluded $\widetilde{\chi}_1^0$ mass range is obtained.
- ¹⁷ ABAZOV 08F looked in 1.1 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for diphoton events with large $\not\!\!E_T$. They may originate from the production of $\widetilde{\chi}^\pm$ in pairs or associated to a $\widetilde{\chi}^0_2$, decaying to a $\widetilde{\chi}^0_1$ which itself decays promptly in GMSB to $\widetilde{\chi}^0_1 \to \gamma \, \widetilde{G}$. No significant excess was found compared to the background expectation. A limit is derived on the masses of SUSY particles in the GMSB framework for $M=2\Lambda$, N=1, $\tan\beta=15$ and $\mu>0$, see Figure 2. It also excludes $\Lambda<91.5$ TeV. Supersedes the results of ABAZOV 05A. Superseded by ABAZOV 10P.
- 18 ABULENCIA 07H searched in 346 pb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with at least three leptons (e or μ) from the decay of $\widetilde{\chi}_1^0$ via $LL\overline{E}$ couplings. The results are consistent with the hypothesis of no signal. Upper limits on the cross-section are extracted and a limit is derived in the framework of mSUGRA on the masses of $\widetilde{\chi}_1^0$ and $\widetilde{\chi}_1^{\pm}$, see e.g. their Fig. 3 and Tab. II.
- 19 ABBIENDI 06B use 600 pb $^{-1}$ of data from $\sqrt{s}=189$ –209 GeV. They look for events with diphotons + $\cancel{\mathbb{Z}}$ final states originating from prompt decays of pair-produced neutralinos in a GMSB scenario with $\widetilde{\chi}_1^0$ NLSP. Limits on the cross-section are computed as a function of m($\widetilde{\chi}_1^0$), see their Fig. 14. The limit on the $\widetilde{\chi}_1^0$ mass is for a pure Bino state assuming a prompt decay, with lifetimes up to 10^{-9} s. Supersedes the results of ABBIENDI 04N.
- ²⁰ ABDALLAH 05B use data from $\sqrt{s}=180$ –209 GeV. They look for events with single photons + \cancel{E} final states. Limits are computed in the plane $(\mathsf{m}(\widetilde{G})$, $\mathsf{m}(\widetilde{\chi}_1^0))$, shown in their Fig. 9b for a pure Bino state in the GMSB framework and in Fig. 9c for a no-scale supergravity model. Supersedes the results of ABREU 00Z.
- ²¹ ABDALLAH 05B use data from $\sqrt{s}=130$ –209 GeV. They look for events with diphotons + $\not\!\!E$ final states and single photons not pointing to the vertex, expected in GMSB when the $\widetilde{\chi}_1^0$ is the NLSP. Limits are computed in the plane $(\mathsf{m}(\widetilde{G}),\,\mathsf{m}(\widetilde{\chi}_1^0))$, see their Fig. 10. The lower limit is derived on the $\widetilde{\chi}_1^0$ mass for a pure Bino state assuming a prompt decay and $m_{\widetilde{e}_R}=m_{\widetilde{e}_L}=2$ $m_{\widetilde{\chi}_1^0}$. It improves to 100 GeV for $m_{\widetilde{e}_R}=m_{\widetilde{e}_L}=1.1$ $m_{\widetilde{\chi}_1^0}$. and the limit in the plane $(\mathsf{m}(\widetilde{\chi}_1^0),\,\mathsf{m}(\widetilde{e}_R))$ is shown in Fig. 10b. For long-lived neutralinos, cross-section limits are displayed in their Fig 11. Supersedes the results of ABREU 00Z.

$\widetilde{\chi}_2^0$, $\widetilde{\chi}_3^0$, $\widetilde{\chi}_4^0$ (Neutralinos) MASS LIMITS

Neutralinos are unknown mixtures of photinos, z-inos, and neutral higgsinos (the supersymmetric partners of photons and of Z and Higgs bosons). The limits here apply only to $\widetilde{\chi}_2^0$, $\widetilde{\chi}_3^0$, and $\widetilde{\chi}_4^0$. $\widetilde{\chi}_1^0$ is the lightest supersymmetric particle (LSP); see $\widetilde{\chi}_1^0$ Mass Limits. It is not possible to quote rigorous mass limits because they are extremely model dependent; i.e. they depend on branching ratios of various $\widetilde{\chi}^0$ decay modes, on the masses of decay products $(\widetilde{e},\ \widetilde{\gamma},\ \widetilde{q},\ \widetilde{g})$, and on the \widetilde{e} mass exchanged in $e^+e^-\to\widetilde{\chi}_i^0\widetilde{\chi}_j^0$. Limits arise either from direct searches, or from the MSSM constraints set on the gaugino and higgsino mass parameters M_2 and μ through searches for lighter charginos and neutralinos. Often limits are given as contour plots in the $m_{\widetilde{\chi}^0}-m_{\widetilde{e}}$ plane vs other parameters. When specific assumptions are made, e.g, the

neutralino is a pure photino $(\tilde{\gamma})$, pure z-ino (\tilde{Z}) , or pure neutral higgsino (\tilde{H}^0) , the neutralinos will be labelled as such.

Limits obtained from e^+e^- collisions at energies up to 136 GeV, as well as other limits from different techniques, are now superseded and have not been included in this compilation. They can be found in the 1998 Edition (The European Physical Journal C3 1 (1998)) of this Review. Some later papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C38 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>600	95	¹ AAD	16AA ATLS	$3/4\ell+ ot\!\!\!E_T$, Tn2n3A, $m_{\widetilde{\chi}_1^0}=$ 0GeV
>670	95	¹ AAD		$3/4\ell + E_T$, Tn2n3B, $m_{\widetilde{\chi}_1^0} < 200$ GeV
>250	95	² AAD	15BA ATLS	$m_{\widetilde{\chi}_1^{\pm}}=m_{\widetilde{\chi}_2^0},m_{\widetilde{\chi}_1^0}=0\mathrm{GeV}$
>380	95	³ AAD	14H ATLS	$\widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0 \rightarrow \tau^{\pm} \nu \widetilde{\chi}_1^0 \tau^{\pm} \tau^{\mp} \widetilde{\chi}_1^0$, simplified model, $m_{\widetilde{\chi}_1^{\pm}} = m_{\widetilde{\chi}_2^0}$,
>700	95	³ AAD	14H ATLS	$egin{aligned} m_{\widetilde{\chi}_1^0} &= 0 \; \mathrm{GeV} \ \widetilde{\chi}_1^\pm \widetilde{\chi}_2^0 & ightarrow \; \ell^\pm u \widetilde{\chi}_1^0 \ell^\pm \ell^\mp \widetilde{\chi}_1^0 , \mathrm{simplified model}, m_{\widetilde{\chi}_1^\pm} &= m_{\widetilde{\chi}_2^0}, \end{aligned}$
>345	95	³ AAD	14н ATLS	$egin{aligned} m_{\widetilde{\chi}_1^0} &= 0 \; \mathrm{GeV} \ \widetilde{\chi}_1^\pm \widetilde{\chi}_2^0 & ightarrow \; W \widetilde{\chi}_1^0 Z \widetilde{\chi}_1^0 , \; \mathrm{simplified} \ \mathrm{model}, \; m_{\widetilde{\chi}_1^\pm} &= m_{\widetilde{\chi}_2^0}, \; m_{\widetilde{\chi}_1^0} &= 0 \end{aligned}$
>148	95	³ AAD	14H ATLS	$\begin{array}{c} \operatorname{GeV} \\ \widetilde{\chi}_{1}^{\pm} \widetilde{\chi}_{2}^{0} \rightarrow W \widetilde{\chi}_{1}^{0} H \widetilde{\chi}_{1}^{0}, \operatorname{simplified} \\ \operatorname{model}, m_{\widetilde{\chi}_{1}^{\pm}} = m_{\widetilde{\chi}_{2}^{0}}, m_{\widetilde{\chi}_{1}^{0}} = 0 \end{array}$
>620	95	⁴ AAD		$\overset{GeV}{\geq} \overset{1}{4\ell^{\pm}}, \widetilde{\chi}^0_{2,3} \rightarrow \ \ell^{\pm}\ell^{\mp}\widetilde{\chi}^0_1, m_{\widetilde{\chi}^0_1}$
		5 AAD		$=$ 0 GeV $3\ell^{\pm}+\cancel{E}_{T}$, pMSSM, SMS
>116.0	95	⁷ ABREU		\geq $2~\ell$, jets $+ ot\!\!\!E_T$, pp $ ightarrow \ \widetilde{\chi}_1^\pm \widetilde{\chi}_2^0$ $ \widetilde{\chi}_4^0$, $1~\leq~$ tan $eta~\leq~$ 40, all Δm ,
>110.0	90	ADREU	UUW DLPH	χ_{4}^{\prime} , $1 \leq \tan \beta \leq 40$, all Δm , all m_0
• • • We do	not use t	the following data	for averages,	fits, limits, etc. • • •

not use the following data for averages, fits, limits, etc. • •

none 180–355 8 AAD 14G ATLS
$$\widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0 \rightarrow W \widetilde{\chi}_1^0 Z \widetilde{\chi}_1^0$$
, simplified model, $m_{\widetilde{\chi}_1^{\pm}} = m_{\widetilde{\chi}_2^0}$, $m_{\widetilde{\chi}_1^0} = 0$ 9 KHACHATRY...14I CMS $\widetilde{\chi}_2^0 \rightarrow (Z, H) \widetilde{\chi}_1^0 \ \widetilde{\ell} \ell$, simplified model 10 AAD 12AS ATLS $3\ell^{\pm} + E_T$, pMSSM 11 AAD 12T ATLS $\ell^{\pm}\ell^{\pm} + E_T$, $pp \rightarrow \widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0$

 $^{^1}$ AAD 16AA summarized and extended ATLAS searches for electroweak supersymmetry in final states containing several charged leptons, E_T , with or without hadronic jets, in 20 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on mass-degenerate $\widetilde{\chi}^0_2$ and $\widetilde{\chi}^0_3$ masses in the Tn2n3A and Tn2n3B simplified models. See their Fig. 15.

- ² AAD 15BA searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of charginos and neutralinos decaying to a final state containing a W boson and a 125 GeV Higgs boson, plus missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with the decays $\widetilde{\chi}_1^{\pm} \to W^{\pm}\widetilde{\chi}_1^0$ and $\widetilde{\chi}_2^0 \to H\widetilde{\chi}_1^0$ having 100% branching fraction, see Fig. 8. A combination of the multiple final states for the Higgs decay yields the best limits (Fig. 8d).
- ³ AAD 14H searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of charginos and neutralinos decaying to a final sate with three leptons and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via either all three generations of leptons, staus only, gauge bosons, or Higgs bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 8.
- ⁴ AAD 14X searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the neutralino mass in an R-parity conserving simplified model where the decay $\widetilde{\chi}_{2,3}^0 \to \ell^\pm \ell^\mp \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 10.
- ⁵ AAD 13 searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for charginos and neutralinos decaying to a final state with three leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 and 3, and in simplified models, see Fig.
- 4. For the simplified models with intermediate slepton decays, degenerate $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_2^0$ masses up to 500 GeV are excluded at 95% C.L. for very large mass differences with the $\widetilde{\chi}_1^0$. Supersedes AAD 12AS.
- 6 CHATRCHYAN 12BJ searched in 4.98 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for direct electroweak production of charginos and neutralinos in events with at least two leptons, jets and missing transverse momentum. No significant excesses over the expected SM backgrounds are observed and 95% C.L. limits on the production cross section of $\widetilde{\chi}_1^{\pm}\,\widetilde{\chi}_2^0$ pair production were set in a number of simplified models, see Figs. 7 to 12. Most limits are for exactly 3 jets.
- ABREU 00W combines data collected at $\sqrt{s}{=}189$ GeV with results from lower energies. The mass limit is obtained by constraining the MSSM parameter space with gaugino and sfermion mass universality at the GUT scale, using the results of negative direct searches for neutralinos (including cascade decays and $\tilde{\tau}\tau$ final states) from ABREU 01, for charginos from ABREU 00J and ABREU 00T (for all Δm_{+}), and for charged sleptons from ABREU 01B. The results hold for the full parameter space defined by all values of M_2 and $|\mu| \leq 2$ TeV with the $\tilde{\chi}_1^0$ as LSP.
- ⁸ AAD 14G searched in 20.3 fb⁻¹ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for electroweak production of chargino-neutralino pairs, decaying to a final sate with two leptons (e and μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via gauge bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 10.
- ⁹ KHACHATRYAN 14I searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of charginos and neutralinos decaying to a final state with three leptons (e or μ) and missing transverse momentum, or with a Z-boson, dijets and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Figs. 12–16.
- 10 AAD 12AS searched in 2.06 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for charginos and neutralinos decaying to a final state with three leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits

are derived in the phenomenological MSSM, see Fig. 2 (top), and in simplified models, see Fig. 2 (bottom).

 11 AAD 12 T looked in 1 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for the production of supersymmetric particles decaying into final states with missing transverse momentum and exactly two isolated leptons (e or μ). Same-sign dilepton events were separately studied. Additionally, in opposite-sign events, a search was made for an excess of sameflavor over different-flavor lepton pairs. No excess over the expected background is observed and limits are placed on the effective production cross section of opposite-sign dilepton events with $E_T >$ 250 GeV and on same-sign dilepton events with $E_T >$ 100 GeV. The latter limit is interpreted in a simplified electroweak gaugino production model.

 $\widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_2^{\pm}$ (Charginos) MASS LIMITS Charginos are unknown mixtures of w-inos and charged higgsinos (the supersymmetric partners of W and Higgs bosons). A lower mass limit for the lightest chargino $(\widetilde{\chi}_1^{\pm})$ of approximately 45 GeV, independent of the field composition and of the decay mode, has been obtained by the LEP experiments from the analysis of the Z width and decays. These results, as well as other now superseded limits from e^+e^- collisions at energies below 136 GeV, and from hadronic collisions, can be found in the 1998 Edition (The European Physical Journal **C3** 1 (1998)) of this Review.

Unless otherwise stated, results in this section assume spectra, production rates, decay modes and branching ratios as evaluated in the MSSM, with gaugino and sfermion mass unification at the GUT scale. These papers generally study production of $\tilde{\chi}_1^0 \tilde{\chi}_2^0$, $\widetilde{\chi}_1^+\widetilde{\chi}_1^-$ and (in the case of hadronic collisions) $\widetilde{\chi}_1^+\widetilde{\chi}_2^0$ pairs, including the effects of cascade decays. The mass limits on $\widetilde{\chi}_1^\pm$ are either direct, or follow indirectly from the constraints set by the non-observation of $\widetilde{\chi}_2^0$ states on the gaugino and higgsino MSSM parameters M_2 and μ . For generic values of the MSSM parameters, limits from high-energy e^+e^- collisions coincide with the highest value of the mass allowed by phase-space, namely $m_{\widetilde{\chi}_1^\pm}\lesssim \sqrt{s}/2$. The still unpublished combination of the results of the four LEP collaborations from the 2000 run of LEP2 at \sqrt{s} up to \simeq 209 GeV yields a lower mass limit of 103.5 GeV valid for general MSSM models. The limits become however weaker in certain regions of the MSSM parameter space where the detection efficiencies or production cross sections are suppressed. For example, this may happen when: (i) the mass differences $\Delta m_+=m_{\widetilde{\chi}_1^\pm}-m_{\widetilde{\chi}_1^0}$ or $\Delta m_{\nu}=m_{\widetilde{\chi}_1^\pm}-m_{\widetilde{\nu}}$ are very small, and the detection efficiency is reduced; (ii) the electron sneutrino mass is small, and the $\widetilde{\chi}_1^\pm$ production rate is suppressed due to a destructive interference between s and t channel exchange diagrams. The regions of MSSM parameter space where the

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics C38 070001 (2014) (http://pdg.lbl.gov).

following limits are valid are indicated in the comment lines or in the footnotes.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 500	95	¹ AAD	16AA ATLS	$2\ell^{\pm}+\cancel{E}_T$, Tchi1chi1B, $m_{\widetilde{\chi}_1^0}=0$ GeV
> 220	95	¹ AAD		$2\ell^{\pm}+\cancel{E}_T$, Tchi1chi1C, low Δ m for
> 700	95	² AAD	16AA ATLS	$\widetilde{\chi}_1^\pm$, $\widetilde{\chi}_1^0$ 3/4 ℓ + $ ot\!$

> 700	95	² AAD	16AA ATLS	3/4 ℓ + $ ot\!$
				0.5 (or 0.95) $(m_{\widetilde{\chi}_1^{\pm}} - m_{\widetilde{\chi}_1^0})^{-1}$
> 400	95	² AAD	16AA ATLS	2 hadronic $\tau + \cancel{E}_T$ & $3\ell + \cancel{E}_T$ combination, Tchi1n2D, $m_{\widetilde{\chi}_1^0} = 0$ GeV
> 540	95	³ KHACHATRY.	16R CMS	$\geq 1\gamma + 1$ e or $\mu + ot\!\!E_T$, <code>Tchi1n1A</code>
> 250	95	⁴ AAD	15BA ATLS	$m_{\widetilde{\chi}_1^{\pm}} = m_{\widetilde{\chi}_2^0}, m_{\widetilde{\chi}_1^0} = 0 \text{ GeV}$
> 590	95	⁵ AAD	15CA ATLS	$\geq 2 \ \gamma + E_T$, GGM, bino-like NLSP, any NLSP mass
none	95	⁵ AAD	15CA ATLS	$\geq 1 \; \gamma + e$, $\mu + ot \!$
124–361 > 700	95	⁶ AAD	14H ATLS	like NLSP $\widetilde{\chi}_1^{\pm}\widetilde{\chi}_2^0 \rightarrow \ell^{\pm}\nu\widetilde{\chi}_1^0\ell^{\pm}\ell^{\mp}\widetilde{\chi}_1^0, \text{ simplified model}, \ m_{\widetilde{\chi}_1^{\pm}} = m_{\widetilde{\chi}_2^0},$
				$m_{\widetilde{\chi}^0_1}=$ 0 GeV
> 345	95	⁶ AAD	14H ATLS	$\widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^{0} \rightarrow W \widetilde{\chi}_1^0 Z \widetilde{\chi}_1^0$, simplified model, $m_{\widetilde{\chi}_1^{\pm}} = m_{\widetilde{\chi}_2^0}$, $m_{\widetilde{\chi}_1^0} = 0$
> 148	95	⁶ AAD	14H ATLS	$\begin{array}{c} \operatorname{GeV} \\ \widetilde{\chi}_{1}^{\pm} \widetilde{\chi}_{2}^{0} \rightarrow & W \widetilde{\chi}_{1}^{0} H \widetilde{\chi}_{1}^{0}, \operatorname{simplified} \\ \operatorname{model}, m_{\widetilde{\chi}_{1}^{\pm}} = m_{\widetilde{\chi}_{2}^{0}}, m_{\widetilde{\chi}_{1}^{0}} = 0 \end{array}$
> 380	95	⁶ AAD	14H ATLS	
				$m_{\widetilde{\chi}^0_1}=0$ GeV
> 750	95	⁷ AAD	14X ATLS	$\geq 4\ell^{\pm}, \widetilde{\chi}_{1}^{\pm} \rightarrow W^{(*)\pm}\widetilde{\chi}_{1}^{0}, \widetilde{\chi}_{1}^{0} \rightarrow$
> 210	95	⁸ KHACHATRY	14L CMS	$\begin{array}{c} \ell^{\pm}\ell^{\mp}\nu, \text{ R-parity viol.} \\ \widetilde{\chi}_{2}^{0} \rightarrow H\widetilde{\chi}_{1}^{0} \text{ and } \widetilde{\chi}_{1}^{\pm} \rightarrow W^{\pm}\widetilde{\chi}_{1}^{0} \\ \text{simplified models, } m_{\widetilde{\chi}_{2}^{0}} = m_{\widetilde{\chi}_{1}^{\pm}}, \\ m_{\widetilde{\chi}_{1}^{0}} = 0 \text{ GeV} \end{array}$
		⁹ AAD	13 ATLS	
		10 AAD	13 ATLS	$3\ell^{\pm}+ ot\!\!\!E_T$, pMSSM, SMS $2\ell^{\pm}+ ot\!\!\!E_T$, pMSSM, SMS
> 540	95	11 AAD	12CT ATLS	$\geq 4\ell^{\pm}$, R , $m_{\widetilde{\chi}_1^0} > 300 \text{ GeV}$
		¹² CHATRCHYAI	N 12BJ CMS	\geq 2 ℓ , jets $+ ot\!$
> 94	95	¹³ ABDALLAH	03м DLPH	$\widetilde{\chi}_{1}^{\pm}$, $\tan\beta \leq 40$, $\Delta m_{+} > 3$ GeV, all
• • • We do	o not use	the following data	for averages, f	fits, limits, etc. • • •
> 570	95	14 KHACHATRY		$\geq 1\gamma + jets + ot \!\!\!\!E_T$, <code>Tchi1chi1A</code>
> 680	95	¹⁴ KHACHATRY.		$\geq 1\gamma + {\sf jets} + {\not \! E_T}$, Tchi1n1A
> 710	95	¹⁴ KHACHATRY.	16AA CMS	$\geq 1\gamma + jets + ot\!\!\!E_T$, GGM, $\widetilde{\chi}_2^0 \widetilde{\chi}_1^\pm$
>1000	95	¹⁵ KHACHATRY	16R CMS	pair production, wino-like NLSP $\geq 1\gamma+1$ e or $\mu+E_T$, Tglu1F, $m_{\widetilde{\chi}_1^\pm}=m_{\widetilde{\chi}_2^0}>200$ GeV
> 307	95	¹⁶ KHACHATRY	16Y CMS	1 or 2 soft ℓ^{\pm} +jets+ \cancel{E}_T , Tchi1n2A, $m_{\widetilde{\chi}_1^{\pm}}-m_{\widetilde{\chi}_1^0}=$ 20 GeV

> 410	95	¹⁷ AAD	14AV	ATLS	$\geq 2 au + ot \!$
					$\widetilde{\chi}_1^{\pm}\widetilde{\chi}_1^{\mp}$ production, $m_{\widetilde{\chi}_2^0}=$
					$m_{\widetilde{\chi}_1^{\pm}}, m_{\widetilde{\chi}_1^0} = 0 GeV$
> 345	95	¹⁸ AAD	14AV	ATLS	$\geq 2 \ au + E_T$, direct $\widetilde{\chi}_1^{\pm} \widetilde{\chi}_1^{\mp}$ production, $m_{\widetilde{\chi}_1^0} = 0 \ \text{GeV}$
none 100–105, 120–135,	95	¹⁹ AAD	14 G	ATLS	$\widetilde{\chi}_1^{\pm} \widetilde{\chi}_1^{\mp} \rightarrow W^+ \widetilde{\chi}_1^0 W^- \widetilde{\chi}_1^0, \text{ simplified model}, \ m_{\widetilde{\chi}_1^0} = 0 \text{ GeV}$
145–160 none 140–465	95	¹⁹ AAD	14 G	ATLS	$\begin{array}{ccc} \widetilde{\chi}_1^{\pm} \widetilde{\chi}_1^{\mp} \to \ell^{+} \nu \widetilde{\chi}_1^0 \ell^{-} \overline{\nu} \widetilde{\chi}_1^0 , \text{simplified model}, m_{\widetilde{\chi}_1^0} = 0 \text{GeV} \end{array}$
none 180–355	95	¹⁹ AAD	14 G		$\widetilde{\chi}_1^{\pm} \widetilde{\chi}_2^0 \rightarrow W \widetilde{\chi}_1^0 Z \widetilde{\chi}_1^0$, simplified model, $m_{\widetilde{\chi}_1^{\pm}} = m_{\widetilde{\chi}_2^0}$, $m_{\widetilde{\chi}_1^0} = 0$
> 168	95	²⁰ AALTONEN	14	CDF	GeV $3\ell^{\pm}+\cancel{E}_{T},\ \widetilde{\chi}_{1}^{\pm}\rightarrow\ \ell\nu\widetilde{\chi}_{1}^{0},$ mSUGRA with $m_{0}{=}60~{\rm GeV}$
		²¹ KHACHATRY.	141	CMS	$\widetilde{\chi}_1^{\pm} \to W \widetilde{\chi}_1^0$, $\ell \widetilde{\nu}$, $\widetilde{\ell} \nu$, simplified
		²² AALTONEN	13Q	CDF	model $\widetilde{\chi}_1^\pm o au X$, simplified gravity- and
		²³ AAD	12AS	ATLS	gauge-mediated models $3\ell^\pm + ot\!\!\!E_T$, pMSSM
		²⁴ AAD	12T	ATLS	$\ell^{\pm}\ell^{\mp} + \cancel{E}_T$, $\ell^{\pm}\ell^{\pm} + \cancel{E}_T$, $pp \rightarrow$
> 163	95	²⁵ CHATRCHYAN ²⁶ CHATRCHYAN			$\begin{array}{ccc} \widetilde{\chi}_{1}^{\pm} \widetilde{\chi}_{2}^{0} \\ \widetilde{W}^{0} & \rightarrow & \gamma \widetilde{G}, \widetilde{W}^{\pm} \rightarrow & \ell^{\pm} \widetilde{G}, \text{GMSB} \\ \tan \beta = 3, & m_{0} = 60 \text{ GeV}, & A_{0} = 0, \\ \mu > 0 & \end{array}$

 $^{^1}$ AAD 16AA summarized and extended ATLAS searches for electroweak supersymmetry in final states containing several charged leptons, E_T , with or without hadronic jets, in 20 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on the $\tilde{\chi}_1^\pm$ mass in the Tchi1chi1B and Tchi1chi1C simplified models. See their Fig. 13.

 $^{^2}$ AAD 16AA summarized and extended ATLAS searches for electroweak supersymmetry in final states containing several charged leptons, E_T , with or without hadronic jets, in 20 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on mass-degenerate $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_2^0$ masses in the Tchi1n2B, Tchi1n2C, and Tchi1n2D simplified models. See their Figs. 16, 17, and 18. Interpretations in phenomenological-MSSM, two-parameter Non Universal Higgs Masses (NUHM2), and gauge-mediated symmetry breaking (GMSB) models are also given in their Figs. 20, 21 and 22.

 $^{^3}$ KHACHATRYAN 16R searched in 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with one or more photons, one electron or muon, and $\not\!\!E_T$. No significant excess above the Standard Model expectations is observed. Limits are set on wino masses in a general gauge-mediated SUSY breaking model (GGM), for a wino-like neutralino NLSP scenario, see Fig. 5. Limits are also set in the Tglu1D and Tchi1n1A simplified models, see Fig. 6. The Tchi1n1A limit is reduced to 340 GeV for a branching ratio reduced by the weak mixing angle.

- ⁴ AAD 15BA searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of charginos and neutralinos decaying to a final state containing a W boson and a 125 GeV Higgs boson, plus missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with the decays $\tilde{\chi}_1^{\pm} \rightarrow W^{\pm} \tilde{\chi}_1^0$ and $\tilde{\chi}_2^0 \rightarrow H \tilde{\chi}_1^0$ having 100% branching fraction, see Fig. 8. A combination of the multiple final states for the Higgs decay yields the best limits (Fig. 8d).
- 5 AAD 15CA searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with one or more photons and E_T , with or without leptons (e, μ). No significant excess above the Standard Model expectations is observed. Limits are set on wino masses in the general gauge-mediated SUSY breaking model (GGM), for wino-like NLSP, see Fig. 9, 12
- ⁶ AAD 14H searched in 20.3 fb⁻¹ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for electroweak production of charginos and neutralinos decaying to a final sate with three leptons and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of direct chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via either all three generations of leptons, staus only, gauge bosons, or Higgs bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 8.
- ⁷ AAD 14X searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the wino-like chargino mass in an R-parity violating simplified model where the decay $\tilde{\chi}_1^{\pm} \to W^{(*)\pm} \tilde{\chi}_1^0$, with $\tilde{\chi}_1^0 \to \ell^{\pm}\ell^{\mp}\nu$, takes place with a branching ratio of 100%, see Fig. 8.
- 8 KHACHATRYAN 14L searched in $19.5~{\rm fb}^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8~{\rm TeV}$ for evidence of chargino-neutralino $\widetilde{\chi}_1^\pm\,\widetilde{\chi}_2^0$ pair production with Higgs or W-bosons in the decay chain, leading to $H\,W$ final states with missing transverse energy. The decays of a Higgs boson to a photon pair are considered in conjunction with hadronic and leptonic decay modes of the W bosons. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of simplified models where the decays $\widetilde{\chi}_2^0 \to H\widetilde{\chi}_1^0$ and $\widetilde{\chi}_1^\pm \to W^\pm\,\widetilde{\chi}_1^0$ take place 100% of the time, see Figs. 22–23.
- ⁹ AAD 13 searched in 4.7 fb⁻¹ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for charginos and neutralinos decaying to a final state with three leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 and 3, and in simplified models, see Fig. 4. For the simplified models with intermediate slepton decays, degenerate $\widetilde{\chi}_1^\pm$ and $\widetilde{\chi}_2^0$ masses up to 500 GeV are excluded at 95% C.L. for very large mass differences with the $\widetilde{\chi}_1^0$. Supersedes AAD 12AS.
- 10 AAD 13B searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for gauginos decaying to a final state with two leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Limits are derived in a simplified model of wino-like chargino pair production, where the chargino always decays to the lightest neutralino via an intermediate on-shell charged slepton, see Fig. 2(b). Chargino masses between 110 and 340 GeV are excluded at 95% C.L. for $m_{\widetilde{\chi}_1^0}=10$ GeV. Exclusion limits are also derived in the phenomenological MSSM, see Fig. 3.
- ¹¹ AAD 12CT searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing four or more leptons (electrons or muons) and either moderate values of missing transverse momentum or large effective mass. No significant excess is found in the data. Limits are presented in a simplified model of R-parity violating supersymmetry in which charginos are pair-produced and then decay into a W-boson and a $\widetilde{\chi}_1^0$, which in turn decays through an RPV coupling into two charged leptons ($e^{\pm}e^{\mp}$ or $e^{\pm}\mu^{\mp}$) and a neutrino. In this model, chargino masses up to 540 GeV are excluded at 95% C.L. for $m_{\widetilde{\chi}_1^0}$ above 300

- GeV, see Fig. 3a. The limit deteriorates for lighter $\tilde{\chi}_1^0$. Limits are also set in an R-parity violating mSUGRA model, see Fig. 3b.
- 12 CHATRCHYAN 12BJ searched in 4.98 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for direct electroweak production of charginos and neutralinos in events with at least two leptons, jets and missing transverse momentum. No significant excesses over the expected SM backgrounds are observed and 95% C.L. limits on the production cross section of $\widetilde{\chi}_1^{\pm}\,\widetilde{\chi}_2^0$ pair production were set in a number of simplified models, see Figs. 7 to 12.
- 13 ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass of charginos is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays), for charginos and for sleptons. These limits are valid for values of $M_2 < 1~{\rm TeV}, \ |\mu| \leq 2~{\rm TeV}$ with the $\widetilde{\chi}^0_1$ as LSP. Constraints from the Higgs search in the $m_h^{\rm max}$ scenario assuming $m_t = 174.3~{\rm GeV}$ are included. The quoted limit applies if there is no mixing in the third family or when $m_{\widetilde{\tau}_1} m_{\widetilde{\chi}^0_1} > 6~{\rm GeV}.$ If mixing is included the limit degrades to 90 GeV. See

Fig. 43 for the mass limits as a function of $tan\beta$. These limits update the results of ABREU 00W.

- 14 KHACHATRYAN 16AA searched in 7.4 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with one or more photons, hadronic jets and E_T . No significant excess above the Standard Model expectations is observed. Limits are set on wino masses in the general gauge-mediated SUSY breaking model (GGM), for a wino-like neutralino NLSP scenario and with the wino mass fixed at 10 GeV above the bino mass, see Fig. 4. Limits are also set in the Tchi1chi1A and Tchi1n1A simplified models, see Fig. 3.
- 15 KHACHATRYAN 16 R searched in $^{19.7}$ fb $^{-1}$ of 16 p collisions at 16 searched in $^{19.7}$ fb $^{-1}$ of 16 p collisions at 16 searched in $^{19.7}$ in 19 No significant excess above the Standard Model expectations is observed. Limits are also set in the Tglu1F simplified model, see Fig. 6.
- 16 KHACHATRYAN 16Y searched in 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with one or two soft isolated leptons, hadronic jets, and E_T . No significant excess above the Standard Model expectations is observed. Limits are set on the $\widetilde{\chi}_1^\pm$ mass (which is degenerate with the $\widetilde{\chi}_2^0$) in the Tchi1n2A simplified model, see Fig. 4.
- ¹⁷ AAD 14AV searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for the direct production of charginos, neutralinos and staus in events containing at last two hadronically decaying τ -leptons, large missing transverse momentum and low jet activity. The quoted limit was derived for direct $\widetilde{\chi}_1^{\pm} \, \widetilde{\chi}_2^0$ and $\widetilde{\chi}_1^{\pm} \, \widetilde{\chi}_1^{\mp}$ production with $\widetilde{\chi}_2^0 \to \widetilde{\tau} \tau \to \tau \tau \widetilde{\chi}_1^0$ and $\widetilde{\chi}_1^{\pm} \to \widetilde{\tau} \nu (\widetilde{\nu}_{\tau} \tau) \to \tau \nu \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_2^0} = m_{\widetilde{\chi}_1^{\pm}}$, $m_{\widetilde{\tau}} = 0.5$ ($m_{\widetilde{\chi}_1^{\pm}} + m_{\widetilde{\chi}_1^0}$), $m_{\widetilde{\chi}_1^0} = 0$ GeV. No excess over the expected SM background is observed. Exclusion limits are set in simplified models of $\widetilde{\chi}_1^{\pm} \, \widetilde{\chi}_1^{\mp}$ and $\widetilde{\chi}_1^{\pm} \, \widetilde{\chi}_2^0$ pair production, see their Figure 7. Upper limits on the cross section and signal strength for direct di-stau production are derived, see Figures 8 and 9. Also, limits are derived in a pMSSM model where the only light slepton is the $\widetilde{\tau}_R$, see Figure 10.
- ¹⁸ AAD 14AV searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for the direct production of charginos, neutralinos and staus in events containing at last two hadronically decaying τ -leptons, large missing transverse momentum and low jet activity. The quoted limit was derived for direct $\widetilde{\chi}_1^{\pm}\widetilde{\chi}_1^{\mp}$ production with $\widetilde{\chi}_1^{\pm}\to\widetilde{\tau}\nu(\widetilde{\nu}_{\tau}\tau)\to\tau\nu\widetilde{\chi}_1^0$, $m_{\widetilde{\tau}}=0.5$ ($m_{\widetilde{\chi}_1^{\pm}}+m_{\widetilde{\chi}_1^0}$), $m_{\widetilde{\chi}_1^0}=0$ GeV. No excess over the expected SM background is observed.

Exclusion limits are set in simplified models of $\widetilde{\chi}_1^{\pm}\widetilde{\chi}_1^{\mp}$ and $\widetilde{\chi}_1^{\pm}\widetilde{\chi}_2^{0}$ pair production, see their Figure 7. Upper limits on the cross section and signal strength for direct di-stau production are derived, see Figures 8 and 9. Also, limits are derived in a pMSSM model where the only light slepton is the $\widetilde{\tau}_R$, see Figure 10.

- AAD 14G searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for electroweak production of chargino pairs, or chargino-neutralino pairs, decaying to a final sate with two leptons (e and μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of chargino pair production, with chargino decays to the lightest neutralino via either sleptons or gauge bosons, see Fig 5.; or in simplified models of chargino and next-to-lightest neutralino production, with decays to the lightest neutralino via gauge bosons, see Fig. 7. An interpretation in the pMSSM is also given, see Fig. 10.
- ²⁰ AALTONEN 14 searched in 5.8 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for evidence of chargino and next-to-lightest neutralino associated production in final states consisting of three leptons (electrons, muons or taus) and large missing transverse momentum. The results are consistent with the Standard Model predictions within 1.85 σ . Limits on the chargino mass are derived in an mSUGRA model with $m_0=60$ GeV, $\tan\beta=3$, $A_0=0$ and $\mu>0$, see their Fig. 2.
- ²¹ KHACHATRYAN 14I searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of chargino pairs decaying to a final state with opposite-sign lepton pairs (e or μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Fig. 18.
- AALTONEN 13Q searched in 6.0 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for evidence of chargino-neutralino associated production in like-sign dilepton final states. One lepton is identified as the hadronic decay of a tau lepton, while the other is an electron or muon. Good agreement with the Standard Model predictions is observed and limits are set on the chargino-neutralino cross section for simplified gravity- and gauge-mediated models, see their Figs. 2 and 3.
- ²³ AAD 12AS searched in 2.06 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for charginos and neutralinos decaying to a final state with three leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the phenomenological MSSM, see Fig. 2 (top), and in simplified models, see Fig. 2 (bottom).
- AAD 12T looked in 1 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for the production of supersymmetric particles decaying into final states with missing transverse momentum and exactly two isolated leptons (e or μ). Opposite-sign and same-sign dilepton events were separately studied. Additionally, in opposite-sign events, a search was made for an excess of same-flavor over different-flavor lepton pairs. No excess over the expected background is observed and limits are placed on the effective production cross section of opposite-sign dilepton events with $E_T>250$ GeV and on same-sign dilepton events with $E_T>100$ GeV. The latter limit is interpreted in a simplified electroweak gaugino production model as a lower chargino mass limit.
- production model as a lower chargino mass limit. 25 CHATRCHYAN 11B looked in 35 pb $^{-1}$ of pp collisions at \sqrt{s} =7 TeV for events with an isolated lepton (e or μ), a photon and E_T which may arise in a generalized gauge mediated model from the decay of Wino-like NLSPs. No evidence for an excess over the expected background is observed. Limits are derived in the plane of squark/gluino mass versus Wino mass (see Fig. 4). Mass degeneracy of the produced squarks and gluinos is assumed.
- 26 CHATRCHYAN 11V looked in $35~\text{pb}^{-1}$ of pp collisions at $\sqrt{s}=7~\text{TeV}$ for events with ≥ 3 isolated leptons (e, μ or τ), with or without jets and $\not\!\!E_T$. No evidence for an excess over the expected background is observed. Limits are derived in the CMSSM $(m_0,\,m_{1/2})$ plane for $\tan\beta=3$ (see Fig. 5).

Long-lived $\widetilde{\chi}^{\pm}$ (Chargino) MASS LIMITS

Limits on charginos which leave the detector before decaying.

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>620	95	¹ AAD	15AE ATLS	stable $\widetilde{\chi}^{\pm}$
>534	95	² AAD	15BM ATLS	stable $\widetilde{\chi}^{\pm}$

>239	95	² AAD	15BM ATLS	$\widetilde{\chi}^{\pm} ightarrow \ \widetilde{\chi}^0_1 \pi^{\pm}$, lifetime 1 ns,
				$m_{\widetilde{\chi}^\pm} - m_{\widetilde{\chi}^0_1} = 0.14 \; GeV$
>482	95	² AAD	15BM ATLS	$\widetilde{\chi}^{\pm} ightarrow \ \widetilde{\chi}_1^0 \pi^{\pm}$, lifetime 15 ns,
				$m_{\widetilde{\chi}^\pm}\stackrel{ extstyle -}{ au}m_{\widetilde{\chi}_1^0}= extstyle 0.14 \; GeV$
>103	95	³ AAD	13H ATLS	long-lived $\widetilde{\chi}^{\pm} {\to} \widetilde{\chi}_1^0 \pi^{\pm}$,
				mAMSB, $\Delta m_{\widetilde{\chi}_1^0} = 160$ MeV
> 92	95	⁴ AAD	1201 ATIC	long-lived $\widetilde{\chi}^{\pm} ightarrow \pi^{\pm} \widetilde{\chi}_{1}^{0}$, mAMSB
> 92	95		12DJ ATLS	long-lived $\chi \to \pi - \chi_1$, maivisb
>171	95	⁵ ABAZOV	09м D0	\widetilde{H}
>102	95	⁶ ABBIENDI	03L OPAL	$m_{\widetilde{\nu}} > 500 \text{ GeV}$
none 2-93.0	95	⁷ ABREU	00T DLPH	\widetilde{H}^{\pm} or $m_{\widetilde{\mathcal{V}}} > m_{\widetilde{\chi}^{\pm}}$
				λ

• • We do not use the following data for averages, fits, limits, etc. • •

>260	95	⁸ KHACHATRY	′ 15 АВ СМЅ	$\widetilde{\chi}_1^\pm ightarrow \ \widetilde{\chi}_1^0 \pi^\pm$, $ au_{\widetilde{\chi}_1^\pm} =$ 0.2ns, AMSB
>800	95	⁹ KHACHATRY		long-lived $\widetilde{\chi}_{1}^{\pm}$, mAMSB, $ au$ >100ns
>100	95	⁹ KHACHATRY	15A0 CMS	long-lived $\widetilde{\chi}_{1}^{\pm}$, mAMSB, $ au > 3$ ns
	95	¹⁰ KHACHATRY	15W CMS	long-lived $\widetilde{\chi}^{f 0}$, $\widetilde{q} ightarrow \ q \widetilde{\chi}^{f 0}$, $\widetilde{\chi}^{f 0} ightarrow$
>270	95	¹¹ AAD	13BD ATLS	$\ell^+\ell^- u$, $ ot\!\!R$ disappearing-track signature, AMSB
>278	95	12 ABAZOV	13B D0	long-lived $\widetilde{\chi}^{\pm}$, gaugino-like
>244	95	¹² ABAZOV	13B D0	long-lived $\widetilde{\chi}^\pm$, higgsino-like

- 1 AAD 15AE searched in 19.1 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ALTAS muon system. In the absence of an excess of events above the expected backgrounds, limits are set on stable charginos, see Fig. 10.
- 2 AAD 15BM searched in 18.4 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for stable and metastable non-relativistic charged particles through their anomalous specific ionization energy loss in the ATLAS pixel detector. In absence of an excess of events above the expected backgrounds, limits are set on stable charginos (see Table 5) and on metastable charginos decaying to $\widetilde{\chi}_1^0\,\pi^\pm$, see Fig. 11.
- ³ AAD 13H searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for direct electroweak production of long-lived charginos in the context of AMSB scenarios. The search is based on the signature of a high-momentum isolated track with few associated hits in the outer part of the tracking system, arising from a chargino decay into a neutralino and a low-momentum pion. The p_T spectrum of the tracks was found to be consistent with the SM expectations. Constraints on the lifetime and the production cross section were obtained, see Fig. 6. In the minimal AMSB framework with $\tan\beta=5$, and $\mu>0$, a chargino having a mass below 103 (85) GeV for a chargino-neutralino mass splitting $\Delta m_{\widetilde{\chi}_1^0}$ of 160 (170) MeV is excluded at the 95% C.L. See Fig. 7 for more precise bounds.
- ⁴ AAD 12BJ looked in 1.02 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for signatures of decaying charginos resulting in isolated tracks with few associated hits in the outer region of the tracking system. The p_T spectrum of the tracks was found to be consistent with the SM expectations. Constraints on the lifetime and the production cross section were obtained. In the minimal AMSB framework with $m_{3/2} < 32$ TeV, $m_0 < 1.5$ TeV, $\tan\beta = 5$, and $\mu > 0$, a chargino having a mass below 92 GeV and a lifetime between 0.5 ns and 2 ns is excluded at the 95% C.L. See their Fig. 8 for more precise bounds. ⁵ ABAZOV 09M searched in 1.1 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s} = 1.96$ TeV for events with
- ⁵ ABAZOV 09M searched in 1.1 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s} = 1.96$ TeV for events with direct production of a pair of charged massive stable particles identified by their TOF. The number of the observed events is consistent with the predicted background. The

- data are used to constrain the production cross section as a function of the $\tilde{\chi}_1^{\pm}$ mass, see their Fig. 2. The quoted limit improves to 206 GeV for gaugino-like charginos.
- ⁶ ABBIENDI 03L used e^+e^- data at $\sqrt{s}=130$ –209 GeV to select events with two high momentum tracks with anomalous dE/dx. The excluded cross section is compared to the theoretical expectation as a function of the heavy particle mass in their Fig. 3. The bounds are valid for colorless fermions with lifetime longer than 10^{-6} s. Supersedes the results from ACKERSTAFF 98P.
- ⁷ ABREU 00T searches for the production of heavy stable charged particles, identified by their ionization or Cherenkov radiation, using data from \sqrt{s} = 130 to 189 GeV. These limits include and update the results of ABREU 98P.
- 8 KHACHATRYAN 15AB searched in 19.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing tracks with little or no associated calorimeter energy deposits and with missing hits in the outer layers of the tracking system (disappearing-track signature). Such disappearing tracks can result from the decay of charginos that are nearly mass degenerate with the lightest neutralino. The number of observed events is in agreement with the background expectation. Limits are set on the cross section of electroweak chargino production in terms of the chargino mass and mean proper lifetime, see Fig. 4. In the minimal AMSB model, a chargino mass below 260 GeV is excluded at 95% C.L., see their Fig. 5.
- 9 KHACHATRYAN 150 searched in 18.8 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for evidence of long-lived charginos in the context of AMSB and pMSSM scenarios. The results are based on a previously published search for heavy stable charged particles at 7 and 8 TeV. In the minimal AMSB framework with $\tan\beta=5$ and $\mu\geq0$, constraints on the chargino mass and lifetime were placed, see Fig. 5. Charginos with a mass below 800 (100) GeV are excluded at the 95% C.L. for lifetimes above 100 ns (3 ns). Constraints are also placed on the pMSSM parameter space, see Fig. 3.
- ¹⁰ KHACHATRYAN 15W searched in up to 20.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for evidence of long-lived neutralinos produced through \widetilde{q} -pair production, with $\widetilde{q}\to q\widetilde{\chi}^0$ and $\widetilde{\chi}^0\to \ell^+\ell^-\nu$ ($\not\!R$: λ_{121} , $\lambda_{122}\neq 0$). 95% C.L. exclusion limits on cross section times branching ratio are set as a function of mean proper decay length of the neutralino, see Figs. 6 and 9.
- 11 AAD 13BD searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing tracks with no associated hits in the outer region of the tracking system resulting from the decay of charginos that are nearly mass degenerate with the lightest neutralino, as is often the case in AMSB scenarios. No significant excess above the background expectation is observed for candidate tracks with large transverse momentum. Constraints on chargino properties are obtained and in the minimal AMSB model, a chargino mass below 270 GeV is excluded at 95% C.L., see their Fig. 7.
- 12 ABAZOV 13B looked in 6.3 fb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for charged massive long-lived particles in events with muon-like particles that have both speed and ionization energy loss inconsistent with muons produced in beam collisions. In the absence of an excess, limits are set at 95% C.L. on gaugino- and higgsino-like charginos, see their Table 20 and Fig. 23.

$\widetilde{\nu}$ (Sneutrino) MASS LIMIT

The limits may depend on the number, $N(\widetilde{\nu})$, of sneutrinos assumed to be degenerate in mass. Only $\widetilde{\nu}_L$ (not $\widetilde{\nu}_R$) is assumed to exist. It is possible that $\widetilde{\nu}$ could be the lightest supersymmetric particle (LSP).

We report here, but do not include in the Listings, the limits obtained from the fit of the final results obtained by the LEP Collaborations on the invisible width of the Z boson ($\Delta\Gamma_{\rm inv.} < 2.0$ MeV, LEP-SLC 06): $m_{\widetilde{\nu}} > 43.7$ GeV ($N(\widetilde{\nu}) = 1$) and $m_{\widetilde{\nu}} > 44.7$ GeV ($N(\widetilde{\nu}) = 3$) .

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>2300	95	$^{ m 1}$ AABOUD	16 P	ATLS	$ R_{\!\scriptscriptstyle i},\widetilde{ u}_{\scriptscriptstyle \mathcal{T}} ightarrow e\mu, \lambda_{311}' = 0.11$
>2200	95	$^{ m 1}$ AABOUD	16 P	ATLS	$ R, \tilde{\nu}_{\tau} \rightarrow e \tau, \lambda_{311}^{711} = 0.11 $
>1900	95	¹ AABOUD	16 P	ATLS	$R, \widetilde{\nu}_{\tau} \rightarrow \mu \tau, \lambda_{311}^{711} = 0.11$
> 400	95	² AAD			$ \mathbb{R}, \geq 4\ell^{\pm}, \widetilde{\nu} \rightarrow \nu \widetilde{\chi}_{1}^{0}, \widetilde{\chi}_{1}^{0} \rightarrow $
> 94	95	³ AAD ⁴ ABDALLAH			$egin{aligned} \ell^\pm \ell^\mp u \ R, \widetilde{ u}_ au & ightarrow e \mu \ 1 & \leq & aneta & \leq & 40, \ m_{\widetilde{e}_R} - m_{\widetilde{\chi}_1^0} & > 10 \; ext{GeV} \end{aligned}$
> 84 > 41	95 95	⁵ HEISTER ⁶ DECAMP	02N 92	ALEP ALEP	$\widetilde{ u}_e$, any Δm $\Gamma(Z o ext{ invisible}); N(\widetilde{ u})=3, model independent$

• • • We do not use the following data for averages, fits, limits, etc. • • •

>1280	95	⁷ KHACHATRY.	16BE CMS	$R\!\!\!/,~\widetilde{ u}_{ au} ightarrow~$ e μ , $~\lambda_{132} = \lambda_{231} = 0$
>2300	95	⁷ KHACHATRY.	16BE CMS	$\lambda'_{311} = 0.01$ $\mathcal{R}, \ \tilde{\nu}_{\tau} \rightarrow e \mu, \ \lambda_{132} = \lambda_{231} = 0.07, \ \lambda'_{311} = 0.11$
>2000	95	⁸ AAD	150 ATLS	$R(e\mu)$, $\widetilde{\nu}_{ au}$, $\lambda'_{311}=$ 0.11, $\lambda_{i3k}=$
>1700	95	⁸ AAD	150 ATLS	$ \mathcal{R}(\tau \mu, e \tau), \widetilde{\nu}_{\tau}, \lambda'_{311} = 0.11, $
> 95	95	9 AAD 10 AAD 11 AALTONEN 12 ABAZOV 13 ABDALLAH	11H ATLS 10Z CDF 10M D0	$\lambda_{i3k}=0.07$ $ \mathcal{R}, \widetilde{\nu}_{ au} ightarrow e\mu, e au, \mu au$ $ \mathcal{R}, \widetilde{\nu}_{ au} ightarrow e\mu$ $ \mathcal{R}, \widetilde{\nu}_{ au} ightarrow e\mu, e au, \mu au$ $ \mathcal{R}, \widetilde{\nu}_{ au} ightarrow e\mu$ AMSB, $\mu > 0$
> 37.1	95	¹⁴ ADRIANI		$\Gamma(Z \rightarrow \text{invisible}); N(\widetilde{\nu})=1$
> 36	95	ABREU		$\Gamma(Z \rightarrow \text{invisible}); N(\widetilde{\nu})=1$
> 31.2	95	¹⁵ ALEXANDER		$\Gamma(Z \rightarrow \text{invisible}); N(\widetilde{\nu})=1$

¹ AABOUD 16P searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for events with different flavour dilepton pairs $(e\mu,\,e\tau,\,\mu\tau)$ from the production of $\widetilde{\nu}_{\tau}$ via an \Re λ'_{311} coupling and followed by a decay via $\lambda_{312}=\lambda_{321}=0.07$ for $e+\mu$, via $\lambda_{313}=\lambda_{331}=0.07$ for $e+\tau$ and via $\lambda_{323}=\lambda_{332}=0.07$ for $\mu+\tau$. No evidence for a dilepton resonance over the SM expectation is observed, and limits are derived on $m_{\widetilde{\nu}}$ at 95% CL, see their Figs. 2(b), 3(b), 4(b), and Table 3.

CL, see their Figs. 2(b), 3(b), 4(b), and Table 3.
² AAD 14X searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the sneutrino mass in an R-parity violating simplified model where the decay $\widetilde{\nu} \to \nu \widetilde{\chi}_1^0$, with $\widetilde{\chi}_1^0 \to \ell^\pm \ell^\mp \nu$, takes place with a branching ratio of 100%, see Fig. 9.

³ AAD 11Z looked in 1.07 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with one electron and one muon of opposite charge from the production of $\widetilde{\nu}_{\tau}$ via an \Re λ'_{311} coupling and followed by a decay via λ_{312} into $e+\mu$. No evidence for an (e,μ) resonance over the SM expectation is observed, and a limit is derived in the plane of λ'_{311} versus $m_{\widetilde{\nu}}$ for three values of λ_{312} , see their Fig. 2. Masses $m_{\widetilde{\nu}}<1.32$ (1.45) TeV are excluded for $\lambda'_{311}=0.10$ and $\lambda_{312}=0.05$ ($\lambda'_{311}=0.11$ and $\lambda_{312}=0.07$).

- 4 ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of M $_2 < 1$ TeV, $\left| \mu \right| \leq 1$ TeV with the $\widetilde{\chi}_1^0$ as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of $\tan\beta$. These limits update the results of ABREU 00W.
- 5 HEISTER 02N derives a bound on $m_{\widetilde{\nu}_e}$ by exploiting the mass relation between the $\widetilde{\nu}_e$ and \widetilde{e} , based on the assumption of universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 and the search described in the \widetilde{e} section. In the MSUGRA framework with radiative electroweak symmetry breaking, the limit improves to $m_{\widetilde{\nu}_e} > \!\! 130$ GeV, assuming a trilinear coupling $A_0 \! = \! 0$ at the GUT scale. See Figs. 5 and 7 for the dependence of the limits on $\tan\beta$.
- $^6\, \rm DECAMP$ 92 limit is from $\Gamma(\rm invisible)/\Gamma(\ell\ell) = 5.91 \pm 0.15$ ($N_{\nu} = 2.97 \pm 0.07$).
- ⁷ KHACHATRYAN 16BE searched in 19.7 fb⁻¹ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for evidence of narrow resonances decaying into $e\,\mu$ final states. No significant excess above the Standard Model expectation is observed and 95% C.L. exclusions are placed on the cross section times branching ratio for the production of an R-parity-violating supersymmetric tau sneutrino, see their Fig. 3.
- ⁸ AAD 150 searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for evidence of heavy particles decaying into $e\mu$, $e\tau$ or $\mu\tau$ final states. No significant excess above the Standard Model expectation is observed, and 95% C.L. exclusions are placed on the cross section times branching ratio for the production of an R-parity-violating supersymmetric tau sneutrino, applicable to any sneutrino flavour, see their Fig. 2.
- ⁹AAD 13AI searched in 4.6 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for evidence of heavy particles decaying into $e\mu$, $e\tau$ or $\mu\tau$ final states. No significant excess above the Standard Model expectation is observed, and 95% C.L. exclusions are placed on the cross section times branching ratio for the production of an R-parity-violating supersymmetric tau sneutrino, see their Fig. 2. For couplings $\lambda'_{311}=0.10$ and $\lambda_{i3k}=0.05$, the lower limits on the $\widetilde{\nu}_{\tau}$ mass are 1610, 1110, 1100 GeV in the $e\mu$, $e\tau$, and $\mu\tau$ channels, respectively.
- ¹⁰ AAD 11H looked in 35 pb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with one electron and one muon of opposite charge from the production of $\widetilde{\nu}_{\tau}$ via an k λ'_{311} coupling and followed by a decay via λ_{312} into $e+\mu$. No evidence for an excess over the SM expectation is observed, and a limit is derived in the plane of λ'_{311} versus $m_{\widetilde{\nu}}$ for several values of λ_{312} , see their Fig. 2. Superseded by AAD 11Z.
- ¹¹ AALTONEN 10Z searched in 1 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events from the production $d\overline{d} \to \widetilde{\nu}_{\tau}$ with the subsequent decays $\widetilde{\nu}_{\tau} \to e\mu$, $\mu\tau$, $e\tau$ in the MSSM framework with R. Two isolated leptons of different flavor and opposite charges are required, with τs identified by their hadronic decay. No statistically significant excesses are observed over the SM background. Upper limits on $\lambda_{311}'^2$ times the branching ratio are listed in their Table III for various $\widetilde{\nu}_{\tau}$ masses. Limits on the cross section times branching ratio for $\lambda_{311}' = 0.10$ and $\lambda_{i3k} = 0.05$, displayed in Fig. 2, are used to set limits on the $\widetilde{\nu}_{\tau}$ mass of 558 GeV for the $e\mu$, 441 GeV for the $\mu\tau$ and 442 GeV for the $e\tau$ channels.
- ^{12} ABAZOV 10M looked in 5.3 fb^{-1} of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with exactly one pair of high p_T isolated $e\mu$ and a veto against hard jets. No evidence for an excess over the SM expectation is observed, and a limit at 95% C.L. on the cross section times branching ratio is derived, see their Fig. 3. These limits are translated into limits on couplings as a function of $m_{\widetilde{\nu}_T}$ as shown on their Fig. 4. As an example, for $m_{\widetilde{\nu}_T}=100$ GeV and $\lambda_{312}\leq0.07$, couplings $\lambda'_{311}>7.7\times10^{-4}$ are excluded.
- 13 ABDALLAH 04H use data from LEP 1 and $\sqrt{s}=192$ –208 GeV. They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region

 $1 < m_{3/2} <$ 50 TeV, $0 < m_0 <$ 1000 GeV, 1.5 < tan $\beta <$ 35, both signs of $\mu.$ The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z width of 3.2 MeV. The limit is for $m_t =$ 174.3 GeV (see Table 2 for other m_t values). The limit improves to 114 GeV for $\mu <$ 0.

 14 ADRIANI 93M limit from $\Delta\Gamma(Z)$ (invisible) < 16.2 MeV.

CHARGED SLEPTONS

This section contains limits on charged scalar leptons $(\widetilde{\ell}, \text{ with } \ell = e, \mu, \tau)$. Studies of width and decays of the Z boson (use is made here of $\Delta\Gamma_{\text{inv}} < 2.0 \,\text{MeV}$, LEP 00) conclusively rule out $m_{\widetilde{\ell}_R} < 40 \,\text{GeV}$ (41 GeV for $\widetilde{\ell}_L$) independently of decay modes, for each individual slepton

GeV for $\widetilde{\ell}_L$) , independently of decay modes, for each individual slepton. The limits improve to 43 GeV (43.5 GeV for $\widetilde{\ell}_L$) assuming all 3 flavors to be degenerate. Limits on higher mass sleptons depend on model assumptions and on the mass splitting $\Delta m = m_{\widetilde{\ell}} - m_{\widetilde{\chi}_1^0}$. The mass and composition

of $\widetilde{\chi}_1^0$ may affect the selectron production rate in e^+e^- collisions through t-channel exchange diagrams. Production rates are also affected by the potentially large mixing angle of the lightest mass eigenstate $\widetilde{\ell}_1 = \widetilde{\ell}_R \sin\theta_\ell + \widetilde{\ell}_L \cos\theta_\ell$. It is generally assumed that only $\widetilde{\tau}$ may have significant mixing. The coupling to the Z vanishes for $\theta_\ell = 0.82$. In the high-energy limit of e^+e^- collisions the interference between γ and Z exchange leads to a minimal cross section for $\theta_\ell = 0.91$, a value which is sometimes used in the following entries relative to data taken at LEP2. When limits on $m_{\widetilde{\ell}_R}$ are quoted, it is understood that limits on $m_{\widetilde{\ell}_L}$ are usually at least as strong.

Possibly open decays involving gauginos other than $\widetilde{\chi}^0_1$ will affect the detection efficiencies. Unless otherwise stated, the limits presented here result from the study of $\widetilde{\ell}^+\widetilde{\ell}^-$ production, with production rates and decay properties derived from the MSSM. Limits made obsolete by the recent analyses of e^+e^- collisions at high energies can be found in previous Editions of this Review.

For decays with final state gravitinos (\widetilde{G}) , $m_{\widetilde{G}}$ is assumed to be negligible relative to all other masses.

ẽ (Selectron) MASS LIMIT

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

¹⁵ ALEXANDER 91F limit is for one species of $\tilde{\nu}$ and is derived from Γ(invisible, new)/Γ($\ell\ell$) < 0.38.

> 97.5		⁴ ABBIENDI	04	OPAL	\widetilde{e}_R , $\Delta m > 11$ GeV, $\left \mu ight >$ 100 GeV,
> 94.4		⁵ ACHARD	04	L3	$\tan \beta = 1.5$ $\widetilde{e}_{R}, \Delta m > 10$ GeV, $\left \mu \right > 200$ GeV, $\tan \beta > 2$
> 71.3		⁵ ACHARD	04	L3	\tilde{e}_R , all Δm
none 30-94	95	⁶ ABDALLAH	03м	DLPH	$\Delta m > 15$ GeV, $\widetilde{e}_R^+ \widetilde{e}_R^-$
> 94	95	⁷ ABDALLAH	03м	DLPH	\widetilde{e}_R , $1 \leq \tan \beta \leq 40$, $\Delta m > 10$ GeV
> 95	95	⁸ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $\widetilde{e}_R^+ \widetilde{e}_R^-$
> 73	95	⁹ HEISTER	02N	ALEP	\widetilde{e}_{R} , any Δm
>107	95	⁹ HEISTER	02N	ALEP	\widetilde{e}_L , any Δm
\				۲۰.	. 1

• • We do not use the following data for averages, fits, limits, etc.

none 90–325 95
10
 AAD 14G ATLS $\widetilde{\ell}\widetilde{\ell}\to\ell^+\widetilde{\chi}_1^0\ell^-\widetilde{\chi}_1^0$, simplified model, $m_{\widetilde{\ell}_L}=m_{\widetilde{\ell}_R}, m_{\widetilde{\chi}_1^0}=0$ GeV 11 KHACHATRY...14I CMS $\widetilde{\ell}\to\ell\widetilde{\chi}_1^0$, simplified model $^{>89}$ 95 12 ABBIENDI 04F OPAL $^{\not{R}}$, \widetilde{e}_L $^{>92}$ 95 13 ABDALLAH 04M DLPH $^{\not{R}}$, \widetilde{e}_R , indirect, $\Delta m>$ 5 GeV

- ¹AAD 14X searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in an R-parity violating simplified model where the decay $\tilde{\ell}\to\ell\tilde{\chi}^0_1$, with $\tilde{\chi}^0_1\to\ell^\pm\ell^\mp\nu$, takes place with a branching ratio of 100%, see Fig. 9.
- 2 CHATRCHYAN 14R searched in 19.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in a stau (N)NLSP simplified model (GMSB) where the decay $\tilde{\ell}\to\ell^\pm\tau^\pm\tau^\mp\tilde{G}$ takes place with a branching ratio of 100%, see Fig. 8.
- ³ AAD 13B searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for sleptons decaying to a final state with two leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Limits are derived in a simplified model of direct left-handed slepton pair production, where left-handed slepton masses between 85 and 195 GeV are excluded at 95% C.L. for $m_{\widetilde{\chi}_1^0}=20$ GeV. See also Fig. 2(a). Exclusion limits are also derived in the phenomenological MSSM, see Fig. 3.
- ⁴ ABBIENDI 04 search for $\widetilde{e}_R\widetilde{e}_R$ production in acoplanar di-electron final states in the 183–208 GeV data. See Fig. 13 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$ and for the limit at $\tan\beta$ =35 This limit supersedes ABBIENDI 00G.
- 5 ACHARD 04 search for $\widetilde{e}_R\widetilde{e}_L$ and $\widetilde{e}_R\widetilde{e}_R$ production in single- and acoplanar di-electron final states in the 192–209 GeV data. Absolute limits on $m_{\widetilde{e}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and $m_0,~1 \leq \tan\beta \leq 60$ and $-2 \leq \mu \leq 2$ TeV. See Fig. 4 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. This limit supersedes ACCIARRI 99W.
- ⁶ ABDALLAH 03M looked for acoplanar dielectron $+\cancel{E}$ final states at $\sqrt{s}=189$ –208 GeV. The limit assumes $\mu=-200$ GeV and $\tan\beta=1.5$ in the calculation of the production cross section and B($\widetilde{e} \rightarrow e \widetilde{\chi}_1^0$). See Fig. 15 for limits in the $(m_{\widetilde{e}_R}, m_{\widetilde{\chi}_1^0})$ plane. These limits include and update the results of ABREU 01
- 7 ABDALLAH 03M uses data from $\sqrt{s}=192$ –208 GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of \textit{M}_2 <1 TeV, $|\mu| \leq$ 1 TeV with the $\tilde{\chi}_1^0$ as LSP. The quoted

limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of $tan\beta$. These limits update the results of ABREU 00W.

- ⁸ HEISTER 02E looked for acoplanar dielectron $+ \not\!\!\!E_T$ final states from e^+e^- interactions between 183 and 209 GeV. The mass limit assumes $\mu < -200$ GeV and $\tan\beta = 2$ for the production cross section and B($\check{e} \rightarrow e \widetilde{\chi}_1^0$)=1. See their Fig. 4 for the dependence of the limit on Δm . These limits include and update the results of BARATE 01.
- ⁹ HEISTER 02N search for $\widetilde{e}_R\widetilde{e}_L$ and $\widetilde{e}_R\widetilde{e}_R$ production in single- and acoplanar di-electron final states in the 183–208 GeV data. Absolute limits on $m_{\widetilde{e}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 , $1 \leq \tan\beta \leq 50$ and $-10 \leq \mu \leq 10$ TeV. The region of small $|\mu|$, where cascade decays are important, is covered by a search for $\widetilde{\chi}_1^0\widetilde{\chi}_3^0$ in final states with leptons and possibly photons. Limits on $m_{\widetilde{e}_L}$ are derived by exploiting the mass relation between the \widetilde{e}_L and \widetilde{e}_R , based on universal m_0 and $m_{1/2}$. When the constraint from the mass limit of the lightest Higgs from HEISTER 02 is included, the bounds improve to $m_{\widetilde{e}_R} > 77(75)$ GeV and $m_{\widetilde{e}_L} > 115(115)$ GeV for a top mass of 175(180) GeV. In the MSUGRA framework with radiative electroweak symmetry breaking, the limits improve further to $m_{\widetilde{e}_R} > 95$ GeV and $m_{\widetilde{e}_L} > 152$ GeV, assuming a trilinear coupling $A_0 = 0$ at the GUT scale. See Figs. 4, 5, 7 for the dependence of the limits on $\tan\beta$.
- 10 AAD 14G searched in 20.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of slepton pairs, decaying to a final sate with two leptons (e and μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of slepton pair production, see Fig. 8. An interpretation in the pMSSM is also given, see Fig. 10.
- ¹¹ KHACHATRYAN 14I searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of slepton pairs decaying to a final state with opposite-sign lepton pairs (e or μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Fig. 18.
- ABBIENDI 04F use data from $\sqrt{s}=189$ –209 GeV. They derive limits on sparticle masses under the assumption of $\not\!\!R$ with $LL\overline{E}$ or $LQ\overline{D}$ couplings. The results are valid for $\tan\beta=1.5,\ \mu=-200$ GeV, with, in addition, $\Delta m > 5$ GeV for indirect decays via $LQ\overline{D}$. The limit quoted applies to direct decays via $LL\overline{E}$ or $LQ\overline{D}$ couplings. For indirect decays, the limits on the $ellow{e}_R$ mass are respectively 99 and 92 GeV for $LL\overline{E}$ and $LQ\overline{D}$ couplings and $m_{\widetilde{\chi}0}=10$ GeV and degrade slightly for larger $ellow{\chi}0$ mass. Supersedes the results of ABBIENDI 00.
- ABDALLAH 04M use data from $\sqrt{s}=192$ –208 GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or \overline{UDD} couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5$, $\Delta m>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect \overline{UDD} decays using the neutralino constraint of 39.5 GeV for $LL\overline{E}$ and of 38.0 GeV for \overline{UDD} couplings, also derived in ABDALLAH 04M. For indirect decays via $LL\overline{E}$ the limit improves to 95 GeV if the constraint from the neutralino is used and to 94 GeV if it is not used. For indirect decays via \overline{UDD} couplings it remains unchanged when the neutralino constraint is not used. Supersedes the result of ABREU 00U.

$\widetilde{\mu}$ (Smuon) MASS LIMIT

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>410	95	¹ AAD			\geq 4 ℓ^{\pm} , $\widetilde{\ell}$ \rightarrow $\ell \widetilde{\chi}_1^0$, $\widetilde{\chi}_1^0$ \rightarrow
		² CHATRCHYAI	V 14 R	CMS	$\ell^{\pm}\ell^{\mp}\nu$, R-parity viol. $\geq 3\ell^{\pm}$, $\tilde{\ell} \rightarrow \ell^{\pm}\tau^{\mp}\tau^{\mp}\tilde{G}$ simplified model, GMSB, stau (N)NLSP scenario
		³ AAD	13 B	ATLS	$2\ell^{\pm} + \cancel{E}_T$, SMS, pMSSM

> 91.0		⁴ ABBIENDI	04	OPAL	$\Delta m > 3 \text{ GeV}, \ \widetilde{\mu}_R^+ \widetilde{\mu}_R^-,$
					$\left \mu ight >$ 100 GeV, tan $eta=$ 1.5
> 86.7		⁵ ACHARD	04	L3	Δm >10 GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$,
					$\left \mu ight >$ 200 GeV, $ aneta\geq 2$
none 30-88	95	⁶ ABDALLAH	03M	DLPH	$\Delta m > 5$ GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$
> 94	95	⁷ ABDALLAH	03м	DLPH	$\widetilde{\mu}_{R}, 1 \leq \tan\beta \leq 40,$ $\Delta m > 10 \text{ GeV}$
					$\Delta m > 10 \; GeV$
> 88	95	⁸ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $\widetilde{\mu}_R^+ \widetilde{\mu}_R^-$
• • • We do n	ot use the	following data fo	r aver	ages, fit	s, limits, etc. • • •
none 90-325	95	⁹ AAD	14 G	ATLS	$\widetilde{\ell}\widetilde{\ell} \to \ell^+\widetilde{\chi}_1^0\ell^-\widetilde{\chi}_1^0$, simplified
					model, $m_{\widetilde{\ell}_L} = m_{\widetilde{\ell}_R}$, $m_{\widetilde{\chi}_1^0} =$

 1 AAD 14x searched in 20.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in an R-parity violating simplified model where the decay $\tilde{\ell} \to \ell \tilde{\chi}_1^0$, with $\tilde{\chi}_1^0 \to \ell^\pm \ell^\mp \nu$, takes place with a branching ratio of 100%, see Fig. 9.

 2 CHATRCHYAN 14R searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the slepton mass in a stau (N)NLSP simplified model (GMSB) where the decay $\tilde{\ell} \to \ell^{\pm} \tau^{\mp} \tilde{G}$ takes place with a branching ratio of 100%, see Fig. 8.

³ AAD 13B searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for sleptons decaying to a final state with two leptons (e and μ) and missing transverse energy. No excess beyond the Standard Model expectation is observed. Limits are derived in a simplified model of direct left-handed slepton pair production, where left-handed slepton masses between 85 and 195 GeV are excluded at 95% C.L. for $m_{\widetilde{\chi}_1^0}=20$ GeV. See also Fig. 2(a). Exclusion

limit at $\tan\!\beta\!=\!35$. Under the assumption of 100% branching ratio for $\widetilde{\mu}_R\to \mu \ \widetilde{\chi}_1^0$, the limit improves to 94.0 GeV for $\Delta m>4$ GeV. See Fig. 11 for the dependence of the limits on $\mathbf{m}_{\widetilde{\chi}_1^0}$ at several values of the branching ratio. This limit supersedes ABBIENDI 00G.

⁵ ACHARD 04 search for $\widetilde{\mu}_R\widetilde{\mu}_R$ production in acoplanar di-muon final states in the 192–209 GeV data. Limits on $m_{\widetilde{\mu}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 , $1 \leq \tan\beta \leq 60$ and $-2 \leq \mu \leq 2$ TeV. See Fig. 4 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$. This limit supersedes ACCIARRI 99W.

⁶ ABDALLAH 03M looked for acoplanar dimuon $+\cancel{E}$ final states at $\sqrt{s}=189$ –208 GeV. The limit assumes B $(\widetilde{\mu}\to \mu\widetilde{\chi}^0_1)=100\%$. See Fig. 16 for limits on the $(m_{\widetilde{\mu}_R}, m_{\widetilde{\chi}^0_1})$ plane. These limits include and update the results of ABREU 01.

 7 ABDALLAH 03M uses data from $\sqrt{s}=192\text{--}208$ GeV to obtain limits in the framework of the MSSM with gaugino and sfermion mass universality at the GUT scale. An indirect limit on the mass is derived by constraining the MSSM parameter space by the results

from direct searches for neutralinos (including cascade decays) and for sleptons. These limits are valid for values of M $_2<1$ TeV, $\left|\mu\right|\leq 1$ TeV with the $\widetilde{\chi}^0_1$ as LSP. The quoted limit is obtained when there is no mixing in the third family. See Fig. 43 for the mass limits as a function of $\tan\beta$. These limits update the results of ABREU 00W.

⁸ HEISTER 02E looked for acoplanar dimuon $+ \not\!\!E_T$ final states from e^+e^- interactions between 183 and 209 GeV. The mass limit assumes B($\widetilde{\mu} \to \mu \widetilde{\chi}_1^0$)=1. See their Fig. 4 for the dependence of the limit on Δm . These limits include and update the results of BARATE 01.

⁹AAD 14G searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of slepton pairs, decaying to a final sate with two leptons (e and μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models of slepton pair production, see Fig. 8. An interpretation in the pMSSM is also given, see Fig. 10.

¹⁰ KHACHATRYAN 14I searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for electroweak production of slepton pairs decaying to a final state with opposite-sign lepton pairs (e or μ) and missing transverse momentum. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in simplified models, see Fig. 18.

 11 ABDALLAH 04M use data from $\sqrt{s}=192\text{--}208$ GeV to derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or \overline{UDD} couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5$, $\Delta m>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect \overline{UDD} decays using the neutralino constraint of 39.5 GeV for $LL\overline{E}$ and of 38.0 GeV for \overline{UDD} couplings, also derived in ABDALLAH 04M. For indirect decays via $LL\overline{E}$ the limit improves to 90 GeV if the constraint from the neutralino is used and remains at 87 GeV if it is not used. For indirect decays via \overline{UDD} couplings it degrades to 85 GeV when the neutralino constraint is not used. Supersedes the result of ABREU 00U.

12 HEISTER 03G searches for the production of smuons in the case of R prompt decays with $LL\overline{E},\ LQ\overline{D}$ or \overline{UDD} couplings at $\sqrt{s}=189$ –209 GeV. The search is performed for direct and indirect decays, assuming one coupling at a time to be non-zero. The limit holds for direct decays mediated by R $LQ\overline{D}$ couplings and improves to 90 GeV for indirect decays (for $\Delta m>10$ GeV). Limits are also given for $LL\overline{E}$ direct ($m_{\widetilde{\mu}R}>87$ GeV) and indirect decays ($m_{\widetilde{\mu}R}>96$ GeV for $m(\widetilde{\chi}_1^0)>23$ GeV from BARATE 98S) and for \overline{UDD} indirect decays ($m_{\widetilde{\mu}R}>85$ GeV for $\Delta m>10$ GeV). Supersedes the results from BARATE 01B.

¹³ ABREU 00V use data from $\sqrt{s}=130-189$ GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. For limits at different $m_{\widetilde{G}}$, see their Fig. 12.

$\widetilde{\tau}$ (Stau) MASS LIMIT

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>85.2		¹ ABBIENDI	04	OPAL	$\Delta m >$ 6 GeV, $\theta_{\tau}{=}\pi/2, \ \left \mu\right > 100$ GeV, $\tan\!\beta{=}1.5$
>78.3		² ACHARD	04		$\Delta m > 15$ GeV, $ heta_{ au} {=} \pi/2$,
		•			$\left \mu ight >$ 200 GeV, $ aneta\geq 2$
>81.9	95	³ ABDALLAH	03M	DLPH	$\Delta m >$ 15 GeV, all $ heta_{ au}$
>79	95	⁴ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $ heta_{ au} = \pi/2$
>76	95	⁴ HEISTER	02E	ALEP	$\Delta m > 15$ GeV, $ heta_{ au} = 0.91$

• • • We do not use the following data for averages, fits, limits, etc. • • •

none 109	95	⁵ AAD	16AA ATLS	2 hadronic $ au+ ot\!$
				$ au \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} = 0$ GeV
		⁶ AAD	12AF ATLS	$2 au+jets+\cancel{\cancel{E}}_T$, GMSB
		⁷ AAD	12AG ATLS	$\geq 1\tau_h + \text{jets} + \cancel{E}_T$, GMSB
		⁸ AAD	12CM ATLS	$\geq 1 au + jets + ot\!$
>87.4	95	⁹ ABBIENDI		$\widetilde{ au}_{m{R}} ightarrow \; au \widetilde{m{G}}$, all $ au (\widetilde{ au}_{m{R}})$
>74	95	¹⁰ ABBIENDI	04F OPAL	$R, \widetilde{\tau}_{l}$
>68	95	¹¹ ABDALLAH	04H DLPH	AMSB, $\mu > 0$
>90	95	¹² ABDALLAH	04м DLPH	R , $\widetilde{ au}_R$, indirect, $\Delta m >$ 5 GeV
none $m_{ au}-$ 26.3	95	³ ABDALLAH	03M DLPH	$\Delta m > m_{_{T}}$, all $ heta_{_{T}}$

¹ ABBIENDI 04 search for $\widetilde{\tau}\widetilde{\tau}$ production in acoplanar di-tau final states in the 183–208 GeV data. See Fig. 15 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$ and for the limit

at $\tan\beta$ =35. Under the assumption of 100% branching ratio for $\widetilde{\tau}_R \to \tau \ \widetilde{\chi}_1^0$, the limit improves to 89.8 GeV for $\Delta m >$ 8 GeV. See Fig. 12 for the dependence of the limits on $\mathbf{m}_{\widetilde{\chi}_1^0}$ at several values of the branching ratio and for their dependence on $\theta_{\mathcal{T}}$. This limit supersedes ABBIENDI 00G.

² ACHARD 04 search for $\widetilde{\tau}\widetilde{\tau}$ production in acoplanar di-tau final states in the 192–209 GeV data. Limits on $m_{\widetilde{\tau}_R}$ are derived from a scan over the MSSM parameter space with universal GUT scale gaugino and scalar masses $m_{1/2}$ and m_0 , $1 \leq \tan\beta \leq 60$ and $-2 \leq \mu \leq 2$ TeV. See Fig. 4 for the dependence of the limits on $m_{\widetilde{\chi}_1^0}$.

³ ABDALLAH 03M looked for acoplanar ditaus $+\cancel{E}$ final states at $\sqrt{s}=130$ –208 GeV. A dedicated search was made for low mass $\widetilde{\tau}$ s decoupling from the Z^0 . The limit assumes B($\widetilde{\tau} \to \tau \widetilde{\chi}^0_1$) = 100%. See Fig. 20 for limits on the $(m_{\widetilde{\tau}}, m_{\widetilde{\chi}^0_1})$ plane and as function

of the $\widetilde{\chi}_1^0$ mass and of the branching ratio. The limit in the low-mass region improves to 29.6 and 31.1 GeV for $\widetilde{\tau}_R$ and $\widetilde{\tau}_L$, respectively, at $\Delta m > m_{\mathcal{T}}$. The limit in the high-mass region improves to 84.7 GeV for $\widetilde{\tau}_R$ and $\Delta m > 15$ GeV. These limits include and update the results of ABREU 01.

⁴ HEISTER 02E looked for acoplanar ditau $+ \not\!\! E_T$ final states from e^+e^- interactions between 183 and 209 GeV. The mass limit assumes B($\tilde{\tau} \to \tau \tilde{\chi}_1^0$)=1. See their Fig. 4 for the dependence of the limit on Δm . These limits include and update the results of BARATE 01.

SAAD 16AA summarized and extended ATLAS searches for electroweak supersymmetry in final states containing several charged leptons, $\not\!\!E_T$, with or without hadronic jets, in 20 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV. The paper reports 95% C.L. exclusion limits on the cross-section for production of $\tilde{\tau}_R$ and $\tilde{\tau}_L$ pairs for various $m_{\widetilde{\chi}_1^0}$, using the 2 hadronic $\tau+\not\!\!E_T$ analysis. The $m_{\widetilde{\tau}_R/L}=109$ GeV is excluded for $m_{\widetilde{\chi}_1^0}=0$ GeV, with the constraints being stronger for $\tilde{\tau}_R$. See their Fig. 12.

⁶ AAD 12AF searched in 2 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with two tau leptons, jets and large $\not\!\!E_T$ in a GMSB framework. No significant excess above the expected background was found and an upper limit on the visible cross section for new phenomena is set. A 95% C.L. lower limit of 32 TeV on the mGMSB breaking scale Λ is set for $M_{mess}=250$ TeV, $N_S=3$, $\mu>0$ and $C_{qrav}=1$, independent of $\tan\beta$.

 7 AAD 12AG searched in 2.05 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with at least one hadronically decaying tau lepton, jets, and large $\not\!\!E_T$ in a GMSB framework. No significant excess above the expected background was found and an upper limit on the visible cross section for new phenomena is set. A 95% C.L. lower limit of 30 TeV on

the mGMSB breaking scale Λ is set for $M_{mess}=250$ TeV, $N_S=3$, $\mu>0$ and $C_{grav}=1$, independent of $\tan\beta$. For large values of $\tan\beta$, the limit on Λ increases to 43 TeV.

- ⁸ AAD 12CM searched in 4.7 fb⁻¹ of pp collisions at \sqrt{s} =7 TeV for events with at least one tau lepton, zero or one additional light lepton (e/μ) jets, and large E_T in a GMSB framework. No significant excess above the expected background was found and upper limit on the visible cross section for new phenomena is set. A 95% C. L. lower limit of 54 TeV on the mGMSB breaking scale Λ is set for $M_{mess}=250$ TeV, $N_S=3$, $\mu>0$ and $C_{grav}=1$, for $\tan\beta>20$. Here the $\widetilde{\tau}_1$ is the NLSP.
- ⁹ ABBIENDI 06B use 600 pb⁻¹ of data from $\sqrt{s}=189$ –209 GeV. They look for events from pair-produced staus in a GMSB scenario with $\widetilde{\tau}$ NLSP including prompt $\widetilde{\tau}$ decays to ditaus + $\not\!\!\!E$ final states, large impact parameters, kinked tracks and heavy stable charged particles. Limits on the cross-section are computed as a function of m($\widetilde{\tau}$) and the lifetime, see their Fig. 7. The limit is compared to the $\sigma \cdot BR^2$ from a scan over the GMSB parameter space.
- 10 ABBIENDI 04F use data from $\sqrt{s}=189-209$ GeV. They derive limits on sparticle masses under the assumption of R with $LL\overline{E}$ or $LQ\overline{D}$ couplings. The results are valid for $\tan\beta=1.5,~\mu=-200$ GeV, with, in addition, $\Delta m>5$ GeV for indirect decays via $LQ\overline{D}$. The limit quoted applies to direct decays with $LL\overline{E}$ couplings and improves to 75 GeV for $LQ\overline{D}$ couplings. The limit on the $\widetilde{\tau}_R$ mass for indirect decays is 92 GeV for $LL\overline{E}$ couplings at $m_{\widetilde{\chi}0}=10$ GeV and no exclusion is obtained for $LQ\overline{D}$ couplings. Supersedes the results of ABBIENDI 00.
- ABDALLAH 04H use data from LEP 1 and $\sqrt{s}=192$ –208 GeV. They re-use results or re-analyze the data from ABDALLAH 03M to put limits on the parameter space of anomaly-mediated supersymmetry breaking (AMSB), which is scanned in the region $1 < m_{3/2} < 50$ TeV, $0 < m_0 < 1000$ GeV, $1.5 < \tan\beta < 35$, both signs of μ . The constraints are obtained from the searches for mass degenerate chargino and neutralino, for SM-like and invisible Higgs, for leptonically decaying charginos and from the limit on non-SM Z width of 3.2 MeV. The limit is for $m_t=174.3$ GeV (see Table 2 for other m_t values). The limit improves to 75 GeV for $\mu<0$.
- 12 ABDALLAH 04M use data from $\sqrt{s}=192-208$ GeV to derive limits on sparticle masses under the assumption of $I\!\!R$ with $LL\overline{E}$ couplings. The results are valid for $\mu=-200$ GeV, $\tan\beta=1.5,~\Delta m~>5$ GeV and assuming a BR of 1 for the given decay. The limit quoted is for indirect decays using the neutralino constraint of 39.5 GeV, also derived in ABDALLAH 04M. For indirect decays via $LL\overline{E}$ the limit decreases to 86 GeV if the constraint from the neutralino is not used. Supersedes the result of ABREU 00U.

Degenerate Charged Sleptons

Unless stated otherwise in the comment lines or in the footnotes, the following limits assume 3 families of degenerate charged sleptons.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>93	95	¹ BARATE	01	ALEP	$\Delta m > 10$ GeV, $\widetilde{\ell}_R^+ \widetilde{\ell}_R^-$
>70	95	¹ BARATE	01	ALEP	all Δm , $\widetilde{\ell}_R^+ \widetilde{\ell}_R^-$
• • • We do not use the	following	data for averages	s, fits,	limits, e	etc. • • •
>91.9	95	² ABBIENDI	06 B	OPAL	$\widetilde{\ell}_R ightarrow \ \ell \widetilde{G}$, all $\ell(\widetilde{\ell}_R)$
>88		³ ABDALLAH			$\widetilde{\ell}_R o \ \ell \widetilde{G}$, all $\ell(\widetilde{\ell}_R)$
>82.7	95	⁴ ACHARD	02	L3	$\widetilde{\ell}_{R}$, R decays,
>83	95	⁵ ABBIENDI	01		$e^+e^- \rightarrow \widetilde{\ell}_1\widetilde{\ell}_1,$ GMSB. $\tan\beta=2$
		⁶ ABREU	01	DLPH	GMSB, $tan \beta = 2$ $\widetilde{\ell} \rightarrow \ell \widetilde{\chi}_{2}^{0}, \ \widetilde{\chi}_{2}^{0} \rightarrow \gamma \widetilde{\chi}_{1}^{0},$
>68.8 >84	95 95	⁷ ACCIARRI ⁸ ABREU	01 00∨	L3 DLPH	$\ell=e,\mu$ $\widetilde{\ell}_{R},\mathcal{R}, 0.7 \leq \tan\beta \leq 40$ $\widetilde{\ell}_{R}\widetilde{\ell}_{R}(\widetilde{\ell}_{R} \to \ell \widetilde{G}),$ $m\widetilde{G} > 9 \text{ eV}$

- 1 BARATE 01 looked for acoplanar dilepton $+ \not\!\! E_T$ and single electron (for $\widetilde e_R \widetilde e_L)$ final states at 189 to 202 GeV. The limit assumes $\mu = -200$ GeV and $\tan\beta = 2$ for the production cross section and decay branching ratios, evaluated within the MSSM, and zero efficiency for decays other than $\widetilde\ell \to \ell \widetilde\chi_1^0$. The slepton masses are determined from the GUT relations without stau mixing. See their Fig. 1 for the dependence of the limit on Δm .
- ³ ABDALLAH 03D use data from $\sqrt{s}=130$ –208 GeV to search for tracks with large impact parameter or visible decay vertices and for heavy charged stable particles. Limits are obtained as function of $m(\widetilde{G})$, after combining these results with the search for slepton pair production in the SUGRA framework from ABDALLAH 03M to cover prompt decays. The above limit is reached for prompt decays and assumes the degeneracy of the sleptons. For limits at different $m(\widetilde{G})$, see their Fig. 9. Supersedes the results of ABREU 01G.
- ⁴ ACHARD 02 searches for the production of sparticles in the case of $\not R$ prompt decays with $LL\overline{E}$ or \overline{UDD} couplings at \sqrt{s} =189–208 GeV. The search is performed for direct and indirect decays, assuming one coupling at the time to be nonzero. The MSUGRA limit results from a scan over the MSSM parameter space with the assumption of gaugino and scalar mass unification at the GUT scale and no mixing in the slepton sector, imposing simultaneously the exclusions from neutralino, chargino, sleptons, and squarks analyses. The limit holds for $LL\overline{E}$ couplings and increases to 88.7 GeV for \overline{UDD} couplings. For L3 limits from $LQ\overline{D}$ couplings, see ACCIARRI 01.
- ⁵ ABBIENDI 01 looked for final states with $\gamma\gamma E$, $\ell\ell E$, with possibly additional activity and four leptons + E to search for prompt decays of $\widetilde{\chi}_1^0$ or $\widetilde{\ell}_1$ in GMSB. They derive limits in the plane $(m_{\widetilde{\chi}_1^0}, m_{\widetilde{\tau}_1})$, see Fig. 6, allowing either the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}_1$ to be the NLSP. Two scenarios are considered: $\tan\beta{=}2$ with the 3 sleptons degenerate in mass and $\tan\beta{=}20$ where the $\widetilde{\tau}_1$ is lighter than the other sleptons. Data taken at $\sqrt{s}{=}189~{\rm GeV}$. For $\tan\beta{=}20$, the obtained limits are $m_{\widetilde{\tau}_1}>69~{\rm GeV}$ and $m_{\widetilde{e}_1,\widetilde{\mu}_1}>88~{\rm GeV}$.
- ⁶ ABREU 01 looked for acoplanar dilepton + diphoton + $\not\!\!E$ final states from $\widetilde{\ell}$ cascade decays at \sqrt{s} =130–189 GeV. See Fig. 9 for limits on the (μ,M_2) plane for $m_{\widetilde{\ell}}$ =80 GeV, $\tan\beta$ =1.0, and assuming degeneracy of $\widetilde{\mu}$ and \widetilde{e} .
- ⁷ ACCIARRI 01 searches for multi-lepton and/or multi-jet final states from R prompt decays with $LL\overline{E}$, $LQ\overline{D}$, or \overline{UDD} couplings at $\sqrt{s}{=}189$ GeV. The search is performed for direct and indirect decays of neutralinos, charginos, and scalar leptons, with the $\widetilde{\chi}_1^0$ or a $\widetilde{\ell}$ as LSP and assuming one coupling to be nonzero at a time. Mass limits are derived using simultaneously the constraints from the neutralino, chargino, and slepton analyses; and the Z^0 width measurements from ACCIARRI 00C in a scan of the parameter space assuming MSUGRA with gaugino and scalar mass universality. Updates and supersedes the results from ACCIARRI 99I.
- ⁸ ABREU 00v use data from $\sqrt{s}=130-189$ GeV to search for tracks with large impact parameter or visible decay vertices. Limits are obtained as function of $m_{\widetilde{G}}$, after combining these results with the search for slepton pair production in the SUGRA framework from ABREU 01 to cover prompt decays and on stable particle searches from ABREU 00Q. For limits at different $m_{\widetilde{G}}$, see their Fig. 12. The above limit assumes the degeneracy of stau and smuon.

Long-lived $\widetilde{\ell}$ (Slepton) MASS LIMIT

Limits on scalar leptons which leave detector before decaying. Limits from Z decays are independent of lepton flavor. Limits from continuum e^+e^- annihilation are also independent of flavor for smuons and staus. Selectron limits from e^+e^- collisions in the continuum depend on MSSM parameters because of the additional neutralino exchange contribution.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>490	95	¹ KHACHATRY	16 BV	vCMS	long-lived $\widetilde{ au}$ from inclusive production, mGMSB SPS line 7
>240	95	¹ KHACHATRY.	16 BV	vCMS	scenario long-lived $\widetilde{\tau}$ from direct pair production, mGMSB SPS line 7
>440	95	² AAD	15 AE	ATLS	scenario mGMSB, $M_{mess} = 250$ TeV, $N_5 = 3$, $\mu > 0$, $C_{grav} = 5000$,
>385	95	² AAD	15 AE	ATLS	$\begin{array}{l} \tan\beta = 10 \\ \mathrm{mGMSB}, \ M_{mess} = 250 \ \mathrm{TeV}, \ N_5 \\ = 3, \ \mu \ > 0, \ C_{grav} = 5000, \end{array}$
>286	95	² AAD	15.5	ATLS	an eta = 50 direct $\widetilde{ au}$ production
none 124–309	95 95	³ AAIJ	-	LHCB	long-lived $\widetilde{\tau}$, mGMSB, SPS7
> 98	95 95	⁴ ABBIENDI		OPAL	$\widetilde{\mu}_R$, $\widetilde{\tau}_R$
none 2–87.5	95 95	⁵ ABREU		DLPH	
> 81.2	95 95	⁶ ACCIARRI	99H		$\widetilde{\mu}_{R}$, $\widetilde{\tau}_{R}$
* -		⁷ BARATE		-	$\widetilde{\widetilde{\mu}}_{R}$, $\widetilde{\widetilde{\tau}}_{R}$
> 81	95			ALEP	$\widetilde{\mu}_{R}$, $\widetilde{\tau}_{R}$
● ● ● We do n	ot use	the following data fo	r aver	rages, fit	s, limits, etc. • • •
>300	95	⁸ AAD	13AA	ATLS	long-lived $\widetilde{ au}$, GMSB, $ an\!eta=5$ –20
		⁹ ABAZOV	13 B	D0	long-lived $\widetilde{ au}$, 100 $<$ $m_{\widetilde{ au}}$ $<$ 300 GeV
>339	95	^{10,11} CHATRCHYAN	I 13 AB	CMS	long-lived $\widetilde{\tau}$, direct $\widetilde{\tau}_1$ pair prod.,
					minimal GMSB, SPS line 7
>500	95	^{10,12} CHATRCHYAN	l 13 AB	CMS	long-lived $\tilde{\tau}$, $\tilde{\tau}_1$ from direct pair prod. and from decay of heavier SUSY particles, minimal
>314	95	¹³ CHATRCHYAN	l 12L	CMS	GMSB, SPS line 7 long-lived $\tilde{\tau}$, $\tilde{\tau}_1$ from decay of heavier SUSY particles, minimal GMSB, SPS line 7
>136	95	¹⁴ AAD	11P	ATLS	stable $\tilde{\tau}$, GMSB scenario, $\tan \beta = 5$

 $^{^1}$ KHACHATRYAN 16BW searched in 2.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with heavy stable charged particles, identified by their anomalously high energy deposits in the silicon tracker and/or long time-of-flight measurements by the muon system. No evidence for an excess over the expected background is observed. Limits are derived for pair production of tau sleptons as a function of mass, depending on their direct or inclusive production in a minimal GMSB scenario along the Snowmass Points and Slopes (SPS) line 7, see Fig. 4 and Table 7.

 $^{^2}$ AAD $^{'}$ 15AE searched in 19.1 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ALTAS muon system. In the absence of an excess of events above the expected backgrounds, limits are set on stable $\widetilde{\tau}$ sleptons in various scenarios, see Figs. 5-7.

 $^{^3}$ AAIJ 15BD searched in 3.0 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ and 8 TeV for evidence of Drell-Yan pair production of long-lived $\widetilde{\tau}$ particles. No evidence for such particles is observed and 95% C.L. upper limits on the cross section of $\widetilde{\tau}$ pair production are derived, see Fig. 7. In the mGMSB, assuming the SPS7 benchmark scenario $\widetilde{\tau}$ masses between 124 and 309 GeV are excluded at 95% C.L.

- ⁴ ABBIENDI 03L used e^+e^- data at $\sqrt{s}=130$ –209 GeV to select events with two high momentum tracks with anomalous dE/dx. The excluded cross section is compared to the theoretical expectation as a function of the heavy particle mass in their Fig. 3. The limit improves to 98.5 GeV for $\widetilde{\mu}_L$ and $\widetilde{\tau}_L$. The bounds are valid for colorless spin 0 particles with lifetimes longer than 10^{-6} s. Supersedes the results from ACKERSTAFF 98P.
- ⁵ ABREU 00Q searches for the production of pairs of heavy, charged stable particles in e^+e^- annihilation at \sqrt{s} = 130–189 GeV. The upper bound improves to 88 GeV for $\widetilde{\mu}_L$, $\widetilde{\tau}_I$. These limits include and update the results of ABREU 98P.
- ⁶ ACCIARRI 99H searched for production of pairs of back-to-back heavy charged particles at \sqrt{s} =130–183 GeV. The upper bound improves to 82.2 GeV for $\widetilde{\mu}_I$, $\widetilde{\tau}_I$.
- ⁷ The BARATE 98K mass limit improves to 82 GeV for $\widetilde{\mu}_L$, $\widetilde{\tau}_L$. Data collected at \sqrt{s} =161–184 GeV.
- 8 AAD 13AA searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events containing long-lived massive particles in a GMSB framework. No significant excess above the expected background was found. A 95% C.L. lower limit of 300 GeV is placed on long-lived $\widetilde{\tau}$'s in the GMSB model with $M_{mess}=250$ TeV, $N_S=3,\,\mu>0,$ for $\tan\beta=5-20.$ The lower limit on the GMSB breaking scale \varLambda was found to be 99–110 TeV, for $\tan\beta$ values between 5 and 40, see Fig. 4 (top). Also, directly produced long-lived sleptons, or sleptons decaying to long-lived ones, are excluded at 95% C.L. up to a $\widetilde{\tau}$ mass of 278 GeV for models with slepton splittings smaller than 50 GeV.
- ⁹ ABAZOV 13B looked in 6.3 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for charged massive long-lived particles in events with muon-like particles that have both speed and ionization energy loss inconsistent with muons produced in beam collisions. In the absence of an excess, limits are set at 95% C.L. on the production cross section of stau leptons in the mass range 100–300 GeV, see their Table 20 and Fig. 23.
- 10 CHATRCHYAN 13AB looked in 5.0 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV and in 18.8 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of $\widetilde{\tau}_1$'s. No evidence for an excess over the expected background is observed. Supersedes CHATRCHYAN 12L.
- 11 CHATRCHYAN 13AB limits are derived for pair production of $\widetilde{\tau}_1$ as a function of mass in minimal GMSB scenarios along the Snowmass Points and Slopes (SPS) line 7 (see Fig. 8 and Table 7). The limit given here is valid for direct pair $\widetilde{\tau}_1$ production.
- 12 CHATRCHYAN 13AB limits are derived for the production of $\widetilde{\tau}_1$ as a function of mass in minimal GMSB scenarios along the Snowmass Points and Slopes (SPS) line 7 (see Fig. 8 and Table 7). The limit given here is valid for the production of $\widetilde{\tau}_1$ from both direct pair production and from the decay of heavier supersymmetric particles.
- 13 CHATRCHYAN 12 L looked in $5.0~{\rm fb}^{-1}$ of pp collisions at $\sqrt{s}=7~{\rm TeV}$ for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of $\widetilde{\tau}_1$'s. No evidence for an excess over the expected background is observed. Limits are derived for the production of $\widetilde{\tau}_1$ as a function of mass in minimal GMSB scenarios along the Snowmass Points and Slopes (SPS) line 7 (see Fig. 3). The limit given here is valid for the production of $\widetilde{\tau}_1$ in the decay of heavier supersymmetric particles.
- 14 AAD 11P looked in $37~\text{pb}^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7~\text{TeV}$ for events with two heavy stable particles, reconstructed in the Inner tracker and the Muon System and identified by their time of flight in the Muon System. No evidence for an excess over the SM expectation is observed. Limits on the mass are derived, see Fig. 3, for $\tilde{\tau}$ in a GMSB scenario and for sleptons produced by electroweak processes only, in which case the limit degrades to 110~GeV.

q̃ (Squark) MASS LIMIT

For $m_{\widetilde{q}} >$ 60–70 GeV, it is expected that squarks would undergo a cascade decay via a number of neutralinos and/or charginos rather than undergo a direct decay to photinos as assumed by some papers. Limits obtained when direct decay is assumed are usually higher than limits when cascade decays are included.

Limits from e^+e^- collisions depend on the mixing angle of the lightest mass eigenstate $\widetilde{q}_1{=}\widetilde{q}_R{\sin}\theta_q{+}\widetilde{q}_L{\cos}\theta_q$. It is usually assumed that only the sbottom and stop squarks have non-trivial mixing angles (see the stop and sbottom sections). Here, unless otherwise noted, squarks are always taken to be either left/right degenerate, or purely of left or right type. Data from Z decays have set squark mass limits above 40 GeV, in the case of $\widetilde{q} \rightarrow q\widetilde{\chi}_1$ decays if $\Delta m{=}m_{\widetilde{q}}-m_{\widetilde{\chi}_1^0}{\gtrsim}5$ GeV. For smaller values of Δm , current constraints on the invisible width of the Z ($\Delta\Gamma_{\rm inv}<2.0$ MeV, LEP 00) exclude $m_{\widetilde{u}_L,R}<$ 44 GeV, $m_{\widetilde{d}_R}<$ 33 GeV, $m_{\widetilde{d}_L}<$ 44 GeV and, assuming all squarks degenerate, $m_{\widetilde{q}}<$ 45 GeV.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

<i>VALUE</i> (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
>1450 (CL =	95%) OI	JR EVALUATION	CMSSM, ta	$\sin \beta = 30$, $\mu > 0$
> 608–1260	$(CL = 95^\circ)$	%) OUR EVALUAT	FION Mass	degenerate squarks
> 490-600 (0	CL = 95%) OUR EVALUAT	ION Single	light squark bounds
> 608	95	¹ AABOUD	16D ATLS	≥ 1 jet $+ ot \!$
>1030	95	² AABOUD	16N ATLS	$=$ 5 GeV \geq 2 jets $+$ $ ot\!$
> 600	95	³ KHACHATRY.	16BS CMS	$egin{aligned} GeV \ jets + ot \!$
>1260	95	³ KHACHATRY.	16BS CMS	jets $+ E_T$, Tsqk1, 8 degenerate light squarks, $m_{\widetilde{\chi}^0_1} = 0$ GeV
> 850	95	⁴ AAD	15BV ATLS	jets $+ \not\!\!E_T$, $\widetilde{q} \to q \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} =$
> 250	95	⁵ AAD	15CS ATLS	100 GeV photon $+ E_T$, $pp \rightarrow \widetilde{q} \widetilde{q}^* \gamma$, $\widetilde{q} \rightarrow q \widetilde{\chi}_1^0$, $m_{\widetilde{q}} - m_{\widetilde{\chi}_1^0} = m_c$
> 490	95	⁶ AAD	15к ATLS	$\widetilde{c} \rightarrow c\widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}} < 200 \text{ GeV}$
> 875	95	⁷ KHACHATRY.		$\widetilde{q} ightarrow q \widetilde{\chi}_1^0$, simplified model, 8 degenerate light \widetilde{q} , $m_{\widetilde{\chi}_1^0} = 0$
> 520	95	⁷ KHACHATRY.	15AF CMS	$\widetilde{q} ightarrow q \widetilde{\chi}_1^0$, simplified model, single light squark, $m_{\widetilde{\chi}_1^0} = 0$
>1450	95	⁷ KHACHATRY.	15AF CMS	CMSSM, $tan\beta = 30$, $A_0 = -2max(m_0, m_{1/2})$, $\mu > 0$
> 850	95	⁸ AAD	14AE ATLS	$\begin{array}{l} {\rm jets} + E_T, \ \widetilde{q} \rightarrow \ q \ \widetilde{\chi}_1^0 \ {\rm simplified} \\ {\rm model, \ mass \ degenerate \ first} \\ {\rm and \ second \ generation \ squarks,} \\ m_{\widetilde{\chi}_1^0} = 0 \ {\rm GeV} \end{array}$

> 440	95	⁸ AAD	14 AE	ATLS	jets $+ \not\!\!E_T$, $\stackrel{.}{q} \to q \stackrel{.}{\chi}^0_1$ simplified model, single light-flavour squark, $m_{\stackrel{.}{\chi}^0_1} = 0$ GeV
>1700	95	⁸ AAD	14 AE	ATLS	
> 800	95	⁹ CHATRCHYAN	14 AH	CMS	$ \begin{array}{ccc} jets + \not\!\!\!E_T, \stackrel{\mathcal{B}}{q} \to q \widetilde{\chi}_1^0 simplified \\ model, m_{\widetilde{\chi}_1^0} = 50 GeV \end{array} $
> 780	95	¹⁰ CHATRCHYAN	141	CMS	multijets $+ \not\!\!E_T$, $\stackrel{\frown}{q} \rightarrow q \stackrel{\frown}{\chi}^0_1$ simplified model, $m_{\stackrel{\frown}{\chi}^0_1} < 200$
>1360	95	¹¹ AAD	13L	ATLS	GeV jets $+ \not\!\!E_T$, CMSSM, $m_{\widetilde{g}} = m_{\widetilde{q}}$
>1200	95	¹² AAD	13Q	ATLS	$\gamma+b+\cancel{E}_T$, higgsino-like neutralino, $m_{\widetilde{\chi}^0_1}>220$ GeV, GMSB
		¹³ CHATRCHYAN	13	CMS	$\ell^{\pm}\ell^{\mp}$ + jets + $ ot\!\!\!E_T$, CMSSM
>1250	95	¹⁴ CHATRCHYAN		CMS	$0.1.2. \geq 3$ $b ext{-jets} + ot \!$
>1430	95	¹⁵ CHATRCHYAN	13H	CMS	$2\gamma + \geq$ 4 jets $+$ low $ ot \!$
> 750	95	¹⁶ CHATRCHYAN	13T	CMS	jets $+ \not\!\!E_T$, $\widetilde{q} \to q \widetilde{\chi}_1^0$ simplified
					model, $m_{\widetilde{\chi}_1^0} = 0 \text{ GeV}$
> 820	95	¹⁷ AAD	12AX	ATLS	ℓ +jets + $\not\!\!E_T$, CMSSM, $m_{\widetilde{q}} = m_{\widetilde{g}}$
>1200	95	¹⁸ AAD	12CJ	ATLS	ℓ^{\pm} +jets+ $\not\!\!E_T$, CMSSM, $m_{\widetilde{q}} = m_{\widetilde{g}}$
> 870	95	¹⁹ AAD		ATLS	$2\gamma + \cancel{E}_T$, GMSB, bino NLSP,
					$m_{\widetilde{\chi}_1^0} > 50 \text{ GeV}$
> 950	95	²⁰ AAD		ATLS	jets $+ ot \!$
		²¹ CHATRCHYAN		CMS	e, μ , jets, razor, CMSSM
> 760	95	²² CHATRCHYAN	12AE	CMS	jets $+ ot \!\!\!\!E_T$, $\widetilde{q} ightarrow q \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} < \infty$
		²³ CHATRCHYAN	1241	CMC	200 GeV $\geq 3\ell^{\pm}$, R
>1110	95	²⁴ CHATRCHYAN	12AL 12ΔT	CMS	\geq 5 ℓ^- , $\not\!$
>1110	95	²⁴ CHATRCHYAN	12AT	CMS	$jets + \not\!\!\!E_T, CMSSM, m_{\widetilde{q}} = m_{\widetilde{g}}$
	ot use th	e following data for			
> 300	95	²⁵ KHACHATRY			19-parameter pMSSM model, global Bayesian analysis, flat
>1600	95	²⁶ KHACHATRY	. 16 BX	CMS	prior $\widetilde{q} \rightarrow q \widetilde{\chi}_1^0, \ \widetilde{\chi}_1^0 \rightarrow \ell\ell\nu, \ RPV, \ \lambda_{121} \ or \ \lambda_{122} \neq 0, \ m_{\widetilde{g}} = 2400 \ CeV$
	95	²⁷ AAD	15AI	ATLS	ℓ^{\pm} + jets + $ ot\!\!\!E_T$
>1650	95	⁴ AAD	15 BV	ATLS	jets $+ \not\!\!E_T$, $m_{\widetilde{g}} = m_{\widetilde{q}}$, $m_{\widetilde{\chi}_1^0} = 1$
					GeV $\frac{1}{2}$
> 790	95	⁴ AAD	15 BV	ATLS	$jets + \not\!\!E_T, \; \widetilde{q} \to \; q W \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} =$
> 820	95	⁴ AAD	15 BV	ATLS	100 GeV 2 or 3 leptons $+$ jets, \widetilde{q} decays via sleptons, $m_{\widetilde{\chi}_1^0}=100$ GeV
> 850	95	⁴ AAD	15 BV	ATLS	$ au$, \widetilde{q} decays via staus, $m_{\widetilde{\chi}_1^0} = 50$
					$GeV^{\overset{\chi_1}{}}$
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>1000	95	²⁸ AAD	15CB ATLS	jets, $\widetilde{q} ightarrow q \widetilde{\chi}_1^0$, $\widetilde{\chi}_1^0 ightarrow \ell q q$, RPV, $m_{\widetilde{\chi}_1^0} = 108$ GeV and $2.5 < c au_{\widetilde{\chi}_1^0} < 200$ mm
> 700	95	²⁹ KHACHATRY	15AR CMS	$2.5 < c au_{\widetilde{\chi}_1^0}^{-1} < 200 \; \mathrm{mm}$ $\widetilde{q} ightarrow q \widetilde{\chi}_1^0, \; \widetilde{\chi}_1^0 ightarrow \; \widetilde{S} g, \; \widetilde{S} ightarrow S \widetilde{G}, \; S ightarrow g g, \; m_{\widetilde{S}} = 100$ $\; \mathrm{GeV}, \; m_S = 90 \; \mathrm{GeV}$
> 550	95	²⁹ KHACHATRY	15AR CMS	ℓ^{\pm} , $\widetilde{q} \rightarrow q \widetilde{\chi}_{1}^{\pm}$, $\widetilde{\chi}_{1}^{\pm} \rightarrow \widetilde{S} W^{\pm}$, $\widetilde{S} \rightarrow S \widetilde{G}$, $S \rightarrow g g$, $m_{\widetilde{S}} =$
>1500	95	³⁰ KHACHATRY	15AZ CMS	100 GeV, $m_S = 90$ GeV $\geq 2 \ \gamma$, ≥ 1 jet, (Razor), binolike NLSP, $m_{\widetilde{\chi}_1^0} = 375$ GeV
>1000	95	³⁰ KHACHATRY	15AZ CMS	$\geq 1 \ \gamma$, $\geq 2 \ \mathrm{jet}$, wino-like NLSP, $m_{\widetilde{\chi}_1^0} = 375 \ \mathrm{GeV}$
> 670	95	³¹ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp}) + jets, \ \widetilde{q} ightarrow \ q' \widetilde{\chi}_1^{\pm},$
		04		$\widetilde{\chi}_1^{\pm} ightarrow W^{(*)\pm} \widetilde{\chi}_2^0, \ \widetilde{\chi}_2^0 ightarrow Z^{(*)} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} < 300 \text{ GeV}$
> 780	95	³¹ AAD	14E ATLS	$q'\widetilde{\chi}_1^{\pm}/\widetilde{\chi}_2^0, \widetilde{\chi}_1^{\pm} ightarrow \ell^{\pm} u \widetilde{\chi}_1^0, \ \widetilde{\chi}_2^0 ightarrow \ell^{\pm} \ell^{\mp} (u u) \widetilde{\chi}_1^0 ext{simpli-}$
> 700	95	³² CHATRCHYA	N 13AO CMS	fied model $\ell^{\pm}\ell^{\mp}+{ m jets}+E_T$, CMSSM, $m_0<700~{ m GeV}$
>1350	95	33 CHATRCHYA	N 13AV CMS	
> 800	95	³⁴ CHATRCHYA	N 13W CMS	$\mathbb{Z}_T = \mathbb{Z}_T$ photons \mathbb{Z}_T , GGM, wino-like NLSP, \mathbb{Z}_T
>1000	95	³⁴ CHATRCHYA	N 13w CMS	$= 375 \text{ GeV} \\ \geq 2 \text{ photons} + \text{jets} + \cancel{E}_T, \\ \text{GGM, bino-like NLSP, } m_{\widetilde{\chi}_1^0}$
> 340	95	³⁵ DREINER	12A THE	$=$ 375 GeV $m_{\widetilde{q}} \sim m_{\widetilde{\chi}_1^0}$
> 650	95	³⁶ DREINER	12A THE	O $m_{\widetilde{q}} = m_{\widetilde{g}}^{\lambda_1} \sim m_{\widetilde{\chi}_1^0}$

 $^{^1}$ AABOUD 16D searched in 3.2 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with an energetic jet and large missing transverse momentum. The results are interpreted as 95% C.L. limits on masses of first and second generation squarks decaying into a quark and the lightest neutralino in scenarios with $m_{\widetilde{q}}-m_{\widetilde{\chi}_1^0}<25$ GeV. See their Fig. 6.

 $^{^2}$ AABOUD 16N searched in 3.2 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events containing hadronic jets, large E_T , and no electrons or muons. No significant excess above the Standard Model expectations is observed. First- and second-generation squark masses below 1030 GeV are excluded at the 95% C.L. decaying to quarks and a massless lightest neutralino. See their Fig. 7a.

- 3 KHACHATRYAN 16BS searched in 2.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with at least one energetic jet , no isolated leptons, and significant $\not\!\!E_T$, using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the squark mass in the Tskq1 simplified model, both in the assumption of a single light squark and of 8 degenerate squarks, see Fig. 11 and Table 3.
- ⁴ AAD 15BV summarized and extended ATLAS searches for gluinos and first- and second-generation squarks in final states containing jets and missing transverse momentum, with or without leptons or *b*-jets in the $\sqrt{s}=8$ TeV data set collected in 2012. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on the squark mass in several R-parity conserving models. See their Figs. 9, 11, 18, 22, 24, 27, 28.
- 5 AAD 15CS searched in 20.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for evidence of pair production of squarks, decaying into a quark and a neutralino, where a photon was radiated either from an initial-state quark, from an intermediate squark, or from a final-state quark. No evidence was found for an excess above the expected level of Standard Model background and a 95% C.L. exclusion limit was set on the squark mass as a function of the squark-neutralino mass difference, see Fig. 19.
- ⁶ AAD 15K searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing at least two jets, where the two leading jets are each identified as originating from c-quarks, and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the mass of superpartners of charm quarks (\tilde{c}) . Assuming that the decay $\tilde{c} \to c \tilde{\chi}_1^0$ takes place 100% of the time, a scalar charm mass below 490 GeV is excluded for $m_{\tilde{\chi}_1^0} <$ 200 GeV. For more details, see their Fig. 2.
- 7 KHACHATRYAN 15AF searched in $19.5~{\rm fb}^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets and significant E_T , using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the squark mass in simplified models where the decay $\widetilde{q}\to q\widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, both for the case of a single light squark or 8 degenerate squarks, see Fig. 12. See also Table 5. Exclusions in the CMSSM, assuming $\tan\beta=30,\,A_0=-2~{\rm max}(m_0,m_{1/2})$ and $\mu>0$, are also presented, see Fig. 15.
- ⁸ AAD 14AE searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for strongly produced supersymmetric particles in events containing jets and large missing transverse momentum, and no electrons or muons. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing squarks that decay via $\tilde{q} \rightarrow q \tilde{\chi}_1^0$, where either a single light state or two degenerate generations of squarks are assumed, see Fig. 10.
- ⁹ CHATRCHYAN 14AH searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with at least two energetic jets and significant $\not\!\!E_T$, using the razor variables (M_R and R^2) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on squark masses in simplified models where the decay $\widetilde{q} \to q \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 28. Exclusions in the CMSSM, assuming $\tan\beta=10$, $A_0=0$ and $\mu>0$, are also presented, see Fig. 26.
- 10 CHATRCHYAN 14I searched in 19.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing multijets and large E_T . No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing squarks that decay via $\widetilde{q} \to q\,\widetilde{\chi}_1^0$, where either a single light state or two degenerate generations of squarks are assumed, see Fig. 7a.
- ¹¹ AAD 13L searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for the production of squarks and gluinos in events containing jets, missing transverse momentum and no high- p_T electrons or muons. No excess over the expected SM background is observed. In

- mSUGRA/CMSSM models with $\tan\beta=10$, $A_0=0$ and $\mu>0$, squarks and gluinos of equal mass are excluded for masses below 1360 GeV at 95% C.L. In a simplified model containing only squarks of the first two generations, a gluino octet and a massless neutralino, squark masses below 1320 GeV are excluded at 95% C.L. for gluino masses below 2 TeV. See Figures 10–15 for more precise bounds.
- 12 AAD 13Q searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events containing a high- p_T isolated photon, at least one jet identified as originating from a bottom quark, and high missing transverse momentum. Such signatures may originate from supersymmetric models with gauge-mediated supersymmetry breaking in events in which one of a pair of higgsino-like neutralinos decays into a photon and a gravitino while the other decays into a Higgs boson and a gravitino. No significant excess above the expected background was found and limits were set on the squark mass as a function of the neutralino mass in a generalized GMSB model (GGM) with a higgsino-like neutralino NLSP, see their Fig. 4. For neutralino masses greater than 220 GeV, squark masses below 1020 GeV are excluded at 95% C.L.
- 13 CHATRCHYAN 13 looked in 4.98 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with two opposite-sign leptons (e, $\mu,~\tau$), jets and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in the mSUGRA/CMSSM model with tan $\beta=10,~A_0=0$ and $\mu>0$, see Fig. 6.
- 14 CHATRCHYAN 13 G searched in 4.98 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for the production of squarks and gluinos in events containing 0,1,2, $\,\geq 3$ b-jets, missing transverse momentum and no electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with $\tan\!\beta=10,\,A_0=0,\,$ and $\mu>0,\,$ squarks and gluinos of equal mass are excluded for masses below 1250 GeV at 95% C.L. Exclusions are also derived in various simplified models, see Fig. 7.
- 15 CHATRCHYAN 13H searched in 4.96 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events with two photons, \geq 4 jets and low E_T due to $\widetilde{q}\to\gamma\widetilde{\chi}^0_1$ decays in a stealth SUSY framework, where the $\widetilde{\chi}^0_1$ decays through a singlino (\widetilde{S}) intermediate state to $\gamma\,S\,\widetilde{G}$, with the singlet state S decaying to two jets. No significant excess above the expected background was found and limits were set in a particular R-parity conserving stealth SUSY model. The model assumes $m_{\widetilde{\chi}^0_1}=0.5$ $m_{\widetilde{q}}$, $m_{\widetilde{S}}=100$ GeV and $m_{\widetilde{S}}=90$ GeV.
 - Under these assumptions, squark masses less than 1430 GeV were excluded at the 95% C.L.
- CHATRCHYAN 13T searched in 11.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets and significant $\not\!\!E_T$, using the α_T variable to discriminate between processes with genuine and misreconstructed $\not\!\!E_T$. No significant excess above the Standard Model expectations is observed. Limits are set on squark masses in simplified models where the decay $q \to q \chi_1^0$ takes place with a branching ratio of 100%, assuming an eightfold degeneracy of the masses of the first two generation squarks, see Fig. 8 and Table 9. Also limits in the case of a single light squark are given.
- 17 AAD 12AX searched in 1.04 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for supersymmetry in events containing jets, missing transverse momentum and one isolated electron or muon. No excess over the expected SM background is observed and model-independent limits are set on the cross section of new physics contributions to the signal regions. In mSUGRA/CMSSM models with $\tan\beta=10,\ A_0=0$ and $\mu>0$, squarks and gluinos of equal mass are excluded for masses below 820 GeV at 95% C.L. Limits are also set on simplified models for squark production and decay via an intermediate chargino and on supersymmetric models with bilinear R-parity violation. Supersedes AAD 11G.
- 18 AAD 12 CJ searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events containing one or more isolated leptons (electrons or muons), jets and E_T . The observations are in good agreement with the SM expectations and exclusion limits have been set in number of SUSY models. In the mSUGRA/CMSSM model with $\tan\beta=10,\,A_0=0,\,$ and $\mu>0,\,$ 95% C.L. exclusion limits have been derived for $m_{\widetilde{q}}<1200$ GeV, assuming equal squark and gluino masses. In minimal GMSB, values of the effective SUSY breaking scale $\Lambda<50$ TeV are excluded at 95% C.L. for $\tan\beta<45$. Also exclusion limits in a number of simplified models have been presented, see Figs. 10 and 12.

- 19 AAD 12CP searched in 4.8 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with two photons and large E_T due to $\widetilde{\chi}^0_1\to \gamma\,\widetilde{G}$ decays in a GMSB framework. No significant excess above the expected background was found and limits were set on the squark mass as a function of the neutralino mass in a generalized GMSB model (GGM) with a bino-like neutralino NLSP. The other sparticle masses were decoupled, $\tan\beta=2$ and $c\tau_{NLSP}<0.1$ mm. Also, in the framework of the SPS8 model, a 95% C.L. lower limit was set on the breaking scale Λ of 196 TeV.
- 20 AAD 12 W searched in $1.04~{\rm fb}^{-1}$ of pp collisions at $\sqrt{s}=7~{\rm TeV}$ for the production of squarks and gluinos in events containing jets, missing transverse momentum and no electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with $\tan\beta=10,\ A_0=0$ and $\mu>0$, squarks and gluinos of equal mass are excluded for masses below 950 GeV at 95% C.L. In a simplified model containing only squarks of the first two generations, a gluino octet and a massless neutralino, squark masses below 875 GeV are excluded at 95% C.L.
- 21 CHATRCHYAN 12 looked in 35 pb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events with e and/or μ and/or jets, a large total transverse energy, and E_T . The event selection is based on the dimensionless razor variable R, related to the E_T and M_R , an indicator of the heavy particle mass scale. No evidence for an excess over the expected background is observed. Limits are derived in the CMSSM $(m_0,\,m_{1/2})$ plane for $\tan\beta=3,\,10$ and 50 (see Fig. 7 and 8). Limits are also obtained for Simplified Model Spectra.
- 22 CHATRCHYAN 12AE searched in 4.98 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events with at least three jets and large missing transverse momentum. No significant excesses over the expected SM backgrounds are observed and 95% C.L. limits on the production cross section of squarks in a scenario where $\widetilde{q}\to q\widetilde{\chi}_1^0$ with a 100% branching ratio, see Fig. 3. For $m_{\widetilde{\chi}_1^0}<200$ GeV, values of $m_{\widetilde{q}}$ below 760 GeV are excluded at 95% C.L. Also limits in the CMSSM are presented, see Fig. 2.
- 23 CHATRCHYAN 12AL looked in 4.98 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for anomalous production of events with three or more isolated leptons. Limits on squark and gluino masses are set in R SUSY models with leptonic $LL\overline{E}$ couplings, $\lambda_{123} > 0.05$, and hadronic \overline{UDD} couplings, $\lambda_{112}'' > 0.05$, see their Fig. 5. In the \overline{UDD} case the leptons arise from supersymmetric cascade decays. A very specific supersymmetric spectrum is assumed. All decays are prompt.
- 24 CHATRCHYAN 12 AT searched in 4.73 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for the production of squarks and gluinos in events containing jets, missing transverse momentum and no electrons or muons. No excess over the expected SM background is observed. In mSUGRA/CMSSM models with $\tan\beta=10,\ A_0=0$ and $\mu>0,$ squarks with masses below 1110 GeV are excluded at 95% C.L. Squarks and gluinos of equal mass are excluded for masses below 1180 GeV at 95% C.L. Exclusions are also derived in various simplified models, see Fig. 6.
- 25 KHACHATRYAN 16 BT performed a global Bayesian analysis of a wide range of CMS results obtained with data samples corresponding to $5.0~{\rm fb}^{-1}$ of pp collisions at $\sqrt{s}=7~{\rm TeV}$ and in $19.5~{\rm fb}^{-1}$ of pp collisions at $\sqrt{s}=8~{\rm TeV}$. The set of searches considered, both individually and in combination, includes those with all-hadronic final states, samesign and opposite-sign dileptons, and multi-lepton final states. An interpretation was given in a scan of the 19-parameter pMSSM. No scan points with a gluino mass less than 500 GeV survived and 98% of models with a squark mass less than 300 GeV were excluded.
- 26 KHACHATRYAN 16BX searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing 4 leptons coming from R-parity-violating decays of $\widetilde{\chi}_1^0 \to \ell\ell\nu$ with $\lambda_{121} \neq 0$ or $\lambda_{122} \neq 0$. No excess over the expected background is observed. Limits are derived on the gluino, squark and stop masses, see Fig. 23.
- on the gluino, squark and stop masses, see Fig. 23.
 27 AAD 15AI searched in 20 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing at least one isolated lepton (electron or muon), jets, and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the squark masses in the

- CMSSM/mSUGRA, see Fig. 15, in the NUHMG, see Fig. 16, and in various simplified models, see Figs. 19-21.
- 28 AAD 15CB searched for events containing at least one long-lived particle that decays at a significant distance from its production point (displaced vertex, DV) into two leptons or into five or more charged particles in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV. The dilepton signature is characterised by DV formed from at least two lepton candidates. Four different final states were considered for the multitrack signature, in which the DV must be accompanied by a high-transverse momentum muon or electron candidate that originates from the DV, jets or missing transverse momentum. No events were observed in any of the signal regions. Results were interpreted in SUSY scenarios involving *R*-parity violation, split supersymmetry, and gauge mediation. See their Fig. 14–20.
- ²⁹ KHACHATRYAN 15AR searched in 19.7 of fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing jets, either a charged lepton or a photon, and low missing transverse momentum. No significant excess above the Standard Model expectations is observed. Limits are set on the squark mass in a stealth SUSY model where the decays $\widetilde{q} \rightarrow q \widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_1^{\pm} \rightarrow \widetilde{S} W^{\pm}$, $\widetilde{S} \rightarrow S \widetilde{G}$ and $S \rightarrow g g$, with $m_{\widetilde{S}}=100$ GeV and $m_{\widetilde{S}}=90$ GeV, take
- place with a branching ratio of 100%. See Fig. 6 for γ or Fig. 7 for ℓ^\pm analyses. 30 KHACHATRYAN 15AZ searched in 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with either at least one photon, hadronic jets and E_T (single photon channel) or with at least two photons and at least one jet and using the razor variables. No significant excess above the Standard Model expectations is observed. Limits are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for both a bino-like and wino-like neutralino NLSP scenario, see Fig. 8 and 9.
- 31 AAD 14E searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for strongly produced supersymmetric particles in events containing jets and two same-sign leptons or three leptons. The search also utilises jets originating from b-quarks, missing transverse momentum and other variables. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos and squarks, see Figures 5 and 6. In the $\tilde{q} \to q' \tilde{\chi}_1^\pm$, $\tilde{\chi}_1^\pm \to W^{(*)\pm} \tilde{\chi}_2^0$, $\tilde{\chi}_2^0 \to Z^{(*)} \tilde{\chi}_1^0$ simplified model, the following assumptions have been made: $m_{\tilde{\chi}_1^\pm} = 0.5 \ m_{\tilde{\chi}_1^0} + m_{\tilde{g}}$, $m_{\tilde{\chi}_2^0} = 0.5$ ($m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_1^0} + m_{\tilde$
- 32 CHATRCHYAN 2013AO searched in 4.98 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with two opposite-sign isolated leptons accompanied by hadronic jets and E_T . No significant excesses over the expected SM backgrounds are observed and 95% C.L. exclusion limits are derived in the mSUGRA/CMSSM model with tan $\beta=10,\ A_0=0$ and $\mu>0,$ see Fig. 8.
- 33 CHATRCHYAN 13 AV searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for new heavy particle pairs decaying into jets (possibly b-tagged), leptons and E_T using the Razor variables. No significant excesses over the expected SM backgrounds are observed and 95% C.L. exclusion limits are derived in the mSUGRA/CMSSM model with $\tan\beta=10,\ A_0=0$ and $\mu>0$, see Fig. 3. The results are also interpreted in various simplified models, see Fig. 4.
- 34 CHATRCHYAN 13W searched in 4.93 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with one or more photons, hadronic jets and $\not\!\!E_T$. No significant excess above the Standard Model expectations is observed. Limits are set on squark masses in the general gauge-mediated SUSY breaking model (GGM), for both a wino-like and bino-like neutralino NLSP scenario, see Fig. 5.
- 35 DREINER 12A reassesses constraints from CMS (at 7 TeV, \sim 4.4 fb $^{-1}$) under the assumption that the fist and second generation squarks and the lightest SUSY particle are quasi-degenerate in mass (compressed spectrum).

 $^{36}\,\text{DREINER}$ 12A reassesses constraints from CMS (at 7 TeV, \sim 4.4 fb $^{-1})$ under the assumption that the first and second generation squarks, the gluino, and the lightest SUSY particle are quasi-degenerate in mass (compressed spectrum).

Long-lived \tilde{q} (Squark) MASS LIMIT

The following are bounds on long-lived scalar quarks, assumed to hadronise into hadrons with lifetime long enough to escape the detector prior to a possible decay. Limits may depend on the mixing angle of mass eigenstates: $\widetilde{q}_1 = \widetilde{q}_L \cos\theta_q + \widetilde{q}_R \sin\theta_q$.

The coupling to the Z^0 boson vanishes for up-type squarks when θ_u =0.98, and for down type squarks when θ_d =1.17.

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
> 805	95	¹ AABOUD	16 B	ATLS	\widetilde{b} <i>R</i> -hadrons
> 890	95	² AABOUD	16 B	ATLS	\widetilde{t} R -hadrons
>1040	95	³ KHACHATRY.	16 BW	/CMS	\widetilde{t} R-hadrons, cloud interaction
>1000	95	³ KHACHATRY.	16 BW	CMS	model \tilde{t} R-hadrons, charge-suppressed interaction model
> 845	95	⁴ AAD	15AE	ATLS	\widetilde{b} R-hadron, stable, Regge model
> 900	95	⁴ AAD	15AE	ATLS	\widetilde{t} R-hadron, stable, Regge model
>1500	95	⁴ AAD	15 AE	ATLS	\widetilde{g} decaying to 300 GeV stable sleptons, LeptoSUSY model
> 751	95	⁵ AAD	15 BM	1ATLS	\widetilde{b} R-hadron, stable, Regge model
> 766	95	⁵ AAD	15 BM	1ATLS	\widetilde{t} R-hadron, stable, Regge model
> 525	95	⁶ KHACHATRY.	15 AK	CMS	\widetilde{t} R-hadrons, 10 μ s $< au <$ 1000 s
> 470	95	⁶ KHACHATRY.	15 AK	CMS	\widetilde{t} R-hadrons, 1 μ s< $ au$ <1000 s
• • • We do no	ot use the	e following data for	aver	ages, fits	s, limits, etc. • • •
> 683	95	⁷ AAD	13AA	ATLS	\tilde{t} , R-hadrons, generic interaction model
> 612	95	⁸ AAD	13AA	ATLS	\widetilde{b} , R -hadrons, generic interaction model
> 344	95	⁹ AAD	13 BC	ATLS	R-hadrons, $\widetilde{t} \to b\widetilde{\chi}_1^0$, Regge
					model, lifetime between 10^{-5} and 10^3 s, $m_{\widetilde{\chi}^0_1}=100$ GeV
> 379	95	¹⁰ AAD	13 BC	ATLS	R-hadrons, $\widetilde{t} ightarrow t \widetilde{\chi}_1^0$, Regge
					model, lifetime between 10^{-5} and 10^3 s, $m_{\widetilde{\chi}^0_1}=100~{ m GeV}$
> 935	95	¹¹ CHATRCHYAN	I 13 AB	CMS	long-lived \widetilde{t} forming R-hadrons, cloud interaction model

 $^{^1}$ AABOUD 16B searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for long-lived R-hadrons using observables related to large ionization losses and slow propagation velocities, which are signatures of heavy charged particles traveling significantly slower than the speed of light. Exclusion limits at 95% C.L. are set on the long-lived sbottom masses exceeding 805 GeV. See their Fig. 5.

 $^{^2}$ AABOUD 16B searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for long-lived R-hadrons using observables related to large ionization losses and slow propagation velocities, which are signatures of heavy charged particles traveling significantly slower than the speed of light. Exclusion limits at 95% C.L. are set on the long-lived stop masses exceeding 890 GeV. See their Fig. 5.

- 3 KHACHATRYAN 16BW searched in 2.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with heavy stable charged particles, identified by their anomalously high energy deposits in the silicon tracker and/or long time-of-flight measurements by the muon system. No evidence for an excess over the expected background is observed. Limits are derived for pair production of top squarks as a function of mass, depending on the interaction model, see Fig. 4 and Table 7.
- ⁴ AAD 15AE searched in 19.1 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ALTAS muon system. In the absence of an excess of events above the expected backgrounds, limits are set R-hadrons in various scenarios, see Fig. 11. Limits are also set in LeptoSUSY models where the gluino decays to stable 300 GeV leptons, see Fig. 9.
- 5 AAD 15BM searched in 18.4 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for stable and metastable non-relativistic charged particles through their anomalous specific ionization energy loss in the ATLAS pixel detector. In absence of an excess of events above the expected backgrounds, limits are set on stable bottom and top squark R-hadrons, see Table 5.
- 6 KHACHATRYAN 15AK looked in a data set corresponding to ${\rm fb}^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV, and a search interval corresponding to 281 h of trigger lifetime, for long-lived particles that have stopped in the CMS detector. No evidence for an excess over the expected background in a cloud interaction model is observed. Assuming the decay $\widetilde{t}\to t\widetilde{\chi}^0_1$ and lifetimes between 1 $\mu{\rm s}$ and 1000 s, limits are derived on \widetilde{t} production as a function of $m_{\widetilde{\chi}^0_1}$, see Figs. 4 and 7. The exclusions require that $m_{\widetilde{\chi}^0_1}$ is kinematically consistent with the minimum values of the jet energy thresholds used.
- ⁷ AAD 13AA searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing colored long-lived particles that hadronize forming R-hadrons. No significant excess above the expected background was found. Long-lived R-hadrons containing a \tilde{t} are excluded for masses up to 683 GeV at 95% C.L in a general interaction model. Also, limits independent of the fraction of R-hadrons that arrive charged in the muon system were derived, see Fig. 6.
- ⁸ AAD 13AA searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing colored long-lived particles that hadronize forming R-hadrons. No significant excess above the expected background was found. Long-lived R-hadrons containing a \tilde{b} are excluded for masses up to 612 GeV at 95% C.L in a general interaction model. Also, limits independent of the fraction of R-hadrons that arrive charged in the muon system were derived, see Fig. 6.
- ⁹AAD 13BC searched in 5.0 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV and in 22.9 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for bottom squark R-hadrons that have come to rest within the ATLAS calorimeter and decay at some later time to hadronic jets and a neutralino. In absence of an excess of events above the expected backgrounds, limits are set on sbottom masses for the decay $\tilde{b} \to b \tilde{\chi}_1^0$, for different lifetimes, and for a neutralino mass of 100 GeV, see their Table 6 and Fig 10.
- 10 AAD 13BC searched in 5.0 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV and in 22.9 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for bottom squark R-hadrons that have come to rest within the ATLAS calorimeter and decay at some later time to hadronic jets and a neutralino. In absence of an excess of events above the expected backgrounds, limits are set on stop masses for the decay $\tilde{t}\to t\tilde{\chi}_1^0$, for different lifetimes, and for a neutralino mass of 100 GeV, see their Table 6 and Fig 10.
- 11 CHATRCHYAN 13AB looked in 5.0 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV and in 18.8 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of \tilde{t}_1 's. No evidence for an excess over the expected background is observed. Limits are derived for pair production of stops as a function of mass in the cloud interaction model (see Fig. 8 and Table 6). In the charge-suppressed model, the limit decreases to 818 GeV.

\tilde{b} (Sbottom) MASS LIMIT

Limits in e^+e^- depend on the mixing angle of the mass eigenstate $\widetilde{b}_1=\widetilde{b}_L\cos\theta_b+\widetilde{b}_R\sin\theta_b$. Coupling to the Z vanishes for $\theta_b\sim 1.17$. As a consequence, no absolute constraint in the mass region \lesssim 40 GeV is available in the literature at this time from e^+e^- collisions. In the Listings below, we use $\Delta m=m_{\widetilde{b}_1}-m_{\widetilde{\chi}_1^0}$.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 323–880	(CL =	95%) OUR EVALUATIO	ON Depe	ndent on mass difference $\widetilde{b} extsf{-}LSP$
>315	95	¹ KHACHATRY17	a CMS	2 VBF jets $+ \cancel{E}_T$, Tsbot1, $m_{\widetilde{b}}$
		•		$m_{\widetilde{\chi}_1^0}=$ 5 GeV
>323	95	² AABOUD 16	D ATLS	\geq 1 jet $+ ot \!$
>840	95	³ AABOUD 16	Q ATLS	$=$ 5 GeV 2 $b ext{-jets}+ ot\!$
>540	95	⁴ AAD 16	BB ATLS	GeV 2 same-sign/ 3ℓ + jets + E_T , Tsbot2, $m_{\widetilde{\chi}^0_1} < 55$ GeV
>680	95	⁵ KHACHATRY16	BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tsbot2, $m_{\widetilde{\chi}^{\pm}_{-}}$ <
				550 GeV, $m_{\widetilde{\chi}_1^0}=$ 50 GeV $^{^{\sim}1}$
>500	95	⁵ KHACHATRY16	BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tsbot2, $m_{\widetilde{b}}$ –
				$m_{\widetilde{\chi}_1^\pm}$ $<$ 100 GeV, $m_{\widetilde{\chi}_1^0}$ $=$ 50 GeV
>880	95	⁶ KHACHATRY16	BS CMS	jets $+ E_T$, Tsbot1, $m_{\widetilde{\chi}_1^0} = 0$ GeV
>307	95	⁷ KHACHATRY16	BX CMS	$\widetilde{b} \rightarrow td \text{ or } ts, \text{ RPV, } \lambda_{332}^{\prime\prime} \text{ or } \lambda_{331}^{\prime\prime}$
>550	95	⁸ KHACHATRY16	BY CMS	opposite-sign $\ell^{\pm}\ell^{\pm}$, Tsbot3, $m_{\widetilde{\chi}_1^0}$
>600	95	⁹ AAD 15	CJ ATLS	$\widetilde{b} ightarrow b \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} < 250 \text{ GeV}$
>440	95	⁹ AAD 15	CJ ATLS	
				= 60 GeV, $m_{\widetilde{b}} - m_{\widetilde{\chi}_1^{\pm}}^{\pm} < m_t^{\chi_1}$
none 300-65	0 95	⁹ AAD 15	CJ ATLS	$\widetilde{b} \rightarrow \widetilde{b} b \widetilde{\chi}_{2}^{0}, \widetilde{\chi}_{2}^{0} \rightarrow h \widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}} =$
				60 GeV, $m_{\widetilde{\chi}_{0}^{0}} > 250 \text{ GeV}^{1}$
>640	95	¹⁰ KHACHATRY15	AF CMS	$\widetilde{b} \rightarrow b\widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{2}^{0}} = 0$
>650	95	¹¹ KHACHATRY15	ан CMS	$\widetilde{b} \rightarrow b\widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}}^{\alpha_{1}} = 0$
>250	95	¹¹ KHACHATRY15	ан CMS	
>570	95	¹² KHACHATRY15	ı CMS	
				=50 GeV, 150< $m_{\widetilde{\chi}_1^\pm}$ <300 GeV
>680 >500 >880 >307 >550 >600 >440 none 300–650 >650 >250	95 95 95 95 95 95 95 95	5 KHACHATRY16 5 KHACHATRY16 6 KHACHATRY16 7 KHACHATRY16 8 KHACHATRY16 9 AAD 15 9 AAD 15 9 AAD 15 10 KHACHATRY15 11 KHACHATRY15 11 KHACHATRY15	BJ CMS BJ CMS BS CMS BX CMS BY CMS CJ ATLS CJ ATLS AF CMS AH CMS AH CMS	$\begin{array}{l} 2\; \mathrm{same\text{-}sign}/3\ell + \mathrm{jets} + E_T, \; \mathrm{Ts-bot2}, \; m_{\widetilde{\chi}_1^0} < 55\; \mathrm{GeV} \\ \mathrm{same\text{-}sign}\; \ell^\pm \ell^\pm, \; \mathrm{Tsbot2}, \; m_{\widetilde{\chi}_1^\pm} < \\ 550\; \mathrm{GeV}, \; m_{\widetilde{\chi}_1^0} = 50\; \mathrm{GeV} \\ \mathrm{same\text{-}sign}\; \ell^\pm \ell^\pm, \; \mathrm{Tsbot2}, \; m_{\widetilde{b}} - \\ m_{\widetilde{\chi}_1^\pm} < 100\; \mathrm{GeV}, \; m_{\widetilde{\chi}_1^0} = 50\; \mathrm{GeV} \\ \mathrm{jets} + E_T, \; \mathrm{Tsbot1}, \; m_{\widetilde{\chi}_1^0} = 0\; \mathrm{GeV} \\ \widetilde{b} \to \; td \; \mathrm{or}\; ts, \; \mathrm{RPV}, \; \lambda_{1332}'' \; \mathrm{or}\; \lambda_{133}'' \; \mathrm{coupling} \\ \mathrm{opposite\text{-}sign}\; \ell^\pm \ell^\pm, \; \mathrm{Tsbot3}, \; m_{\widetilde{\chi}_1^0} \\ \widetilde{b} \to \; b\widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 250\; \mathrm{GeV} \\ \widetilde{b} \to \; b\widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 250\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to \; W^{(*)} \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} = 60\; \mathrm{GeV}, \; m_{\widetilde{b}} - m_{\widetilde{\chi}_1^0}, \; m_{\widetilde{\chi}_1^0} = 60\; \mathrm{GeV}, \; m_{\widetilde{\chi}_1^0} > 250\; \mathrm{GeV} \\ \widetilde{b} \to \; b\widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} = 0 \\ \widetilde{b} \to \; b\widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} = 0 \\ \widetilde{b} \to \; b\widetilde{\chi}_1^0, \; m_{\widetilde{b}} - m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} = 0 \\ \widetilde{b} \to \; b\widetilde{\chi}_1^0, \; m_{\widetilde{b}} - m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^\pm, \; \widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0, \; m_{\widetilde{\chi}_1^0} < 10\; \mathrm{GeV} \\ \widetilde{b} \to \; t\widetilde{\chi}_1^0, \; t\widetilde{\chi}_1^0, \; t\widetilde{\chi}_1^0, \; t\widetilde{\chi}_1^0, \; t\widetilde{\chi}_1^0, \; t$

>255	95	¹³ AAD	14T	ATLS	$\widetilde{b}_1 \rightarrow b\widetilde{\chi}_1^0, m_{\widetilde{b}_1} - m_{\widetilde{\chi}_1^0} \approx m_b$
>400	95	¹⁴ CHATRCHYAN	14AH		
		¹⁵ CHATRCHYAN	14 R	CMS	$\geq 3\ell^{\pm}$, $\widetilde{b} \stackrel{\chi_1}{ o} t \widetilde{\chi}_1^{\pm}$, $\widetilde{\chi}_1^{\pm} \rightarrow W^{\pm} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0}$
\\\/ -		Alex fellowing date	c		= 50 GeV
• • • vve do	not use				fits, limits, etc. • • •
		¹⁶ KHACHATRY	.15AD	CMS	$\ell^{\pm}\ell^{\mp}$ + jets + $ ot\!$
none 340-600	95	¹⁷ AAD	14AX	ATLS	\geq 3 <i>b</i> -jets $+$ $ ot\!\!\!E_T$, $\widetilde{b} ightarrow b \widetilde{\chi}_2^0$ sim-
					plified model with $\widetilde{\chi}_2^0 o h\widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0}$ =60 GeV, $m_{\widetilde{\chi}_2^0}$ =300 GeV
. 440	0.5	¹⁸ AAD	14-	ATLC	$\chi_1^{\prime\prime}$ $\chi_2^{\prime\prime}$ $\chi_2^{\prime\prime}$ $\chi_2^{\prime\prime}$
>440	95	10 AAD	14E	AILS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp})$ + jets, $\widetilde{b}_1 \rightarrow t\widetilde{\chi}_1^{\pm}$
					with $\widetilde{\chi}_1^\pm o W^{(*)} \pm \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^\pm} = 2 \; m_{\widetilde{\chi}_1^0}$
>500	95	¹⁹ CHATRCHYAN	14H	CMS	same-sign $\ell^{\pm}\ell^{\pm}$, $\widetilde{b} \rightarrow t\widetilde{\chi}_{1}^{\pm}$,
					$\widetilde{\chi}_1^\pm o \ W^\pm \widetilde{\chi}_1^0$ simplified
					model, $m_{\widetilde{\chi}_1^{\pm}} = 2$ GeV, $m_{\widetilde{\chi}_1^0} =$
>620	95	²⁰ AAD	13 AU	ATLS	100 GeV $_2$ b-jets $+ \not\!\!E_T$, $\widetilde{b}_1 o b\widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} <$
>550	95	²¹ CHATRCHYAN	13AT	CMS	120 GeV jets $+ \not\!\!E_T, \; \widetilde{b} \to b \widetilde{\chi}^0_1 \; { m simplified}$ model, $m_{\widetilde{\chi}^0_1} = 50 \; { m GeV}$
>600	95	²² CHATRCHYAN	13T	CMS	
>450	95	²³ CHATRCHYAN	13∨	CMS	same-sign $\ell^{\pm}\bar{\ell^{\pm}}+\geq 2$ <i>b</i> -jets,
					$\widetilde{b} ightarrow t \widetilde{\chi}_1^\pm$, $\widetilde{\chi}_1^\pm ightarrow W^\pm \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} = 50~{ m GeV}$
>390		²⁴ AAD	12AN	ATLS	$\widetilde{b}_1 ightarrow b \widetilde{\chi}_1^0$, simplified model, $m_{\widetilde{\chi}_1^0} < 60 \; ext{GeV}$
		²⁵ CHATRCHYAN	12AI	CMS	$\ell^{\pm}\ell^{\pm} + h_{\text{-iets}} + E_{\text{m}}$
>410	95	²⁶ CHATRCHYAN	12 BO	CMS	$\widetilde{b}_1 \rightarrow b\widetilde{\chi}_1^0$, simplified model, $m_{\widetilde{\chi}_1^0}$
>294	95	²⁷ AAD	11K	ATLS	$= 50 \text{ GeV}$ stable \tilde{b}
		²⁸ AAD	110	ATLS	$\widetilde{g} \rightarrow \widetilde{b}_1 b, \widetilde{b}_1 \rightarrow b \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 60$
		²⁹ CHATRCHYAN	11 D	CMS	$\widetilde{b}.\widetilde{t} o b$
>230	95	³⁰ AALTONEN	10R	CDF	$\widetilde{b}_1 \rightarrow b \widetilde{\chi}_1^0, \ m_{\widetilde{\chi}_1^0} < 70 \text{ GeV}$
>247	95	³¹ ABAZOV	10L	D0	$\widetilde{b}_1 ightarrow b\widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 0 GeV$
, -				- v	χ_1, \dots, χ_1

- 1 KHACHATRYAN 17A searched in 18.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with two forward jets, produced through vector boson fusion, and missing transverse momentum. No significant excess above the Standard Model expectations is observed. A limit is set on sbottom masses in the Tsbot1 simplified model, see Fig. 3.
- ² AABOUD 16D searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for events with an energetic jet and large missing transverse momentum. The results are interpreted as 95%C.L. limits on mass of sbottom decaying into a b-quark and the lightest neutralino in scenarios with $m_{\widetilde{b}_1} m_{\widetilde{\chi}_1^0}$ between 5 and 20 GeV. See their Fig. 6.
- 3 AABOUD 16Q searched in 3.2 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events containing two jets identified as originating from b-quarks and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks. Assuming that the decay $\tilde{b}_1 \to b \tilde{\chi}_1^0$ (Tsbot1) takes place 100% of the time, a \tilde{b}_1 mass below 840 (800) GeV is excluded for $m_{\tilde{\chi}_1^0} <$ 100 (360) GeV. Differences in mass above 100 GeV

between the \widetilde{b}_1 and the $\widetilde{\chi}_1^0$ are excluded up to a \widetilde{b}_1 mass of 500 GeV. For more details, see their Fig. 4.

- ⁴ AAD 16BB searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for events with exactly two same-sign leptons or at least three leptons, multiple hadronic jets, b-jets, and $\not\!\!\!E_T$. No significant excess over the Standard Model expectation is found. Exclusion limits at 95% C.L. are set on the sbottom mass for the Tsbot2 model, assuming $m_{\widetilde{\chi}_1^\pm} = m_{\widetilde{\chi}_1^0} + 100$ GeV. See their Fig. 4c.
- 5 KHACHATRYAN $_{16\mathrm{BJ}}$ searched in $2.3~\mathrm{fb}^{-1}$ of pp collisions at $\sqrt{s}=13~\mathrm{TeV}$ for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in the Tsbot2 simplified model, see Fig. 6.
- 6 KHACHATRYAN 16BS searched in 2.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with at least one energetic jet , no isolated leptons, and significant $\not\!\!E_T$, using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in the Tsbot1 simplified model, see Fig. 11 and Table 3.
- ⁷ KHACHATRYAN 16BX searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing 2 leptons coming from R-parity-violating decays of supersymmetric particles. No excess over the expected background is observed. Limits are derived on the sbottom mass, assuming the RPV $\tilde{b} \rightarrow td$ or $\tilde{b} \rightarrow ts$ decay, see Fig. 15.
- 8 KHACHATRYAN 16BY searched in 2.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with two opposite-sign, same-flavour leptons, jets, and missing transverse momentum. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the Tglu4C simplified model, see Fig. 4, and on sbottom masses in the Tsbot3 simplified model, see Fig. 5.
- ⁹ AAD 15CJ searched in 20 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for evidence of third generation squarks by combining a large number of searches covering various final states. Limits on the sbottom mass are shown, either assuming the $\tilde{b} \to b \tilde{\chi}_1^0$ decay, see Fig.
- 11, or assuming the $\widetilde{b} \to t\widetilde{\chi}_1^\pm$ decay, with $\widetilde{\chi}_1^\pm \to W^{(*)}\widetilde{\chi}_1^0$, see Fig. 12a, or assuming the $\widetilde{b} \to b\widetilde{\chi}_2^0$ decay, with $\widetilde{\chi}_2^0 \to h\widetilde{\chi}_1^0$, see Fig. 12b. Interpretations in the pMSSM are also discussed, see Figures 13–15.
- 10 KHACHATRYAN 15AF searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets and significant E_T , using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in simplified models where the decay $\tilde{b} \to b \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 12. See also Table 5. Exclusions in the CMSSM, assuming $\tan\beta=30$, $A_0=-2$ max(m_0 , $m_{1/2}$) and $\mu>0$, are also presented, see Fig. 15.

- In KHACHATRYAN 15AH searched in 19.4 or 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing either a fully reconstructed top quark, or events containing dijets requiring one or both jets to originate from b-quarks, or events containing a mono-jet. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in simplified models where the decay $\widetilde{b} \to b \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 12. Limits are also set in a simplified model where the decay $\widetilde{b} \to c \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 12.
- 12 KHACHATRYAN 151 searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events in which b-jets and four W-bosons are produced. Five individual search channels are combined (fully hadronic, single lepton, same-sign dilepton, opposite-sign dilepton, multilepton). No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in a simplified model where the decay $\tilde{b} \to t \tilde{\chi}_1^\pm$, with $\tilde{\chi}_1^\pm \to W^\pm \tilde{\chi}_1^0$, takes place with a branching ratio of 100%, see Fig. 7.
- 13 AAD 14T searched in 20.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for monojet-like events. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which assume that the decay $\tilde{b}_1 \rightarrow b \tilde{\chi}_1^0$ takes place 100% of the time, see Fig. 12.
- CHATRCHYAN 14AH searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with at least two energetic jets and significant E_T , using the razor variables (M_R and R^2) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a b-quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay $\tilde{b} \to b \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming $\tan\beta=10$, $A_0=0$ and $\mu>0$, are also presented, see Fig. 26.
- ¹⁵CHATRCHYAN 14R searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a simplified model where the decay $\tilde{b} \to t \tilde{\chi}_1^\pm$, with $\tilde{\chi}_1^\pm \to W^\pm \tilde{\chi}_1^0$, takes place with a branching ratio of 100%, see Fig. 11.
- 16 KHACHATRYAN 15AD searched in 19.4 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with two opposite-sign same flavor isolated leptons featuring either a kinematic edge, or a peak at the Z-boson mass, in the invariant mass spectrum. No evidence for a statistically significant excess over the expected SM backgrounds is observed and 95% C.L. exclusion limits are derived in a simplified model of sbottom pair production where the sbottom decays into a b-quark, two opposite-sign dileptons and a neutralino LSP, through an intermediate state containing either an off-shell Z-boson or a slepton, see Fig. 8.
- 17 AAD 14AX searched in $20.1~{\rm fb}^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for the strong production of supersymmetric particles in events containing either zero or at last one high high- p_T lepton, large missing transverse momentum, high jet multiplicity and at least three jets identified as originating from b-quarks. No excess over the expected SM background is observed. Limits are derived in mSUGRA/CMSSM models with $\tan\beta=30,~A_0=-2~m_0$ and $\mu>0$, see their Fig. 14. Also, exclusion limits are set in simplified models containing scalar bottom quarks, where the decay $\widetilde{b}\to b\widetilde{\chi}_2^0$ and $\widetilde{\chi}_2^0\to h\widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see their Figures 11.
- 18 AAD 14E searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for strongly produced supersymmetric particles in events containing jets and two same-sign leptons or three leptons. The search also utilises jets originating from b-quarks, missing transverse momentum and other variables. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing bottom, see Fig. 7. Limits are also derived in the mSUGRA/CMSSM, bRPV and GMSB models, see their Fig. 8.

- ¹⁹ CHATRCHYAN 14H searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the sbottom mass in a simplified models where the decay $\widetilde{b} \to t \widetilde{\chi}_1^\pm$, $\widetilde{\chi}_1^\pm \to W^\pm \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, with varying mass of the $\widetilde{\chi}_1^\pm$, for $m_{\widetilde{\chi}_1^0}=50$ GeV, see Fig. 6.
- 20 AAD 13AU searched in 20.1 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing two jets identified as originating from b-quarks and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks. Assuming that the decay $\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$ takes place 100% of the time, a \tilde{b}_1 mass below 620 GeV is excluded for $m_{\tilde{\chi}_1^0} <$ 120 GeV. For more details, see their Fig. 5.
- ²¹ CHATRCHYAN 13AT provides interpretations of various searches for supersymmetry by the CMS experiment based on 4.73–4.98 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV in the framework of simplified models. Limits are set on the sbottom mass in a simplified models where sbottom quarks are pair-produced and the decay $\tilde{b} \to b \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 4.
- ²² CHATRCHYAN 13T searched in 11.7 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets and significant E_T , using the α_T variable to discriminate between processes with genuine and misreconstructed E_T . No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay $\tilde{b} \to b \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 8 and Table 9.
- ²³ CHATRCHYAN 13V searched in 10.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with two isolated same-sign dileptons and at least two b-jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the bottom mass in a simplified models where the decay $\tilde{b} \to t \tilde{\chi}_1^\pm$, $\tilde{\chi}_1^\pm \to W^\pm \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, with varying mass of the $\tilde{\chi}_1^\pm$, for $m_{\tilde{\chi}_1^0}=50$ GeV, see Fig. 4.
- 24 AAD 12AN searched in 2.05 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for scalar bottom quarks in events with large missing transverse momentum and two b-jets in the final state. The data are found to be consistent with the Standard Model expectations. Limits are set in an R-parity conserving minimal supersymmetric scenario, assuming ${\rm B}(\widetilde{b}_1\to b\widetilde{\chi}_1^0)=100\%$, see their Fig. 2.
- 25 CHATRCHYAN 12AI looked in 4.98 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with two same-sign leptons $(e,\ \mu),$ but not necessarily same flavor, at least 2 b-jets and missing transverse energy. No excess beyond the Standard Model expectation is observed. Exclusion limits are derived in a simplified model for sbottom pair production, where the sbottom decays through $\widetilde{b}_1 \rightarrow t \widetilde{\chi}_1 \, W,$ see Fig. 8.
- 26 CHATRCHYAN 12BO searched in 4.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for scalar bottom quarks in events with large missing transverse momentum and two b-jets in the final state. The data are found to be consistent with the Standard Model expectations. Limits are set in an R-parity conserving minimal supersymmetric scenario, assuming $\mathrm{B}(\widetilde{b}_1 \to b \, \widetilde{\chi}_1^0) = 100\%$, see their Fig. 2.
- ²⁷ AAD 11K looked in 34 pb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or time of flight in the tile calorimeter, from pair production of \widetilde{b} . No evidence for an excess over the SM expectation is observed and limits on the mass are derived for pair production of sbottom, see Fig. 4.
- 28 AAD 110 looked in 35 pb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with jets, of which at least one is a b-jet, and E_T . No excess above the Standard Model was found. Limits are derived in the $(m_{\widetilde{g}},\,m_{\widetilde{b}_1})$ plane (see Fig. 2) under the assumption of 100%

branching ratios and \widetilde{b}_1 being the lightest squark. The quoted limit is valid for $m_{\widetilde{b}_1} < 500$ GeV. A similar approach for \widetilde{t}_1 as the lightest squark with $\widetilde{g} \to \widetilde{t}_1 t$ and $\widetilde{t}_1 \to b \widetilde{\chi}_1^\pm$ with 100% branching ratios leads to a gluino mass limit of 520 GeV for 130 $< m_{\widetilde{t}_1} < 300$ GeV. Limits are also derived in the CMSSM $(m_0, m_{1/2})$ plane for $\tan\beta = 40$, see Fig. 4, and in scenarios based on the gauge group SO(10).

- 29 CHATRCHYAN 11D looked in $35~\text{pb}^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7~\text{TeV}$ for events with ≥ 2 jets, at least one of which is b-tagged, and E_T , where the b-jets are decay products of \widetilde{t} or \widetilde{b} . No evidence for an excess over the expected background is observed. Limits are derived in the CMSSM $(m_0,\ m_{1/2})$ plane for $\tan\beta=50$ (see Fig. 2).
- 30 AALTONEN 10R searched in 2.65 fb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with E_T and exactly two jets, at least one of which is b-tagged. The results are in agreement with the SM prediction, and a limit on the cross section of 0.1 pb is obtained for the range of masses $80 < m_{\widetilde{b}_1} < 280$ GeV assuming that the sbottom decays exclusively to $b\widetilde{\chi}_1^0$. The excluded mass region in the framework of conserved R_p is shown in a plane
- of $(m_{\widetilde{b}_1}, m_{\widetilde{\chi}_1^0})$, see their Fig.2. 31 ABAZOV 10L looked in 5.2 fb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events with at least 2 b-jets and E_T from the production of $\widetilde{b}_1\,\widetilde{b}_1$. No evidence for an excess over the SM expectation is observed, and a limit on the cross section is derived under the assumption of 100% branching ratio. The excluded mass region in the framework of conserved R_p is shown in a plane of $(m_{\widetilde{b}_1}, m_{\widetilde{\chi}_1^0})$, see their Fig. 3b. The exclusion also extends to $m_{\widetilde{\chi}_1^0}=110$ GeV for $160 < m_{\widetilde{b}_1} < 200$ GeV.

\tilde{t} (Stop) MASS LIMIT

Limits depend on the decay mode. In e^+e^- collisions they also depend on the mixing angle of the mass eigenstate $\widetilde{t}_1=\widetilde{t}_L\cos\theta_t+\widetilde{t}_R\sin\theta_t$. The coupling to the Z vanishes when $\theta_t=0.98$. In the Listings below, we use $\Delta m\equiv m_{\widetilde{t}_1}-m_{\widetilde{\chi}_1^0}$ or $\Delta m\equiv m_{\widetilde{t}_1}-m_{\widetilde{\nu}}$, depending on relevant decay mode. See also bounds in " \widetilde{q} (Squark) MASS LIMIT."

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID	TECN	COMMENT
> 323-800 (CL	_ = 95%)	OUR EVALUATION	ON Range fo	or decay via charm or top
> 750	95	¹ KHACHATRY	17 CMS	$jets + \not\!\! E_T, Tstop1, m_{\widetilde{\chi}_1^0} = 100GeV$
> 323	95	² AABOUD	16D ATLS	≥ 1 jet $+ ot \!$
none, 745–780	95	³ AABOUD	16J ATLS	$1~\ell^{\pm} + \sum_{T=0}^{\infty} 4~ ext{jets} + ot\!$
none, 100-315	95	⁴ AAD	16AM ATLS	2 large-radius jets, Tstop1RPV
> 490–650	95	⁵ AAD	16AY ATLS	2ℓ (including hadronic $ au$) + $ ot\!$
> 700	95	⁶ KHACHATRY	16AV CMS	1 or 2 ℓ^{\pm} +jets+ b -jets+ $ ot\!$

> 700	95	⁶ KHACHATRY.	16AV CMS	1 or 2 ℓ^{\pm} +jets+ b -jets $ ot\!$
				$= 0.75 \ m_{\widetilde{t}_1} + 0.25 \ m_{\widetilde{\chi}_1^0}$
> 775	95	⁷ KHACHATRY.	16BK CMS	${\sf jets+} E_T, {\sf Tstop1}, m_{\widetilde{\chi}_1^0} < 200 {\sf GeV}$
> 620	95	⁷ KHACHATRY.	16BK CMS	jets+ E_T , Tstop2, $m_{\widetilde{\chi}_1^0}=0$ GeV
> 800	95	⁸ KHACHATRY.	16BS CMS	jets+ E_T , Tstop1, $m_{\widetilde{\chi}_1^0} = 0$ GeV
> 316	95	⁹ KHACHATRY.	16Y CMS	1 or 2 soft ℓ^{\pm} + jets $+ \not\!\!\!E_T$, Tstop3, $m_{\widetilde{t}} - m_{\widetilde{\chi}^0} = 25$ GeV
> 250	95	¹⁰ AAD	15CJ ATLS	$B(\widetilde{t} \to c \widetilde{\chi}_1^0) + B(\widetilde{t} \to bff' \widetilde{\chi}_1^0)$ $= 1, \ m_{\widetilde{t}} - m_{\widetilde{\chi}_1^0} = 10 \text{ GeV}$
> 270	95	¹⁰ AAD	15CJ ATLS	$\widetilde{t} \rightarrow c\widetilde{\chi}_1^0, m_{\widetilde{t}} - m_{\widetilde{\chi}_1^0} = 80 \text{ GeV}$
none, 200-700	95	¹⁰ AAD		$\widetilde{t} \rightarrow t \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 0$
> 500	95	¹⁰ AAD		$B(\widetilde{t} \to t\widetilde{\chi}_1^0) + B(\widetilde{t} \to b\widetilde{\chi}_1^{\pm})$
				$= 1, \widetilde{\chi}_{1}^{\pm} \rightarrow W^{(*)} \widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{\pm}}$
				$=2m_{\widetilde{\chi}_1^0}$, $m_{\widetilde{\chi}_1^0}$ $< 160~{ m GeV}^{\Lambda_1}$
> 600	95	¹⁰ AAD	15CJ ATLS	$\widetilde{t}_2 \rightarrow Z\widetilde{t}_1, \ m_{\widetilde{t}_1} - m_{\widetilde{\chi}_1^0} = 180$
				GeV, $m_{\widetilde{\chi}_1^0}=0$
> 600	95	¹⁰ AAD	15CJ ATLS	$\widetilde{t}_2 \rightarrow h\widetilde{t}_1, m_{\widetilde{t}_1} - m_{\widetilde{\chi}_1^0} = 180$
				GeV, $m_{\widetilde{\chi}_1^0}=0$
none, 172.5-191	95	¹¹ AAD	15J ATLS	$\widetilde{t} ightarrow t \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 1 GeV$
> 450	95	¹² KHACHATRY.	15AF CMS	$\widetilde{t} \rightarrow t \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 0, m_{\widetilde{t}} > m_t$
				$+ m_{\widetilde{\chi}_1^0}$
> 560	95	¹³ KHACHATRY.	15AH CMS	$\widetilde{t} \rightarrow t \widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}} = 0, m_{\widetilde{t}} > m_{t}$
				$+ m_{\widetilde{\chi}_1^0}$
> 250	95	¹⁴ KHACHATRY.	15AH CMS	$\widetilde{t} \rightarrow c \widetilde{\chi}_{1}^{0}, m_{\widetilde{t}} - m_{\widetilde{\chi}_{1}^{0}} < 10 \text{ GeV}$
none, 200-350	95	¹⁵ KHACHATRY.	15L CMS	$\widetilde{t} \rightarrow qq, \not R, \lambda_{312}'' \neq 0$
none, 200–385	95	¹⁵ KHACHATRY.	15L CMS	$\widetilde{t} \rightarrow qb, \not R, \lambda''_{323} \neq 0$
> 730	95	¹⁶ KHACHATRY.	15X CMS	$\widetilde{t} \rightarrow q q, \ k, \ \lambda_{312} \neq 0$ $\widetilde{t} \rightarrow q b, \ k, \ \lambda_{323} \neq 0$ $\widetilde{t} \rightarrow t \widetilde{\chi}_1^0, \ m_{\widetilde{\chi}_1^0} = 100 \text{ GeV},$
				$m_{\widetilde{t}} > m_t + m_{\widetilde{\gamma}0}$
none 400–645	95	¹⁶ KHACHATRY.	15x CMS	$\widetilde{t} ightarrow t \widetilde{\chi}_1^0 { m or} \widetilde{t} ightarrow b \widetilde{\chi}_1^\pm, m_{\widetilde{\chi}_1^0}$
				$=$ 100 GeV, $m_{\widetilde{\chi}_1^\pm} - m_{\widetilde{\chi}_1^0} =$
none 270-645	95	¹⁷ AAD	14AJ ATLS	$5 \; GeV \ \geq \; 4 \; jets + ot \!$
				$m_{\widetilde{\chi}_1^0} < 30 \text{ GeV}$

none 250-550	95	¹⁷ AAD	14AJ ATLS	\geq 4 jets $+ ot \!$
				$m_{\widetilde{\chi}_1^0} <$ 60 GeV
none 210-640	95	¹⁸ AAD	14BD ATLS	$\ell^{\pm} + jets + ot \!$
> 500	95	¹⁸ AAD	14BD ATLS	ℓ^{\pm} + jets + $\not\!\!E_T$, $\widetilde t_1 o b\widetilde{\chi}_1^{\pm}$, $m_{\widetilde{\chi}_1^{\pm}} = 2 \; m_{\widetilde{\chi}_1^0}$, 100 GeV $<$
none 150–445	95	¹⁹ AAD	14F ATLS	$m_{\widetilde{\chi}_1^0} < 150 \mathrm{GeV}$ $\ell^\pm \ell^\mp$ final state, $\widetilde{t}_1 \to b \widetilde{\chi}_1^\pm$, $m_{\widetilde{t}_1} - m_{\widetilde{\chi}_1^\pm} = 10 \mathrm{GeV}$, $m_{\widetilde{\chi}_1^0}$
none 215–530	95	¹⁹ AAD	14F ATLS	t_1 χ_1 χ_1 χ_1 χ_1 χ_1 $\ell^\pm\ell^\mp$ final state, $\tilde{t}_1 ightarrow t \tilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} = 1 \; ext{GeV}$
> 270	95	²⁰ AAD	14T ATLS	$\widetilde{t}_1 ightarrow c \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} = 200$ GeV
> 240	95	²⁰ AAD		$\widetilde{t}_1 \rightarrow c \widetilde{\chi}_1^0, m_{\widetilde{t}_1}^{\lambda_1} - m_{\widetilde{\chi}_1^0} < 85 \text{ GeV}$
> 255	95	²⁰ AAD		$\widetilde{t}_1 \rightarrow bff'\widetilde{\chi}_1^0, m_{\widetilde{t}_1} - m_{\widetilde{\chi}_1^0} \approx$
> 400	95	²¹ CHATRCHYA	N 14AH CMS	m_{b} jets $+ \not\!\!E_{T}$, $\widetilde{t} ightarrow t \widetilde{\chi}_{1}^{0}$ simplified model, $m_{\widetilde{\chi}_{1}^{0}} = 50 \; \mathrm{GeV}$
		²² CHATRCHYAI	N 14R CMS	$\geq 3\ell^{\pm}$, $\widetilde{t} \rightarrow (b\widetilde{\chi}_{1}^{\pm}/t\widetilde{\chi}_{1}^{0})$, $\widetilde{\chi}_{1}^{\pm} \rightarrow (qq'/\ell\nu)\widetilde{\chi}_{1}^{0}$, $\widetilde{\chi}_{1}^{0} \rightarrow (H/Z)\widetilde{G}$, GMSB, natural higgsino NLSP scenario
> 740	95	²³ KHACHATRY	14T CMS	$ au+$ b -jets, R , $LQ\overline{D}$, $\lambda_{333}' eq 0$,
> 580	95	²³ KHACHATRY	14T CMS	$\widetilde{t} ightarrow au b$ simplified model $ au + b$ -jets, R , $LQ\overline{D}$, $\lambda'_{3jk} \neq 0$ $(j \neq =3)$, $\widetilde{t} ightarrow \widetilde{\chi}^{\pm} b$, $\widetilde{\chi}^{\pm} ightarrow$
				$qq au^\pm$ simplified model
• • • We do no	ot use th	e following data for	•	·
> 890	95	²⁴ KHACHATRY	16AC CMS	$e^+e^-+ \geq 5 \text{ jets; } \widetilde{t} \rightarrow b\widetilde{\chi}_1^{\pm};$
>1000	95	²⁴ KHACHATRY	16AC CMS	$\widetilde{\chi}_{1}^{\pm} \rightarrow \ell^{\pm} jj$, RPV, λ'_{ijk} $\mu^{+}\mu^{-}+ \geq 5$ jets; $\widetilde{t} \rightarrow b\widetilde{\chi}_{1}^{\pm}$; $\widetilde{\chi}_{1}^{\pm} \rightarrow \ell^{\pm} jj$, RPV, λ'_{ijk}
> 950	95	²⁵ KHACHATRY	16BX CMS	$\widetilde{t} \rightarrow t \widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \rightarrow \ell \ell \nu, RPV,$
> 790 > 230	95	²⁶ KHACHATRY ROLBIECKI	15E CMS 15 THEO	λ_{121} or $\lambda_{122} \neq 0$ $\widetilde{t}_1 \rightarrow b\ell$, RPV, $c\tau = 2$ cm WW xsection, $\widetilde{t}_1 \rightarrow bW\widetilde{\chi}_1^0$, $m_{\widetilde{t}_1} \simeq m_b + m_W + m_{\widetilde{\chi}_1^0}$
> 600				14 D VV 2/0

> 540	95	27 AAD 14B	ATLS	$Z+b \not\!\!E_T$, $\widetilde t_1 ightarrow t \widetilde \chi_1^0$, $\widetilde \chi_1^0 ightarrow$
				$Z\widetilde{G}$, natural GMSB, 100 GeV $< m_{\widetilde{\chi}^0_1} < m_{\widetilde{t}_1} - 10$ GeV
> 360	95	²⁸ CHATRCHYAN 140	CMS	$\widetilde{t}_1 \rightarrow b\widetilde{\chi}_1^{\pm} r, \widetilde{\chi}_1^{\pm} \rightarrow f f' \widetilde{\chi}_1^0,$
				$\widetilde{\chi}_1^0 \to H\widetilde{G}$ simplified model,
				$m_{\widetilde{\chi}_1^{\pm}}^{-} - m_{\widetilde{\chi}_1^0} = 5 \text{ GeV,GMSB}$
> 215	95	CZAKON 14		$\widetilde{t} ightarrow t \chi_1^0$, $m_{\chi_1^0} < 10$ GeV
		²⁹ KHACHATRY14C	CMS	$\widetilde{t}_2 ightarrow H\widetilde{t}_1 \text{ or } \widetilde{t}_2 ightarrow Z\widetilde{t}_1 \text{ simplified model}$

- 1 KHACHATRYAN 17 searched in 2.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events containing four or more jets, no more than one lepton, and missing transverse momentum, using the razor variables (M_R and R^2) to discriminate between signal and background processes. No evidence for an excess over the expected background is observed. Limits are derived on the stop mass in the Tstop1 simplified model, see Fig. 17.
- 2 AABOUD 16D searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV in events with an energetic jet and large missing transverse momentum. The results are interpreted as 95% C.L. limits on mass of stop decaying into a charm-quark and the lightest neutralino in scenarios with $m_{\widetilde{t}_1}-m_{\widetilde{\chi}_1^0}$ between 5 and 20 GeV. See their Fig. 5.
- ³ AABOUD 16J searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV in final states with one isolated electron or muon, jets, and missing transverse momentum. For the direct stop pair production model where the stop decays via top and lightest neutralino, the results exclude at 95% C.L. stop masses between 745 GeV and 780 GeV for a massless $\tilde{\chi}_1^0$. See their Fig. 8.
- ⁴ AAD 16AM searched in 17.4 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing two large-radius hadronic jets. No deviation from the background prediction is observed. Top squarks with masses between 100 and 315 GeV are excluded at 95% C.L. in the hypothesis that they both decay via R-parity violating coupling $\lambda_{323}^{"}$ to b- and s-quarks. See their Fig. 10.
- 5 AAD 16AY searched in 20 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with either two hadronically decaying tau leptons, one hadronically decaying tau and one light lepton, or two light leptons. No significant excess over the Standard Model expectation is found. Exclusion limits at 95% C.L. on the mass of top squarks decaying via $\widetilde{\tau}$ to a nearly massless gravitino are placed depending on $m_{\widetilde{\tau}}$ which is ranging from the 87 GeV LEP limit to $m_{\widetilde{t}_1}$. See their Figs. 9 and 10.
- 6 KHACHATRYAN 16AV searched in 19.7 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with one or two isolated leptons, hadronic jets, b-jets and $\not\!\!E_T$. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in the Tstop1 and Tstop2 simplified models, see Fig. 11.
- 7 KHACHATRYAN 16BK searched in 18.9 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with hadronic jets and $\not\!\!E_T$. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in the Tstop1 and Tstop2 simplified models, see Fig. 16.
- 8 KHACHATRYAN 16BS searched in 2.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with at least one energetic jet , no isolated leptons, and significant $\not\!\!E_T$, using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in the Tstop1 simplified model, see Fig. 11 and Table 3.
- 9 KHACHATRYAN 16Y searched in 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with one or two soft isolated leptons, hadronic jets, and $\not\!\!E_T$. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in the Tstop3 simplified model, see Fig. 3.

- 10 AAD 15CJ searched in 20 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for evidence of third generation squarks by combining a large number of searches covering various final states. Stop decays with and without charginos in the decay chain are considered and summaries of all ATLAS Run 1 searches for direct stop production can be found in Fig. 4 (no intermediate charginos) and Fig. 7 (intermediate charginos). Limits are set on stop masses in compressed mass regions regions, with $B(\tilde{t}\to c\,\tilde{\chi}_1^0)+B(\tilde{t}\to bf\,f'\,\tilde{\chi}_1^0)=1$, see Fig. 5. Limits are also set on stop masses assuming that both the decay $\tilde{t}\to t\,\tilde{\chi}_1^0$ and $\tilde{t}\to b\,\tilde{\chi}_1^\pm$ are possible, with both their branching rations summing up to 1, assuming $\tilde{\chi}_1^\pm\to W^{(*)}\tilde{\chi}_1^0$ and $m_{\tilde{\chi}_1^\pm}=2\,m_{\tilde{\chi}_1^0}$, see Fig. 6. Limits on the mass of the next-to-lightest stop \tilde{t}_2 , decaying either to $Z\,\tilde{t}_1$, $h\,\tilde{t}_1$ or $t\,\tilde{\chi}_1^0$, are also presented, see Figs. 9 and 10.Interpretations in the pMSSM are also discussed, see Figs 13–15.
- 11 AAD 15J interpreted the measurement of spin correlations in $t\overline{t}$ production using 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV in exclusion limits on the pair production of light \widetilde{t}_1 squarks with masses similar to the top quark mass. The \widetilde{t}_1 is assumed to decay through $\widetilde{t}_1 \to t\,\widetilde{\chi}_1^0$ with predominantly right-handed top and a 100% branching ratio. The data are found to be consistent with the Standard Model expectations and masses between the top quark mass and 191 GeV are excluded, see their Fig. 2
- 12 KHACHATRYAN 15AF searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets and significant E_T , using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay $\widetilde{t} \to t \, \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 12. See also Table 5. Exclusions in the CMSSM, assuming $\tan\beta=30,\,A_0=-2\,\max(m_0,\,m_{1/2})$ and $\mu>0$, are also presented, see Fig. 15.
- 13 KHACHATRYAN 15AH searched in 19.4 or 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing either a fully reconstructed top quark, or events containing dijets requiring one or both jets to originate from b-quarks, or events containing a mono-jet. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay $\tilde{t} \to t \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 9. Limits are also set in simplified models where the decays $\tilde{t} \to t \tilde{\chi}_1^0$ and $\tilde{t} \to b \tilde{\chi}_1^\pm$, with $m_{\tilde{\chi}_1^\pm}$ $m_{\tilde{\chi}_1^0}=5$ GeV, each take place with a branching ratio of 50%, see Fig. 10, or with other fractions, see Fig. 11. Finally, limits are set in a simplified model where the decay $\tilde{t} \to c \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Figs. 9, 10 and 11.
- 14 KHACHATRYAN 15AH searched in 19.4 or 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing either a fully reconstructed top quark, or events containing dijets requiring one or both jets to originate from b-quarks, or events containing a mono-jet. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay $\tilde{t} \to t \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 9. Limits are also set in simplified models where the decays $\tilde{t} \to t \tilde{\chi}_1^0$ and $\tilde{t} \to b \tilde{\chi}_1^\pm$, with $m_{\tilde{\chi}_1^\pm} m_{\tilde{\chi}_1^0} = 5$ GeV, each take place with a branching ratio of 50%, see Fig. 10, or with other fractions, see Fig. 11. Finally, limits are set in a simplified model where the decay $\tilde{t} \to c \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Figs. 9, 10, and 11.
- ¹⁵ KHACHATRYAN 15L searched in 19.4 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for pair production of heavy resonances decaying to pairs of jets in four jet events. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in R-parity-violating supersymmetry models where $\tilde{t} \to qq \left(\lambda_{312}'' \neq 0\right)$, see Fig. 6 (top) and $\tilde{t} \to qb \left(\lambda_{323}'' \neq 0\right)$, see Fig. 6 (bottom).

- ¹⁶ KHACHATRYAN 15x searched in 19.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets, at least one of which is required to originate from a b quark, possibly a lepton, and significant E_T , using the razor variables (M_R and R^2) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in simplified models where the decay $\widetilde{t} \to t \widetilde{\chi}_1^0$ and the decay $\widetilde{t} \to b \widetilde{\chi}_1^\pm$, with $m_{\widetilde{\chi}_1^\pm} m_{\widetilde{\chi}_1^0} = 5$ GeV, take place with branching ratios varying between 0 and 100%, see Figs. 15, 16 and 17
- ¹⁷ AAD 14AJ searched in 20.1 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing four or more jets and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which either assume that the decay $\tilde{t}_1 \to t \tilde{\chi}_1^0$ takes place 100% of the time, see Fig. 8, or that this decay takes place 50% of the time, while the decay $\tilde{t}_1 \to b \tilde{\chi}_1^\pm$ takes place the other 50% of the time, see Fig. 9.
- 18 AAD 14BD searched in 20 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing one isolated lepton, jets and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which either assume that the decay $\tilde{t}_1 \to t \tilde{\chi}_1^0$ takes place 100% of the time, see Fig. 15, or the decay $\tilde{t}_1 \to b \tilde{\chi}_1^\pm$ takes place 100% of the time, see Fig. 16–22. For the mixed decay scenario, see Fig. 23.
- AAD 14F searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing two leptons (e or μ), and possibly jets and missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which either assume that the decay $\tilde{t}_1 \to b \tilde{\chi}_1^\pm$ takes place 100% of the time, see Figs. 14–17 and 20, or that the decay $\tilde{t}_1 \to t \tilde{\chi}_1^0$ takes place 100% of the time, see Figs. 18 and 19.
- AAD 14T searched in 20.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for monojet-like and c-tagged events. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the masses of third-generation squarks in simplified models which assume that the decay $\widetilde{t}_1 \to c \widetilde{\chi}_1^0$ takes place 100% of the time, see Fig. 9 and 10. The results of the monojet-like analysis are also interpreted in terms of stop pair production in the four-body decay $\widetilde{t}_1 \to bff'\widetilde{\chi}_1^0$, see Fig. 11.
- ²¹ CHATRCHYAN 14AH searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with at least two energetic jets and significant E_T , using the razor variables (M_R and R^2) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a b-quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay $\tilde{t} \to t \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming $\tan \beta = 10$, $A_0 = 0$ and $\mu > 0$, are also presented, see Fig. 26.
- ²²CHATRCHYAN 14R searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in a natural higgsino NLSP simplified model (GMSB) where the decay $\widetilde{t} \to b\widetilde{\chi}_1^\pm$, with $\widetilde{\chi}_1^\pm \to (qq'/\ell\nu)H$, $Z\widetilde{G}$, takes place with a branching ratio of 100% (the particles between brackets have a soft p_T spectrum), see Figs. 4–6.
- ²³ KHACHATRYAN 14T searched in 19.7 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with τ -leptons and b-quark jets, possibly with extra light-flavour jets. No excess above the Standard Model expectations is observed. Limits are set on stop masses in R SUSY

models with $LQ\overline{D}$ couplings, in two simplified models. In the first model, the decay $\widetilde{t} \to \tau b$ is considered, with $\lambda'_{333} \neq 0$, see Fig. 3. In the second model, the decay $\widetilde{t} \to \widetilde{\chi}^{\pm} b$, with the subsequent decay $\widetilde{\chi}^{\pm} \to qq\tau^{\pm}$ is considered, with $\lambda'_{3jk} \neq 0$ and the mass splitting between the top squark and the charging chosen to be 100 GeV, see Fig. 4.

- ²⁴ KHACHATRYAN 16AC searched in 19.7 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with low missing transverse momentum, two oppositely charged electrons or muons, and at least five jets, at least one of which is a b-jet, for evidence of R-parity violating, charging-mediated decays of the top squark. No significant excess above the Standard Model expectations is observed. Limits are set on the stop mass in R-parity-violating supersymmetry models where $\tilde{t} \to b \tilde{\chi}_1^\pm$ with $\tilde{\chi}_1^\pm \to \ell^\pm jj$, $\lambda'_{ijk} \neq 0$ ($i,j,k \leq 2$), and with $m_{\tilde{t}} m_{\tilde{\chi}_1^\pm} = 100$ GeV, see Fig. 3.
- 25 KHACHATRYAN 16BX searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing 4 leptons coming from R-parity-violating decays of $\widetilde{\chi}_1^0 \to \ell\ell\nu$ with $\lambda_{121} \neq 0$ or $\lambda_{122} \neq 0$. No excess over the expected background is observed. Limits are derived on the gluino, squark and stop masses, see Fig. 23.
- 26 KHACHATRYAN 15E searched for long-lived particles decaying to leptons in 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV. Events were selected with an electron and muon with opposite charges and each with transverse impact parameter values between 0.02 and 2 cm. Limits are set on SUSY benchmark models with pair production of top squarks decaying into an $e\,\mu$ final state via RPV interactions. See their Fig. 2
- 27 AAD 14 B searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing a Z boson, with or without additional leptons, plus jets originating from b-quarks and significant missing transverse momentum. No excess over the expected SM background is observed. Limits are derived in simplified models featuring \tilde{t}_2 production, with $\tilde{t}_2 \to Z\,\tilde{t}_1,\;\tilde{t}_1 \to t\,\tilde{\chi}_1^0$ with a 100% branching ratio, see Fig. 4, and in the framework of natural GMSB, see Fig. 6.
- 28 CHATRCHYAN 14U searched in $19.7~{\rm fb}^{-1}$ of pp collisions at $\sqrt{s}=8~{\rm TeV}$ for evidence of direct pair production of top squarks, with Higgs bosons in the decay chain. The search is performed using a selection of events containing two Higgs bosons, each decaying to a photon pair, missing transverse energy and possibly b-quark jets. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of a "natural SUSY" simplified model where the decays $\tilde{t}_1 \to b \tilde{\chi}_1^\pm$, with $\tilde{\chi}_1^\pm \to f f' \tilde{\chi}_1^0$, and $\tilde{\chi}_1^0 \to H \tilde{G}$, all happen with 100% branching ratio, see Fig. 4.
- ²⁹ KHACHATRYAN 14C searched in 19.5 fb⁻¹ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for evidence of direct pair production of top squarks, with Higgs or Z-bosons in the decay chain. The search is performed using a selection of events containing leptons and b-quark jets. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in the context of a simplified model with pair production of a heavier top-squark mass eigenstate \tilde{t}_2 decaying to a lighter top-squark eigenstate \tilde{t}_1 via either $\tilde{t}_2 \to H\tilde{t}_1$ or $\tilde{t}_2 \to Z\tilde{t}_1$, followed in both cases by $\tilde{t}_1 \to t\,\tilde{\chi}_1^0$. The interpretation is performed in the region where the mass difference between the \tilde{t}_1 and $\tilde{\chi}_1^0$ is approximately equal to the top-quark mass, which is not probed by searches for direct \tilde{t}_1 pair production, see Figs. 5 and 6. The analysis excludes top squarks with masses $m_{\tilde{t}_2} < 575$ GeV and $m_{\tilde{t}_1} < 400$ GeV at 95% C.L.

Heavy \tilde{g} (Gluino) MASS LIMIT

For $m_{\widetilde{g}} > 60$ –70 GeV, it is expected that gluinos would undergo a cascade decay via a number of neutralinos and/or charginos rather than undergo a direct decay to photinos as assumed by some papers. Limits obtained when direct decay is assumed are usually higher than limits when cascade decays are included.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
> 700–1780 ($CL = 95^{\circ}$	%) OUR EVALUATI	ON	${\sf Range}$	reflects model dependence
>1400	95	¹ KHACHATRY1	17 (CMS	$jets + \not\!\! E_T, Tglu1A, m_{\widetilde{\chi}_1^0} = 200 GeV$
>1650	95	¹ KHACHATRY1	17 (CMS	$\text{jets}+\cancel{E}_T, \text{Tglu2A}, m_{\widetilde{\chi}_1^0}=200 \text{ GeV}$
>1600	95	¹ KHACHATRY1	17 (CMS	$\text{jets}+\cancel{E}_T, \text{Tglu3A}, m_{\widetilde{\chi}_1^0}=200 \text{GeV}$
>1570	95	² AABOUD	16AC /	ATLS	\geq 2 jets $+$ 1 or 2 $ au$ + $ ot\!\!\!E_T$, Tglu1F, $m_{\widetilde{\chi}^0_1}=$ 100 GeV
>1460	95	³ AABOUD	16J /	ATLS	$1 \ell^{\pm} + \geq 4 \text{ jets} + E_T$, Tglu3C, $m_{\widetilde{t}_1} - m_{\widetilde{\chi}_1^0} = 5 \text{ GeV}$
>1650	95	⁴ AABOUD	16м /	ATLS	$2 \gamma + \cancel{E}_T$, Tglu1D, any NLSP
>1510	95	⁵ AABOUD	16N /	ATLS	\geq 4 jets $+ \not\!\!E_T$, Tglu1A, $m_{\widetilde{\chi}_1^0} = \blacksquare$
>1500	95	6 AABOUD	16N /	ATLS	0 GeV \geq 4 jets $+$ $ ot\!$
>1780	95	⁷ AAD	16 AD <i>i</i>	ATLS	$(m_{\widetilde{g}}+m_{\widetilde{\chi}_1^0})/2,\ m_{\widetilde{\chi}_1^0}=$ 200 ${ m GeV}$ $0\ell, \geq 3\ b ext{-jets} + E_T, { m Tglu2A}, \ m_{\widetilde{\chi}_1^0} < 800 { m GeV}$
>1760	95	⁸ AAD	16 AD <i>i</i>	ATLS	1ℓ , $\overset{\chi_1}{\geq}$ 3 b -jets $+ \not\!\!E_T$, Tglu3A, $m_{\widetilde{\chi}_1^0} < 700 \; { m GeV}$
>1300	95	⁹ AAD	16 BB /	ATLS	$\begin{array}{c} \chi_1 \\ 2 \text{ same-sign}/3\ell + \text{jets} + \cancel{E}_T, \\ \text{Tglu1D, } m_{\widetilde{\chi}_1^0} < 600 \text{ GeV} \end{array}$
>1100	95	⁹ AAD	16 BB /	ATLS	2 same-sign/ 3ℓ + jets + E_T , Tglu1E, $m_{\widetilde{\chi}_1^0} < 300 \text{ GeV}$
>1200	95	⁹ AAD	16 BB /	ATLS	2 same-sign $/3\ell$ + jets + $\not\!\!E_T$, Tglu3A, $m_{\widetilde{\chi}_1^0} <$ 600 GeV
>1600		10 AAD	16 BG /	ATLS	1ℓ , \geq 4 jets, $\not\!\!E_T$, Tglu1B, $m_{\widetilde{\chi}_1^\pm} = (m_{\widetilde{g}} + m_{\widetilde{\chi}_1^0})/2$, $m_{\widetilde{\chi}_1^0} = 100~{ m GeV}$
>1400	95	11 AAD	16∨ /	ATLS	\geq 7 to \geq 10 jets $+$ $ ot\!\!E_T$, Tglu1E, $m_{\sim 0} <$ 200 GeV
>1400	95	¹¹ AAD	16∨ /	ATLS	$\stackrel{\chi_1}{\geq}$ 7 to ${\geq}$ 10 jets $+ \cancel{E}_T$, pMSSM $M_1 = 60$ GeV, M_2
>1100	95	¹² KHACHATRYI	16AM (CMS	$= 3 \text{ TeV, } \tan\beta = 10, \ \mu < 0$ boosted $W+b$, Tglu3C, $m_{\widetilde{t}_1} - m_{\widetilde{\chi}_1^0} < 80 \text{GeV}, m_{\widetilde{\chi}_1^0} < 400 \text{GeV}$
					1 1

> 700	95	¹² KHACHATRY16AM CMS	boosted $W+b$, Tglu3B, $m_{\widetilde t_1}$ –
			$m_{\widetilde{\chi}_1^0}$ =175 GeV, $m_{\widetilde{\chi}_1^0}$ =0 GeV
>1050	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^\pm\ell^\pm$, Tglu3A, $m_{\widetilde{\chi}_1^0} < 800$ GeV
>1300	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu3A, $m_{\widetilde{\chi}_1^0} = 0$
>1140	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu3B, $m_{\widetilde{t}}$ –
			$m_{\widetilde{\chi}_1^0}=$ 20 GeV, $m_{\widetilde{\chi}_1^0}=0$
> 850	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu3B, $m_{\widetilde{t}}-m_{\widetilde{\chi}_1^0}=$ 20 GeV, $m_{\widetilde{\chi}_1^0}<$ 700 GeV
> 950	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu3D, $m_{\widetilde{\chi}_{1}^{\pm}}$
			$=m_{\widetilde{\chi}^0_1}+ 5~{\sf GeV}$
>1100	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu1B, $m_{\widetilde{\chi}_{2}^{\pm}} = \blacksquare$
			$0.5(m_{\widetilde{g}}+m_{\widetilde{\chi}_1^0}), m_{\widetilde{\chi}_1^0} < 400 \text{GeV}$
> 830	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu1B, $m_{\widetilde{\chi}_{1}^{\pm}}$
			$0.5(m_{\widetilde{g}}+m_{\widetilde{\chi}^0_1}), m_{\widetilde{\chi}^0_1} < 700 { m GeV}$
>1300	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu3B, $m_{\widetilde{t}}$ –
		10	$m_{\widetilde{\chi}_1^0} = m_t, m_{\widetilde{\chi}_1^0} = 0$
>1050	95	¹³ KHACHATRY16BJ CMS	same-sign $\ell^{\pm}\ell^{\pm}$, Tglu3B, $m_{\tilde{t}}$ -
> 170F	O.E.	¹⁴ KHACHATRY16BS CMS	$m_{\widetilde{\chi}_1^0} = m_t, m_{\widetilde{\chi}_1^0} < 800 \text{ GeV}$
>1725	95		jets $+ E_T$, Tglu1A, $m_{\widetilde{\chi}_1^0} = 0$
>1750	95	14 KHACHATRY16BS CMS	jets $+ E_T$, Tglu2A, $m_{\widetilde{\chi}_1^0} = 0$
>1550	95	¹⁴ KHACHATRY16BS CMS	jets $+ \not\!\!E_T$, Tglu3A, $m_{\widetilde{\chi}_1^0} = 0$
>1030	95	15 KHACHATRY16BX CMS	$\widetilde{g} \rightarrow tbs$, RPV, $\lambda_{332}^{\prime\prime}$ coupling
>1280	95	¹⁶ KHACHATRY16BY CMS	opposite-sign $\ell^{\pm}\ell^{\pm}$, Tglu4C, $m_{\widetilde{\chi}_1^0}=1000~{ m GeV}$
>1030	95	¹⁶ KHACHATRY16BY CMS	opposite-sign $\ell^{\pm}\ell^{\pm}$, Tglu4C, $m_{\widetilde{\chi}_1^0}=0$ GeV
>1440	95	¹⁷ KHACHATRY16V CMS	$\operatorname{jets} + E_T$, Tglu1A, $m_{\widetilde{\chi}_1^0} = 0$
>1600	95	¹⁷ KHACHATRY16v CMS	jets + $ ot\!\!\!E_T$, Tglu2A, $m_{\widetilde{\chi}_1^0} = 0$
>1550	95	¹⁷ KHACHATRY16V CMS	jets + $ ot\!\!\!E_T$, Tglu3A, $m_{\widetilde{\chi}_1^0}^{\lambda_1}=0$
>1450	95	¹⁷ KHACHATRY16V CMS	jets + $ ot\!\!\!E_T$, Tglu1C, $m_{\widetilde{\chi}_1^0}^{\lambda_1}=0$
> 820	95	¹⁸ AAD 15BG ATLS	GGM, $\widetilde{g} \rightarrow q\widetilde{q}Z\widetilde{G}$, $\tan\beta = 30$,
> 850	95	¹⁸ AAD 15BG ATLS	$\mu > 600 \ { m GeV}$ ${ m GGM}, \ \widetilde{g} ightarrow q \widetilde{q} Z \widetilde{G}, \ { m tan} eta = 1.5,$
>1150	95	¹⁹ AAD 15BV ATLS	$\mu >$ 450 GeV general RPC \widetilde{g} decays, $m_{\widetilde{\chi}^0_1} <$
> 700	95	²⁰ AAD 15BX ATLS	$\begin{array}{c} 100 \; \text{GeV} \\ \widetilde{g} \to \; X \widetilde{\chi}_1^0, \; \text{independent of} \; m_{\widetilde{\chi}_1^0} \end{array}$

>1290	95	²¹ AAD	15CA ATLS	\geq 2 γ + $ ot\!$
>1260	95	²¹ AAD	15CA ATLS	$\geq 1 \ \gamma + b$ -jets $+ E_T$, GGM, higgsino-bino admix. NLSP and $\mu < 0$, m(NLSP) > 450 GeV
>1140	95	²¹ AAD	15CA ATLS	$\geq 1 \ \gamma + { m jets} + E_T$, GGM, higgsino-bino admixture NLSP,
>1225	95	²² KHACHATRY	15AF CMS	all $\mu > 0$ $\widetilde{g} \rightarrow q \overline{q} \widetilde{\chi}_1^0, \ m_{\widetilde{\chi}_1^0} = 0$
>1300	95	²² KHACHATRY	15AF CMS	$\widetilde{g} \rightarrow b\overline{b}\widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}}^{\lambda_{1}} = 0$
>1225	95	²² KHACHATRY	15AF CMS	$\widetilde{g} \rightarrow t \overline{t} \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 0$
>1550	95	²² KHACHATRY	15AF CMS	CMSSM, $\tan \beta = 30$, $m_{\widetilde{g}} = m_{\widetilde{q}}$, $A_0 = -2\max(m_0, m_{1/2})$, $\mu > 0$
>1150	95	²² KHACHATRY	15AF CMS	CMSSM, $\tan \beta = 30$, $A_0 = -2 \max(m_0, m_{1/2}), \ \mu > 0$
>1280	95	²³ KHACHATRY	15I CMS	$\widetilde{g} \rightarrow t\widetilde{t}\widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}} = 0$
>1310	95	²⁴ KHACHATRY	15x CMS	$\widetilde{g} ightarrow b \overline{b} \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} = 100 \text{ GeV}$
>1175	95	²⁴ KHACHATRY	15X CMS	$\widetilde{g} ightarrow t \overline{t} \widetilde{\chi}_1^0, m_{\widetilde{\chi}_1^0} = 100 { m GeV}$
>1330	95	²⁵ AAD	14AE ATLS	jets $+ \not\!\!E_T$, $\stackrel{\sim}{g} \stackrel{\sim}{\to} q \overline{q} \stackrel{\sim}{\chi}^0_1$ simplified model, $m_{\widetilde{\chi}^0_1} = 0 \; {\sf GeV}$
>1700	95	²⁵ AAD	14AE ATLS	$egin{aligned} egin{aligned} & \chi_1 \ & \text{jets} + ot \!$
>1090	95	²⁶ AAD	14AG ATLS	$ au+jets+\cancel{\cancel{E}}_T$, natural Gauge
>1600	95	²⁶ AAD	14AG ATLS	Mediation $ au + {\rm jets} + E_T$, mGMSB, M $_{mess} = 250$ GeV, $N_5 = 3$, $\mu > 0$,
>1350	95	²⁷ AAD	14X ATLS	$egin{align} \mathcal{C}_{grav} &= 1 \ &\geq 4\ell^{\pm},\widetilde{g} ightarrow $
> 640	95	²⁸ AAD	14X ATLS	$ \geq 4\ell^{\pm}, \widetilde{g} \rightarrow q\overline{q}\widetilde{\chi}_{1}^{0}, \widetilde{\chi}_{1}^{0} \rightarrow \ell^{\pm}\ell^{\mp}\widetilde{G}, \tan\beta = 30, \text{GGM} $
>1000	95	²⁹ CHATRCHYA	N 14AH CMS	$\mathrm{jets}+E_T,\ \widetilde{g} ightarrow q \overline{q} \widetilde{\chi}_1^0 \mathrm{\ simplified}$ model, $m_{\widetilde{\chi}_1^0}=\mathrm{50\ GeV}$
>1350	95	²⁹ CHATRCHYA	N 14AH CMS	jets $+ ot \!$
>1000	95	³⁰ CHATRCHYA	N 14AH CMS	jets $+ \not\!\!E_T$, $\widetilde{g} \rightarrow b \overline{b} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} = 50 \mathrm{GeV}$
>1000	95	³¹ CHATRCHYA	N 14AH CMS	jets $+ \cancel{E}_T$, $\widetilde{g} \rightarrow t\overline{t}\widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} = 50 \text{ GeV}$
>1160	95	³² CHATRCHYA	N 14ı CMS	ets $+ E_T$, $\widetilde{g} \rightarrow q \overline{q} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} < 100 \text{ GeV}$
>1130	95	³² CHATRCHYA	N 14ı CMS	multijets $+ \not\!\!\!E_T$, $\not\!\!\!\!E \to t \overline{t} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} < 100$ GeV

>1210	95	³² CHATRCHYAN	N 14I CMS	multijets $+ \not\!\!E_T$, $\widetilde{g} \to q \overline{q} W/Z \widetilde{\chi}_1^0$ simplified model,
>1260	95	³³ CHATRCHYAN	N14n CMS	$egin{aligned} &m_{\widetilde{\chi}_1^0} &< 10 \tilde{0} ext{ GeV} \ &1 \ell^{\pm} + ext{ jets} + \geq 2 b ext{-jets}, ilde{g} ightarrow \ &t \overline{t} \chi_1^0 ext{ simplified model}, \ &m_{\widetilde{\chi}_1^0} = 0 ext{ GeV}, m_{\widetilde{t}} > m_{\widetilde{g}} \end{aligned}$
> 650 none 200–835	95 95	34 CHATRCHYAN 34 CHATRCHYAN 35 CHATRCHYAN	N14P CMS	χ_{1} $\widetilde{g} \rightarrow jjj, \not R$ $\widetilde{g} \rightarrow bjj, \not R$ $\geq 3\ell^{\pm}, (\widetilde{g}/\widetilde{q}) \rightarrow q\ell^{\pm}\ell^{\mp}\widetilde{G}$ simplified model, GMSB, slep-
		³⁶ CHATRCHYAN	N 14R CMS	ton co-NLSP scenario $\geq 3\ell^\pm$, $\widetilde{g} \to t \overline{t} \widetilde{\chi}_1^0$ simplified model
• • • We do n	ot use tl	he following data fo	or averages, fi	ts, limits, etc. • • •
>1600	95	³⁷ KHACHATRY.		$1\ell^{\pm}+{ m jets}+{\it b} ext{-jets}+{ ot}\!$
> 500	95	³⁸ KHACHATRY.	16BT CMS	19-parameter pMSSM model, global Bayesian analysis, flat
>1400	95	³⁹ KHACHATRY.	16BX CMS	$ \begin{array}{c} prior \\ \widetilde{g} \to q q \widetilde{\chi}_1^0, \widetilde{\chi}_1^0 \to \ell \ell \nu, RPV, \\ \lambda_{121} or \lambda_{122} \neq 0, m_{\widetilde{\chi}_1^0} > \end{array} $
	95	⁴⁰ AAD	15AB ATLS	$\widetilde{g} o \widetilde{S} g$, $c au = 1$ m, $\widetilde{S} o S \widetilde{G}$ and $S o g g$, BR $= 100\%$
	95	⁴¹ AAD	15AI ATLS	$\ell^\pm+{ m jets}+ ot\!$
>1600	95	¹⁹ AAD	15BV ATLS	pMSSM, M $_1=$ 60 GeV, $m_{\widetilde{q}}<1500$ GeV
>1280	95	¹⁹ AAD	15 _{BV} ATLS	mSUGRA, $m_0 > 2 \text{ TeV}$
>1100	95	¹⁹ AAD	15 _{BV} ATLS	via $\widetilde{ au}$, natural GMSB, all $m_{\widetilde{ au}}$
>1330	95	¹⁹ AAD	15 _{BV} ATLS	jets $+ \not\!\!E_T$, $\widetilde{g} \to q \overline{q} \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} =$
				1 GeV
>1500	95	¹⁹ AAD	15BV ATLS	$egin{aligned} & \operatorname{T} \ \operatorname{GeV} \ & \operatorname{jets} + ot\!$
>1650	95	¹⁹ AAD	15BV ATLS	jets $+ E_T$, $m_{\widetilde{g}} = m_{\widetilde{q}}$, $m_{\widetilde{\chi}^0_1} = 1$
> 850	95	¹⁹ AAD	15BV ATLS	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$
>1270	95	¹⁹ AAD	15BV ATLS	jets $+ \cancel{E}_T$, $\widetilde{g} \rightarrow q\overline{q} W \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0}$
>1150	95	¹⁹ AAD	15BV ATLS	$=100~ ext{GeV} \ ext{jets} + \ell^{\pm}\ell^{\pm}, \widetilde{g} ightarrow q \overline{q} W Z \widetilde{\chi}_1^0, \ m_{\widetilde{\chi}_1^0} = 100~ ext{GeV}$
>1320	95	¹⁹ AAD	15BV ATLS	$\chi_1^{\widetilde{\iota}}$ jets $+$ $\ell^{\pm}\ell^{\pm}$, \widetilde{g} decays via sleptons, $m_{\widetilde{\chi}_1^0}=100~{ m GeV}$
>1220	95	¹⁹ AAD	15BV ATLS	$\chi_1^{}$ $ au$, \widetilde{q} decays via staus, $m_{\widetilde{\chi}_1^0}=100$
>1310	95	¹⁹ AAD	15BV ATLS	GeV b-jets, $\widetilde{g} \rightarrow t\overline{t}\widetilde{\chi}_{1}^{0}, m_{\widetilde{\chi}_{1}^{0}} < 400$
				GeV

>1220	95	¹⁹ AAD	15BV ATLS	b -jets, $\widetilde{g} ightarrow \widetilde{t}_1 t$ and $\widetilde{t}_1 ightarrow t \widetilde{\chi}_1^0$, $m_{\mathcal{T}_1} < 1000 \; ext{GeV}$
>1180	95	¹⁹ AAD	15BV ATLS	b -jets, $\widetilde{g} ightarrow \widetilde{t}_1 t$ and $\widetilde{t}_1 ightarrow$ $b\widetilde{\chi}_1^\pm$, $m_{{\mathcal T}_1} < 1000$ GeV, $m_{\widetilde{\chi}_1^0} = 60$ GeV
>1260	95	¹⁹ AAD	15 _{BV} ATLS	<i>b</i> -jets, $\widetilde{g} \rightarrow \widetilde{t}_1 t$ and $\widetilde{g} \rightarrow c \widetilde{\chi}_1^0$
> 880	95	¹⁹ AAD	15BV ATLS	jets, $\widetilde{g} \rightarrow \widetilde{t}_1 t$ and $\widetilde{t}_1 \rightarrow s b$, RPV, $400 < m_{\widetilde{t}_1} < 1000 \; \text{GeV}$
>1200	95	¹⁹ AAD	15BV ATLS	b -jets, $\widetilde{g} ightarrow \widetilde{b}_1 b$ and $\widetilde{b}_1 ightarrow b \widetilde{\chi}_1^0$, $m_{\widetilde{b}_1} < 1$ 000 GeV
>1250	95	¹⁹ AAD	15BV ATLS	b -jets, $\widetilde{g} \rightarrow b \overline{b} \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} < 400$
none, 750–1250	95	¹⁹ AAD	15BV ATLS	GeV b-jets, \widetilde{g} decay via offshell \widetilde{t}_1 and \widetilde{b}_1 , $m_{\widetilde{\chi}_1^0} < 500$ GeV
		⁴² AAD	15CB ATLS	ℓ , $\widetilde{g} \rightarrow (e/\mu)qq$, RPV, bench-
> 600	95	⁴² AAD	15CB ATLS	mark gluino, neutralino masses $\ell\ell/Z$, $\widetilde{g} \rightarrow (ee/\mu\mu/e\mu)qq$, RPV, $m_{\widetilde{\chi}_1^0} = 400$ GeV and 0.7
>1100	95	⁴² AAD	15CB ATLS	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
>1400	95	⁴² AAD	15CB ATLS	$<$ c $ au_{\widetilde{\chi}_1^0}$ $<$ 500 mm jets or $ ot\!$
>1500	95	⁴² AAD	15CB ATLS	E_T , $\widetilde{g} \rightarrow q q \widetilde{\chi}_1^0$, Split SUSY, $m_{\widetilde{\chi}_1^0} = 100$ GeV and $20 < 100$
>1000	95	⁴³ AAD	15X ATLS	c au < 250 mm \geq 10 jets, $\widetilde{g} \rightarrow q\overline{q}\widetilde{\chi}_{1}^{0}$, $\widetilde{\chi}_{1}^{0} \rightarrow qqq$ (RPV), $m_{\widetilde{\chi}_{2}^{0}}$ =500 GeV
> 917	95	⁴³ AAD	15X ATLS	\geq 6,7 jets, $\widetilde{g} \rightarrow qqq$, (light-
> 929	95	⁴³ AAD	15X ATLS	quark, λ'' couplings, RPV) ≥ 6.7 jets, $\tilde{g} \rightarrow qqq$, (b-quark, λ'' couplings, RPV)
		⁴⁴ KHACHATRY	15AD CMS	$\ell^{\pm}\ell^{\mp}+{\sf jets}+{\not\!\!E_T},{\sf GMSB},\widetilde{g} ightarrow$
>1300	95	⁴⁵ KHACHATRY	15AZ CMS	$q\overline{q}Z\widetilde{G}$ $\geq 2\gamma, \ \geq 1$ jet, (Razor), binolike NLSP, $m_{\widetilde{\chi}^0_1}=$ 375 GeV
> 800	95	⁴⁵ KHACHATRY	15AZ CMS	$\geq 1 \ \gamma$, $\geq 2 \ \mathrm{jet}$, wino-like NLSP, $m_{\widetilde{\chi}_1^0} = 375 \ \mathrm{GeV}$
>1280	95	⁴⁶ AAD	14AX ATLS	\geq 3 b -jets $+ ot \!$

>1250	95	⁴⁶ AAD	14AX ATLS	$ \geq 3 \text{ b-jets} + \not\!\!E_T, \ \widetilde{g} \to \ \widetilde{b}_1 b \widetilde{\chi}_1^0 \\ \text{simplified model, } \widetilde{b}_1 \to b \widetilde{\chi}_1^0, \\ m_{\widetilde{\chi}_1^0} = \text{60 GeV, } m_{\widetilde{b}_1} < \text{900} $
>1190	95	⁴⁶ AAD	14AX ATLS	$\begin{array}{l} \text{GeV} \\ \geq 3 \text{ b-jets} + \not\!\!E_T$, $\widetilde{g} \to $\widetilde{t}_1 t \widetilde{\chi}_1^0$ \\ \text{simplified model, } \widetilde{t}_1 \to t \widetilde{\chi}_1^0$, \\ m_{\widetilde{\chi}_1^0} = \text{60 GeV, } m_{\widetilde{t}_1} < 1000 \end{array}$
>1180	95	⁴⁶ AAD	14AX ATLS	$\begin{array}{l} \text{GeV} \\ \geq 3 \text{ b-jets} + \not\!\!E_T$, $\widetilde{g} \to \widetilde{t}_1 t \widetilde{\chi}_1^0 \\ \text{simplified model, } \widetilde{t}_1 \to b \widetilde{\chi}_1^\pm, \\ m_{\widetilde{\chi}_1^\pm} = 2m_{\widetilde{\chi}_1^0}, m_{\widetilde{\chi}_1^0} = 60 \text{ GeV}, \\ m_{\widetilde{t}_1} < 1000 \text{ GeV} \end{array}$
>1250	95	⁴⁶ AAD	14AX ATLS	\geq 3 <i>b</i> -jets $+ \not\!\!E_T$, $\widetilde{g} \to b \overline{b} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} <$ 400
>1340	95	⁴⁶ AAD	14AX ATLS	GeV \geq 3 <i>b</i> -jets $+ \not\!\!E_T$, $\widetilde{g} \to t \overline{t} \widetilde{\chi}_1^0$ simplified model, $m_{\widetilde{\chi}_1^0} < 400$
>1300	95	⁴⁶ AAD	14AX ATLS	GeV ≥ 3 b-jets $+ E_T$, $\widetilde{g} \rightarrow t \overline{b} \widetilde{\chi}_1^{\pm}$ simplified model, $\widetilde{\chi}_1^{\pm} \rightarrow f f' \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^{\pm}} - m_{\widetilde{\chi}_1^0} = 2$ GeV,
050	0.5	⁴⁷ AAD	14- ATIC	$m_{\widetilde{\chi}_1^0} < 300 \text{ GeV}$
> 950	95		14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp}) + ext{jets}, \ \widetilde{g} ightarrow t \overline{t} \widetilde{\chi}_1^0$ simplified model
>1000	95	⁴⁷ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp})$ + jets, $\widetilde{g} \rightarrow t\widetilde{t}_1$ with $\widetilde{t}_1 \rightarrow b\widetilde{\chi}_1^{\pm}$ simplified model, $m_{\widetilde{t}_1} <$ 200 GeV, $m_{\widetilde{\chi}_1^{\pm}}$ = 118 GeV, $m_{\widetilde{\chi}_1^0} =$ 60 GeV
> 640	95	⁴⁷ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp}) + \mathrm{jets}, \ \widetilde{g} \rightarrow t \ \widetilde{t}_1$ with $\widetilde{t}_1 \rightarrow c \ \widetilde{\chi}_1^0 \ \mathrm{simplified}$ model, $m_{\widetilde{t}_1} = m_{\widetilde{\chi}_1^0} + 20 \ \mathrm{GeV}$
> 850	95	⁴⁷ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp})$ + jets, $\widetilde{g} \rightarrow t\widetilde{t}_1$ with $\widetilde{t}_1 \rightarrow bs$ simplified model, R
> 860	95	⁴⁷ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp}) + \text{jets, } \widetilde{g} \rightarrow q q' \widetilde{\chi}_{1}^{\pm},$ $\widetilde{\chi}_{1}^{\pm} \rightarrow W^{(*)} \pm \widetilde{\chi}_{1}^{0} \text{ simplified model, } m_{\widetilde{\chi}_{1}^{\pm}} = 2 m_{\widetilde{\chi}_{1}^{0}},$ $m_{\widetilde{\chi}_{1}^{0}} < 400 \text{ GeV}$
>1040	95	⁴⁷ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp})$ + jets, $\widetilde{g} \rightarrow q q' \widetilde{\chi}_{1}^{\pm}$, $\widetilde{\chi}_{1}^{\pm} \rightarrow W^{(*)\pm}\widetilde{\chi}_{2}^{0}$, $\widetilde{\chi}_{2}^{0} \rightarrow Z^{(*)}\widetilde{\chi}_{1}^{0}$ simplified model, $m_{\widetilde{\chi}_{1}^{0}} < 520 \text{ GeV}$

>1200	95	⁴⁷ AAD	14E ATLS	$\ell^{\pm}\ell^{\pm}(\ell^{\mp}) + \mathrm{jets}, \ \widetilde{g} ightarrow q \ q' \widetilde{\chi}_{1}^{\pm}/\widetilde{\chi}_{2}^{0}, \ \widetilde{\chi}_{1}^{\pm} ightarrow \ell^{\pm} u \widetilde{\chi}_{1}^{0}, \ \widetilde{\chi}_{2}^{0} ightarrow \ell^{\pm} (u u) \widetilde{\chi}_{1}^{0} \ \mathrm{simplification}$
>1050	95	48 CHATRCHYAN	N 14H CMS	fied model same-sign $\ell^\pm\ell^\pm$, $\widetilde{g} ightarrow t \overline{t} \widetilde{\chi}_1^0$
				simplified model, massless $\widetilde{\chi}_1^0$
> 900	95	⁴⁹ CHATRCHYAN	N 14H CMS	same-sign $\ell^{\pm}\ell^{\pm}$, $\widetilde{g} ightarrow q q' \widetilde{\chi}_{1}^{\pm}$,
				$\widetilde{\chi}_1^\pm o \ {\it W}^\pm \widetilde{\chi}_1^0$ simplified
				model, $m_{\widetilde{\chi}_1^{\pm}} = 0.5 \ m_{\widetilde{g}}$, mass-
				less $\widetilde{\chi}_1^0$
>1050	95	⁵⁰ CHATRCHYAI	N 14H CMS	same-sign $\ell^\pm\ell^\pm$, $\widetilde{g} ightarrow b \overline{t} \widetilde{\chi}_1^\pm$,
				$\widetilde{\chi}_1^\pm ightarrow \ W^\pm \widetilde{\chi}_1^0$ simplified
				model, $m_{\widetilde{\chi}_1^{\pm}} = 300$ GeV, $m_{\widetilde{\chi}_1^0}$
		F1		$= 50 \text{ GeV}$ same-sign $\ell^{\pm}\ell^{\pm}$, $\widetilde{g} \rightarrow tbs$ sim-
> 900	95	⁵¹ CHATRCHYAI	N 14H	same-sign $\ell^{\pm}\ell^{\pm}$, $\widetilde{g} \rightarrow tbs$ simplified model, R

 $^{^1}$ KHACHATRYAN 17 searched in 2.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events containing four or more jets, no more than one lepton, and missing transverse momentum, using the razor variables (M_R and R^2) to discriminate between signal and background processes. No evidence for an excess over the expected background is observed. Limits are derived on the gluino mass in the Tglu1A, Tglu2A and Tglu3A simplified models, see Figs. 16 and 17. Also, assuming gluinos decay only via three-body processes involving third-generation quarks plus a neutralino/chargino, and assuming $m_{\widetilde{\chi}_1^\pm}=m_{\widetilde{\chi}_1^0}+5$ GeV,

a branching ratio-independent limit on the gluino mass is given, see Fig. 16.

 $^{^2}$ AABOUD 16AC searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV in final states with hadronic jets, 1 or two hadronically decaying τ and E_T . In Tglu1F, gluino masses are excluded at 95% C.L. up to 1570 GeV for neutralino masses of 100 GeV or below. Neutralino masses up to 700 GeV are excluded for all gluino masses between 800 GeV and 1500 GeV, while the strongest neutralino-mass exclusion of 750 GeV is achieved for gluino masses around 1400 GeV. See their Fig. 8. Limits are also presented in the context of Gauge-Mediated Symmetry Breaking models: in this case, values of Λ below 92 TeV are excluded at the 95% CL, corresponding to gluino masses below 2000 GeV. See their Fig. 9.

³ AABOUD 16J searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV in final states with one isolated electron or muon, hadronic jets, and E_T . Gluino-mediated pair production of stops with a nearly mass-degenerate stop and neutralino are targeted and gluino masses are excluded at 95% C.L. up to 1460 GeV. A 100% of stops decaying via charm + neutralino is assumed. The results are also valid in case of 4-body decays $t_1 \rightarrow ff'b\tilde{\chi}_1^0$. See their Fig. 8.

⁴AABOUD 16M searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for events with two photons, hadronic jets and E_T . No significant excess above the Standard Model expectations is observed. Exclusion limits at 95% C.L. are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for bino-like NLSP. See their Fig. 3.

 $^{^5}$ AABOUD 16N searched in 3.2 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events containing hadronic jets, large E_T , and no electrons or muons. No significant excess above the Standard Model expectations is observed. Gluino masses below 1510 GeV are excluded at the 95% C.L. in a simplified model with only gluinos and the lightest neutralino. See their Fig. 7b.

- 7 AAD 16AD searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events containing several energetic jets, of which at least three must be identified as b-jets, large $\not\!\!E_T$ and no electrons or muons. No significant excess above the Standard Model expectations is observed. For $\widetilde{\chi}^0_1$ below 800 GeV, gluino masses below 1780 GeV are excluded at 95% C.L. for gluinos decaying via bottom squarks. See their Fig. 7a.
- ⁸ AAD 16AD searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for events containing several energetic jets, of which at least three must be identified as b-jets, large $\not\!\!\!E_T$ and one electron or muon. Large-radius jets with a high mass are also used to identify highly boosted top quarks. No significant excess above the Standard Model expectations is observed. For $\widetilde{\chi}_1^0$ below 700 GeV, gluino masses below 1760 GeV are excluded at 95% C.L. for gluinos decaying via top squarks. See their Fig. 7b.
- ⁹AAD 16BB searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for events with exactly two same-sign leptons or at least three leptons, multiple hadronic jets, b-jets, and E_T . No significant excess over the Standard Model expectation is found. Exclusion limits at 95% C.L. are set on the gluino mass in various simplified models (Tglu1D, Tglu1E, Tglu3A). See their Figs. 4.a, 4.b, and 4.d.
- 10 AAD 16BG searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV in final states with one isolated electron or muon, hadronic jets, and E_T . The data agree with the SM background expectation in the six signal selections defined in the search, and the largest deviation is a 2.1 standard deviation excess. Gluinos are excluded at 95% C.L. up to 1600 GeV assuming they decay via the lightest chargino to the lightest neutralino as in the model Tglu1B for $m_{\widetilde{\chi}_1^0}{=}100$ GeV, assuming $m_{\widetilde{\chi}_1^\pm}{=}(m_{\widetilde{g}}+m_{\widetilde{\chi}_1^0})/2$. See their Fig. 6.
- 11 AAD 16V searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with E_T various hadronic jet multiplicities from ≥ 7 to ≥ 10 and with various b-jet multiplicity requirements. No significant excess over the Standard Model expectation is found. Exclusion limits at 95% C.L. are set on the gluino mass in one simplified model (Tglu1E) and a pMSSM-inspired model. See their Fig. 5.
- 12 KHACHATRYAN 16AM searched in 19.7 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with highly boosted W-bosons and b-jets, using the razor variables (M_R and R^2) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the Tglu3C and Tglu3B simplified models, see Fig. 12.
- 13 KHACHATRYAN 16BJ searched in 2.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the following simplified models: Tglu3A and Tglu3D, see Fig. 4, Tglu3B and Tglu3C, see Fig. 5, and Tglu1B, see Fig. 7.
- 14 KHACHATRYAN 16BS searched in 2.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with at least one energetic jet , no isolated leptons, and significant $\not\!\! E_T$, using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the Tglu1A, Tglu2A and Tglu3A simplified models, see Fig. 10 and Table 3.
- ¹⁵ KHACHATRYAN 16BX searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing 0 or 1 leptons and b-tagged jets, coming from R-parity-violating decays of supersymmetric particles. No excess over the expected background is observed. Limits are derived on the gluino mass, assuming the RPV $\tilde{g} \rightarrow tbs$ decay, see Fig. 7 and 10.

- 16 KHACHATRYAN 16 Searched in $^{2.3}$ fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with two opposite-sign, same-flavour leptons, jets, and missing transverse momentum. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the Tglu4C simplified model, see Fig. 4, and on sbottom masses in the Tsbot3 simplified model, see Fig. 5.
- 17 KHACHATRYAN 16V searched in 2.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with at least four energetic jets and significant E_T , no identified isolated electron or muon or charged track. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the Tglu1A, Tglu1C, Tglu2A, and Tglu3A simplified models, see Fig. 8.
- 18 AAD 15 BG searched in $^{20.3}$ fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with jets, missing E_T , and two opposite-sign same flavor isolated leptons featuring either a kinematic edge, or a peak at the Z-boson mass, in the invariant mass spectrum. No evidence for a statistically significant excess over the expected SM backgrounds are observed and 95% C.L. exclusion limits are derived in a GGM simplified model of gluino pair production where the gluino decays into quarks, a Z-boson, and a massless gravitino LSP, see Fig. 12. Also, limits are set in simplified models with slepton/sneutrino intermediate states, see Fig. 13.
- 19 AAD 15BV summarized and extended ATLAS searches for gluinos and first- and second-generation squarks in final states containing jets and missing transverse momentum, with or without leptons or b-jets in the \sqrt{s} =8 TeV data set collected in 2012. The paper reports the results of new interpretations and statistical combinations of previously published analyses, as well as new analyses. Exclusion limits at 95% C.L. are set on the gluino mass in several R-parity conserving models, leading to a generalized constraint on gluino masses exceeding 1150 GeV for lightest supersymmetric particle masses below 100 GeV. See their Figs. 10, 19, 20, 21, 23, 25, 26, 29-37.
- AAD 15BX interpreted the results of a wide range of ATLAS direct searches for supersymmetry, during the first run of the LHC using the \sqrt{s} =7 TeV and \sqrt{s} = 8 TeV data set collected in 2012, within the wider framework of the phenomenological MSSM (pMSSM). The integrated luminosity was up to 20.3 fb⁻¹. From an initial random sampling of 500 million pMSSM points, generated from the 19-parameter pMSSM, a total of 310,327 model points with $\widetilde{\chi}_1^0$ LSP were selected each of which satisfies constraints from previous collider searches, precision measurements, cold dark matter energy density measurements and direct dark matter searches. The impact of the ATLAS Run 1 searches on this space was presented, considering the fraction of model points surviving, after projection into two-dimensional spaces of sparticle masses. Good complementarity is observed between different ATLAS analyses, with almost all showing regions of unique sensitivity. ATLAS searches have good sensitivity at LSP mass below 800 GeV.
- 21 AAD 15CA searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with one or more photons, hadronic jets or b-jets and E_T . No significant excess above the Standard Model expectations is observed. Limits are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for bino-like or higgsino-bino admixtures NLSP, see Fig. 8, 10, 11
- ²² KHACHATRYAN 15AF searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets and significant E_T , using the transverse mass variable M_{T2} to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay $\tilde{g} \to q \bar{q} \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 13(a), or where the decay $\tilde{g} \to b \bar{b} \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 13(b), or where the decay $\tilde{g} \to t \bar{t} \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 13(c). See also Table 5. Exclusions in the CMSSM, assuming tan $\beta=30$, $A_0=-2$ max $(m_0,m_{1/2})$ and $\mu>0$, are also presented, see Fig. 15.
- 23 KHACHATRYAN 151 searched in 19.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events in which b-jets and four W-bosons are produced. Five individual search channels are combined (fully hadronic, single lepton, same-sign dilepton, opposite-sign dilepton, multilepton). No significant excess above the Standard Model expectations is observed.

Limits are set on the gluino mass in a simplified model where the decay $\tilde{g} \to t \bar{t} \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 5. Also a simplified model with gluinos decaying into on-shell top squarks is considered, see Fig. 6.

- ²⁴ KHACHATRYAN 15X searched in 19.3fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least two energetic jets, at least one of which is required to originate from a b quark, and significant $\not\!\!E_T$, using the razor variables (M_R) and R^2) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay $g \to b \, \overline{b} \, \chi_1^0$ and the decay $g \to t \, \overline{t} \, \chi_1^0$ take place with branching ratios varying between 0, 50 and 100%, see Figs. 13 and 14.
- ²⁵ AAD 14AE searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for strongly produced supersymmetric particles in events containing jets and large missing transverse momentum, and no electrons or muons. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos and squarks, see Figures 5, 6 and 7. Limits are also derived in the mSUGRA/CMSSM with parameters $\tan\beta=30$, $A_0=-2$ m_0 and $\mu>0$, see their Fig. 8.
- 26 AAD 14AG searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing one hadronically decaying τ -lepton, zero or one additional light leptons (electrons or muons), jets and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set in several SUSY scenarios. For an interpretation in the minimal GMSB model, see their Fig. 8. For an interpretation in the mSUGRA/CMSSM with parameters $\tan\beta=30,\,A_0=-2\,m_0$ and $\mu>0$, see their Fig. 9. For an interpretation in the framework of natural Gauge Mediation, see Fig. 10. For an interpretation in the bRPV scenario, see their Fig. 11.
- AAD 14X searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in an R-parity violating simplified model where the decay $\tilde{g}\to q\overline{q}\tilde{\chi}_1^0$, with $\tilde{\chi}_1^0\to \ell^\pm\ell^\mp\nu$, takes place with a branching ratio of 100%, see Fig. 8.
- 28 AAD 14X searched in 20.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least four leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a general gauge-mediation model (GGM) where the decay $\tilde{g} \to q \bar{q} \tilde{\chi}_1^0$, with $\tilde{\chi}_1^0 \to \ell^\pm \ell^\mp \tilde{G}$, takes place with a branching ratio of 100%, for two choices of $\tan\beta=1.5$ and 30, see Fig. 11. Also some constraints on the higgsino mass parameter μ are discussed.
- 29 CHATRCHYAN 14AH searched in 4.7 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events with at least two energetic jets and significant E_T , using the razor variables (M_R and R^2) to discriminate between signal and background processes. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay $\tilde{g} \to q \bar{q} \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 28. Exclusions in the CMSSM, assuming $\tan\beta=10$, $A_0=0$ and $\mu>0$, are also presented, see Fig. 26.
- 30 CHATRCHYAN 14AH searched in 4.7 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events with at least two energetic jets and significant E_T , using the razor variables (M_R and R^2) to discriminate between signal and background processes. A second analysis requires at least one of the jets to be originating from a b-quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay $\tilde{g} \to b\bar{b}\tilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming $\tan\beta=10,\,A_0=0$ and $\mu>0,$ are also presented, see Fig. 26.
- ³¹ CHATRCHYAN 14AH searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with at least two energetic jets and significant $\not\!\!E_T$, using the razor variables (M_R and R^2) to discriminate between signal and background processes. A second analysis requires

- at least one of the jets to be originating from a b-quark. No significant excess above the Standard Model expectations is observed. Limits are set on sbottom masses in simplified models where the decay $\widetilde{g} \to t \overline{t} \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Figs. 28 and 29. Exclusions in the CMSSM, assuming $\tan\beta=10$, $A_0=0$ and $\mu>0$, are also presented, see Fig. 26.
- ³²CHATRCHYAN 14I searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing multijets and large E_T . No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos that decay via $\widetilde{g} \to q \overline{q} \widetilde{\chi}_1^0$ with a 100% branching ratio, see Fig. 7b, or via $\widetilde{g} \to t \overline{t} \widetilde{\chi}_1^0$ with a 100% branching ratio, see Fig. 7c, or via $\widetilde{g} \to q \overline{q} W/Z \widetilde{\chi}_1^0$, see Fig. 7d.
- 33 CHATRCHYAN 14N searched in 19.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing a single isolated electron or muon and multiple jets, at least two of which are identified as originating from a b-quark. No significant excesses over the expected SM backgrounds are observed. The results are interpreted in three simplified models of gluino pair production with subsequent decay into virtual or on-shell top squarks, where each of the top squarks decays in turn into a top quark and a $\widetilde{\chi}_1^0$, see Fig. 4. The models differ in which masses are allowed to vary.
- 34 CHATRCHYAN 14P searched in 19.4 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for three-jet resonances produced in the decay of a gluino in R-parity violating supersymmetric models. No excess over the expected SM background is observed. Assuming a 100% branching ratio for the gluino decay into three light-flavour jets, limits are set on the cross section of gluino pair production, see Fig. 7, and gluino masses below 650 GeV are excluded at 95% C.L. Assuming a 100% branching ratio for the gluino decaying to one b-quark jet and two light-flavour jets, gluino masses between 200 GeV and 835 GeV are excluded at 95% C L.
- 35 CHATRCHYAN 14R searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a slepton co-NLSP simplified model (GMSB) where the decay $\tilde{g} \to q \ell^{\pm} \ell^{\mp} \tilde{G}$ takes place with a branching ratio of 100%, see Fig. 8.
- takes place with a branching ratio of 100%, see Fig. 8. $^{36} \text{ CHATRCHYAN 14R searched in 19.5 fb}^{-1} \text{ of } pp \text{ collisions at } \sqrt{s} = 8 \text{ TeV for events with at least three leptons (electrons, muons, taus) in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in a simplified model where the decay <math>\widetilde{g} \to t \overline{t} \widetilde{\chi}_1^0$ takes place with a branching ratio of 100%, see Fig. 11.
- 37 KHACHATRYAN 16AY searched in 2.3 fb $^{-1}$ of pp collisions at $\sqrt{s}=13$ TeV for events with one isolated high transverse momentum lepton (e or μ), hadronic jets of which at least one is identified as coming from a b-quark, and large E_T . No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in the Tglu3A simplified model, see Fig. 10, and in the Tglu3B model, see Fig. 11.
- 38 KHACHATRYAN 16BT performed a global Bayesian analysis of a wide range of CMS results obtained with data samples corresponding to 5.0 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV and in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV. The set of searches considered, both individually and in combination, includes those with all-hadronic final states, same-sign and opposite-sign dileptons, and multi-lepton final states. An interpretation was given in a scan of the 19-parameter pMSSM. No scan points with a gluino mass less than 500 GeV survived and 98% of models with a squark mass less than 300 GeV were excluded.
- 39 KHACHATRYAN 16BX searched in 19.5 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events containing 4 leptons coming from R-parity-violating decays of $\widetilde{\chi}_1^0 \to \ell\ell\nu$ with $\lambda_{121} \neq 0$ or $\lambda_{122} \neq 0$. No excess over the expected background is observed. Limits are derived on the gluino, squark and stop masses, see Fig. 23.
- ⁴⁰ AAD 15AB searched for the decay of neutral, weakly interacting, long-lived particles in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV. Signal events require at least two reconstructed vertices possibly originating from long-lived particles decaying to jets in the inner tracking

detector and muon spectrometer. No significant excess of events over the expected background was found. Results were interpreted in Stealth SUSY benchmark models where a pair of gluinos decay to long-lived singlinos, \widetilde{S} , which in turn each decay to a low-mass gravitino and a pair of jets. The 95% confidence-level limits are set on the cross section \times branching ratio for the decay $\widetilde{g} \to \widetilde{S} g$, as a function of the singlino proper lifetime ($c\tau$). See their Fig. 10(f)

- ⁴¹ AAD 15AI searched in 20 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events containing at least one isolated lepton (electron or muon), jets, and large missing transverse momentum. No excess of events above the expected level of Standard Model background was found. Exclusion limits at 95% C.L. are set on the gluino mass in the CMSSM/mSUGRA, see Fig. 15, in the NUHMG, see Fig. 16, and in various simplified models, see Figs. 18–22.
- 42 AAD 15CB searched for events containing at least one long-lived particle that decays at a significant distance from its production point (displaced vertex, DV) into two leptons or into five or more charged particles in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV. The dilepton signature is characterised by DV formed from at least two lepton candidates. Four different final states were considered for the multitrak signature, in which the DV must be accompanied by a high-transverse momentum muon or electron candidate that originates from the DV, jets or missing transverse momentum. No events were observed in any of the signal regions. Results were interpreted in SUSY scenarios involving R-parity violation, split supersymmetry, and gauge mediation. See their Fig. 12–20.
- 43 AAD 15X searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events containing large number of jets, no requirements on missing transverse momentum and no isolated electrons or muons. The sensitivity of the search is enhanced by considering the number of b-tagged jets and the scalar sum of masses of large-radius jets in an event. No evidence was found for excesses above the expected level of Standard Model background. Exclusion limits at 95% C.L. are set on the gluino mass assuming the gluino decays to various quark flavors, and for various neutralino masses. See their Fig. 11–16.
- 44 KHACHATRYAN 15AD searched in 19.4 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with two opposite-sign same flavor isolated leptons featuring either a kinematic edge, or a peak at the Z-boson mass, in the invariant mass spectrum. No evidence for a statistically significant excess over the expected SM backgrounds is observed and 95% C.L. exclusion limits are derived in a simplified model of gluino pair production where the gluino decays into quarks, a Z-boson, and a massless gravitino LSP, see Fig. 9.
- 45 KHACHATRYAN 15AZ searched in 19.7 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV for events with either at least one photon, hadronic jets and E_T (single photon channel) or with at least two photons and at least one jet and using the razor variables. No significant excess above the Standard Model expectations is observed. Limits are set on gluino masses in the general gauge-mediated SUSY breaking model (GGM), for both a bino-like and wino-like neutralino NLSP scenario, see Fig. 8 and 9.
- 46 AAD 14AX searched in 20.1 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for the strong production of supersymmetric particles in events containing either zero or at last one high high- p_T lepton, large missing transverse momentum, high jet multiplicity and at least three jets identified as originating from b-quarks. No excess over the expected SM background is observed. Limits are derived in mSUGRA/CMSSM models with $\tan\beta=30,\,A_0=-2m_0$ and $\mu>0$, see their Fig. 14. Also, exclusion limits in simplified models containing gluinos and scalar top and bottom quarks are set, see their Figures 12, 13.
- 47 AAD 14E searched in 20.3 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for strongly produced supersymmetric particles in events containing jets and two same-sign leptons or three leptons. The search also utilises jets originating from b-quarks, missing transverse momentum and other variables. No excess over the expected SM background is observed. Exclusion limits are derived in simplified models containing gluinos and squarks, see Figures 5 and 6. In the $\tilde{g} \rightarrow qq'\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow W^{(*)\pm}\tilde{\chi}_2^0, \tilde{\chi}_2^0 \rightarrow Z^{(*)}\tilde{\chi}_1^0$ simplified model, the following assumptions have been made: $m_{\tilde{\chi}_1^{\pm}}=0.5 \ m_{\tilde{\chi}_1^0}+m_{\tilde{g}}, m_{\tilde{\chi}_2^0}=0.5 \ (m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^\pm}), m_{\tilde{\chi}_1^0}<520$ GeV. In the $\tilde{g} \rightarrow qq'\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow \ell^{\pm}\nu\tilde{\chi}_1^0$ or $\tilde{g} \rightarrow 0.5 \ (m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^\pm}), m_{\tilde{\chi}_1^0}<520$ GeV. In the $\tilde{g} \rightarrow qq'\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow \ell^{\pm}\nu\tilde{\chi}_1^0$ or $\tilde{g} \rightarrow 0.5 \ (m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}), m_{\tilde{\chi}_1^0}=0.5 \ (m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}), m_{\tilde{\chi}_1^0}=0.5 \ (m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi}_1^0}+m_{\tilde{\chi$

- $q\,q'\,\widetilde{\chi}_2^0,\ \widetilde{\chi}_2^0 \to \ell^\pm\ell^\mp(\nu\nu)\widetilde{\chi}_1^0$ simplified model, the following assumptions have been made: $m_{\widetilde{\chi}_1^\pm} = m_{\widetilde{\chi}_2^0} = 0.5\ (m_{\widetilde{\chi}_1^0} + m_{\widetilde{g}}),\ m_{\widetilde{\chi}_1^0} < 660\ {\rm GeV}.$ Limits are also derived in the mSUGRA/CMSSM, bRPV and GMSB models, see their Fig. 8.
- ⁴⁸ CHATRCHYAN 14H searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay $\tilde{g} \to t \bar{t} \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, or where the decay $\tilde{g} \to \tilde{t}t$, $\tilde{t} \to t \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, with varying mass of the $\tilde{\chi}_1^0$, or where the decay $\tilde{g} \to \tilde{b}b$, $\tilde{b} \to t \tilde{\chi}_1^\pm$, $\tilde{\chi}_1^\pm \to W^\pm \tilde{\chi}_1^0$ takes place with a branching ratio of 100%, with varying mass of the $\tilde{\chi}_1^\pm$, see Fig. 5.
- ⁴⁹ CHATRCHYAN 14H searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay $\tilde{g} \rightarrow qq'\tilde{\chi}_1^{\pm}$, $\tilde{\chi}_1^{\pm} \rightarrow W^{\pm}\tilde{\chi}_1^0$ takes place with a branching ratio of 100%, with varying mass of the $\tilde{\chi}_1^{\pm}$ and $\tilde{\chi}_1^0$, see Fig. 7.
- 50 CHATRCHYAN 14H searched in 19.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the decay $\tilde{g}\to b\bar{t}\tilde{\chi}_1^\pm,\,\tilde{\chi}_1^\pm\to\,W^\pm\tilde{\chi}_1^0$ takes place with a branching ratio of 100%, for two choices of $m_{\tilde{\chi}_1^\pm}$ and fixed $m_{\tilde{\chi}_1^0}$, see Fig. 6.
- ⁵¹ CHATRCHYAN 14H searched in 19.5 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with two isolated same-sign dileptons and jets in the final state. No significant excess above the Standard Model expectations is observed. Limits are set on the gluino mass in simplified models where the R-parity violating decay $\tilde{g} \rightarrow tbs$ takes place with a branching ratio of 100%, see Fig. 8.

Long-lived \widetilde{g} (Gluino) MASS LIMIT

Limits on light gluinos ($m_{\widetilde{g}} < 5 \text{ GeV}$) were last listed in our PDG 14 edition: K. Olive, et al. (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (GeV)	CL%	DOCUMENT ID		TECN	COMMENT
>1580	95	¹ AABOUD	16 B	ATLS	long-lived R-hadrons
> 740–1590	95	² AABOUD	16 C	ATLS	R-hadrons, Tglu1A, $ au \geq 0.4$ ns, $m_{\widetilde{\chi}^0_1} = 100~{ m GeV}$
>1570	95	² AABOUD		ATLS	R-hadrons, Tglu1A, stable
>1610	95	³ KHACHATRY.			long-lived \widetilde{g} forming R-hadrons, $f=0.1$, cloud interaction model
>1580	95	³ KHACHATRY.			long-lived \widetilde{g} forming R-hadrons, $f=0.1$, charge-suppressed interaction model
>1520	95	³ KHACHATRY.			long-lived \tilde{g} forming R-hadrons, $f = 0.5$, cloud interaction model
>1540	95	³ KHACHATRY.	16 BV	vCMS	long-lived \tilde{g} forming R-hadrons, f = 0.5, charge-suppressed interaction model
>1270	95	⁴ AAD	15AE	ATLS	\widetilde{g} R-hadron, generic R-hadron model
HTTP://PDG.	LBL.GO	V Page	e 79		Created: 5/30/2017 17:22

>1360	95	⁴ AAD	15AE ATLS	
>1115	95	⁵ AAD	15BM ATLS	sleptons, LeptoSUSY model \widetilde{g} R-hadron, stable
>1185	95	⁵ AAD	15BM ATLS	_
				ns, $m_{\widetilde{\chi}^0_1}=1\bar{0}0$ GeV
>1099	95	⁵ AAD	15BM ATLS	^
				ns, $m_{\widetilde{g}} - m_{\widetilde{\chi}_1^0}^- = 100$ GeV
>1182	95	⁵ AAD	15BM ATLS	$\widetilde{g} ightarrow t \overline{t} \widetilde{\chi}_1^0$, lifetime 10 ns,
				$\widetilde{g} ightarrow t \overline{t} \widetilde{\chi}_1^0$, lifetime 10 ns, $m_{\widetilde{\chi}_1^0} = 100 \; ext{GeV}$
>1157	95	⁵ AAD	15BM ATLS	$\widetilde{g} \rightarrow t \overline{t} \widetilde{\chi}_1^0$, lifetime 10 ns,
				$m_{\widetilde{g}} - m_{\widetilde{\chi}_1^0} = 480 \text{ GeV}$
> 869	95	⁵ AAD	15BM ATLS	$\widetilde{\varphi} \to (\varphi/q\overline{q})\widetilde{\chi}_{0}^{0}$ lifetime 1
<i>y</i> 303		,	1057 (1. 20	$\widetilde{g} ightarrow (g/q\overline{q})\widetilde{\chi}_1^0$, lifetime 1 ns, $m_{\widetilde{\chi}_1^0}=100~{ m GeV}$
> 821	95	⁵ AAD	15BM ATLS	$\widetilde{g} \rightarrow (g/q\overline{q})\widetilde{\chi}_1^0$, lifetime
				$\widetilde{g} ightarrow (g/q\overline{q})\widetilde{\chi}^0_1$, lifetime 1 ns, $m_{\widetilde{g}}-m_{\widetilde{\chi}^0_1}=100$
		E		GeV
> 836	95	⁵ AAD	15BM ATLS	GeV $\widetilde{g} \rightarrow t \overline{t} \widetilde{\chi}_1^0$, lifetime 1 ns,
		_		$m_{\widetilde{\chi}_1^0} = 100 \text{ GeV}$
> 836	95	⁵ AAD	15BM ATLS	$\widetilde{g} \rightarrow t \overline{t} \widetilde{\chi}_1^0$, lifetime 10 ns,
				$m_{\widetilde{\chi}_1^0} = 100 \text{ GeV}$ $\widetilde{g} \rightarrow t \overline{t} \widetilde{\chi}_1^0$, lifetime 10 ns, $m_{\widetilde{g}} - m_{\widetilde{\chi}_1^0} = 480 \text{ GeV}$
>1000	95	⁶ KHACHATRY	15AK CMS	\widetilde{g} R-hadrons, $10~\mu \mathrm{s} < au < 1000$
> 880	95	⁶ KHACHATRY	15AK CMS	\widetilde{g} R-hadrons, 1 μ s $< au$ $<$ 1000 s
\bullet \bullet We do not	use the f	ollowing data for a	verages, fits,	
> 985	95	⁷ AAD	13AA ATLS	
> 832	95	⁸ AAD	13BC ATLS	tion model $\widetilde{g} ightarrow \left. ar{g} / q \overline{q} \widetilde{\chi}_1^0 , ight.$
,				generic R-hadron model,
				lifetime between 10^{-5} and
				10^3 s, $m_{\widetilde{\chi}^0_1}=100$ GeV
>1322	95	⁹ CHATRCHYA	N 13AB CMS	long-lived \widetilde{g} forming R-hadrons, f = 0.1, cloud
		10		interaction model
none 200–341	95	¹⁰ AAD	12P ATLS	long-lived $\widetilde{g} \rightarrow g \widetilde{\chi}_1^0$, $m_{\widetilde{\chi}_1^0} =$
> 640	95	¹¹ CHATRCHYA	N 12AN CMS	$100 \; GeV$ long-lived $\widetilde{g} o \; g \widetilde{\chi}_1^0$
>1098	95 95	12 CHATRCHYAI		long-lived $\widetilde{g} \rightarrow g \chi_1$ long-lived \widetilde{g} forming R -
× 1030				hadrons, $f = 0.1$
> 586	95	¹³ AAD	11K ATLS	S
> 544	95	¹⁴ AAD	11P ATLS	stable \widetilde{g} , GMSB scenario, $tan\beta = 5$
> 370	95	¹⁵ KHACHATRY		long lived \widetilde{g}
> 398	95	¹⁶ KHACHATRY	11c CMS	stable \widetilde{g}

 $^{^1}$ AABOUD 16B searched in 3.2 fb $^{-1}$ of pp collisions at $\sqrt{s}=$ 13 TeV for long-lived R-hadrons using observables related to large ionization losses and slow propagation velocities are considered to the search of locities, which are signatures of heavy charged particles traveling significantly slower than the speed of light. Exclusion limits at 95% C.L. are set on the long-lived gluino masses exceeding 1580 GeV. See their Fig. 5.

- ² AABOUD 16C searched in 3.2 fb⁻¹ of pp collisions at $\sqrt{s}=13$ TeV for long-lived and stable R-hadrons identified by anomalous specific ionization energy loss in the ATLAS Pixel detector. Gluino R-hadrons with lifetimes above 0.4 ns are excluded at 95% C.L. with lower mass limit range between 740 GeV and 1590 GeV. In the case of stable R-hadrons, the lower mass limit is 1570 GeV. See their Figs. 5 and 6.
- 3 KHACHATRYAN 16BW searched in 2.5 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=13$ TeV for events with heavy stable charged particles, identified by their anomalously high energy deposits in the silicon tracker and/or long time-of-flight measurements by the muon system. No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass, depending on the interaction model and on the fraction f, of produced gluinos hadronizing into a \widetilde{g} gluon state, see Fig. 4 and Table 7.
- ⁴ AAD 15AE searched in 19.1 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for heavy long-lived charged particles, measured through their specific ionization energy loss in the ATLAS pixel detector or their time-of-flight in the ALTAS muon system. In the absence of an excess of events above the expected backgrounds, limits are set R-hadrons in various scenarios, see Fig. 11. Limits are also set in LeptoSUSY models where the gluino decays to stable 300 GeV leptons, see Fig. 9.
- ⁵ AAD 15BM searched in 18.4 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for stable and metastable non-relativistic charged particles through their anomalous specific ionization energy loss in the ATLAS pixel detector. In absence of an excess of events above the expected backgrounds, limits are set within a generic R-hadron model, on stable gluino R-hadrons (see Table 5) and on metastable gluino R-hadrons decaying to $(g/q\overline{q})$ plus a light $\widetilde{\chi}_1^0$ (see Fig. 7) and decaying to $t\overline{t}$ plus a light $\widetilde{\chi}_1^0$ (see Fig. 9).
- 6 KHACHATRYAN 15AK looked in a data set corresponding to 18.6 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV, and a search interval corresponding to 281 h of trigger lifetime, for long-lived particles that have stopped in the CMS detector. No evidence for an excess over the expected background in a cloud interaction model is observed. Assuming the decay $\widetilde{g}\to g\,\widetilde{\chi}_1^0$ and lifetimes between 1 $\mu{\rm s}$ and 1000 s, limits are derived on \widetilde{g} production as a function of $m_{\widetilde{\chi}_1^0}$, see Figs. 4 and 6. The exclusions require that $m_{\widetilde{\chi}_1^0}$ is kinematically consistent with the minimum values of the jet energy thresholds used.
- ⁷ AAD 13AA searched in 4.7 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events containing colored long-lived particles that hadronize forming R-hadrons. No significant excess above the expected background was found. Long-lived R-hadrons containing a \tilde{g} are excluded for masses up to 985 GeV at 95% C.L in a general interaction model. Also, limits independent of the fraction of R-hadrons that arrive charged in the muon system were derived, see Fig. 6.
- ⁸ AAD 13BC searched in 5.0 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV and in 22.9 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for bottom squark R-hadrons that have come to rest within the ATLAS calorimeter and decay at some later time to hadronic jets and a neutralino. In absence of an excess of events above the expected backgrounds, limits are set on gluino masses for different decays, lifetimes, and neutralino masses, see their Table 6 and Fig. 10.
- ⁹CHATRCHYAN 13AB looked in 5.0 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV and in 18.8 fb⁻¹ of pp collisions at $\sqrt{s}=8$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of \tilde{g} 's. No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 8 and Table 5), depending on the fraction, f, of formation of \tilde{g} -g (R-gluonball) states. The quoted limit is for f=0.1, while for f=0.5 it degrades to 1276 GeV. In the conservative scenario where every hadronic interaction causes it to become neutral, the limit decreases to 928 GeV for f=0.1.
- 10 AAD 12P looked in 31 pb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with pair production of long-lived gluinos. The hadronization of the gluinos leads to R-hadrons which may stop inside the detector and later decay via $\widetilde{g} \to g \, \widetilde{\chi}_1^0$ during gaps between the proton bunches. No significant excess over the expected background is observed. From

a counting experiment, a limit at 95% C.L. on the cross section as a function of $m_{\widetilde{g}}$ is derived for $m_{\widetilde{\chi}_1^0}=100$ GeV, see Fig. 4. The limit is valid for lifetimes between 10^{-5}

and 10^3 seconds and assumes the *Generic* matter interaction model for the production cross section.

- 11 CHATRCHYAN 12AN looked in 4.0 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=7$ TeV for events with pair production of long-lived gluinos. The hadronization of the gluinos leads to R-hadrons which may stop inside the detector and later decay via $\widetilde{g}\to g\,\widetilde{\chi}_1^0$ during gaps between the proton bunches. No significant excess over the expected background is observed. From a counting experiment, a limit at 95% C.L. on the cross section as a function of $m_{\widetilde{g}}$ is derived, see Fig. 3. The mass limit is valid for lifetimes between 10^{-5}
 - and 10^3 seconds, for what they call "the daughter gluon energy $E_g >$ " 100 GeV and assuming the *cloud* interaction model for *R*-hadrons. Supersedes KHACHATRYAN 11.
- ¹² CHATRCHYAN 12L looked in 5.0 fb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of \tilde{g} 's. No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 3), depending on the fraction, f, of formation of \tilde{g} -g (R-glueball) states. The quoted limit is for f = 0.1, while for f = 0.5 it degrades to 1046 GeV. In the conservative scenario where every hadronic interaction causes it to become neutral, the limit decreases to 928 GeV for f=0.1. Supersedes KHACHATRYAN 11C.
- 13 AAD 11 K looked in 34 pb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or time of flight in the tile calorimeter, from pair production of \widetilde{g} . No evidence for an excess over the SM expectation is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 4), for a fraction, f = 10%, of formation of $\widetilde{g}-g$ (R-gluonball). If instead of a phase space driven approach for the hadronic scattering of the R-hadrons, a triple-Regge model or a bag-model is used, the limit degrades to 566 and 562 GeV, respectively.
- ¹⁴ AAD 11P looked in 37 pb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with heavy stable particles, reconstructed and identified by their time of flight in the Muon System. There is no requirement on their observation in the tracker to increase the sensitivity to cases where gluinos have a large fraction, f, of formation of neutral $\tilde{g}-g$ (R-gluonball). No evidence for an excess over the SM expectation is observed. Limits are derived as a function of mass (see Fig. 4), for f=0.1. For fractions f = 0.5 and 1.0 the limit degrades to 537 and 530 GeV, respectively.
- 15 KHACHATRYAN 11 looked in 10 pb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for events with pair production of long-lived gluinos. The hadronization of the gluinos leads to R-hadrons which may stop inside the detector and later decay via $\widetilde{g}\to g\,\widetilde{\chi}_1^0$ during gaps between the proton bunches. No significant excess over the expected background is observed. From a counting experiment, a limit at 95% C.L. on the cross section times branching ratio is derived for $m_{\widetilde{g}}-m_{\widetilde{\chi}_1^0}>100$ GeV, see their Fig. 2. Assuming 100% branching
 - ratio, lifetimes between 75 ns and 3×10^5 s are excluded for $m_{\widetilde{g}}=300$ GeV. The \widetilde{g} mass exclusion is obtained with the same assumptions for lifetimes between 10 μs and 1000 s, but shows some dependence on the model for R-hadron interactions with matter, illustrated in Fig. 3. From a time-profile analysis, the mass exclusion is 382 GeV for a lifetime of 10 μs under the same assumptions as above.
- ¹⁶ KHACHATRYAN 11C looked in 3.1 pb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with heavy stable particles, identified by their anomalous dE/dx in the tracker or additionally requiring that it be identified as muon in the muon chambers, from pair production of \tilde{g} . No evidence for an excess over the expected background is observed. Limits are derived for pair production of gluinos as a function of mass (see Fig. 3), depending on the fraction, f, of formation of $\tilde{g}-g$ (R-gluonball). The quoted limit is for f=0.1, while

for f=0.5 it degrades to 357 GeV. In the conservative scenario where every hadronic interaction causes it to become neutral, the limit decreases to 311 GeV for f=0.1.

LIGHT \widetilde{G} (Gravitino) MASS LIMITS FROM COLLIDER EXPERIMENTS

The following are bounds on light ($\ll 1\,\text{eV}$) gravitino indirectly inferred from its coupling to matter suppressed by the gravitino decay constant.

Unless otherwise stated, all limits assume that other supersymmetric particles besides the gravitino are too heavy to be produced. The gravitino is assumed to be undetected and to give rise to a missing energy (\cancel{E}) signature.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE (eV)	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not	use the fo	llowing data for a	verages, fits, l	imits, etc. • • •
$> 3.5 \times 10^{-4}$	95	¹ AAD	15BH ATLS	$egin{aligned} \operatorname{jet} + ot \!$
> 3 × 10 ⁻⁴	95	¹ AAD	15вн ATLS	$jet + \not\!\!E_T, \ pp o (\widetilde{q}/\widetilde{g})\widetilde{G}, \ m_{\widetilde{q}} = m_{\widetilde{g}} = 1000 \ GeV$
> 2 × 10 ⁻⁴	95	¹ AAD	15BH ATLS	$jet + \not\!\!E_T, \ pp o (\widetilde{q}/\widetilde{g})\widetilde{G}, \ m_{\widetilde{g}} = m_{\widetilde{g}} = 1500 \ GeV$
$> 1.09 \times 10^{-5}$	95	² ABDALLAH	05в DLPH	$e^+e^- ightarrow \ \widetilde{\widetilde{G}}\ \widetilde{G}\ \gamma$
$> 1.35 \times 10^{-5}$	95	³ ACHARD		$e^+e^- ightarrow\widetilde{G}\widetilde{G}\gamma$
$> 1.3 \times 10^{-5}$		⁴ HEISTER		$e^+e^- o \widetilde{\it G}\widetilde{\it G}\gamma$
$>11.7 \times 10^{-6}$	95	⁵ ACOSTA	02н CDF	$p\overline{p} ightarrow \widetilde{G}\widetilde{G}\gamma$
$> 8.7 \times 10^{-6}$	95	⁶ ABBIENDI,G	00D OPAL	$e^+e^- ightarrow\ \widetilde{G}\widetilde{G}\gamma$

 $^{^1}$ AAD 15BH searched in 20.3 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV for associated production of a light gravitino and a squark or gluino. The squark (gluino) is assumed to decay exclusively to a quark (gluon) and a gravitino. No evidence was found for an excess above the expected level of Standard Model background and 95% C.L. lower limits were set on the gravitino mass as a function of the squark/gluino mass, both in the case of degenerate and non-degenerate squark/gluino masses, see Figs. 14 and 15.

² ABDALLAH 05B use data from $\sqrt{s}=180$ –208 GeV. They look for events with a single photon + $\not\!\!E$ final states from which a cross section limit of $\sigma<0.18~pb$ at 208 GeV is obtained, allowing a limit on the mass to be set. Supersedes the results of ABREU 00Z.

 $^{^3}$ ACHARD 04E use data from $\sqrt{s}=189\text{--}209$ GeV. They look for events with a single photon $+ \not\!\! E$ final states from which a limit on the Gravitino mass is set corresponding to $\sqrt{F}~>238$ GeV. Supersedes the results of ACCIARRI 99R.

 $^{^4\, {\}rm HEISTER}$ 03C use the data from $\sqrt{s}=$ 189–209 GeV to search for γE_T final states.

⁵ ACOSTA 02H looked in 87 pb^{-1} of $p\overline{p}$ collisions at $\sqrt{s}{=}1.8$ TeV for events with a high- E_T photon and E_T . They compared the data with a GMSB model where the final state could arise from $q\overline{q} \to \widetilde{G}\widetilde{G}\gamma$. Since the cross section for this process scales as $1/|F|^4$, a limit at 95% CL is derived on $|F|^{1/2} >$ 221 GeV. A model independent limit for the above topology is also given in the paper.

⁶ ABBIENDI,G 00D searches for γE final states from \sqrt{s} =189 GeV.

Supersymmetry Miscellaneous Results

Results that do not appear under other headings or that make nonminimal assumptions.

Some earlier papers are now obsolete and have been omitted. They were last listed in our PDG 14 edition: K. Olive, *et al.* (Particle Data Group), Chinese Physics **C38** 070001 (2014) (http://pdg.lbl.gov).

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use	the follow	ing data for avera	ges, fits, limit	s, etc. • • •
>65	95	¹ AABOUD	16AF ATLS	selected ATLAS searches on EWK sector
none 0–2	95	² AAD	16AG ATLS	dark photon, γ_d , in SUSY- and Higgs-portal models
		³ AAD	13P ATLS	
		⁴ AALTONEN	12AB CDF	hidden-valley Higgs
none 100-185	95	⁵ AAD	11AA ATLS	scalar gluons
		⁶ CHATRCHYAN	I11E CMS	$\mu\mu$ resonances
		⁷ ABAZOV	10N D0	γ_D , hidden valley

 $^{^1}$ AABOUD 16AF uses a selection of searches by ATLAS for the electroweak production of SUSY particles studying resulting constraints on dark matter candidates. They use 20 fb $^{-1}$ of $p\,p$ collisions at $\sqrt{s}=8$ TeV. A likelihood-driven scan of an effective model focusing on the gaugino-higgsino and Higgs sector of the pMSSM is performed. The ATLAS searches impact models where $m_{\chi_1^0}<65$ GeV, excluding 86% of them. See their Figs. 2, 4, and 6.

 $^{^2}$ AAD 16AG searches for prompt lepton-jets using 20 fb $^{-1}$ of pp collisions at $\sqrt{s}=8$ TeV collected with the ATLAS detector. Lepton-jets are expected from decays of low-mass dark photons in SUSY-portal and Higgs-portal models. No significant excess of events is observed and 95% CL upper limits are computed on the production cross section times branching ratio for two prompt lepton-jets in models predicting 2 or 4 γ_d via SUSY-portal topologies, for γ_d mass values between 0 and 2 GeV. See their Figs 9 and 10. The results are also interpreted in terms of a 90% CL exclusion region in kinetic mixing and dark-photon mass parameter space. See their Fig. 13.

 $^{^3}$ AAD 13P searched in 5 fb $^{-1}$ of pp collisions at $\sqrt{s}=7$ TeV for single lepton-jets with at least four muons; pairs of lepton-jets, each with two or more muons; and pairs of lepton-jets with two or more electrons. All of these could be signatures of Hidden Valley supersymmetric models. No statistically significant deviations from the Standard Model expectations are found. 95% C.L. limits are placed on the production cross section times branching ratio of dark photons for several parameter sets of a Hidden Valley model.

⁴ AALTONEN 12AB looked in 5.1 fb⁻¹ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for anomalous production of multiple low-energy leptons in association with a W or Z boson. Such events may occur in hidden valley models in which a supersymmetric Higgs boson is produced in association with a W or Z boson, with $H \to \widetilde{\chi}_1^0 \widetilde{\chi}_1^0$ pair and with the $\widetilde{\chi}_1^0$ further decaying into a dark photon (γ_D) and the unobservable lightest SUSY particle of the hidden sector. As the γ_D is expected to be light, it may decay into a lepton pair. No significant excess over the SM expectation is observed and a limit at 95% C.L. is set on the cross section for a benchmark model of supersymmetric hidden-valley Higgs production.

⁵AAD 11AA looked in 34 pb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with ≥ 4 jets originating from pair production of scalar gluons, each decaying to two gluons. No two-jet resonances are observed over the SM background. Limits are derived on the cross section times branching ratio (see Fig. 3). Assuming 100% branching ratio for the decay to two gluons, the quoted exclusion range is obtained, except for a 5 GeV mass window around 140 GeV.

⁶ CHATRCHYAN 11E looked in 35 pb⁻¹ of pp collisions at $\sqrt{s}=7$ TeV for events with collimated μ pairs (leptonic jets) from the decay of hidden sector states. No evidence for

new resonance production is found. Limits are derived and compared to various SUSY models (see Fig. 4) where the LSP, either the $\tilde{\chi}_1^0$ or a \tilde{q} , decays to dark sector particles.

 7 ABAZOV 10N looked in 5.8 fb $^{-1}$ of $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV for events from hidden valley models in which a $\widetilde{\chi}^0_1$ decays into a dark photon, γ_D , and the unobservable lightest SUSY particle of the hidden sector. As the γ_D is expected to be light, it may decay into a tightly collimated lepton pair, called lepton jet. They searched for events with E_T and two isolated lepton jets observable by an opposite charged lepton pair $e\,e,\,e\,\mu$ or $\mu\mu$. No significant excess over the SM expectation is observed, and a limit at 95% C.L. on the cross section times branching ratio is derived, see their Table I. They also examined the invariant mass of the lepton jets for a narrow resonance, see their Fig. 4, but found no evidence for a signal.

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KHACHATRY AACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15C 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035	V. Khachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY ROLBIECKI XIAO AAD AAD AAD AAD AAD AAD AAD AAD AAD A	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14BD 14BH 14EH	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY ROLBIECKI XIAO AAD AAD AAD AAD AAD AAD AAD AAD AAD A	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1406 124	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY	15AH 15AK 15AO 15AR 15AZ 15E 15I 15L 15O 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 096 JHEP 1410 1024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 124 JHEP 1406 124 JHEP 1405 071	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY ROLBIECKI XIAO AAD AAD AAD AAD AAD AAD AAD AAD AAD A	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1406 124	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F 14G 14H	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 124 JHEP 1405 071 JHEP 1405 071 JHEP 1405 071 JHEP 1405 071	V. Khachatryan et al. C. Aadet al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15C 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F 14G 14H 14K	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 176 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1405 071 JHEP 1405 071 JHEP 1404 169 PR D90 012004	V. Khachatryan et al. C. Aadet al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15C 15W 15X 15 14AE 14AG 14AV 14AX 14BD 14BH 14E 14F 14G 14H 14K 14T	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 176 JHEP 1409 176 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1405 071 JHEP 1405 071 JHEP 1404 169 PR D90 012004 PR D90 012004	V. Khachatryan et al. C. Aadet al. G. Aad et al.	(CMS Collab.) (AMADE, HEID) (PandaX Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15C 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F 14G 14H 14K	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 176 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1405 071 JHEP 1405 071 JHEP 1404 169 PR D90 012004	V. Khachatryan et al. C. Aadet al. G. Aad et al.	(CMS Collab.) (AMADE, HEID) (PandaX Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15C 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F 14G 14H 14K 14T	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1410 096 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1405 071 JHEP 1404 169 PR D90 012004 PR D90 052008 PR D90 052001	V. Khachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (MADE, HEID) (PandaX Collab.) (ATLAS Collab.)
KHACHATRY KACHATRY AACHATRY AACHATRY ROLBIECKI XIAO AAD AAD AAD AAD AAD AAD AAD AAD AAD A	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14BD 14BH 14E 14F 14G 14H 14F 14G 14H 14F 14G 14H 14H 14H 14H 14K 14H 14K 14H 14K 14H 14K 14H 14K 14H 14K 14K 14K 14K 14K 14K 14K 14K 14K 14K	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1410 096 JHEP 1410 096 JHEP 1410 118 PR D90 112005 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 124 JHEP 1405 071 JHEP 1404 169 PR D90 012004 PR D90 052001 PR D90 052001	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15C 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14E 14F 14G 14H 14K 14T	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1410 096 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1405 071 JHEP 1404 169 PR D90 012004 PR D90 052008 PR D90 052001	V. Khachatryan et al. C. Aad et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14BD 14BH 14E 14F 14G 14H 14F 14G 14H 14F 14G 14H 14H 14H 14H 14K 14H 14K 14H 14K 14H 14K 14H 14K 14H 14K 14K 14K 14K 14K 14K 14K 14K 14K 14K	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 124 JHEP 1406 124 JHEP 1404 169 PR D90 012004 PR D90 012004 PR D90 052008 PR D90 052001 PR D90 012011 PR D89 042001	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15L 15O 15W 15X 15 14AE 14AG 14AJ 14AV 14BD 14BH 14E 14F 14G 14H 14K 14H 14K 14T 14X	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1406 124 JHEP 1405 071 JHEP 1404 169 PR D90 012004 PR D90 052008 PR D90 052008 PR D90 012011 PR D89 042001 PRL 112 091303	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AACHACHATRY AACHACHACHATRY AACHACHATRY AAC	15AH 15AK 15AO 15AR 15AZ 15E 15I 15L 15O 15W 15X 15 14AE 14AG 14AJ 14AV 14AX 14B 14BD 14BH 14F 14G 14H 14K 14T 14K 14T 14X	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1406 124 JHEP 1405 071 JHEP 1404 169 PR D90 052008 PR D90 052001 PR D90 012011 PR D89 042001 PRL 112 091303 JCAP 1402 008	V. Khachatryan et al. C. Kachatryan et al. C. Aad et	(CMS Collab.) (ATLAS Collab.) (CDF Collab.) (EFermi-LAT Collab.) (LUX Collab.)
KHACHATRY AAD AAD AAD AAD AAD AAD AAD AAD AAD	15AH 15AK 15AO 15AR 15AZ 15E 15I 15L 15O 15W 15X 15 14AE 14AG 14AJ 14AV 14BD 14BH 14E 14F 14G 14H 14K 14H 14K 14T 14X	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1406 124 JHEP 1405 071 JHEP 1404 169 PR D90 012004 PR D90 052008 PR D90 052008 PR D90 012011 PR D89 042001 PRL 112 091303	V. Khachatryan et al. C. Kachatryan et al. C. Aad et al.	(CMS Collab.) (ATLAS Collab.)
KHACHATRY AACHATRY AACHACHATRY AACHACHACHATRY AACHACHATRY AAC	15AH 15AK 15AO 15AR 15AZ 15E 15I 15U 15W 15X 15 14AE 14AG 14AJ 14AX 14B 14BD 14BH 14F 14F 14G 14H 14K 14T 14K 14T 14X 14 14 14 14 14 14 14 14 14 14 14 14 14	EPJ C75 151 EPJ C75 325 PL B743 503 PR D92 072006 PRL 114 061801 PL B745 5 PL B747 98 PL B748 255 PR D91 052012 PR D91 052012 PR D91 052018 PL B750 247 PR D92 052004 JHEP 1409 103 JHEP 1409 103 JHEP 1409 015 JHEP 1410 096 JHEP 1410 024 EPJ C74 2883 JHEP 1411 118 PR D90 112005 JHEP 1406 035 JHEP 1406 035 JHEP 1406 124 JHEP 1405 071 JHEP 1404 169 PR D90 052008 PR D90 052001 PR D90 012011 PR D89 042001 PRL 112 091303 JCAP 1402 008	V. Khachatryan et al. C. Kachatryan et al. C. Aad et	(CMS Collab.) (ATLAS Collab.) (CDF Collab.) (EFermi-LAT Collab.) (LUX Collab.)

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CHATRCHYAN		PR D90 032006	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
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CZAKON	14	PRL 113 201803		ACH, CAMB, UCB, LBL+)
FELIZARDO	14	PR D89 072013	M. Felizardo <i>et al.</i>	(SIMPLE Collab.)
KHACHATRY		PL B736 371	V. Khachatryan <i>et al.</i>	(CMS Collab.)
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PDG	14	CP C38 070001	K. Olive et al.	(PDG Collab.)
ROSZKOWSKI	14	JHEP 1408 067	L. Roszkowski, E.M. Sessolo,	
AAD	13	PL B718 841	G. Aad et al.	(ATLAS Collab.)
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AAD	13AI	PL B723 15	G. Aad et al.	(ATLAS Collab.)
AAD	13AP	PR D88 012001	G. Aad <i>et al.</i>	(ATLAS Collab.)
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AAD	13B	PL B718 879	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD		PR D88 112003	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	13BD	PR D88 112006	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	13H	JHEP 1301 131	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	13L	PR D87 012008	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	13P	PL B719 299	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD	13Q	PL B719 261	G. Aad et al.	(ATLAS Collab.)
AAD	13R	PL B719 280	G. Aad et al.	(ATLAS Collab.)
AALTONEN	13I	PR D88 031103	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN	13Q	PRL 110 201802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AARTSEN	13	PRL 110 131302	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)
AARTSEN	13C	PR D88 122001	M.G. Aartsen <i>et al.</i>	(IceCube Collab.)
ABAZOV	13B	PR D87 052011	V.M. Abazov et al.	(D0 Collab.)
ABRAMOWSKI		PRL 110 041301	A. Abramowski <i>et al.</i>	(H.E.S.S. Collab.)
ACKERMANN		PR D88 082002	M. Ackermann <i>et al.</i> S. Adrian-Martinez <i>et al.</i>	(Fermi-LAT Collab.)
ADRIAN-MAR AGNESE	13	JCAP 1311 032 PR D88 031104		(ANTARES Collab.) (CDMS Collab.)
AGNESE	13A	PRL 111 251301	R. Agnese <i>et al.</i> R. Agnese <i>et al.</i>	(CDMS Collab.)
APRILE	13	PRL 111 021301	E. Aprile <i>et al.</i>	(XENON100 Collab.)
BERGSTROM	13	PRL 111 171101	L. Bergstrom et al.	(XENONIOU CONAD.)
BOLIEV	13	JCAP 1309 019	M. Boliev <i>et al.</i>	
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	-		S. Chatrchyan <i>et al.</i> S. Chatrchyan <i>et al.</i>	(CMS Collab.)
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	13AB 13AH	JHEP 1307 122	S. Chatrchyan <i>et al.</i> S. Chatrchyan <i>et al.</i>	(CMS Collab.) (CMS Collab.) (CMS Collab.)
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CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN	13AB 13AH 13AO 13AT 13AV 13G 13H	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077	 S. Chatrchyan et al. 	(CMS Collab.) (CMS Collab.) (CMS Collab.) (CMS Collab.) (CMS Collab.) (CMS Collab.)
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CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zi	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN AISO CHATRCHYAN ELLIS JIN KOPP	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zl. J. Kopp	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN AISO CHATRCHYAN ELLIS JIN KOPP STREGE	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013	S. Chatrchyan et al. J. Ellis et al. J. Ellis et al. J. Kopp C. Strege et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 13 12AF	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zl. J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 13 12AF 12AG	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zl. J. Kopp C. Strege et al. G. Aad et al. G. Aad et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zl. J. Kopp C. Strege et al. G. Aad et al. G. Aad et al. G. Aad et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zi J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zi J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS 12AX	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1303 037 JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006 PR D87 099903 (errat.)	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Z. J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.) (ATLAS Collab.) (ATLAS Collab.) (ATLAS Collab.) (ATLAS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD AAD AAD AAD AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 12AF 12AG 12AN 12AS 12AS	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1310 26 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006 PR D87 099903 (errat.) EPJ C72 1993	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Z. J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD AAD AAD AAD AAD AAD AA	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS 12AX	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006 PR D87 099903 (errat.) EPJ C72 1993 PR D86 092002	S. Chatrchyan et al. G. Chatrchyan et al. G. Chatrchyan et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD AAD AAD AAD AAD	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS 12AX	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006 PR D87 099903 (errat.) EPJ C72 1993 PR D86 092002 EPJ C72 2215	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Z. J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD AAD AAD AAD AAD AAD AA	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS 12AX 12AX	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006 PR D87 099903 (errat.) EPJ C72 1993 PR D86 092002	S. Chatrchyan et al. J. Ellis et al. HB. Jin, YL. Wu, YF. Zl. J. Kopp C. Strege et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN CHATRCHYAN Also CHATRCHYAN ELLIS JIN KOPP STREGE AAD AAD AAD AAD AIso AIso AAD AAD AAD AAD AAD AAD AAD AAD AAD AA	13AB 13AH 13AO 13AT 13AV 13G 13H 13T 13V 13W 13B 13 13 12AF 12AG 12AN 12AS 12AX 12AX	JHEP 1307 122 PL B722 273 PR D87 072001 PR D88 052017 PRL 111 081802 JHEP 1301 077 PL B719 42 EPJ C73 2568 JHEP 1303 037 JHEP 1307 041 (errat.) JHEP 1303 111 EPJ C73 2403 JCAP 1311 026 PR D88 076013 JCAP 1304 013 PL B714 180 PL B714 197 PRL 108 181802 PRL 108 261804 PR D85 012006 PR D87 099903 (errat.) EPJ C72 1993 PR D86 092002 EPJ C72 2215 PL B718 411	S. Chatrchyan et al. G. Chatrchyan et al. G. Aad et al.	(CMS Collab.) (ATLAS Collab.)

AAD AAD	12R 12T	PL B707 478 PL B709 137	G. Aad <i>et al.</i> G. Aad <i>et al.</i>	(ATLAS Collab.) (ATLAS Collab.)
AAD	12W		G. Aad et al.	(ATLAS Collab.)
AALTONEN		PR D85 092001	T. Aaltonen et al.	(CDF Collab.)
ABAZOV ABBASI	12AD 12	PR D86 071701 PR D85 042002	V.M. Abazov <i>et al.</i> R. Abbasi <i>et al.</i>	(D0 Collab.)
AKIMOV	12	PL B709 14	D.Yu. Akimov <i>et al.</i>	(IceCube Collab.) (ZEPLIN-III Collab.)
AKULA	12	PR D85 075001	S. Akula <i>et al.</i>	(NEAS, MICH)
ANGLOHER	12	EPJ C72 1971	G. Angloher et al.	(CRESST-II Collab.)
APRILE ARBEY	12 12A	PRL 109 181301 PL B708 162	E. Aprile <i>et al.</i> A. Arbey <i>et al.</i>	(XENON100 Collab.)
ARCHAMBAU		PL B711 153	S. Archambault <i>et al.</i>	(PICASSO Collab.)
BAER	12	JHEP 1205 091	H. Baer, V. Barger, A. M	
BALAZS	12	EPJ C73 2563	C. Balazs et al.	
BECHTLE BEHNKE	12 12	JHEP 1206 098 PR D86 052001	P. Bechtle <i>et al.</i> E. Behnke <i>et al.</i>	(COUPP Collab.)
Also		PR D90 079902 (errat.)		(COUPP Collab.)
BESKIDT	12	EPJ C72 2166	C. Beskidt et al.	(KARLE, JINR, ITEP)
BOTTINO BUCHMUEL	12	PR D85 095013 EPJ C72 2020	A. Bottino, N. Fornengo, O. Buchmueller <i>et al.</i>	S. Scopel (TORI, S0GA)
CAO	12A	PL B710 665	J. Cao <i>et al.</i>	
CHATRCHYAN	12	PR D85 012004	S. Chatrchyan et al.	(CMS Collab.)
		PRL 109 171803	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN		JHEP 1208 110 JHEP 1206 169	S. Chatrohyan et al.	(CMS Collab.)
		JHEP 1200 109 JHEP 1208 026	S. Chatrchyan <i>et al.</i> S. Chatrchyan <i>et al.</i>	(CMS Collab.) (CMS Collab.)
		JHEP 1210 018	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN	12BJ	JHEP 1211 147	S. Chatrchyan et al.	(CMS Collab.)
		JHEP 1211 172	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN	-	JHEP 1212 055 PL B713 408	S. Chatrchyan <i>et al.</i> S. Chatrchyan <i>et al.</i>	(CMS Collab.) (CMS Collab.)
DAW	12	ASP 35 397	E. Daw et al.	(DRIFT-IId Collab.)
DREINER	12A	EPL 99 61001	H.K. Dreiner, M. Kramer	
ELLIS	12B	EPJ C72 2005	J. Ellis, K. Olive	(61) 451 5 6 11 1)
FELIZARDO FENG	12 12B	PRL 108 201302 PR D85 075007	M. Felizardo <i>et al.</i> J. Feng, K. Matchev, D.	(SIMPLE Collab.)
KADASTIK	125	JHEP 1205 061	M. Kadastik <i>et al.</i>	Samord
KIM	12	PRL 108 181301	S.C. Kim et al.	(KIMS Collab.)
STREGE	12	JCAP 1203 030		(LOIC, AMST, MADU, GRAN+)
AAD AAD	11AA 11G	EPJ C71 1828	G. Aad <i>et al.</i> G. Aad <i>et al.</i>	(ATLAS Collab.) (ATLAS Collab.)
AAD	11G 11H	PRL 106 131802 PRL 106 251801	G. Aad et al.	(ATLAS Collab.)
AAD	11K	PL B701 1	G. Aad et al.	(ATLAS Collab.)
AAD	110	PL B701 398	G. Aad et al.	(ATLAS Collab.)
AAD AAD	11P	PL B703 428	G. Aad et al.	(ATLAS Collab.)
ABRAMOWSKI	11Z 11	EPJ C71 1809 PRL 106 161301	G. Aad <i>et al.</i> A. Abramowski <i>et al.</i>	(ATLAS Collab.) (H.E.S.S. Collab.)
AHMED	11A	PR D84 011102		CDMS and EDELWEISS Collabs.)
ARMENGAUD		PL B702 329	E. Armengaud et al.	(EDELWEISS II Collab.)
BUCHMUEL BUCHMUEL		EPJ C71 1583	O. Buchmueller <i>et al.</i>	
CHATRCHYAN		EPJ C71 1722 JHEP 1106 093	O. Buchmueller <i>et al.</i> S. Chatrchyan <i>et al.</i>	(CMS Collab.)
CHATRCHYAN		JHEP 1107 113	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN		JHEP 1107 098	S. Chatrchyan et al.	(CMS Collab.)
CHATRCHYAN		PL B704 411	S. Chatrchyan <i>et al.</i>	(CMS Collab.)
KHACHATRY KHACHATRY		PRL 106 011801 JHEP 1103 024	V. Khachatryan <i>et al.</i> V. Khachatryan <i>et al.</i>	(CMS Collab.) (CMS Collab.)
ROSZKOWSKI		PR D83 015014	L. Roszkowski <i>et al.</i>	(CIVIS CONIAD.)
AALTONEN	10	PRL 104 011801	T. Aaltonen et al.	(CDF Collab.)
AALTONEN	10R	PRL 105 081802	T. Aaltonen <i>et al.</i>	(CDF Collab.)
AALTONEN ABAZOV	10Z 10L	PRL 105 191801 PL B693 95	T. Aaltonen <i>et al.</i> V.M. Abazov <i>et al.</i>	(CDF Collab.) (D0 Collab.)
ABAZOV	10M	PRL 105 191802	V.M. Abazov et al.	(D0 Collab.)
ABAZOV	10N	PRL 105 211802	V.M. Abazov et al.	(D0 Collab.)
ABAZOV ABDO	10P 10	PRL 105 221802 JCAP 1004 014	V.M. Abazov <i>et al.</i> A.A. Abdo <i>et al.</i>	(D0 Collab.) (Fermi-LAT Collab.)
ACKERMANN	10	JCAP 1004 014 JCAP 1005 025	M. Ackermann	(Fermi-LAT Collab.)
AHMED	10	SCI 327 1619	Z. Ahmed <i>et al.</i>	`(CDMS II Collab.)
ARMENGAUD	10	PL B687 294	E. Armengaud et al.	(EDELWEISS II Collab.)
ELLIS ABAZOV	10 09M	EPJ C69 201 PRL 102 161802	J. Ellis, A. Mustafayev, ł V.M. Abazov <i>et al.</i>	C. Olive (D0 Collab.)
	02141	102 101002	ADUZOV CL al.	(Do Collab.)

ABBASI				
	09B	PRL 102 201302	R. Abbasi <i>et al.</i>	(IceCube Collab.)
AHMED	09	PRL 102 011301	Z. Ahmed et al.	`(CDMS Collab.)
ANGLOHER	09	ASP 31 270	G. Angloher <i>et al.</i>	(CRESST Collab.)
ARCHAMBAU.		PL B682 185	S. Archambault <i>et al.</i>	(PICASSO Collab.)
BUCHMUEL		EPJ C64 391	O. Buchmueller <i>et al.</i>	(LOIC, FNAL, CERN $+$)
DREINER	09	EPJ C62 547	H. Dreiner <i>et al.</i>	
LEBEDENKO	09	PR D80 052010	V.N. Lebedenko <i>et al.</i>	(ZEPLIN-III Collab.)
LEBEDENKO	09A	PRL 103 151302	V.N. Lebedenko et al.	(ZEPLIN-III Collab.)
SORENSEN	09	NIM A601 339	P. Sorensen et al.	(XENON10 Collab.)
ABAZOV	08F	PL B659 856	V.M. Abazov et al.	(D0 Collab.)
ANGLE	80	PRL 100 021303	J. Angle <i>et al.</i>	(XENON10 Collab.)
ANGLE	08A	PRL 101 091301	J. Angle <i>et al.</i>	(XENON10 Collab.)
BEDNYAKOV	80	PAN 71 111 V	'.A. Bednyakov, H.P. Klapdor-I	Kleingrothaus, I.V. Krivosheina
		Translated from YAF 71		
BEHNKE	80	SCI 319 933	E. Behnke	(COUPP Collab.)
BENETTI	08	ASP 28 495	P. Benetti <i>et al.</i>	(WARP Collab.)
BUCHMUEL				(vviiti conab.)
	80	JHEP 0809 117	O. Buchmueller et al.	(CEDAL MAINIAL)
ELLIS	08	PR D78 075012	J. Ellis, K. Olive, P. Sandick	
ABULENCIA	07H	PRL 98 131804	A. Abulencia <i>et al.</i>	(CDF Collab.)
ALNER	07A	ASP 28 287	G.J. Alner <i>et al.</i>	(ZEPLIN-II Collab.)
CALIBBI	07	JHEP 0709 081	L. Calibbi et al.	,
ELLIS	07	JHEP 0706 079	J. Ellis, K. Olive, P. Sandick	(CERN, MINN)
LEE	07A	PRL 99 091301	H.S. Lee <i>et al.</i>	(KIMS Collab.)
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ABBIENDI	06B	EPJ C46 307	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ACHTERBERG	06	ASP 26 129	A. Achterberg <i>et al.</i>	(AMANDA Collab.)
ACKERMANN	06	ASP 24 459	M. Ackermann et al.	(AMANDA Collab.)
AKERIB	06	PR D73 011102	D.S. Akerib et al.	` (CDMS Collab.)
AKERIB	06A	PRL 96 011302	D.S. Akerib <i>et al.</i>	(CDMS Collab.)
				(CDIVIS COIIAD.)
ALLANACH	06	PR D73 015013	B.C. Allanach et al.	
BENOIT	06	PL B637 156	A. Benoit <i>et al.</i>	
DE-AUSTRI	06	JHEP 0605 002	R.R. de Austri, R. Trotta, L.	. Roszkowski
DEBOER	06	PL B636 13	W. de Boer <i>et al.</i>	
LEP-SLC	06	PRPL 427 257	ALEPH, DELPHI, L3, OPAL,	SLD and working groups
SHIMIZU	06A	PL B633 195	Y. Shimizu et al.	, old and months groups
				T C
SMITH	06	PL B642 567	N.J.T. Smith, A.S. Murphy,	
ABAZOV	05A	PRL 94 041801	V.M. Abazov <i>et al.</i>	(D0 Collab.)
ABDALLAH	05B	EPJ C38 395	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
AKERIB	05	PR D72 052009	D.S. Akerib et al.	(CDMS Collab.)
ALNER	05	PL B616 17	G.J. Alner et al.	(UK Dark Matter Collab.)
ALNER	05A	ASP 23 444	G.J. Alner <i>et al.</i>	(UK Dark Matter Collab.)
ALIVLIN	05/1	A31 23 TTT	G.J. Aillei et al.	
ANCLOUED	$\Delta \Gamma$		C A - -	
ANGLOHER	05	ASP 23 325	G. Angloher et al.	(CRESST-II Collab.)
BAER	05		H. Baer <i>et al.</i>	(CRESST-II Collab.) (FSU, MSU, HAWA)
	05	ASP 23 325	9	(CRESST-II Collab.) (FSU, MSU, HAWA)
BAER	05	ASP 23 325 JHEP 0507 065	H. Baer <i>et al.</i>	(CRESST-II Collab.)
BAER BARNABE-HE. ELLIS	05 05 05	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007	H. Baer <i>et al.</i> M. Barnabe-Heider <i>et al.</i> J. Ellis <i>et al.</i>	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.)
BAER BARNABE-HE. ELLIS SANGLARD	05 05 05 05	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI	05 05 05 05 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI	05 05 05 05 04 04F	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI	05 05 05 05 04 04F 04H	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI	05 05 05 05 04 04F	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI	05 05 05 05 04 04F 04H	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH	05 05 05 05 04 04F 04H 04N 04H	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH	05 05 05 05 04 04F 04H 04N	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH AISO	05 05 05 05 04 04F 04H 04N 04H 04M	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.)	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH Also ACHARD	05 05 05 05 04 04F 04H 04N 04H	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH AISO	05 05 05 05 04 04F 04H 04N 04H 04M	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.)	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al. P. Achard et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH Also ACHARD	05 05 05 05 04 04F 04H 04N 04H	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH Also ACHARD ACHARD AKERIB	05 05 05 05 04 04F 04H 04N 04H 04M	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B587 16 PRL 93 211301	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al. P. Achard et al. D.S. Akerib et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AIso ACHARD ACHARD AKERIB BALTZ	05 05 05 04 04F 04H 04N 04H 04M 04 04E 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Achard et al. P. Achard et al. D.S. Akerib et al. E. Baltz, P. Gondolo	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH Also ACHARD ACHARD ACHARD AKERIB BALTZ BELANGER	05 05 05 04 04F 04H 04N 04H 04M 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH Also ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO	05 05 05 05 04 04F 04H 04N 04H 04M 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Achard et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (CDMSII Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AISO ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI	05 05 05 05 04 04F 04H 04M 04H 04M 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Achard et al. P. Achard et al. P. Achard et al. D.S. Akerib et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. S. Desai et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AISO ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI ELLIS	05 05 05 05 04 04F 04H 04N 04H 04M 04 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C35 1 PL B602 167 EPJ C36 1 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523 PR D69 015005	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Achard et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. S. Desai et al. J. Ellis et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (CDMSII Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AISO ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI	05 05 05 05 04 04F 04H 04M 04H 04M 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Achard et al. P. Achard et al. P. Achard et al. D.S. Akerib et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. S. Desai et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (CDMSII Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AISO ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI ELLIS	05 05 05 05 04 04F 04H 04N 04H 04M 04 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C35 1 PL B602 167 EPJ C36 1 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523 PR D69 015005	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. J. Achard et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. S. Desai et al. J. Ellis et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.) (CDMSII Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AISO ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI ELLIS ELLIS HEISTER	05 05 05 05 04 04F 04H 04H 04H 04H 04H 04 04 04 04 04 04 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523 PR D69 015005 PR D70 055005 PL B583 247	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. J. Ellis et al. J. Ellis et al. J. Ellis et al. A. Heister et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (CDMSII Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH AIso ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI ELLIS ELLIS ELLIS HEISTER PIERCE	05 05 05 04 04F 04H 04H 04H 04H 04 04 04 04 04 04 04 04 04 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C34 145 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523 PR D69 015005 PR D70 055005 PL B583 247 PR D70 075006	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. S. Desai et al. J. Ellis et al. J. Ellis et al. A. Heister et al. A. Pierce	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.) (CDMSII Collab.) (Super-Kamiokande Collab.)
BAER BARNABE-HE. ELLIS SANGLARD ABBIENDI ABBIENDI ABBIENDI ABBIENDI ABDALLAH ABDALLAH ALSO ACHARD ACHARD ACHARD AKERIB BALTZ BELANGER BOTTINO DESAI ELLIS ELLIS HEISTER PIERCE ABBIENDI	05 05 05 05 04 04F 04H 04N 04H 04M 04 04 04 04 04 04 04 04 04 04 04 04 04	ASP 23 325 JHEP 0507 065 PL B624 186 PR D71 095007 PR D71 122002 EPJ C32 453 EPJ C33 149 EPJ C35 1 PL B602 167 EPJ C36 1 EPJ C37 129 (errat.) PL B580 37 PL B580 37 PL B587 16 PRL 93 211301 JHEP 0410 052 JHEP 0403 012 PR D69 037302 PR D70 083523 PR D69 015005 PL B583 247 PR D70 075006 PL B572 8	H. Baer et al. M. Barnabe-Heider et al. J. Ellis et al. V. Sanglard et al. G. Abbiendi et al. G. Abbiendi et al. G. Abbiendi et al. J. Abdallah et al. J. Abdallah et al. J. Abdallah et al. P. Achard et al. P. Achard et al. E. Baltz, P. Gondolo G. Belanger et al. A. Bottino et al. J. Ellis et al. J. Ellis et al. A. Heister et al. A. Pierce G. Abbiendi et al.	(CRESST-II Collab.) (FSU, MSU, HAWA) (PICASSO Collab.) (EDELWEISS Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (L3 Collab.) (CDMSII Collab.) (CDMSII Collab.) (Super-Kamiokande Collab.)
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ELLIS 03D PL B573 162 J. Ellis et al. ELLIS 03E PR D67 123502 J. Ellis et al. HEISTER 03C EPJ C28 1 A. Heister et al. (ALEPH HEISTER 03G EPJ C31 1 A. Heister et al. KLAPDOR-K 03 ASP 18 525 H.V. Klapdor-Kleingrothaus et al. LAHANAS 03 PL B568 55 A. Lahanas, D. Nanopoulos	
KLAPDOR-K 03 ASP 18 525 H.V. Klapdor-Kleingrothaus et al.	Collab.) Collab.) Collab.)
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TAKEDA 03 PL B572 145 A. Takeda <i>et al.</i> ABRAMS 02 PR D66 122003 D. Abrams <i>et al.</i> (CDMS)	Collab.) Collab.)
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	Collab.)
ANGLOHER 02 ASP 18 43 G. Angloher <i>et al.</i> (CRESST ARNOWITT 02 hep-ph/0211417 R. Arnowitt, B. Dutta	
BENOIT 02 PL B545 43 A. Benoit <i>et al.</i> (EDELWEISS	Collab.)
ELLIS 02B PL B532 318 J. Ellis, A. Ferstl, K.A. Olive	C \
HEISTER 02 PL B526 191 A. Heister et al. (ALEPH HEISTER HEISTER 02E PL B526 206 A. Heister et al. (ALEPH HEISTER)	- :
HEISTER 02J PL B533 223 A. Heister et al. (ALEPH	- :
HEISTER 02N PL B544 73 A. Heister et al. (ALEPH	Collab.)
KIM 02 PL B527 18 H.B. Kim <i>et al.</i> KIM 02B JHEP 0212 034 Y.G. Kim <i>et al.</i>	
KIM 02B JHEP 0212 034 Y.G. Kim <i>et al.</i> LAHANAS 02 EPJ C23 185 A. Lahanas, V.C. Spanos	
MORALES 02B ASP 16 325 A. Morales <i>et al.</i> (COSME	Collab.)
	Collab.)
ABBIENDI 01 PL B501 12 G. Abbiendi et al. (OPAL	- :
ABREU 01 EPJ C19 29 P. Abreu <i>et al.</i> (DELPHI of ABREU 01B EPJ C19 201 P. Abreu <i>et al.</i> (DELPHI of DELPHI	
ABREU 01G PL B503 34 P. Abreu et al. (DELPHI)	- :
	Collab.)
BALTZ 01 PRL 86 5004 E. Baltz, P. Gondolo	
BARATE 01 PL B499 67 R. Barate et al. (ALEPH CALERY	
BARATE 01B EPJ C19 415 R. Barate <i>et al.</i> (ALEPH BARGER 01C PL B518 117 V. Barger, C. Kao	Collab.)
BAUDIS 01 PR D63 022001 L. Baudis <i>et al.</i> (Heidelberg-Moscow	Collab.)
BENOIT 01 PL B513 15 A. Benoit et al. (EDELWEISS	
BERNABEI 01 PL B509 197 R. Bernabei et al. (DAMA	Collab.)
BOTTINO 01 PR D63 125003 A. Bottino <i>et al.</i> CORSETTI 01 PR D64 125010 A. Corsetti, P. Nath	
CORSETTI 01 PR D64 125010 A. Corsetti, P. Nath ELLIS 01B PL B510 236 J. Ellis <i>et al.</i>	
ELLIS 01C PR D63 065016 J. Ellis, A. Ferstl, K.A. Olive	
GOMEZ 01 PL B512 252 M.E. Gomez, J.D. Vergados	
LAHANAS 01 PL B518 94 A. Lahanas, D.V. Nanopoulos, V. Spanos	C II I)
ABBIENDI 00 EPJ C12 1 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. ABBIENDI 00G EPJ C14 51 G. Abbiendi et al. (OPAL et al. ABBIENDI 00G EPJ C14 51 G. ABBIENDI 00G E	
ABBIENDI 00H EPJ C14 187 G. Abbiendi et al. (OPAL	- :
Also EPJ C16 707 (errat.) G. Abbiendi <i>et al.</i> (OPAL	Collab.)
ABBIENDI,G 00D EPJ C18 253 G. Abbiendi et al. (OPAL	- :
ABREU 00J PL B479 129 P. Abreu <i>et al.</i> (DELPHI of ABREU 00Q PL B478 65 P. Abreu <i>et al.</i> (DELPHI of DELPHI of DELPH	
ABREU 00T PL B485 95 P. Abreu et al. (DELPHI)	
ABREU 00U PL B487 36 P. Abreu et al. (DELPHI	
ABREU 00V EPJ C16 211 P. Abreu et al. (DELPHI	,
ABREU 00W PL B489 38 P. Abreu <i>et al.</i> (DELPHI of ABREU 00Z EPJ C17 53 P. Abreu <i>et al.</i> (DELPHI of DELPHI	
ABREU 00Z EPJ C17 53 P. Abreu <i>et al.</i> (DELPHI et al. (CDMS et al. (CD	
	Collab.)
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ACCOMANDO 00 NP B585 124 E. Accomando et al.	C II I)
BERNABEI 00 PL B480 23 R. Bernabei et al. (DAMA e BERNABEI 00C EPJ C18 283 R. Bernabei et al. (DAMA e	
BERNABEI 00D NJP 2 15 R. Bernabei et al. (DAMA	- :
BOEHM 00B PR D62 035012 C. Boehm, A. Djouadi, M. Drees	- /
ELLIS 00 PR D62 075010 J. Ellis <i>et al.</i>	
FENG 00 PL B482 388 J.L. Feng, K.T. Matchev, F. Wilczek LEP 00 CERN-EP-2000-016 LEP Collabs. (ALEPH, DELPHI, L3, OPAL,	SIDi)
MORALES 00 PL B489 268 A. Morales et al. (IGEX)	Collab.)
	Collab.)

SPOONER	00	PL B473 330	N.J.C. Spooner et al.	(UK Dark Matter Col.)
ACCIARRI	99H	PL B456 283	M. Acciarri <i>et al.</i>	` (L3 Collab.)
ACCIARRI	991	PL B459 354	M. Acciarri <i>et al.</i>	(L3 Collab.)
	99R		M. Acciarri <i>et al.</i>	• • • • • • • • • • • • • • • • • • • •
ACCIARRI		PL B470 268		(L3 Collab.)
ACCIARRI	99W	PL B471 280	M. Acciarri <i>et al.</i>	(L3 Collab.)
AMBROSIO	99	PR D60 082002	M. Ambrosio <i>et al.</i>	(Macro Collab.)
BAUDIS	99	PR D59 022001	L. Baudis <i>et al.</i>	(Heidelberg-Moscow Collab.)
BELLI	99C	NP B563 97	P. Belli <i>et al.</i>	(DAMA Collab.)
OOTANI	99	PL B461 371	W. Ootani et al.	,
ABREU	98P	PL B444 491	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	98F	EPJ C4 207	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACKERSTAFF	98P	PL B433 195	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BARATE	98K	PL B433 176	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	98S	EPJ C4 433	R. Barate <i>et al.</i>	(ALEPH Collab.)
BERNABEI	98C	PL B436 379	R. Bernabei <i>et al.</i>	(DAMA Collab.)
ELLIS	98	PR D58 095002	J. Ellis <i>et al.</i>	
ELLIS	98B	PL B444 367	J. Ellis, T. Falk, K. Olive	
PDG	98	EPJ C3 1	C. Caso et al.	(PDG Collab.)
BAER	97	PR D57 567	H. Baer, M. Brhlik	(1 DG Conub.)
				(DAMA C-II-L)
BERNABEI	97	ASP 7 73	R. Bernabei <i>et al.</i>	(DAMA Collab.)
EDSJO	97	PR D56 1879	J. Edsjo, P. Gondolo	
ARNOWITT	96	PR D54 2374	R. Arnowitt, P. Nath	
BAER	96	PR D53 597	H. Baer, M. Brhlik	
BERGSTROM	96	ASP 5 263	L. Bergstrom, P. Gondolo	
LEWIN	96	ASP 6 87	J.D. Lewin, P.F. Smith	
BEREZINSKY	95	ASP 5 1	V. Berezinsky <i>et al.</i>	
FALK	95		•	dnieki (MININ LICCE)
		PL B354 99	T. Falk, K.A. Olive, M. Sre	`
LOSECCO	95	PL B342 392	J.M. LoSecco	(NDAM)
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)
DREES	93	PR D47 376	M. Drees, M.M. Nojiri	(DESY, SLAC)
DREES	93B	PR D48 3483	M. Drees, M.M. Nojiri	
FALK	93	PL B318 354	T. Falk et al.	(UCB, UCSB, MINN)
KELLEY	93	PR D47 2461	S. Kelley et al.	` (TAMU, ALAH)
MIZUTA	93	PL B298 120	S. Mizuta, M. Yamaguchi	(TOHO)
MORI	93	PR D48 5505		KEK, NIIG, TOKY, TOKA+)
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BOTTINO	92	MPL A7 733	A. Bottino <i>et al.</i>	(TORI, ZARA)
Also		PL B265 57	A. Bottino et al.	(TORI, INFN)
DECAMP	92	PRPL 216 253	D. Decamp <i>et al.</i>	(ALEPH Collab.)
LOPEZ	92	NP B370 445	J.L. Lopez, D.V. Nanopoulos	s, K.J. Yuan (TAMU)
MCDONALD	92	PL B283 80	J. McDonald, K.A. Olive, M	
ABREU	91F	NP B367 511	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ALEXANDER	91F	ZPHY C52 175	G. Alexander et al.	(OPAL Collab.)
BOTTINO	91	PL B265 57	A. Bottino <i>et al.</i>	(TORI, INFN)
GELMINI	91	NP B351 623	G.B. Gelmini, P. Gondolo, E	
				Noulet (OCLA, TNST)
GRIEST	91	PR D43 3191	K. Griest, D. Seckel	(61116 51141)
KAMIONKOW.		PR D44 3021	M. Kamionkowski	(CHIC, FNAL)
MORI	91B	PL B270 89	M. Mori <i>et al.</i>	(Kamiokande Collab.)
NOJIRI	91	PL B261 76	M.M. Nojiri	(KEK)
OLIVE	91	NP B355 208	K.A. Olive, M. Srednicki	(MINN, ÚCSB)
ROSZKOWSKI	91	PL B262 59	L. Roszkowski	` (CERN)
GRIEST	90	PR D41 3565	K. Griest, M. Kamionkowski	
BARBIERI	89C	NP B313 725	R. Barbieri, M. Frigeni, G. (
OLIVE	89	PL B230 78	K.A. Olive, M. Srednicki	(MINN, UCSB)
ELLIS	88D	NP B307 883	J. Ellis, R. Flores	
GRIEST	88B	PR D38 2357	K. Griest	
OLIVE	88	PL B205 553	K.A. Olive, M. Srednicki	(MINN, UCSB)
SREDNICKI	88	NP B310 693	M. Srednicki, R. Watkins, K	(.A. Olive (MINN, UCSB)
ELLIS	84	NP B238 453	J. Ellis <i>et al.</i>	(CERN)
GOLDBERG	83	PRL 50 1419	H. Goldberg	(NEAS)
KRAUSS	83	NP B227 556	L.M. Krauss	(HARV)
VYSOTSKII				` .
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