

$D_{s1}(2536)^{\pm}$

$I(J^P) = 0(1^+)$
 J, P need confirmation.

Seen in $D^*(2010)^+ K^0$, $D^*(2007)^0 K^+$, and $D_s^+ \pi^+ \pi^-$. Not seen
in $D^+ K^0$ or $D^0 K^+$. $J^P = 1^+$ assignment strongly favored.

 $D_{s1}(2536)^{\pm}$ MASS

The fit includes D^{\pm} , D^0 , D_s^{\pm} , $D^{*\pm}$, D^{*0} , $D_s^{*\pm}$, $D_1(2420)^0$, $D_2^{*(2460)}{}^0$,
and $D_{s1}(2536)^{\pm}$ mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
2535.11±0.06 OUR FIT				
2535.21±0.28 OUR AVERAGE				
2537.7 ± 0.5 ± 3.1	24	¹ ABLIKIM	19P BES3	$4.6 e^+ e^- \rightarrow D_s^+ \bar{D}^0 K^-$
2535.7 ± 0.6 ± 0.5	46	² ABAZOV	09G D0	$B_s^0 \rightarrow D_{s1}^- \mu^+ \nu_\mu X$
2534.78±0.31±0.40	182	AUBERT	08B BABR	$B \rightarrow \bar{D}^{(*)} D^* K$
2534.6 ± 0.3 ± 0.7	193	AUBERT	06P BABR	$10.6 e^+ e^- \rightarrow D_s^+ \pi^+ \pi^- X$
2535.3 ± 0.7	92	³ HEISTER	02B ALEP	$e^+ e^- \rightarrow D^{*+} K^0 X, D^{*0} K^+ X$
2534.2 ± 1.2	9	ASRATYAN	94 BEBC	$\nu N \rightarrow D^* K^0 X, D^{*0} K^\pm X$
2535 ± 0.6 ± 1	75	FRABETTI	94B E687	$\gamma Be \rightarrow D^{*+} K^0 X, D^{*0} K^+ X$
2535.2 ± 0.5 ± 1.5	28	ALBRECHT	92R ARG	$10.4 e^+ e^- \rightarrow D^{*0} K^+ X$
2536.6 ± 0.7 ± 0.4		AVERY	90 CLEO	$e^+ e^- \rightarrow D^{*+} K^0 X$
2535.9 ± 0.6 ± 2.0		ALBRECHT	89E ARG	$D_{s1}^* \rightarrow D^*(2010) K^0$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
2534.1 ± 0.6	116	⁴ AUSHEV	11 BELL	$B \rightarrow D_{s1}(2536)^+ D^{(*)}$
2535.08±0.01±0.15	8038	⁵ LEES	11B BABR	$10.6 e^+ e^- \rightarrow D^{*+} K_S^0 X$
2535.57 ^{+0.44} _{-0.41} ± 0.10	236	⁶ CHEKANOV	09 ZEUS	$e^\pm p \rightarrow D^{*+} K_S^0 X, D^{*0} K^+ X$
2535.3 ± 0.2 ± 0.5	134	⁷ ALEXANDER	93 CLE2	$e^+ e^- \rightarrow D^{*0} K^+ X$
2534.8 ± 0.6 ± 0.6	44	⁸ ALEXANDER	93 CLE2	$e^+ e^- \rightarrow D^{*+} K^0 X$
2535 ± 28		⁹ ASRATYAN	88 HLBC	$\nu N \rightarrow D_s \gamma\gamma X$

¹ From a fit of the D_s^+ recoil mass distribution with an incoherent sum of the S -wave and D -wave Breit-Wigner line shapes.

² Using the $D^*(2010)^{\pm}$ mass of 2010.0 ± 0.4 MeV from PDG 06.

³ Calculated using $m(D^*(2010)^{\pm}) = 2010.0 \pm 0.5$ MeV, $m(D^*(2007)^0) = 2006.7 \pm 0.5$ MeV, and the mass difference below.

⁴ Systematic uncertainties not evaluated.

⁵ Calculated using the mass difference $m(D_{s1}^+) - m(D^{*+})_{PDG}$ below and $m(D^{*+})_{PDG} = 2010.25 \pm 0.14$ MeV. Assuming S -wave decay of the $D_{s1}(2536)$ to $D^{*+} K_S^0$, using a Breit-Wigner line shape corresponding to $L=0$.

- ⁶ Calculated using the mass difference $m(D_{s1}^+) - m(D^{*+})_{PDG}$ reported below and $m(D^{*+})_{PDG} = 2010.27 \pm 0.17$ MeV.
⁷ Calculated using $m(D^*(2007)^0) = 2006.6 \pm 0.5$ MeV and the mass difference below.
⁸ Calculated using $m(D^*(2010)^{\pm}) = 2010.1 \pm 0.6$ MeV and the mass difference below.
⁹ Not seen in $D^* K$.

$m_{D_{s1}(2536)^{\pm}} - m_{D_s^*(2111)}$

The fit includes D^{\pm} , D^0 , D_s^{\pm} , $D^{*\pm}$, D^{*0} , $D_s^{*\pm}$, $D_1(2420)^0$, $D_2^*(2460)^0$, and $D_{s1}(2536)^{\pm}$ mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
422.9 \pm 0.4 OUR FIT			
424 \pm 28	ASRATYAN	88	HLBC $D_s^{*\pm} \gamma$

$m_{D_{s1}(2536)^{\pm}} - m_{D^*(2010)^{\pm}}$

The fit includes D^{\pm} , D^0 , D_s^{\pm} , $D^{*\pm}$, D^{*0} , $D_s^{*\pm}$, $D_1(2420)^0$, $D_2^*(2460)^0$, and $D_{s1}(2536)^{\pm}$ mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
524.85 \pm 0.04 OUR FIT				
524.84 \pm 0.04 OUR AVERAGE				
524.83 \pm 0.01 \pm 0.04	8038	¹⁰ LEES	11B BABR	$10.6 e^+ e^- \rightarrow D^{*+} K_S^0 X$
$525.30^{+0.44}_{-0.41} \pm 0.10$	236 \pm 30	CHEKANOV 09	ZEUS	$e^{\pm} p \rightarrow D^{*+} K_S^0 X$, $D^{*0} K^+ X$
525.3 \pm 0.6 \pm 0.1	41	HEISTER 02B	ALEP	$e^+ e^- \rightarrow D^{*+} K^0 X$
524.7 \pm 0.6 \pm 0.2	44	ALEXANDER 93	CLE2	$e^+ e^- \rightarrow D^{*+} K_S^0 X$
10 Assuming S-wave decay of the $D_{s1}(2536)$ to $D^{*+} K_S^0$, using a Breit-Wigner line shape corresponding to L=0.				

$m_{D_{s1}(2536)^{\pm}} - m_{D^*(2007)^0}$

The fit includes D^{\pm} , D^0 , D_s^{\pm} , $D^{*\pm}$, D^{*0} , $D_s^{*\pm}$, $D_1(2420)^0$, $D_2^*(2460)^0$, and $D_{s1}(2536)^{\pm}$ mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
528.26 \pm 0.05 OUR FIT	Error includes scale factor of 1.1.			
528.68 \pm 0.28 OUR AVERAGE				
528.7 \pm 1.9 \pm 0.5	51	HEISTER 02B	ALEP	$e^+ e^- \rightarrow D^{*0} K^+ X$
527.3 \pm 2.2	29	ACKERSTAFF 97W	OPAL	$e^+ e^- \rightarrow D^{*0} K^+ X$
528.7 \pm 0.2 \pm 0.2	134	ALEXANDER 93	CLE2	$e^+ e^- \rightarrow D^{*0} K^+ X$

$D_{s1}(2536)^{\pm}$ WIDTH

VALUE (MeV)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
0.92 \pm 0.05 OUR AVERAGE					
1.7 \pm 1.2 \pm 0.6	24	¹¹ ABLIKIM	19P BES3	$4.6 e^+ e^- \rightarrow D_s^+ \bar{D}^0 K^-$	
0.92 \pm 0.03 \pm 0.04	8038	¹² LEES	11B BABR	$10.6 e^+ e^- \rightarrow D^{*+} K_S^0 X$	

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.75 ± 0.23	116	¹³ AUSHEV	11	BELL	$B \rightarrow D_{s1}(2536)^+ D^(*)$
< 2.5	95	193	AUBERT	06P	$BABR$
< 3.2	90	75	FRABETTI	94B	E687
< 2.3	90		ALEXANDER	93	CLEO
< 3.9	90		ALBRECHT	92R	ARG
< 5.44	90		AVERY	90	CLEO
< 4.6	90		ALBRECHT	89E	ARG

¹¹ From a fit of the D_s^+ recoil mass distribution with an incoherent sum of the S -wave and S -wave Breit-Wigner line shapes.

¹² Assuming S -wave decay of the $D_{s1}(2536)$ to $D^{*+} K_S^0$, using a Breit-Wigner line shape corresponding to $L=0$.

¹³ Systematic uncertainties not evaluated.

$D_{s1}(2536)^+$ DECAY MODES

Branching fractions are given relative to the one **DEFINED AS 1**.

$D_{s1}(2536)^-$ modes are charge conjugates of the modes below.

Mode	Fraction (Γ_i/Γ)	Confidence level
$\Gamma_1 D^*(2010)^+ K^0$	(31 ± 7) %	
$\Gamma_2 (D^*(2010)^+ K^0)_{S-wave}$	(22 ± 5) %	
$\Gamma_3 (D^*(2010)^+ K^0)_{D-wave}$		
$\Gamma_4 K_S^0 D^*(2010)^+$	(17 ± 4) %	
$\Gamma_5 D^+ \pi^- K^+$	$(10.0 \pm 2.5) \times 10^{-3}$	
$\Gamma_6 D^*(2007)^0 K^+$	(36 ± 6) %	
$\Gamma_7 D^+ K^0$	< 12 %	90%
$\Gamma_8 D^0 K^+$	< 4 %	90%
$\Gamma_9 D_s^{*+} \gamma$	possibly seen	
$\Gamma_{10} D_s^+ \pi^+ \pi^-$	seen	

$D_{s1}(2536)^+$ BRANCHING RATIOS

$$\Gamma((D^*(2010)^+ K^0)_{S-wave})/\Gamma(D^*(2010)^+ K^0) \quad \Gamma_2/\Gamma_1$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.72 ± 0.05 ± 0.01	5485	BALAGURA	08	BELL $e^+ e^- \rightarrow D^{*+} K^0 X$

$$\Gamma(K_S^0 D^*(2010)^+)/\Gamma(D^*(2007)^0 K^+) \quad \Gamma_4/\Gamma_6$$

VALUE	DOCUMENT ID	TECN	COMMENT
0.48 ± 0.07 ± 0.02	GAO	23	BELL $e^+ e^-$ at 10.52 GeV

$$\Gamma(D^+ \pi^- K^+)/\Gamma(D^*(2010)^+ K^0) \quad \Gamma_5/\Gamma_1$$

VALUE (units 10^{-2})	EVTS	DOCUMENT ID	TECN	COMMENT
3.27 ± 0.18 ± 0.37	1264	BALAGURA	08	BELL $e^+ e^- \rightarrow D^+ \pi^- K^+ X$

$\Gamma(D^*(2007)^0 K^+)/\Gamma_{\text{total}}$	Γ_6/Γ
<u>VALUE (%)</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
$35.9 \pm 4.8 \pm 3.5$	14 ABLIKIM 24BN BES3 $e^+ e^- \rightarrow D_s^+ D_{s1}(2536)^-$

¹⁴ Determined as ratio of exclusive $e^+ e^- \rightarrow D_s^+ [\bar{D}^{*0} K^-]$ and inclusive $e^+ e^- \rightarrow D_s^+ X$ measurements.

$\Gamma(D^*(2007)^0 K^+)/\Gamma(D^*(2010)^+ K^0)$	Γ_6/Γ_1
<u>VALUE</u>	<u>EVTS</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
1.18 ± 0.16 OUR AVERAGE	
$0.88 \pm 0.24 \pm 0.08$	116 AUSHEV 11 BELL $B \rightarrow D_{s1}(2536)^+ D^{(*)}$
$2.3 \pm 0.6 \pm 0.3$	236 \pm 30 CHEKANOV 09 ZEUS $e^\pm p \rightarrow D^{*+} K_S^0 X,$
$1.32 \pm 0.47 \pm 0.23$	92 15 HEISTER 02B ALEP $e^+ e^- \rightarrow D^{*+} K^0 X,$
$1.9 \begin{array}{l} +1.1 \\ -0.9 \end{array} \pm 0.4$	35 15 ACKERSTAFF 97W OPAL $D^{*0} K^+ X$
1.1 ± 0.3	ALEXANDER 93 CLEO $e^+ e^- \rightarrow D^{*0} K^0 X$
$1.4 \pm 0.3 \pm 0.2$	16 ALBRECHT 92R ARG $10.4 e^+ e^- \rightarrow D^{*0} K^+ X, D^{*+} K^0 X$

¹⁵ Ratio of the production rates measured in Z^0 decays.

¹⁶ Evaluated by us from published inclusive cross-sections.

$\Gamma(D^+ K^0)/\Gamma(D^*(2010)^+ K^0)$	Γ_7/Γ_1
<u>VALUE</u>	<u>CL%</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
<0.40	90 ALEXANDER 93 CLEO $e^+ e^- \rightarrow D^{*+} K^0 X$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$	
<0.43	90 ALBRECHT 89E ARG $D_{s1}^* \rightarrow D^*(2010) K^0$

$\Gamma(D^0 K^+)/\Gamma(D^*(2007)^0 K^+)$	Γ_8/Γ_6
<u>VALUE</u>	<u>CL%</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
<0.12	90 ALEXANDER 93 CLEO $e^+ e^- \rightarrow D^{*0} K^+ X$

$\Gamma(D_s^{*+} \gamma)/\Gamma_{\text{total}}$	Γ_9/Γ
<u>VALUE</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
possibly seen	ASRATYAN 88 HLBC $\nu N \rightarrow D_s \gamma\gamma X$

$\Gamma(D_s^{*+} \gamma)/\Gamma(D^*(2007)^0 K^+)$	Γ_9/Γ_6
<u>VALUE</u>	<u>CL%</u> <u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
<0.42	90 ALEXANDER 93 CLEO $e^+ e^- \rightarrow D^{*0} K^+ X$

$\Gamma(D_s^+ \pi^+ \pi^-)/\Gamma_{\text{total}}$	Γ_{10}/Γ
<u>VALUE</u>	<u>DOCUMENT ID</u> <u>TECN</u> <u>COMMENT</u>
seen	AUBERT 06P BABR $10.6 e^+ e^- \rightarrow D_s^+ \pi^+ \pi^- X$

$D_{s1}(2536)^{\pm}$ REFERENCES

ABLIKIM	24BN	PRL 133 171903	M. Ablikim <i>et al.</i>	(BESIII Collab.)
GAO	23	PR D108 112015	B.S. Gao <i>et al.</i>	(BELLE Collab.)
ABLIKIM	19P	CP C43 031001	M. Ablikim <i>et al.</i>	(BESIII Collab.)
AUSHEV	11	PR D83 051102	T. Aushev <i>et al.</i>	(BELLE Collab.)
LEES	11B	PR D83 072003	J.P. Lees <i>et al.</i>	(BABAR Collab.)
ABAZOV	09G	PRL 102 051801	V.M. Abazov <i>et al.</i>	(D0 Collab.)
CHEKANOV	09	EPJ C60 25	S. Chekanov <i>et al.</i>	(ZEUS Collab.)
AUBERT	08B	PR D77 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)
BALAGURA	08	PR D77 032001	V. Balagura <i>et al.</i>	(BELLE Collab.)
AUBERT	06P	PR D74 032007	B. Aubert <i>et al.</i>	(BABAR Collab.)
PDG	06	JP G33 1	W.-M. Yao <i>et al.</i>	(PDG Collab.)
HEISTER	02B	PL B526 34	A. Heister <i>et al.</i>	(ALEPH Collab.)
ACKERSTAFF	97W	ZPHY C76 425	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ASRATYAN	94	ZPHY C61 563	A.E. Asratyan <i>et al.</i>	(BIRM, BELG, CERN+) (FNAL E687 Collab.)
FRAZETTI	94B	PRL 72 324	P.L. Frabetti <i>et al.</i>	(CLEO Collab.)
ALEXANDER	93	PL B303 377	J. Alexander <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92R	PL B297 425	H. Albrecht <i>et al.</i>	(CLEO Collab.)
EVERY	90	PR D41 774	P. Avery, D. Besson	(ARGUS Collab.)
ALBRECHT	89E	PL B230 162	H. Albrecht <i>et al.</i>	(ITEP, SERP)
ASRATYAN	88	ZPHY C40 483	A.E. Asratyan <i>et al.</i>	
