



$$J = \frac{1}{2}$$

## $\mu$ MASS (atomic mass units u)

The muon's mass is obtained from the muon-electron mass ratio as determined from the measurement of Zeeman transition frequencies in muonium ( $\mu^+ e^-$  atom). Since the electron's mass is most accurately known in u, the muon's mass is also most accurately known in u. The conversion factor to MeV has approximately the same relative uncertainty as the mass of the muon in u. In this datablock we give the result in u, and in the following datablock in MeV.

<u>VALUE (u)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.1134289259 ± 0.0000000025</b>	TIESINGA	21	RVUE 2018 CODATA value
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
0.1134289257 ± 0.0000000025	MOHR	16	RVUE 2014 CODATA value
0.1134289267 ± 0.0000000029	MOHR	12	RVUE 2010 CODATA value
0.1134289256 ± 0.0000000029	MOHR	08	RVUE 2006 CODATA value
0.1134289264 ± 0.0000000030	MOHR	05	RVUE 2002 CODATA value
0.1134289168 ± 0.0000000034	<sup>1</sup> MOHR	99	RVUE 1998 CODATA value
0.113428913 ± 0.0000000017	<sup>2</sup> COHEN	87	RVUE 1986 CODATA value
<sup>1</sup> MOHR 99 make use of other 1998 CODATA entries below.			
<sup>2</sup> COHEN 87 make use of other 1986 CODATA entries below.			

## $\mu$ MASS

The mass is known more precisely in u (atomic mass units) than in MeV. The conversion is:  $1 \text{ u} = 931.494 102 42(28) \text{ MeV}/c^2$  (2018 CODATA value, TIESINGA 21). The conversion error contributes significantly to the uncertainty of the masses given below.

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>105.6583755 ± 0.0000023</b>	TIESINGA	21	RVUE	2018 CODATA value
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
105.6583745 ± 0.0000024	MOHR	16	RVUE	2014 CODATA value
105.6583715 ± 0.0000035	MOHR	12	RVUE	2010 CODATA value
105.6583668 ± 0.0000038	MOHR	08	RVUE	2006 CODATA value
105.6583692 ± 0.0000094	MOHR	05	RVUE	2002 CODATA value
105.6583568 ± 0.0000052	MOHR	99	RVUE	1998 CODATA value
105.658353 ± 0.000016	<sup>1</sup> COHEN	87	RVUE	1986 CODATA value
105.658386 ± 0.000044	<sup>2</sup> MARIAM	82	CNTR +	
105.65836 ± 0.00026	<sup>3</sup> CROWE	72	CNTR	
105.65865 ± 0.00044	<sup>4</sup> CRANE	71	CNTR	
<sup>1</sup> Converted to MeV using the 1998 CODATA value of the conversion constant, $931.494013 \pm 0.000037 \text{ MeV}/u$ .				
<sup>2</sup> MARIAM 82 give $m_\mu/m_e = 206.768259(62)$ .				
<sup>3</sup> CROWE 72 give $m_\mu/m_e = 206.7682(5)$ .				

<sup>4</sup> CRANE 71 give  $m_{\mu}/m_e = 206.76878(85)$ .

## $\mu$ MEAN LIFE $\tau$

Measurements with an error  $> 0.001 \times 10^{-6}$  s have been omitted.

VALUE ( $10^{-6}$ s)	DOCUMENT ID	TECN	CHG	COMMENT
<b>2.1969811 ± 0.0000022 OUR AVERAGE</b>				
2.1969803 ± 0.0000021 ± 0.0000007 <sup>1</sup>	TISHCHENKO 13	CNTR	+	Surface $\mu^+$ at PSI
2.197083 ± 0.000032 ± 0.000015	BARCZYK 08	CNTR	+	Muons from $\pi^+$ decay at rest
2.197013 ± 0.000021 ± 0.000011	CHITWOOD 07	CNTR	+	Surface $\mu^+$ at PSI
2.197078 ± 0.000073	BARDIN 84	CNTR	+	
2.197025 ± 0.000155	BARDIN 84	CNTR	-	
2.19695 ± 0.00006	GIOVANETTI 84	CNTR	+	
2.19711 ± 0.00008	BALANDIN 74	CNTR	+	
2.1973 ± 0.0003	DUCLOS 73	CNTR	+	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
2.1969803 ± 0.0000022	WEBBER 11	CNTR	+	Surface $\mu^+$ at PSI
<sup>1</sup> TISHCHENKO 13 uses $1.6 \times 10^{12}$ $\mu^+$ events and supersedes WEBBER 11.				

## $\tau_{\mu^+}/\tau_{\mu^-}$ MEAN LIFE RATIO

A test of *CPT* invariance.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>1.000024 ± 0.000078</b>	BARDIN 84	CNTR	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
1.0008 ± 0.0010	BAILEY 79	CNTR	Storage ring
1.000 ± 0.001	MEYER 63	CNTR	Mean life $\mu^+ / \mu^-$

## $(\tau_{\mu^+} - \tau_{\mu^-}) / \tau_{\text{average}}$

A test of *CPT* invariance. Calculated from the mean-life ratio, above.

VALUE	DOCUMENT ID
<b><math>(2 \pm 8) \times 10^{-5}</math> OUR EVALUATION</b>	

## $\mu/p$ MAGNETIC MOMENT RATIO

This ratio is used to obtain a precise value of the muon mass and to reduce experimental muon Larmor frequency measurements to the muon magnetic moment anomaly. Measurements with an error  $> 0.00001$  have been omitted. By convention, the minus sign on this ratio is omitted. CODATA values were fitted using their selection of data, plus other data from multiparameter fits.

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>3.183345142 ± 0.000000071</b>	TIESINGA 21	RVUE		2018 CODATA value

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.183345142±0.000000071	MOHR	16	RVUE	2014 CODATA value
3.183345107±0.000000084	MOHR	12	RVUE	2010 CODATA value
3.183345137±0.000000085	MOHR	08	RVUE	2006 CODATA value
3.183345118±0.000000089	MOHR	05	RVUE	2002 CODATA value
3.18334513 ±0.00000039	LIU	99	CNTR +	HFS in muonium
3.18334539 ±0.00000010	MOHR	99	RVUE	1998 CODATA value
3.18334547 ±0.00000047	COHEN	87	RVUE	1986 CODATA value
3.1833441 ±0.0000017	KLEMPPT	82	CNTR +	Precession strob
3.1833461 ±0.0000011	MARIAM	82	CNTR +	HFS splitting
3.1833448 ±0.0000029	CAMANI	78	CNTR +	See KLEMPPT 82
3.1833403 ±0.0000044	CASPERSON	77	CNTR +	HFS splitting
3.1833402 ±0.0000072	COHEN	73	RVUE	1973 CODATA value
3.1833467 ±0.0000082	CROWE	72	CNTR +	Precession phase

See the related review(s):  
[Muon Anomalous Magnetic Moment](#)

### $\mu$ MAGNETIC MOMENT ANOMALY

The parity-violating decay of muons in a storage ring is observed. The difference frequency  $\omega_a$  between the muon spin precession and the orbital angular frequency  $(e/m_\mu c)\langle B \rangle$  is measured, as is the free proton NMR frequency  $\omega_p$ , thus determining the ratio  $R=\omega_a/\omega_p$ . Given the magnetic moment ratio  $\lambda=\mu_\mu/\mu_p$  (from hyperfine structure in muonium),  $(g-2)/2 = R/(\lambda-R)$ .

$$\mu_\mu/(e\hbar/2m_\mu)-1 = (g_\mu-2)/2$$

<u>VALUE (units <math>10^{-10}</math>)</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>11659206.1± 4.1</b>	<sup>1</sup>	ABI	21	MUG2 ±	Combined FNAL and BNL values
• • • We do not use the following data for averages, fits, limits, etc. • • •					
11659204.0± 5.4		ABI	21	MUG2 +	Storage ring
11659208.0± 5.4±3.3		BENNETT	06	MUG2 ±	Average $\mu^+$ and $\mu^-$
11659208 ± 6		BENNETT	04	MUG2 ±	Average $\mu^+$ and $\mu^-$
11659214 ± 8 ±3		BENNETT	04	MUG2 -	Storage ring
11659203 ± 6 ±5		BENNETT	04	MUG2 +	Storage ring
11659204 ± 7 ±5		BENNETT	02	MUG2 +	Storage ring
11659202 ± 14 ±6		BROWN	01	MUG2 +	Storage ring
11659191 ± 59		BROWN	00	MUG2 +	
11659100 ± 110	<sup>2</sup>	BAILEY	79	CNTR +	Storage ring
11659360 ± 120	<sup>2</sup>	BAILEY	79	CNTR -	Storage ring
11659230 ± 85	<sup>2</sup>	BAILEY	79	CNTR ±	Storage ring
11620000 ±5000		CHARPAK	62	CNTR +	

<sup>1</sup> ABI 21 combined their value with the previous independent BNL measurement of BENNETT 06. ABI 21 also report that the difference of this combination with the standard model value of  $(11659181.0 \pm 4.3) \times 10^{-10}$  (AOYAMA 20) has a significance of  $4.2 \sigma$ .

<sup>2</sup> BAILEY 79 values recalculated by HUGHES 99 using the COHEN 87  $\mu/p$  magnetic moment. The improved MOHR 99 value does not change the result.

$$(g_{\mu^+} - g_{\mu^-}) / g_{\text{average}}$$

A test of *CPT* invariance.

<u>VALUE (units <math>10^{-8}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
<b><math>-0.11 \pm 0.12</math></b>	BENNETT 04	MUG2
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●		
$-2.6 \pm 1.6$	BAILEY 79	CNTR

## $\mu$ ELECTRIC DIPOLE MOMENT (d)

A nonzero value is forbidden by both *T* invariance and *P* invariance.

<u>VALUE (<math>10^{-19}</math> ecm)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>&lt; 1.8</b>	95	<sup>1</sup> BENNETT 09	MUG2	±	Storage ring
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$-0.1 \pm 1.0$		<sup>2</sup> BENNETT 09	MUG2	+	Storage ring
$-0.1 \pm 0.7$		<sup>3</sup> BENNETT 09	MUG2	-	Storage ring
$-3.7 \pm 3.4$		<sup>4</sup> BAILEY 78	CNTR	±	Storage ring
$8.6 \pm 4.5$		BAILEY 78	CNTR	+	Storage ring
$0.8 \pm 4.3$		BAILEY 78	CNTR	-	Storage ring

<sup>1</sup> This is the combination of the two BENNETT 09 measurements quoted here separately for  $\mu^+$  and  $\mu^-$ . The result is also presented as a measurement of  $(0.0 \pm 0.9) \times 10^{-19}$  e cm.

<sup>2</sup> Also reported as the limit of  $|d(\mu^+)| < 2.1 \times 10^{-19}$  e cm at 95% CL.

<sup>3</sup> Also reported as the limit of  $|d(\mu^-)| < 1.5 \times 10^{-19}$  e cm at 95% CL.

<sup>4</sup> This is the combination of the two BAILEY 78 results quoted here separately for  $\mu^+$  and  $\mu^-$ . BAILEY 78 uses the convention  $d = 1/2 \cdot (d_{\mu^+} - d_{\mu^-})$  and reports  $3.7 \pm 3.4$ . We convert their result to use the same convention as BENNETT 09.

## MUON-ELECTRON CHARGE RATIO ANOMALY $q_{\mu^+}/q_{e^-} + 1$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b><math>(1.1 \pm 2.1) \times 10^{-9}</math></b>	<sup>1</sup> MEYER 00	CNTR	+	1s-2s muonium interval

<sup>1</sup> MEYER 00 measure the 1s-2s muonium interval, and then interpret the result in terms of muon-electron charge ratio  $q_{\mu^+}/q_{e^-}$ .

## $\mu^-$ DECAY MODES

$\mu^+$  modes are charge conjugates of the modes below.

<u>Mode</u>	<u>Fraction (<math>\Gamma_i/\Gamma</math>)</u>	<u>Confidence level</u>
$\Gamma_1$ $e^- \bar{\nu}_e \nu_\mu$	$\approx 100\%$	
$\Gamma_2$ $e^- \bar{\nu}_e \nu_\mu \gamma$	[a] $(6.0 \pm 0.5) \times 10^{-8}$	
$\Gamma_3$ $e^- \bar{\nu}_e \nu_\mu e^+ e^-$	[b] $(3.4 \pm 0.4) \times 10^{-5}$	

### Lepton Family number (*LF*) violating modes

$\Gamma_4$	$e^- \nu_e \bar{\nu}_\mu$	<i>LF</i>	$[c] < 1.2$	%	90%
$\Gamma_5$	$e^- \gamma$	<i>LF</i>	$< 4.2$	$\times 10^{-13}$	90%
$\Gamma_6$	$e^- e^+ e^-$	<i>LF</i>	$< 1.0$	$\times 10^{-12}$	90%
$\Gamma_7$	$e^- 2\gamma$	<i>LF</i>	$< 7.2$	$\times 10^{-11}$	90%

[a] This only includes events with energy of  $e > 45$  MeV and energy of  $\gamma > 40$  MeV. Since the  $e^- \bar{\nu}_e \nu_\mu$  and  $e^- \bar{\nu}_e \nu_\mu \gamma$  modes cannot be clearly separated, we regard the latter mode as a subset of the former.

[b] See the Particle Listings below for the energy limits used in this measurement.

[c] A test of additive vs. multiplicative lepton family number conservation.

### $\mu^-$ BRANCHING RATIOS

$\Gamma(e^- \bar{\nu}_e \nu_\mu \gamma) / \Gamma_{\text{total}}$						$\Gamma_2 / \Gamma$
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>	
<b><math>(6.03 \pm 0.14 \pm 0.53) \times 10^{-8}</math></b>	13k	<sup>1</sup> BALDINI	16A	SPEC	$\gamma$ KE > 40 MeV	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
	862	BOGART	67	CNTR	$\gamma$ KE > 14.5 MeV	
$(3.3 \pm 1.3) \times 10^{-3}$		CRITTENDEN	61	CNTR	$\gamma$ KE > 20 MeV	
$(1.4 \pm 0.4) \times 10^{-2}$		CRITTENDEN	61	CNTR	$\gamma$ KE > 10 MeV	
	27	ASHKIN	59	CNTR		

<sup>1</sup> BALDINI 16 measurement refers to  $\mu^+ \rightarrow e^+ \nu \bar{\nu} \gamma$  decay and requires energy of  $e^+ > 45$  MeV and energy  $\gamma > 40$  MeV.

$\Gamma(e^- \bar{\nu}_e \nu_\mu e^+ e^-) / \Gamma_{\text{total}}$						$\Gamma_3 / \Gamma$
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>	
<b><math>3.4 \pm 0.2 \pm 0.3</math></b>	7443	<sup>1</sup> BERTL	85	SPEC	+	SINDRUM
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●						
$2.2 \pm 1.5$	7	<sup>2</sup> CRITTENDEN	61	HLBC	+	$E(e^+ e^-) > 10$ MeV
2	1	<sup>3</sup> GUREVICH	60	EMUL	+	
$1.5 \pm 1.0$	3	<sup>4</sup> LEE	59	HBC	+	

<sup>1</sup> BERTL 85 has transverse momentum cut  $p_T > 17$  MeV/c. Systematic error was increased by us.

<sup>2</sup> CRITTENDEN 61 count only those decays where total energy of either ( $e^+$ ,  $e^-$ ) combination is >10 MeV.

<sup>3</sup> GUREVICH 60 interpret their event as either virtual or real photon conversion.  $e^+$  and  $e^-$  energies not measured.

<sup>4</sup> In the three LEE 59 events, the sum of energies  $E(e^+) + E(e^-) + E(e^+)$  was 51 MeV, 55 MeV, and 33 MeV.

$\Gamma(e^- \nu_e \bar{\nu}_\mu) / \Gamma_{\text{total}}$						$\Gamma_4 / \Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>	
<b>&lt; 0.012</b>	90	<sup>1</sup> FREEDMAN	93	CNTR	+	$\nu$ oscillation search

Forbidden by the additive conservation law for lepton family number. A multiplicative law predicts this branching ratio to be 1/2. For a review see NEMETHY 81.

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.018	90	KRAKAUER	91B	CALO	+	
< 0.05	90	<sup>2</sup> BERGSMA	83	CALO		$\bar{\nu}_\mu e \rightarrow \mu^- \bar{\nu}_e$
< 0.09	90	JONKER	80	CALO		See BERGSMA 83
-0.001 ± 0.061		WILLIS	80	CNTR	+	
0.13 ± 0.15		BLIETSCHAU	78	HLBC	±	Avg. of 4 values
< 0.25	90	EICHTEN	73	HLBC	+	

<sup>1</sup> FREEDMAN 93 limit on  $\bar{\nu}_e$  observation is here interpreted as a limit on lepton family number violation.

<sup>2</sup> BERGSMA 83 gives a limit on the inverse muon decay cross-section ratio  $\sigma(\bar{\nu}_\mu e^- \rightarrow \mu^- \bar{\nu}_e) / \sigma(\nu_\mu e^- \rightarrow \mu^- \nu_e)$ , which is essentially equivalent to  $\Gamma(e^- \nu_e \bar{\nu}_\mu) / \Gamma_{\text{total}}$  for small values like that quoted.

### $\Gamma(e^- \gamma) / \Gamma_{\text{total}}$

$\Gamma_5 / \Gamma$

Forbidden by lepton family number conservation.

VALUE (units $10^{-11}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< <b>0.042</b>	90	BALDINI	16	SPEC	+ MEG at PSI

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.057	90	ADAM	13B	SPEC	+ MEG at PSI
< 0.24	90	ADAM	11	SPEC	+ MEG at PSI
< 2.8	90	ADAM	10	SPEC	+ MEG at PSI
< 1.2	90	AHMED	02	SPEC	+ MEGA
< 1.2	90	BROOKS	99	SPEC	+ LAMPF
< 4.9	90	BOLTON	88	CBOX	+ LAMPF
< 100	90	AZUELOS	83	CNTR	+ TRIUMF
< 17	90	KINNISON	82	SPEC	+ LAMPF
< 100	90	SCHAAF	80	ELEC	+ SIN

### $\Gamma(e^- e^+ e^-) / \Gamma_{\text{total}}$

$\Gamma_6 / \Gamma$

Forbidden by lepton family number conservation.

VALUE (units $10^{-12}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< <b>1.0</b>	90	<sup>1</sup> BELLGARDT	88	SPEC	+ SINDRUM

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 36	90	BARANOV	91	SPEC	+ ARES
< 35	90	BOLTON	88	CBOX	+ LAMPF
< 2.4	90	<sup>1</sup> BERTL	85	SPEC	+ SINDRUM
< 160	90	<sup>1</sup> BERTL	84	SPEC	+ SINDRUM
< 130	90	<sup>1</sup> BOLTON	84	CNTR	+ LAMPF

<sup>1</sup> These experiments assume a constant matrix element.

### $\Gamma(e^- 2\gamma) / \Gamma_{\text{total}}$

$\Gamma_7 / \Gamma$

Forbidden by lepton family number conservation.

VALUE (units $10^{-11}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
< <b>7.2</b>	90	BOLTON	88	CBOX	+ LAMPF

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 1	90	<sup>1</sup> BALDINI	20	MEG	PSI
< 840	90	<sup>2</sup> AZUELOS	83	CNTR	+ TRIUMF
< 5000	90	<sup>3</sup> BOWMAN	78	CNTR	DEPOMMIER 77 data

<sup>1</sup> BALDINI 20 uses  $7.5 \times 10^{14}$  stopped muons to obtain limits on  $\mu \rightarrow eX$  decay mediated by a new light particle  $X$  with lifetimes  $< 40$  ps, which decays to  $\gamma\gamma$  for  $X$ -mass ranges 20–45 MeV. The limit of the order  $< 10^{-11}$  at 90% CL is for the mass range 20–30 MeV.

<sup>2</sup> AZUELOS 83 uses the phase space distribution of BOWMAN 78.

<sup>3</sup> BOWMAN 78 assumes an interaction Lagrangian local on the scale of the inverse  $\mu$  mass.

## LIMIT ON $\mu^- \rightarrow e^-$ CONVERSION

Forbidden by lepton family number conservation.

### $\sigma(\mu^- {}^{32}\text{S} \rightarrow e^- {}^{32}\text{S}) / \sigma(\mu^- {}^{32}\text{S} \rightarrow \nu_\mu {}^{32}\text{P}^*)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 7 \times 10^{-11}$	90	BADERT...	80	STRC SIN
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4 \times 10^{-10}$	90	BADERT...	77	STRC SIN

### $\sigma(\mu^- \text{Cu} \rightarrow e^- \text{Cu}) / \sigma(\mu^- \text{Cu} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.6 \times 10^{-8}$	90	BRYMAN	72	SPEC

### $\sigma(\mu^- \text{Ti} \rightarrow e^- \text{Ti}) / \sigma(\mu^- \text{Ti} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.3 \times 10^{-12}$	90	<sup>1</sup> DOHMEN	93	SPEC SINDRUM II
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.6 \times 10^{-12}$	90	AHMAD	88	TPC TRIUMF
$< 1.6 \times 10^{-11}$	90	BRYMAN	85	TPC TRIUMF

<sup>1</sup> DOHMEN 93 assumes  $\mu^- \rightarrow e^-$  conversion leaves the nucleus in its ground state, a process enhanced by coherence and expected to dominate.

### $\sigma(\mu^- \text{Pb} \rightarrow e^- \text{Pb}) / \sigma(\mu^- \text{Pb} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 4.6 \times 10^{-11}$	90	HONECKER	96	SPEC SINDRUM II
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.9 \times 10^{-10}$	90	AHMAD	88	TPC TRIUMF

### $\sigma(\mu^- \text{Au} \rightarrow e^- \text{Au}) / \sigma(\mu^- \text{Au} \rightarrow \text{capture})$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
$< 7 \times 10^{-13}$	90	BERTL	06	SPEC –	SINDRUM II

## LIMIT ON $\mu^- \rightarrow e^+$ CONVERSION

Forbidden by total lepton number conservation.

### $\sigma(\mu^- {}^{32}\text{S} \rightarrow e^+ {}^{32}\text{Si}^*) / \sigma(\mu^- {}^{32}\text{S} \rightarrow \nu_\mu {}^{32}\text{P}^*)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 9 \times 10^{-10}$	90	BADERT...	80	STRC SIN
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.5 \times 10^{-9}$	90	BADERT...	78	STRC SIN

**$\sigma(\mu^- 127\text{I} \rightarrow e^+ 127\text{Sb}^*) / \sigma(\mu^- 127\text{I} \rightarrow \text{anything})$**

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3 \times 10^{-10}$	90	<sup>1</sup> ABELA 80	CNTR	Radiochemical tech.

<sup>1</sup> ABELA 80 is upper limit for  $\mu^- e^+$  conversion leading to particle-stable states of  $127\text{Sb}$ . Limit for total conversion rate is higher by a factor less than 4 (G. Backenstoss, private communication).

**$\sigma(\mu^- \text{Cu} \rightarrow e^+ \text{Co}) / \sigma(\mu^- \text{Cu} \rightarrow \nu_\mu \text{Ni})$**

VALUE	CL%	DOCUMENT ID	TECN
$<2.6 \times 10^{-8}$	90	BRYMAN 72	SPEC
$<2.2 \times 10^{-7}$	90	CONFORTO 62	OSPK

**$\sigma(\mu^- \text{Ti} \rightarrow e^+ \text{Ca}) / \sigma(\mu^- \text{Ti} \rightarrow \text{capture})$**

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
$<3.6 \times 10^{-11}$	90	1	<sup>1,2</sup> KAULARD 98	SPEC	–	SINDRUM II
$<1.7 \times 10^{-12}$	90	1	<sup>2,3</sup> KAULARD 98	SPEC	–	SINDRUM II
$<4.3 \times 10^{-12}$	90		<sup>3</sup> DOHMEN 93	SPEC		SINDRUM II
$<8.9 \times 10^{-11}$	90		<sup>1</sup> DOHMEN 93	SPEC		SINDRUM II
$<1.7 \times 10^{-10}$	90		<sup>4</sup> AHMAD 88	TPC		TRIUMF

• • • We do not use the following data for averages, fits, limits, etc. • • •  
<sup>1</sup> This limit assumes a giant resonance excitation of the daughter Ca nucleus (mean energy and width both 20 MeV).  
<sup>2</sup> KAULARD 98 obtained these same limits using the unified classical analysis of FELDMAN 98.  
<sup>3</sup> This limit assumes the daughter Ca nucleus is left in the ground state. However, the probability of this is unknown.  
<sup>4</sup> Assuming a giant-resonance-excitation model.

**LIMIT ON MUONIUM  $\rightarrow$  ANTIMUONIUM CONVERSION**

Forbidden by lepton family number conservation.

**$R_g = G_C / G_F$**

The effective Lagrangian for the  $\mu^+ e^- \rightarrow \mu^- e^+$  conversion is assumed to be

$$\mathcal{L} = 2^{-1/2} G_C [\bar{\psi}_\mu \gamma_\lambda (1 - \gamma_5) \psi_e] [\bar{\psi}_\mu \gamma_\lambda (1 - \gamma_5) \psi_e] + \text{h.c.}$$

The experimental result is then an upper limit on  $G_C/G_F$ , where  $G_F$  is the Fermi coupling constant.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b><math>&lt; 0.0030</math></b>	90	1	<sup>1</sup> WILLMANN 99	SPEC	+	$\mu^+$ at 26 GeV/c
$< 0.14$	90	1	<sup>2</sup> GORDEEV 97	SPEC	+	JINR phasotron
$< 0.018$	90	0	<sup>3</sup> ABELA 96	SPEC	+	$\mu^+$ at 24 MeV
$< 6.9$	90		NI 93	CBOX		LAMPF
$< 0.16$	90		MATTHIAS 91	SPEC		LAMPF
$< 0.29$	90		HUBER 90B	CNTR		TRIUMF
$< 20$	95		BEER 86	CNTR		TRIUMF
$< 42$	95		MARSHALL 82	CNTR		

<sup>1</sup> WILLMANN 99 quote both probability  $P_{M\bar{M}} < 8.3 \times 10^{-11}$  at 90%CL in a 0.1 T field and  $R_g = G_C/G_F$ .



<sup>2</sup> GORDEEV 97 quote limits on both  $f=G_{MM}/G_F$  and the probability  $W_{MM} < 4.7 \times 10^{-7}$  (90% CL).

<sup>3</sup> ABELA 96 quote both probability  $P_{M\bar{M}} < 8 \times 10^{-9}$  at 90% CL and  $R_g = G_C/G_F$ .

See the related review(s):  
[Muon Decay Parameters](#)

**$\mu$  DECAY PARAMETERS**

**$\rho$  PARAMETER**

( $V-A$ ) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.74979 ± 0.00026 OUR AVERAGE</b>					
0.74977 ± 0.00012 ± 0.00023		<sup>1</sup> BAYES	11 TWST	+	Surface $\mu^+$
0.7518 ± 0.0026		DERENZO	69 RVUE		
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
0.75014 ± 0.00017 ± 0.00045		<sup>2</sup> MACDONALD	08 TWST	+	Surface $\mu^+$
0.75080 ± 0.00032 ± 0.00100	6G	<sup>3</sup> MUSSER	05 TWST	+	Surface $\mu^+$
0.72 ± 0.06 ± 0.08		AMORUSO	04 ICAR		Liquid Ar TPC
0.762 ± 0.008	170k	<sup>4</sup> FRYBERGER	68 ASPK	+	25–53 MeV $e^+$
0.760 ± 0.009	280k	<sup>4</sup> SHERWOOD	67 ASPK	+	25–53 MeV $e^+$
0.7503 ± 0.0026	800k	<sup>4</sup> PEOPLES	66 ASPK	+	20–53 MeV $e^+$

<sup>1</sup> The quoted systematic error includes a contribution of 0.00013 (added in quadrature) from uncertainties on radiative corrections and on the Michel parameter  $\eta$ .

<sup>2</sup> The quoted systematic error includes a contribution of 0.00011 (added in quadrature) from the dependence on the Michel parameter  $\eta$ .

<sup>3</sup> The quoted systematic error includes a contribution of 0.00023 (added in quadrature) from the dependence on the Michel parameter  $\eta$ .

<sup>4</sup>  $\eta$  constrained = 0. These values incorporated into a two parameter fit to  $\rho$  and  $\eta$  by DERENZO 69.

**$\eta$  PARAMETER**

( $V-A$ ) theory predicts  $\eta = 0$ .

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.057 ± 0.034 OUR AVERAGE</b>					
0.071 ± 0.037 ± 0.005	30M	DANNEBERG	05 CNTR	+	7–53 MeV $e^+$
0.011 ± 0.081 ± 0.026	5.3M	<sup>1</sup> BURKARD	85BCNTR	+	9–53 MeV $e^+$
−0.12 ± 0.21	6346	DERENZO	69 HBC	+	1.6–6.8 MeV $e^+$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
−0.0021 ± 0.0070 ± 0.0010	30M	<sup>2</sup> DANNEBERG	05 CNTR	+	7–53 MeV $e^+$
−0.012 ± 0.015 ± 0.003	5.3M	<sup>2</sup> BURKARD	85BCNTR	+	9–53 MeV $e^+$
−0.007 ± 0.013	5.3M	<sup>3</sup> BURKARD	85BFIT	+	9–53 MeV $e^+$
−0.7 ± 0.5	170k	<sup>4</sup> FRYBERGER	68 ASPK	+	25–53 MeV $e^+$
−0.7 ± 0.6	280k	<sup>4</sup> SHERWOOD	67 ASPK	+	25–53 MeV $e^+$
0.05 ± 0.5	800k	<sup>4</sup> PEOPLES	66 ASPK	+	20–53 MeV $e^+$
−2.0 ± 0.9	9213	<sup>5</sup> PLANO	60 HBC	+	Whole spectrum

<sup>1</sup> Previously we used the global fit result from BURKARD 85B in OUR AVERAGE, we now only include their actual measurement.

<sup>2</sup>  $\alpha = \alpha' = 0$  assumed.

<sup>3</sup> Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

<sup>4</sup>  $\rho$  constrained = 0.75.

<sup>5</sup> Two parameter fit to  $\rho$  and  $\eta$ ; PLANO 60 discounts value for  $\eta$ .

**$\delta$  PARAMETER** $(V-A)$  theory predicts  $\delta = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>0.75047 ± 0.00034 OUR AVERAGE</b>					
0.75049 ± 0.00021 ± 0.00027		<sup>1</sup> BAYES	11	TWST +	Surface $\mu^+$
0.7486 ± 0.0026 ± 0.0028		<sup>2</sup> BALKE	88	SPEC +	Surface $\mu^+$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
0.75067 ± 0.00030 ± 0.00067		MACDONALD	08	TWST +	Surface $\mu^+$
0.74964 ± 0.00066 ± 0.00112	6G	GAPONENKO	05	TWST +	Surface $\mu^+$
		<sup>3</sup> VOSSLER	69		
0.752 ± 0.009	490k	FRYBERGER	68	ASPK +	25–53 MeV $e^+$
0.782 ± 0.031		KRUGER	61		
0.78 ± 0.05	8354	PLANO	60	HBC +	Whole spectrum

<sup>1</sup> The quoted systematic error includes a contribution of 0.00006 (added in quadrature) from uncertainties on radiative corrections and on the Michel parameter  $\eta$ .<sup>2</sup> BALKE 88 uses  $\rho = 0.752 \pm 0.003$ .<sup>3</sup> VOSSLER 69 has measured the asymmetry below 10 MeV. See comments about radiative corrections in VOSSLER 69. **$(\xi \text{ PARAMETER}) \times (\mu \text{ LONGITUDINAL POLARIZATION})$**  $(V-A)$  theory predicts  $\xi = 1$ , longitudinal polarization = 1.

VALUE	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.0009 <math>\begin{smallmatrix} +0.0016 \\ -0.0007 \end{smallmatrix}</math> OUR AVERAGE</b>				
1.00084 ± 0.00029 $\begin{smallmatrix} +0.00165 \\ -0.00063 \end{smallmatrix}$	BUENO	11	TWST	Surface $\mu^+$ beam
1.0027 ± 0.0079 ± 0.0030	BELTRAMI	87	CNTR	SIN, $\pi$ decay in flight
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1.0003 ± 0.0006 ± 0.0038	JAMIESON	06	TWST +	surface $\mu^+$ beam
1.0013 ± 0.0030 ± 0.0053	<sup>1</sup> IMAZATO	92	SPEC +	$K^+ \rightarrow \mu^+ \nu_\mu$
0.975 ± 0.015	AKHMANOV	68	EMUL	140 kG
0.975 ± 0.030	GUREVICH	64	EMUL	See AKHMANOV 68
0.903 ± 0.027	<sup>2</sup> ALI-ZADE	61	EMUL +	27 kG
0.93 ± 0.06	PLANO	60	HBC +	8.8 kG
0.97 ± 0.05	BARDON	59	CNTR	Bromoform target

<sup>1</sup> The corresponding 90% confidence limit from IMAZATO 92 is  $|\xi P_\mu| > 0.990$ . This measurement is of  $K^+$  decay, not  $\pi^+$  decay, so we do not include it in an average, nor do we yet set up a separate data block for  $K$  results.<sup>2</sup> Depolarization by medium not known sufficiently well. **$\xi \times (\mu \text{ LONGITUDINAL POLARIZATION}) \times \delta / \rho$** 

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>1.00179 <math>\begin{smallmatrix} +0.00156 \\ -0.00071 \end{smallmatrix}</math></b>					
		<sup>1</sup> BAYES	11	TWST +	Surface $\mu^+$ beam
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
>0.99682	90	<sup>2</sup> JODIDIO	86	SPEC +	TRIUMF
>0.9966	90	<sup>3</sup> STOKER	85	SPEC +	$\mu$ -spin rotation
>0.9959	90	CARR	83	SPEC +	11 kG

<sup>1</sup> BAYES 11 obtains the limit  $> 0.99909$  (90% CL) with the constraint that  $\xi \times (\mu \text{ LONGITUDINAL POLARIZATION}) \times \delta / \rho \leq 1.0$ .

<sup>2</sup> JODIDIO 86 includes data from CARR 83 and STOKER 85. The value here is from the erratum.

<sup>3</sup> STOKER 85 find  $(\xi P_{\mu} \delta/\rho) > 0.9955$  and  $> 0.9966$ , where the first limit is from new  $\mu$  spin-rotation data and the second is from combination with CARR 83 data. In  $V-A$  theory,  $(\delta/\rho) = 1.0$ .

### $\xi' = \text{LONGITUDINAL POLARIZATION OF } e^+$

( $V-A$ ) theory predicts the longitudinal polarization =  $\pm 1$  for  $e^{\pm}$ , respectively. We have flipped the sign for  $e^-$  so our programs can average.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>1.00 ± 0.04 OUR AVERAGE</b>					
0.998 ± 0.045	1M	BURKARD	85	CNTR +	Bhabha + annihil
0.89 ± 0.28	29k	SCHWARTZ	67	OSPK -	Moller scattering
0.94 ± 0.38		BLOOM	64	CNTR +	Brems. transmiss.
1.04 ± 0.18		DUCLOS	64	CNTR +	Bhabha scattering
1.05 ± 0.30		BUHLER	63	CNTR +	Annihilation

### $\xi''$ PARAMETER

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>0.98 ± 0.04 OUR AVERAGE</b>					
0.981 ± 0.045 ± 0.003	3.87M	PRIEELS	14	CNTR +	Bhabha + annihil
0.65 ± 0.36	326k	<sup>1</sup> BURKARD	85	CNTR +	Bhabha + annihil

<sup>1</sup> BURKARD 85 measure  $(\xi'' - \xi\xi')/\xi$  and  $\xi'$  and set  $\xi = 1$ .

### TRANSVERSE $e^+$ POLARIZATION IN PLANE OF $\mu$ SPIN, $e^+$ MOMENTUM

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>7 ± 8 OUR AVERAGE</b>					
6.3 ± 7.7 ± 3.4	30M	DANNEBERG	05	CNTR +	7-53 MeV $e^+$
16 ± 21 ± 10	5.3M	BURKARD	85B	CNTR +	Annihil 9-53 MeV

### TRANSVERSE $e^+$ POLARIZATION NORMAL TO PLANE OF $\mu$ SPIN, $e^+$ MOMENTUM

Zero if  $T$  invariance holds.

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>-2 ± 8 OUR AVERAGE</b>					
-3.7 ± 7.7 ± 3.4	30M	DANNEBERG	05	CNTR +	7-53 MeV $e^+$
7 ± 22 ± 7	5.3M	BURKARD	85B	CNTR +	Annihil 9-53 MeV

### $\alpha/A$

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>0.4 ± 4.3</b>					
		<sup>1</sup> BURKARD	85B	FIT	

• • • We do not use the following data for averages, fits, limits, etc. • • •

15 ± 50 ± 14	5.3M	BURKARD	85B	CNTR +	9-53 MeV $e^+$
--------------	------	---------	-----	--------	----------------

<sup>1</sup> Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

**$\alpha'/A$** Zero if  $T$  invariance holds.

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>-10 ±20 OUR AVERAGE</b>					
- 3.4±21.3± 4.9	30M	DANNEBERG 05	CNTR	+	7-53 MeV $e^+$
-47 ±50 ±14	5.3M	<sup>1</sup> BURKARD 85B	CNTR	+	9-53 MeV $e^+$
••• We do not use the following data for averages, fits, limits, etc. •••					
- 0.2± 4.3		<sup>2</sup> BURKARD 85B	FIT		

<sup>1</sup>Previously we used the global fit result from BURKARD 85B in OUR AVERAGE, we now only include their actual measurement. BURKARD 85B measure  $e^+$  polarizations  $P_{T_1}$  and  $P_{T_2}$  versus  $e^+$  energy.

<sup>2</sup>Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

 **$\beta/A$** 

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>3.9± 6.2</b>		<sup>1</sup> BURKARD 85B	FIT		
••• We do not use the following data for averages, fits, limits, etc. •••					
2 ±17 ±6	5.3M	BURKARD 85B	CNTR	+	9-53 MeV $e^+$

<sup>1</sup>Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

 **$\beta'/A$** Zero if  $T$  invariance holds.

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b>2 ± 7 OUR AVERAGE</b>					
- 0.5± 7.8±1.8	30M	DANNEBERG 05	CNTR	+	7-53 MeV $e^+$
17 ±17 ±6	5.3M	<sup>1</sup> BURKARD 85B	CNTR	+	9-53 MeV $e^+$
••• We do not use the following data for averages, fits, limits, etc. •••					
- 1.3± 3.5±0.6	30M	<sup>2</sup> DANNEBERG 05	CNTR	+	7-53 MeV $e^+$
1.5± 6.3		<sup>3</sup> BURKARD 85B	FIT		

<sup>1</sup>Previously we used the global fit result from BURKARD 85B in OUR AVERAGE, we now only include their actual measurement. BURKARD 85B measure  $e^+$  polarizations  $P_{T_1}$  and  $P_{T_2}$  versus  $e^+$  energy.

<sup>2</sup> $\alpha = \alpha' = 0$  assumed.

<sup>3</sup>Global fit to all measured parameters. The fit correlation coefficients are given in BURKARD 85B.

 **$a/A$** 

This comes from an alternative parameterization to that used in the Summary Table (see the "Note on Muon Decay Parameters" above).

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
••• We do not use the following data for averages, fits, limits, etc. •••			
<15.9	90	<sup>1</sup> BURKARD 85B	FIT

<sup>1</sup>Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

**$a'/A$** 

This comes from an alternative parameterization to that used in the Summary Table (see the “Note on Muon Decay Parameters” above).

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
---	--------------------	-------------

• • • We do not use the following data for averages, fits, limits, etc. • • •

$5.3 \pm 4.1$	<sup>1</sup> BURKARD	85B FIT
---------------	----------------------	---------

<sup>1</sup>Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 **$(b'+b)/A$** 

This comes from an alternative parameterization to that used in the Summary Table (see the “Note on Muon Decay Parameters” above).

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
---	------------	--------------------	-------------

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.04	90	<sup>1</sup> BURKARD	85B FIT
-------	----	----------------------	---------

<sup>1</sup>Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 **$c/A$** 

This comes from an alternative parameterization to that used in the Summary Table (see the “Note on Muon Decay Parameters” above).

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
---	------------	--------------------	-------------

• • • We do not use the following data for averages, fits, limits, etc. • • •

<6.4	90	<sup>1</sup> BURKARD	85B FIT
------	----	----------------------	---------

<sup>1</sup>Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 **$c'/A$** 

This comes from an alternative parameterization to that used in the Summary Table (see the “Note on Muon Decay Parameters” above).

<u>VALUE (units <math>10^{-3}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
---	--------------------	-------------

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.5 \pm 2.0$	<sup>1</sup> BURKARD	85B FIT
---------------	----------------------	---------

<sup>1</sup>Global fit to all measured parameters. Correlation coefficients are given in BURKARD 85B.

 **$\bar{\eta}$  PARAMETER**

( $V-A$ ) theory predicts  $\bar{\eta} = 0$ .  $\bar{\eta}$  affects spectrum of radiative muon decay.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
--------------	--------------------	-------------	------------	----------------

**0.02 ± 0.08 OUR AVERAGE**

$-0.014 \pm 0.090$	EICHENBER...	84	ELEC	+	$\rho$ free
--------------------	--------------	----	------	---	-------------

$+0.09 \pm 0.14$	BOGART	67	CNTR	+	
------------------	--------	----	------	---	--

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.035 \pm 0.098$	EICHENBER...	84	ELEC	+	$\rho=0.75$ assumed
--------------------	--------------	----	------	---	---------------------

---

$\mu$  REFERENCES

ABI	21	PRL 126 141801	B. Abi <i>et al.</i>	(Muon g-2 Collab.)
TIESINGA	21	RMP 93 025010	E. Tiesinga <i>et al.</i>	(NIST)
AOYAMA	20	PRPL 887 1	T. Aoyama <i>et al.</i>	
BALDINI	20	EPJ C80 858	A.M. Baldini <i>et al.</i>	(MEG Collab.)
BALDINI	16	EPJ C76 434	A.M. Baldini <i>et al.</i>	(MEG Collab.)
BALDINI	16A	EPJ C76 108	A.M. Baldini <i>et al.</i>	(MEG Collab.)
MOHR	16	RMP 88 035009	P.J. Mohr, D.B. Newell, B.N. Taylor	(NIST)
PRIEELS	14	PR D90 112003	R. Prieels <i>et al.</i>	(LOUV, ETH, PSI+)
ADAM	13B	PRL 110 201801	J. Adam <i>et al.</i>	(MEG Collab.)
TISHCHENKO	13	PR D87 052003	V. Tishchenko <i>et al.</i>	(MuLan Collab.)
MOHR	12	RMP 84 1527	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
ADAM	11	PRL 107 171801	J. Adam <i>et al.</i>	(MEG Collab.)
BAYES	11	PRL 106 041804	R. Bayes <i>et al.</i>	(TWIST Collab.)
Also		PR D85 092013	A. Hillairet <i>et al.</i>	(TWIST Collab.)
BUENO	11	PR D84 032005	J.F. Bueno <i>et al.</i>	(TWIST Collab.)
Also		PR D85 039908 (errata.)	J.F. Bueno <i>et al.</i>	(TWIST Collab.)
WEBBER	11	PRL 106 041803	D.M. Webber <i>et al.</i>	(MuLan Collab.)
Also		PRL 106 079901 (errata.)	D.M. Webber <i>et al.</i>	(MuLan Collab.)
ADAM	10	NP B834 1	J. Adam <i>et al.</i>	(MEG Collab.)
BENNETT	09	PR D80 052008	G.W. Bennett <i>et al.</i>	(MUG-2 Collab.)
BARCZYK	08	PL B663 172	A. Barczyk <i>et al.</i>	(FAST Collab.)
MACDONALD	08	PR D78 032010	R.P. MacDonald <i>et al.</i>	(TWIST Collab.)
MOHR	08	RMP 80 633	P.J. Mohr, B.N. Taylor, D.B. Newell	(NIST)
CHITWOOD	07	PRL 99 032001	D.B. Chitwood <i>et al.</i>	(MULAN Collab.)
BENNETT	06	PR D73 072003	G.W. Bennett <i>et al.</i>	(MUG-2 Collab.)
BERTL	06	EPJ C47 337	W. Bertl <i>et al.</i>	(SINDRUM II Collab.)
JAMIESON	06	PR D74 072007	B. Jamieson <i>et al.</i>	(TWIST Collab.)
DANNEBERG	05	PRL 94 021802	N. Danneberg <i>et al.</i>	(ETH, JAGL, PSI+)
GAPONENKO	05	PR D71 071101	A. Gaponenko <i>et al.</i>	(TWIST Collab.)
MOHR	05	RMP 77 1	P.J. Mohr, B.N. Taylor	(NIST)
MUSSER	05	PRL 94 101805	J.R. Musser <i>et al.</i>	(TWIST Collab.)
AMORUSO	04	EPJ C33 233	S. Amoruso <i>et al.</i>	(ICARUS Collab.)
BENNETT	04	PRL 92 161802	G.W. Bennett <i>et al.</i>	(Muon(g-2) Collab.)
AHMED	02	PR D65 112002	M. Ahmed <i>et al.</i>	(MEGA Collab.)
BENNETT	02	PRL 89 101804	G.W. Bennett <i>et al.</i>	(Muon(g-2) Collab.)
BROWN	01	PRL 86 2227	H.N. Brown <i>et al.</i>	(Muon(g-2) Collab.)
BROWN	00	PR D62 091101	H.N. Brown <i>et al.</i>	(BNL/G-2 Collab.)
MEYER	00	PRL 84 1136	V. Meyer <i>et al.</i>	
BROOKS	99	PRL 83 1521	M.L. Brooks <i>et al.</i>	(MEGA/LAMPF Collab.)
HUGHES	99	RMP 71 5133	V.V. Hughes, T. Kinoshita	
LIU	99	PRL 82 711	W. Liu <i>et al.</i>	(LAMPF Collab.)
MOHR	99	JPCRD 28 1713	P.J. Mohr, B.N. Taylor	(NIST)
Also		RMP 72 351	P.J. Mohr, B.N. Taylor	(NIST)
WILLMANN	99	PRL 82 49	L. Willmann <i>et al.</i>	
FELDMAN	98	PR D57 3873	G.J. Feldman, R.D. Cousins	
KAULARD	98	PL B422 334	J. Kaulard <i>et al.</i>	(SINDRUM-II Collab.)
GORDEEV	97	PAN 60 1164	V.A. Gordeev <i>et al.</i>	(PNPI)
		Translated from YAF 60 1291.		
ABELA	96	PRL 77 1950	R. Abela <i>et al.</i>	(PSI, ZURI, HEIDH, TBIL+)
HONECKER	96	PRL 76 200	W. Honecker <i>et al.</i>	(SINDRUM II Collab.)
DOHMEN	93	PL B317 631	C. Dohmen <i>et al.</i>	(PSI SINDRUM-II Collab.)
FREEDMAN	93	PR D47 811	S.J. Freedman <i>et al.</i>	(LAMPF E645 Collab.)
NI	93	PR D48 1976	B. Ni <i>et al.</i>	(LAMPF Crystal-Box Collab.)
IMAZATO	92	PRL 69 877	J. Imazato <i>et al.</i>	(KEK, INUS, TOKY+)
BARANOV	91	SJNP 53 802	V.A. Baranov <i>et al.</i>	(JINR)
		Translated from YAF 53 1302.		
KRAKAUER	91B	PL B263 534	D.A. Krakauer <i>et al.</i>	(UMD, UCI, LANL)
MATTHIAS	91	PRL 66 2716	B.E. Matthias <i>et al.</i>	(YALE, HEIDP, WILL+)
Also		PRL 67 932 (erratum)	B.E. Matthias <i>et al.</i>	(YALE, HEIDP, WILL+)
HUBER	90B	PR D41 2709	T.M. Huber <i>et al.</i>	(WYOM, VICT, ARIZ+)
AHMAD	88	PR D38 2102	S. Ahmad <i>et al.</i>	(TRIU, VICT, VPI, BRCO+)
Also		PR 59 970	S. Ahmad <i>et al.</i>	(TRIU, VPI, VICT, BRCO+)
BALKE	88	PR D37 587	B. Balke <i>et al.</i>	(LBL, UCB, COLO, NWES+)
BELLGARDT	88	NP B299 1	U. Bellgardt <i>et al.</i>	(SINDRUM Collab.)
BOLTON	88	PR D38 2077	R.D. Bolton <i>et al.</i>	(LANL, STAN, CHIC+)
Also		PRL 56 2461	R.D. Bolton <i>et al.</i>	(LANL, STAN, CHIC+)
Also		PRL 57 3241	D. Grosnick <i>et al.</i>	(CHIC, LANL, STAN+)
BELTRAMI	87	PL B194 326	I. Beltrami <i>et al.</i>	(ETH, SIN, MAINZ)
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RISC, NBS)

BEER	86	PRL 57 671	G.A. Beer <i>et al.</i>	(VICT, TRIU, WYOM)
JODIDIO	86	PR D34 1967	A. Jodidio <i>et al.</i>	(LBL, NWES, TRIU)
Also		PR D37 237 (erratum)	A. Jodidio <i>et al.</i>	(LBL, NWES, TRIU)
BERTL	85	NP B260 1	W. Bertl <i>et al.</i>	(SINDRUM Collab.)
BRYMAN	85	PRL 55 465	D.A. Bryman <i>et al.</i>	(TRIU, CNRC, BRCO+)
BURKARD	85	PL 150B 242	H. Burkhardt <i>et al.</i>	(ETH, SIN, MAINZ)
BURKARD	85B	PL 160B 343	H. Burkhardt <i>et al.</i>	(ETH, SIN, MAINZ)
Also		PR D24 2004	F. Corriveau <i>et al.</i>	(ETH, SIN, MAINZ)
Also		PL 129B 260	F. Corriveau <i>et al.</i>	(ETH, SIN, MAINZ)
STOKER	85	PRL 54 1887	D.P. Stoker <i>et al.</i>	(LBL, NWES, TRIU)
BARDIN	84	PL 137B 135	G. Bardin <i>et al.</i>	(SACL, CERN, BGNA, FIRZ)
BERTL	84	PL 140B 299	W. Bertl <i>et al.</i>	(SINDRUM Collab.)
BOLTON	84	PRL 53 1415	R.D. Bolton <i>et al.</i>	(LANL, CHIC, STAN+)
EICHENBER...	84	NP A412 523	W. Eichenberger, R. Engfer, A. van der Schaff	
GIOVANETTI	84	PR D29 343	K.L. Giovanetti <i>et al.</i>	(WILL)
AZUELOS	83	PRL 51 164	G. Azuelos <i>et al.</i>	(MONT, TRIU, BRCO)
Also		PRL 39 1113	P. Depommier <i>et al.</i>	(MONT, BRCO, TRIU+)
BERGSMA	83	PL 122B 465	F. Bergsma <i>et al.</i>	(CHARM Collab.)
CARR	83	PRL 51 627	J. Carr <i>et al.</i>	(LBL, NWES, TRIU)
KINNISON	82	PR D25 2846	W.W. Kinnison <i>et al.</i>	(EFI, STAN, LANL)
Also		PRL 42 556	J.D. Bowman <i>et al.</i>	(LASL, EFI, STAN)
KLEMPPT	82	PR D25 652	E. Klempt <i>et al.</i>	(MAINZ, ETH)
MARIAM	82	PRL 49 993	F.G. Mariam <i>et al.</i>	(YALE, HEIDH, BERN)
MARSHALL	82	PR D25 1174	G.M. Marshall <i>et al.</i>	(BRCO)
NEMETHY	81	CNPP 10 147	P. Nemethy, V.W. Hughes	(LBL, YALE)
ABELA	80	PL 95B 318	R. Abela <i>et al.</i>	(BASL, KARLK, KARLE)
BADERT...	80	LNC 28 401	A. Badertscher <i>et al.</i>	(BERN)
Also		NP A377 406	A. Badertscher <i>et al.</i>	(BERN)
JONKER	80	PL 93B 203	M. Jonker <i>et al.</i>	(CHARM Collab.)
SCHAAF	80	NP A340 249	A. van der Schaaf <i>et al.</i>	(ZURI, ETH+)
Also		PL 72B 183	H.P. Povel <i>et al.</i>	(ZURI, ETH, SIN)
WILLIS	80	PRL 44 522	S.E. Willis <i>et al.</i>	(YALE, LBL, LASL+)
Also		PRL 45 1370	S.E. Willis <i>et al.</i>	(YALE, LBL, LASL+)
BAILEY	79	NP B150 1	J.M. Bailey	(CERN, DARE, MAINZ)
BADERT...	78	PL 79B 371	A. Badertscher <i>et al.</i>	(BERN)
BAILEY	78	JP G4 345	J.M. Bailey	(DARE, BERN, SHEF, MAINZ, RMCS+)
Also		NP B150 1	J.M. Bailey	(CERN, DARE, MAINZ)
BLIETSCHAU	78	NP B133 205	J. Blietschau <i>et al.</i>	(Gargamelle Collab.)
BOWMAN	78	PRL 41 442	J.D. Bowman <i>et al.</i>	(LASL, IAS, CMU+)
CAMANI	78	PL 77B 326	M. Camani <i>et al.</i>	(ETH, MAINZ)
BADERT...	77	PRL 39 1385	A. Badertscher <i>et al.</i>	(BERN)
CASPERSON	77	PRL 38 956	D.E. Casperson <i>et al.</i>	(BERN, HEIDH, LASL+)
DEPOMMIER	77	PRL 39 1113	P. Depommier <i>et al.</i>	(MONT, BRCO, TRIU+)
BALANDIN	74	JETP 40 811	M.P. Balandin <i>et al.</i>	(JINR)
		Translated from ZETF 67 1631.		
COHEN	73	JPCRD 2 664	E.R. Cohen, B.N. Taylor	(RISC, NBS)
DUCLOS	73	PL 47B 491	J. Duclos, A. Magnon, J. Picard	(SACL)
EICHTEN	73	PL 46B 281	T. Eichten <i>et al.</i>	(Gargamelle Collab.)
BRYMAN	72	PRL 28 1469	D.A. Bryman <i>et al.</i>	(VPI)
CROWE	72	PR D5 2145	K.M. Crowe <i>et al.</i>	(LBL, WASH)
CRANE	71	PRL 27 474	T. Crane <i>et al.</i>	(YALE)
DERENZO	69	PR 181 1854	S.E. Derenzo	(EFI)
VOSSLER	69	NC 63A 423	C. Vossler	(EFI)
AKHMANOV	68	SJNP 6 230	V.V. Akhmanov <i>et al.</i>	(KIAE)
		Translated from YAF 6 316.		
FRYBERGER	68	PR 166 1379	D. Fryberger	(EFI)
BOGART	67	PR 156 1405	E. Bogart <i>et al.</i>	(COLU)
SCHWARTZ	67	PR 162 1306	D.M. Schwartz	(EFI)
SHERWOOD	67	PR 156 1475	B.A. Sherwood	(EFI)
PEOPLES	66	Nevis 147 unpub.	J. Peoples	(COLU)
BLOOM	64	PL 8 87	S. Bloom <i>et al.</i>	(CERN)
DUCLOS	64	PL 9 62	J. Duclos <i>et al.</i>	(CERN)
GUREVICH	64	PL 11 185	I.I. Gurevich <i>et al.</i>	(KIAE)
BUHLER	63	PL 7 368	A. Buhler-Broglin <i>et al.</i>	(CERN)
MEYER	63	PR 132 2693	S.L. Meyer <i>et al.</i>	(COLU)
CHARPAK	62	PL 1 16	G. Charpak <i>et al.</i>	(CERN)
CONFORTO	62	NC 26 261	G. Conforto <i>et al.</i>	(INFN, ROMA, CERN)
ALI-ZADE	61	JETP 13 313	S.A. Ali-Zade, I.I. Gurevich, B.A. Nikolsky	
		Translated from ZETF 40 452.		

CRITTENDEN	61	PR 121 1823	R.R. Crittenden, W.D. Walker, J. Ballam	(WISC+)
KRUGER	61	UCRL 9322 unpub.	H. Kruger	(LRL)
GUREVICH	60	JETP 10 225	I.I. Gurevich, B.A. Nikolsky, L.V. Surkova	(ITEP)
		Translated from ZETF 37	318.	
PLANO	60	PR 119 1400	R.J. Plano	(COLU)
ASHKIN	59	NC 14 1266	J. Ashkin <i>et al.</i>	(CERN)
BARDON	59	PRL 2 56	M. Bardon, D. Berley, L.M. Lederman	(COLU)
LEE	59	PRL 3 55	J. Lee, N.P. Samios	(COLU)

---