

**$\rho(770)$**  $I^G(J^{PC}) = 1^+(1^{--})$  **$\rho(770)$  T-MATRIX POLE  $\sqrt{s}$** Note that  $\Gamma = -2 \operatorname{Im}(\sqrt{s})$ .

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>(761–765) – <math>i</math> (71–74) OUR ESTIMATE</b>			
$(762.5 \pm 1.7) - i (73.2 \pm 1.1)$	<sup>1</sup> HOFERICHT... 24	RVUE	Compilation
$(763.7^{+1.7}_{-1.5}) - i (73.2^{+1.0}_{-1.1})$	<sup>2</sup> GARCIA-MAR... 11	RVUE	Compilation
$(754 \pm 18) - i (74 \pm 10)$	<sup>3</sup> PELAEZ 04A	RVUE	$\pi\pi \rightarrow \pi\pi$
$(762.4 \pm 1.8) - i (72.6 \pm 1.4)$	COLANGELO 01	RVUE	$\pi\pi \rightarrow \pi\pi$
<sup>1</sup> Update of GARCIA-MARTIN 11 with input for $\pi\pi$ T-matrix from HOFERICHTER 16.			
<sup>2</sup> Reanalysis of the $K_{e4}$ data of BATLEY 10C and the $\pi N \rightarrow \pi\pi N$ data of HYAMS 73, GRAYER 74, and PROTOPOPESCU 73 using GKPY equations.			
<sup>3</sup> Reanalysis of data from PROTOPOPESCU 73, ESTABROOKS 74, GRAYER 74, and COHEN 80 in the unitarized ChPT model.			

 **$\rho(770)$  MASS**We no longer list  $S$ -wave Breit-Wigner fits, or data with high combinatorial background.**NEUTRAL ONLY,  $e^+ e^-$** 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>775.26±0.23 OUR AVERAGE</b>				
775.3 $\pm 0.5$ $\pm 0.6$		<sup>1</sup> ACHASOV 21	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
775.02 $\pm 0.35$		<sup>2</sup> LEES 12G	BABR	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$
775.97 $\pm 0.46 \pm 0.70$	900k	<sup>3</sup> AKHMETSHIN 07		$e^+ e^- \rightarrow \pi^+ \pi^-$
774.6 $\pm 0.4$ $\pm 0.5$	800k	<sup>4,5</sup> ACHASOV 06	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
775.65 $\pm 0.64 \pm 0.50$	114k	<sup>6,7</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^-$
775.9 $\pm 0.5$ $\pm 0.5$	1.98M	<sup>8</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.8 $\pm 0.9$ $\pm 2.0$	500k	<sup>8</sup> ACHASOV 02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.9 $\pm 1.1$		<sup>9</sup> BARKOV 85	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^-$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
775.4 $\pm 0.1$	34M	<sup>10</sup> IGNATOV 24	CMD3	$e^+ e^- \rightarrow \pi^+ \pi^-$
763.49 $\pm 0.53$		<sup>11</sup> BARTOS 17	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
758.23 $\pm 0.46$		<sup>12</sup> BARTOS 17A	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
775.8 $\pm 0.5$ $\pm 0.3$	1.98M	<sup>13</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.9 $\pm 0.6$ $\pm 0.5$	1.98M	<sup>14</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.0 $\pm 0.6$ $\pm 1.1$	500k	<sup>15</sup> ACHASOV 02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.1 $\pm 0.7$ $\pm 5.3$		<sup>16</sup> BENAYOUN 98	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$ , $\mu^+ \mu^-$
770.5 $\pm 1.9$ $\pm 5.1$		<sup>17</sup> GARDNER 98	RVUE	$0.28-0.92 e^+ e^- \rightarrow \pi^+ \pi^-$
764.1 $\pm 0.7$		<sup>18</sup> O'CONNELL 97	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
757.5 $\pm 1.5$		<sup>19</sup> BERNICHA 94	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
768 $\pm 1$		<sup>20</sup> GESHKEN...	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$

- <sup>1</sup> From a fit of the cross section in the energy range  $0.525 < \sqrt{s} < 0.883$  GeV parameterized by the sum of the Breit-Wigner amplitudes for the  $\rho(770)$ ,  $\omega$  and  $\rho(1450)$  resonances.
- <sup>2</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho-\omega$  interference and leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.
- <sup>3</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.
- <sup>4</sup> Supersedes ACHASOV 05A.
- <sup>5</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.
- <sup>6</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho-\omega$  interference.
- <sup>7</sup> Update of AKHMETSHIN 02.
- <sup>8</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .
- <sup>9</sup> From the GOUNARIS 68 parametrization of the pion form factor.
- <sup>10</sup> From a fit of the pion form factor in the energy range  $0.32 < \sqrt{s} < 1.2$  GeV using the GOUNARIS 68 parametrization with the complex phase of the  $\rho-\omega$  interference with  $\omega$  and  $\phi$  masses and widths constrained by the values and their errors from PDG 22, and leaving  $\rho(1450)$ ,  $\rho(1700)$  resonances as free parameters of the fit. Systematic errors not estimated.
- <sup>11</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of LEES 12G and ABLIKIM 16C.
- <sup>12</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, and AMBROSINO 11A.
- <sup>13</sup> Assuming  $m_{\rho^+} = m_{\rho^-} = m_{\rho^0}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-} = \Gamma_{\rho^0}$ .
- <sup>14</sup> Without limitations on masses and widths.
- <sup>15</sup> Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0 \pi\pi} = g_{\rho^\pm \pi\pi}$ .
- <sup>16</sup> Using the data of BARKOV 85 in the hidden local symmetry model.
- <sup>17</sup> From the fit to  $e^+ e^- \rightarrow \pi^+ \pi^-$  data from the compilations of HEYN 81 and BARKOV 85, including the GOUNARIS 68 parametrization of the pion form factor.
- <sup>18</sup> A fit of BARKOV 85 data assuming the direct  $\omega \pi\pi$  coupling.
- <sup>19</sup> Applying the S-matrix formalism to the BARKOV 85 data.
- <sup>20</sup> Includes BARKOV 85 data. Model-dependent width definition.

## CHARGED ONLY, $\tau$ DECAYS and $e^+ e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>775.11±0.34 OUR AVERAGE</b>					
774.6 $\pm 0.2$ $\pm 0.5$	5.4M	1,2 FUJIKAWA	08	BELL	$\pm$ $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
775.5 $\pm 0.7$		2,3 SCHAEL	05C	ALEP	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
775.5 $\pm 0.5$ $\pm 0.4$	1.98M	4 ALOISIO	03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.1 $\pm 1.1$ $\pm 0.5$	87k	5,6 ANDERSON	00A	CLE2	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
761.60 $\pm 0.95$		7 BARTOS	17A	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
774.8 $\pm 0.6$ $\pm 0.4$	1.98M	8 ALOISIO	03	KLOE	$- 1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
776.3 $\pm 0.6$ $\pm 0.7$	1.98M	8 ALOISIO	03	KLOE	$+ 1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
773.9 $\pm 2.0$ $+ 0.3$ $- 1.0$		9 SANZ-CILLERO03	RVUE		$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
774.5 $\pm 0.7$ $\pm 1.5$	500k	4 ACHASOV	02	SND	$\pm 1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.1 $\pm 0.5$		10 PICH	01	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$

<sup>1</sup>  $|F_\pi(0)|^2$  fixed to 1.<sup>2</sup> From the GOUNARIS 68 parametrization of the pion form factor.<sup>3</sup> The error combines statistical and systematic uncertainties. Supersedes BARATE 97M.<sup>4</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .<sup>5</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.<sup>6</sup> From the GOUNARIS 68 parametrization of the pion form factor. The second error is a model error taking into account different parametrizations of the pion form factor.<sup>7</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUB-NICKA 10 to analyze the data of FUJIKAWA 08.<sup>8</sup> Without limitations on masses and widths.<sup>9</sup> Using the data of BARATE 97M and the effective chiral Lagrangian.<sup>10</sup> From a fit of the model-independent parameterization of the pion form factor to the data of BARATE 97M.

## MIXED CHARGES, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>763.0±0.3±1.2</b>	600k	<sup>1</sup> ABELE	99E	CBAR	0± 0.0 $\bar{p}p \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> Assuming the equality of  $\rho^+$  and  $\rho^-$  masses and widths.

## CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>766.5±1.1 OUR AVERAGE</b>					
763.7±3.2		ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^- \pi^0 \pi^0$
768 ± 9		AGUILAR...	91	EHS	400 $p\bar{p}$
767 ± 3	2935	<sup>1</sup> CAPRARO	87	SPEC	— 200 $\pi^- \text{Cu} \rightarrow \pi^- \pi^0 \text{Cu}$
761 ± 5	967	<sup>1</sup> CAPRARO	87	SPEC	— 200 $\pi^- \text{Pb} \rightarrow \pi^- \pi^0 \text{Pb}$
771 ± 4		HUSTON	86	SPEC	+ 202 $\pi^+ \text{A} \rightarrow \pi^+ \pi^0 \text{A}$
766 ± 7	6500	<sup>2</sup> BYERLY	73	OSPK	— 5 $\pi^- p$
766.8±1.5	9650	<sup>3</sup> PISUT	68	RVUE	— 1.7–3.2 $\pi^- p$ , $t < 10$
767 ± 6	900	<sup>1</sup> EISNER	67	HBC	— 4.2 $\pi^- p$ , $t < 10$

<sup>1</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.<sup>2</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.<sup>3</sup> From fit of 3-parameter relativistic  $P$ -wave Breit-Wigner to total mass distribution. Includes BATON 68, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65 and CARMONY 64.

## NEUTRAL ONLY, PHOTOPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>769.2± 0.9 OUR AVERAGE</b>				
770.8± 1.3 <sup>+2.3</sup> <sub>-2.4</sub>	900k	ANDREEV	20	H1 $e p \rightarrow e \pi^+ \pi^- p$
771 ± 2 <sup>+2</sup> <sub>-1</sub>	63.5k	<sup>1</sup> ABRAMOWICZ12	ZEUS	$e p \rightarrow e \pi^+ \pi^- p$
770 ± 2 ± 1	79k	<sup>2</sup> BREITWEG	98B	ZEUS 50–100 $\gamma p$
767.6± 2.7		BARTALUCCI	78	CNTR $\gamma p \rightarrow e^+ e^- p$
775 ± 5		GLADDING	73	CNTR 2.9–4.7 $\gamma p$
767 ± 4	1930	BALLAM	72	HBC 2.8 $\gamma p$
770 ± 4	2430	BALLAM	72	HBC 4.7 $\gamma p$
765 ± 10		ALVENSLEB...	70	CNTR $\gamma A$ , $t < 0.01$
767.7± 1.9	140k	BIGGS	70	CNTR $< 4.1 \gamma C \rightarrow \pi^+ \pi^- C$
765 ± 5	4000	ASBURY	67B	CNTR $\gamma + Pb$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
771 ± 2	79k	<sup>3</sup> BREITWEG	98B	ZEUS 50–100 $\gamma p$

<sup>1</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho-\omega$  interference.

<sup>2</sup> From the parametrization according to SOEDING 66.

<sup>3</sup> From the parametrization according to ROSS 66.

## NEUTRAL ONLY, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>769.0 ±0.9 OUR AVERAGE</b>		Error includes scale factor of 1.4. See the ideogram below.		
765 ±6		BERTIN 97C	OBLX	$0.0 \bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
773 ±1.6		WEIDENAUER 93	ASTE	$\bar{p}p \rightarrow \pi^+ \pi^- \omega$
762.6 ±2.6		AGUILAR-...	91	EHS 400 $p\bar{p}$
770 ±2		<sup>1</sup> HEYN 81	RVUE	Pion form factor
768 ±4		<sup>2,3</sup> BOHACIK 80	RVUE	
769 ±3		<sup>4</sup> WICKLUND 78	ASPK	$3,4,6 \pi^\pm N$
768 ±1	76k	DEUTSCH...	HBC	$16 \pi^+ p$
767 ±4	4100	ENGLER 74	DBC	$6 \pi^+ n \rightarrow \pi^+ \pi^- p$
775 ±4	32k	<sup>2</sup> PROTOPOP... 73	HBC	$7.1 \pi^+ p, t < 0.4$
764 ±3	6.8k	<sup>5</sup> RATCLIFF 72	ASPK	$15 \pi^- p, t < 0.3$
774 ±3	1.7k	REYNOLDS 69	HBC	$2.26 \pi^- p$
769.2 ±1.5	13.3k	<sup>6</sup> PISUT 68	RVUE	$1.7-3.2 \pi^- p, t < 10$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
774.34 ±0.18 ±0.35	970k	<sup>7</sup> ABLIKIM 18C	BES3	$\eta'(958) \rightarrow \gamma \pi^+ \pi^-$
772.93 ±0.18 ±0.34	970k	<sup>8</sup> ABLIKIM 18C	BES3	$\eta'(958) \rightarrow \gamma \pi^+ \pi^-$
773.5 ±2.5		<sup>9</sup> COLANGELO 01	RVUE	$\pi \pi \rightarrow \pi \pi$
762.3 ±0.5 ±1.2	600k	<sup>10</sup> ABELE 99E	CBAR	$0.0 \bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
777 ±2	4.9k	<sup>11</sup> ADAMS 97	E665	$470 \mu p \rightarrow \mu XB$
770 ±2		<sup>12</sup> BOGOLYUB... 97	MIRA	$32 \bar{p}p \rightarrow \pi^+ \pi^- X$
768 ±8		<sup>12</sup> BOGOLYUB... 97	MIRA	$32 pp \rightarrow \pi^+ \pi^- X$
761.1 ±2.9		DUBNICKA 89	RVUE	$\pi$ form factor
777.4 ±2.0		<sup>13</sup> CHABAUD 83	ASPK	$17 \pi^- p$ polarized
769.5 ±0.7		<sup>2,3</sup> LANG 79	RVUE	
770 ±9		<sup>3</sup> ESTABROOKS 74	RVUE	$17 \pi^- p \rightarrow \pi^+ \pi^- n$
773.5 ±1.7	11.2k	<sup>14</sup> JACOBS 72	HBC	$2.8 \pi^- p$
775 ±3	2.2k	<sup>15</sup> HYAMS 68	OSPK	$11.2 \pi^- p$

<sup>1</sup> HEYN 81 includes all spacelike and timelike  $F_\pi$  values until 1978.

<sup>2</sup> From pole extrapolation.

<sup>3</sup> From phase shift analysis of GRAYER 74 data.

<sup>4</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.

<sup>5</sup> Published values contain misprints. Corrected by private communication RATCLIFF 74.

<sup>6</sup> Includes MALAMUD 69, ARMENISE 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, GOLDHABER 64, ABOLINS 63.

<sup>7</sup> From a fit to  $\pi^+ \pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and box anomaly components.

<sup>8</sup> From a fit to  $\pi^+ \pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and  $\rho(1450)$  components.

<sup>9</sup> Breit-Wigner mass from a phase-shift analysis of HYAMS 73 and PROTOPOPOV 73 data.

<sup>10</sup> Using relativistic Breit-Wigner and taking into account  $\rho-\omega$  interference.

<sup>11</sup> Systematic errors not evaluated.

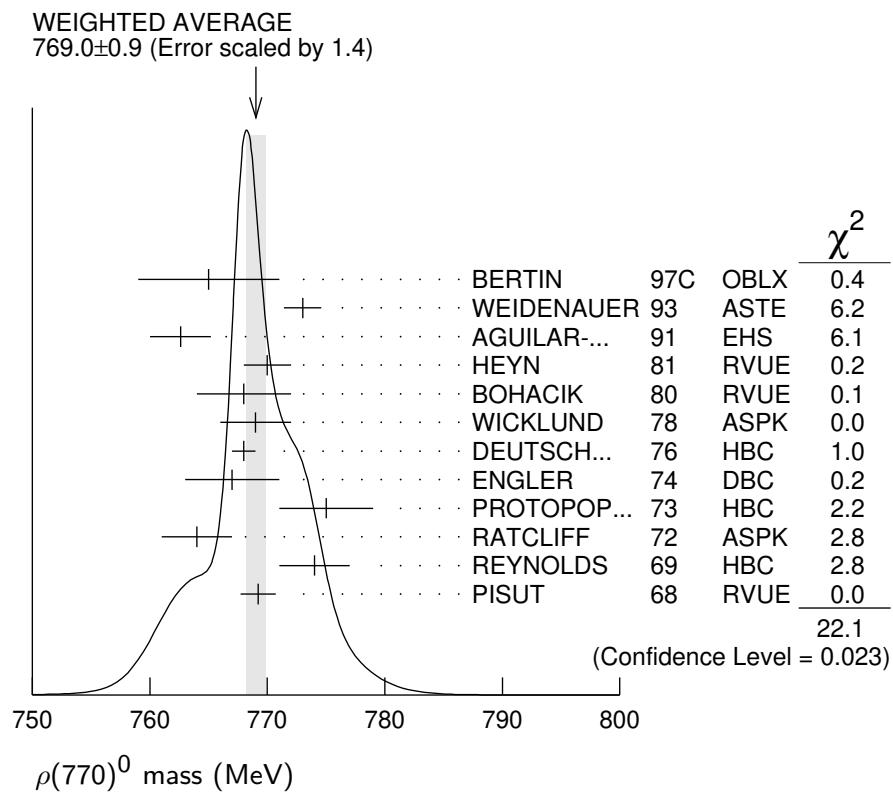
<sup>12</sup> Systematic effects not studied.

<sup>13</sup> From fit of 3-parameter relativistic Breit-Wigner to helicity-zero part of P-wave intensity.

CHABAUD 83 includes data of GRAYER 74.

<sup>14</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>15</sup> Of HYAMS 68 six parametrizations, this is theoretically soundest. MR



### $m_{\rho(770)^0} - m_{\rho(770)^{\pm}}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.7 ±0.8 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.			
-2.4 ±0.8	1 SCHAEL	05C	ALEP		$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
0.4 ±0.7 ±0.6	1.98M 2 ALOISIO	03	KLOE	1.02	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1.3 ±1.1 ±2.0	500k 2 ACHASOV	02	SND	1.02	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1.6 ±0.6 ±1.7	600k ABELE	99E	CBAR	±0	0.0 $\bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
-4 ±4	3000 3 REYNOLDS	69	HBC	-0	2.26 $\pi^- p$
-5 ±5	3600 3 FOSTER	68	HBC	±0	0.0 $\bar{p}p$
2.4 ±2.1	22950 4 PISUT	68	RVUE		$\pi N \rightarrow \rho N$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
-3.37±1.06	5 BARTOS	17A	RVUE		$e^+ e^- \rightarrow \pi^+ \pi^-$ , $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$

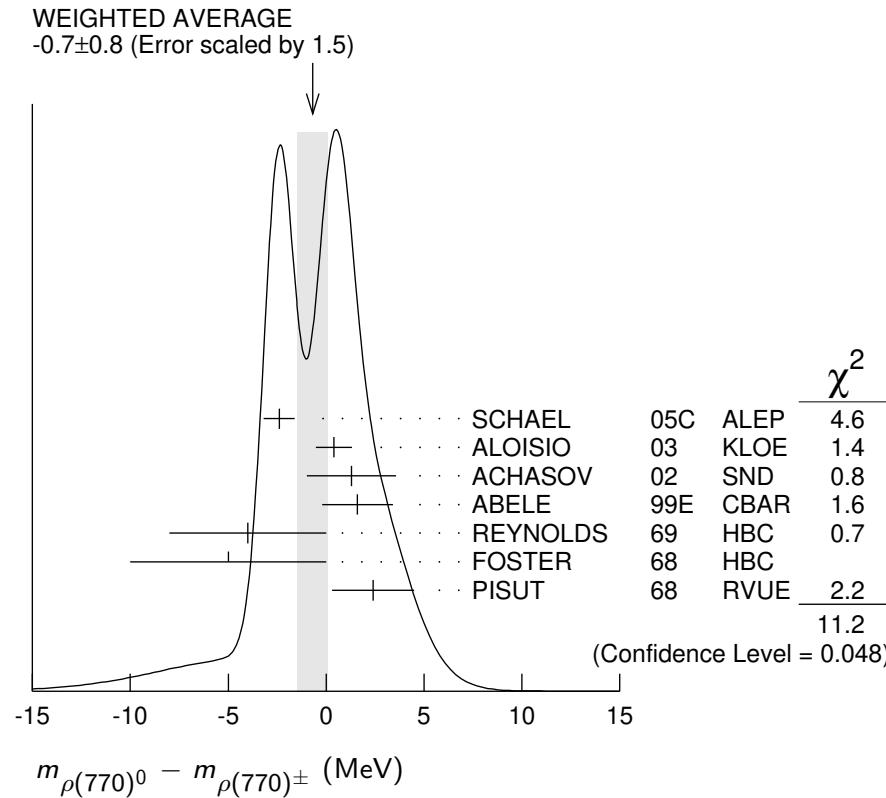
<sup>1</sup> From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEL 05C and  $e^+ e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05. Supersedes BARATE 97M.

<sup>2</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>3</sup> From quoted masses of charged and neutral modes.

<sup>4</sup> Includes MALAMUD 69, ARMENISE 68, BATON 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65, CARMONY 64, GOLDHABER 64, ABOLINS 63.

<sup>5</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, AMBROSINO 11A, and FUJIKAWA 08.



### $m_{\rho(770)^+} - m_{\rho(770)^-}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.5 \pm 0.8 \pm 0.7$       1.98M      <sup>1</sup> ALOISIO      03      KLOE       $1.02 \text{ e}^+ \text{e}^- \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> Without limitations on masses and widths.

### $\rho(770)$ RANGE PARAMETER

The range parameter  $R$  enters an energy-dependent correction to the width, of the form  $(1 + q_r^2 R^2) / (1 + q^2 R^2)$ , where  $q$  is the momentum of one of the pions in the  $\pi\pi$  rest system. At resonance,  $q = q_r$ .

VALUE (GeV $^{-1}$ )	DOCUMENT ID	TECN	CHG	COMMENT
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**5.3 $^{+0.9}_{-0.7}$**       <sup>1</sup> CHABAUD      83      ASPK      0       $17 \pi^- p$  polarized

<sup>1</sup> The old PISUT 68 value, properly corrected, was  $3.2 \pm 0.6$ .

## $\rho(770)$ WIDTH

We no longer list  $S$ -wave Breit-Wigner fits, or data with high combinatorial background.

### NEUTRAL ONLY, $e^+ e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>147.4 <math>\pm 0.8</math> OUR AVERAGE</b>		Error includes scale factor of 2.0.		See the ideogram below.
145.6 $\pm 0.6$ $\pm 0.8$		<sup>1</sup> ACHASOV 21	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
149.59 $\pm 0.67$		<sup>2</sup> LEES 12G	BABR	$e^+ e^- \rightarrow \pi^+ \pi^- \gamma$
145.98 $\pm 0.75$ $\pm 0.50$	900k	<sup>3</sup> AKHMETSHIN 07		$e^+ e^- \rightarrow \pi^+ \pi^-$
146.1 $\pm 0.8$ $\pm 1.5$	800k	<sup>4,5</sup> ACHASOV 06	SND	$e^+ e^- \rightarrow \pi^+ \pi^-$
143.85 $\pm 1.33$ $\pm 0.80$	114k	<sup>6,7</sup> AKHMETSHIN 04	CMD2	$e^+ e^- \rightarrow \pi^+ \pi^-$
147.3 $\pm 1.5$ $\pm 0.7$	1.98M	<sup>8</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
151.1 $\pm 2.6$ $\pm 3.0$	500k	<sup>8</sup> ACHASOV 02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
150.5 $\pm 3.0$		<sup>9</sup> BARKOV 85	OLYA	$e^+ e^- \rightarrow \pi^+ \pi^-$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
148.76 $\pm 0.17$	34M	<sup>10</sup> IGNATOV 24	CMD3	$e^+ e^- \rightarrow \pi^+ \pi^-$
144.06 $\pm 0.85$		<sup>11</sup> BARTOS 17	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
144.56 $\pm 0.80$		<sup>12</sup> BARTOS 17A	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
143.9 $\pm 1.3$ $\pm 1.1$	1.98M	<sup>13</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
147.4 $\pm 1.5$ $\pm 0.7$	1.98M	<sup>14</sup> ALOISIO 03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
149.8 $\pm 2.2$ $\pm 2.0$	500k	<sup>15</sup> ACHASOV 02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
147.9 $\pm 1.5$ $\pm 7.5$		<sup>16</sup> BENAYOUN 98	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$ , $\mu^+ \mu^-$
153.5 $\pm 1.3$ $\pm 4.6$		<sup>17</sup> GARDNER 98	RVUE	$0.28-0.92 e^+ e^- \rightarrow \pi^+ \pi^-$
145.0 $\pm 1.7$		<sup>18</sup> O'CONNELL 97	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
142.5 $\pm 3.5$		<sup>19</sup> BERNICHA 94	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$
138 $\pm 1$		<sup>20</sup> GESHKEN... 89	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$

<sup>1</sup> From a fit of the cross section in the energy range  $0.525 < \sqrt{s} < 0.883$  GeV parameterized by the sum of the Breit-Wigner amplitudes for the  $\rho(770)$ ,  $\omega$  and  $\rho(1450)$  resonances.

<sup>2</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho-\omega$  interference and leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>3</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.

<sup>4</sup> Supersedes ACHASOV 05A.

<sup>5</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.

<sup>6</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho-\omega$  interference.

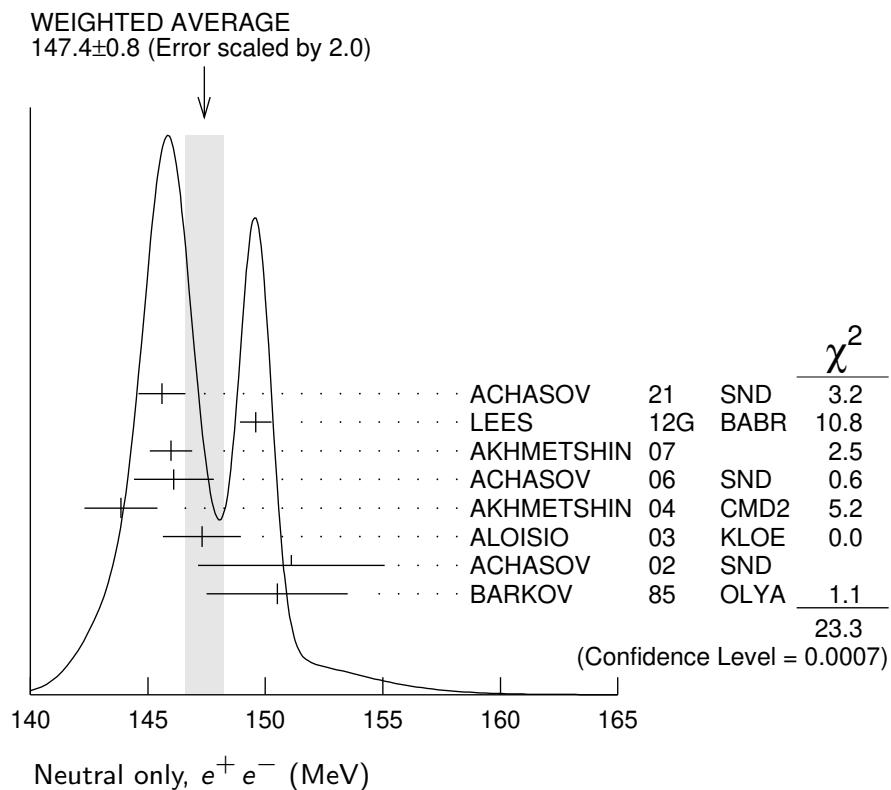
<sup>7</sup> From a fit in the energy range 0.61 to 0.96 GeV. Update of AKHMETSHIN 02.

<sup>8</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>9</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>10</sup> From a fit of the pion form factor in the energy range  $0.32 < \sqrt{s} < 1.2$  GeV using the GOUNARIS 68 parametrization with the complex phase of the  $\rho-\omega$  interference with  $\omega$  and  $\phi$  masses and widths constrained by the values and their errors from PDG 22, and leaving  $\rho(1450)$ ,  $\rho(1700)$  resonances as free parameters of the fit. Systematic errors not estimated.

- 11 Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUB-NICKA 10 to analyze the data of LEES 12G and ABLIKIM 16C.  
 12 Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUB-NICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, and AMBROSINO 11A.  
 13 Assuming  $m_{\rho^+} = m_{\rho^-} = m_{\rho^0}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-} = \Gamma_{\rho^0}$ .  
 14 Without limitations on masses and widths.  
 15 Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0 \pi\pi} = g_{\rho^\pm \pi\pi}$ .  
 16 Using the data of BARKOV 85 in the hidden local symmetry model.  
 17 From the fit to  $e^+ e^- \rightarrow \pi^+ \pi^-$  data from the compilations of HEYN 81 and BARKOV 85, including the GOUNARIS 68 parametrization of the pion form factor.  
 18 A fit of BARKOV 85 data assuming the direct  $\omega \pi\pi$  coupling.  
 19 Applying the S-matrix formalism to the BARKOV 85 data.  
 20 Includes BARKOV 85 data. Model-dependent width definition.



### CHARGED ONLY, $\tau$ DECAYS and $e^+ e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>149.1 ±0.8 OUR FIT</b>					
<b>149.1 ±0.8 OUR AVERAGE</b>					
148.1 ±0.4 ±1.7 5.4M	1,2 FUJIKAWA	08 BELL	±		$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
149.0 ±1.2	2,3 SCHABEL	05C ALEP			$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
149.9 ±2.3 ±2.0 500k	4 ACHASOV	02 SND	±		$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
150.4 ±1.4 ±1.4 87k	5,6 ANDERSON	00A CLE2			$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$139.90 \pm 0.46$	7 BARTOS	17A RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
$143.7 \pm 1.3 \pm 1.2$ 1.98M	4 ALOISIO	03 KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$142.9 \pm 1.3 \pm 1.4$ 1.98M	8 ALOISIO	03 KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$144.7 \pm 1.4 \pm 1.2$ 1.98M	8 ALOISIO	03 KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
$150.2 \pm 2.0 \pm 0.7$ $-1.6$	9 SANZ-CILLERO03	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
$150.9 \pm 2.2 \pm 2.0$ 500k	10 ACHASOV	02 SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup>  $|F_\pi(0)|^2$  fixed to 1.

<sup>2</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>3</sup> The error combines statistical and systematic uncertainties. Supersedes BARATE 97M.

<sup>4</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>5</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.

<sup>6</sup> From the GOUNARIS 68 parametrization of the pion form factor. The second error is a model error taking into account different parametrizations of the pion form factor.

<sup>7</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUB-NICKA 10 to analyze the data of FUJIKAWA 08.

<sup>8</sup> Without limitations on masses and widths.

<sup>9</sup> Using the data of BARATE 97M and the effective chiral Lagrangian.

<sup>10</sup> Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0 \pi \pi} = g_{\rho^\pm \pi \pi}$ .

## MIXED CHARGES, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>149.5 ± 1.3</b>	600k	<sup>1</sup> ABELE	99E	CBAR	$0 \pm 0.0 \bar{p}p \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> Assuming the equality of  $\rho^+$  and  $\rho^-$  masses and widths.

## CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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**150.2 ± 2.4 OUR FIT**

**150.2 ± 2.4 OUR AVERAGE**

$152.8 \pm 4.3$		ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^- \pi^0 \pi^0$
$155 \pm 11$	2.9k	<sup>1</sup> CAPRARO	87	SPEC	$200 \pi^- Cu \rightarrow \pi^- \pi^0 Cu$
$154 \pm 20$	967	<sup>1</sup> CAPRARO	87	SPEC	$200 \pi^- Pb \rightarrow \pi^- \pi^0 Pb$
$150 \pm 5$		HUSTON	86	SPEC	$202 \pi^+ A \rightarrow \pi^+ \pi^0 A$
$146 \pm 12$	6.5k	<sup>2</sup> BYERLY	73	OSPK	$5 \pi^- p$
$148.2 \pm 4.1$	9.6k	<sup>3</sup> PISUT	68	RVUE	$1.7-3.2 \pi^- p, t < 10$
$146 \pm 13$	900	EISNER	67	HBC	$4.2 \pi^- p, t < 10$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$137.0 \pm 0.4$	<sup>4</sup> ABLIKIM	17	BES3	$J/\psi \rightarrow \gamma 3\pi$
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<sup>1</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>2</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.

<sup>3</sup> From fit of 3-parameter relativistic  $P$ -wave Breit-Wigner to total mass distribution. Includes BATON 68, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65 and CARMONY 64.

<sup>4</sup> S-matrix pole at a fixed  $\rho$  meson mass of 775.49 MeV.

**NEUTRAL ONLY, PHOTOPRODUCED**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>151.5<sup>+ 1.9</sup><sub>- 2.1</sub> OUR AVERAGE</b>				
151.3 $\pm$ 2.2 <sup>+ 1.6</sup> <sub>- 2.8</sub>	900k	ANDREEV	20	H1 $e p \rightarrow e \pi^+ \pi^- p$
155 $\pm$ 5 $\pm$ 2	63.5k	<sup>1</sup> ABRAMOWICZ12	ZEUS	$e p \rightarrow e \pi^+ \pi^- p$
146 $\pm$ 3 $\pm$ 13	79k	<sup>2</sup> BREITWEG	98B	ZEUS 50–100 $\gamma p$
150.9 $\pm$ 3.0		BARTALUCCI	78	CNTR $\gamma p \rightarrow e^+ e^- p$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
138 $\pm$ 3	79k	<sup>3</sup> BREITWEG	98B	ZEUS 50–100 $\gamma p$
147 $\pm$ 11		GLADDING	73	CNTR 2.9–4.7 $\gamma p$
155 $\pm$ 12	2430	BALLAM	72	HBC 4.7 $\gamma p$
145 $\pm$ 13	1930	BALLAM	72	HBC 2.8 $\gamma p$
140 $\pm$ 5		ALVENSLEB...	70	CNTR $\gamma A, t < 0.01$
146.1 $\pm$ 2.9	140k	BIGGS	70	CNTR $< 4.1 \gamma C \rightarrow \pi^+ \pi^- C$
160 $\pm$ 10		LANZEROTTI	68	CNTR $\gamma p$
130 $\pm$ 5	4000	ASBURY	67B	CNTR $\gamma + Pb$

<sup>1</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho - \omega$  interference.

<sup>2</sup> From the parametrization according to SOEDING 66.

<sup>3</sup> From the parametrization according to ROSS 66.

**NEUTRAL ONLY, OTHER REACTIONS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>150.9 <math>\pm</math> 1.7 OUR AVERAGE</b>		Error includes scale factor of 1.1.		
122 $\pm$ 20		BERTIN	97C	OBLX $0.0 \bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
145.7 $\pm$ 5.3		WEIDENAUER	93	ASTE $\bar{p}p \rightarrow \pi^+ \pi^- \omega$
144.9 $\pm$ 3.7		DUBNICKA	89	RVUE $\pi$ form factor
148 $\pm$ 6	<sup>1,2</sup> BOHACIK		80	RVUE
152 $\pm$ 9		<sup>3</sup> WICKLUND	78	ASPK 3,4,6 $\pi^\pm p N$
154 $\pm$ 2	76k	DEUTSCH...	76	HBC 16 $\pi^+ p$
157 $\pm$ 8	6.8k	<sup>4</sup> RATCLIFF	72	ASPK 15 $\pi^- p, t < 0.3$
143 $\pm$ 8	1.7k	REYNOLDS	69	HBC 2.26 $\pi^- p$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
150.85 $\pm$ 0.55 $\pm$ 0.67	970k	<sup>5</sup> ABLIKIM	18C	BES3 $\eta'(958) \rightarrow \gamma \pi^+ \pi^-$
150.18 $\pm$ 0.55 $\pm$ 0.65	970k	<sup>6</sup> ABLIKIM	18C	BES3 $\eta'(958) \rightarrow \gamma \pi^+ \pi^-$
147.0 $\pm$ 2.5	600k	<sup>7</sup> ABELE	99E	CBAR $0.0 \bar{p}p \rightarrow \pi^+ \pi^- \pi^0$
146 $\pm$ 3	4.9k	<sup>8</sup> ADAMS	97	E665 470 $\mu p \rightarrow \mu XB$
160.0 $\pm$ 4.1 $\pm$ 4.0		<sup>9</sup> CHABAUD	83	ASPK 17 $\pi^- p$ polarized
155 $\pm$ 1		<sup>10</sup> HEYN	81	RVUE $\pi$ form factor
148.0 $\pm$ 1.3	<sup>1,2</sup> LANG		79	RVUE
146 $\pm$ 14	4.1k	ENGLER	74	DBC 6 $\pi^+ n \rightarrow \pi^+ \pi^- p$
143 $\pm$ 13		<sup>2</sup> ESTABROOKS	74	RVUE 17 $\pi^- p \rightarrow \pi^+ \pi^- n$
160 $\pm$ 10	32k	<sup>1</sup> PROTOPOP...	73	HBC 7.1 $\pi^+ p, t < 0.4$
145 $\pm$ 12	2.2k	<sup>3,11</sup> HYAMS	68	OSPK 11.2 $\pi^- p$
163 $\pm$ 15	13.3k	<sup>12</sup> PISUT	68	RVUE 1.7–3.2 $\pi^- p, t < 10$

- <sup>1</sup> From pole extrapolation.  
<sup>2</sup> From phase shift analysis of GRAYER 74 data.  
<sup>3</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.  
<sup>4</sup> Published values contain misprints. Corrected by private communication RATCLIFF 74.  
<sup>5</sup> From a fit to  $\pi^+\pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and box anomaly components.  
<sup>6</sup> From a fit to  $\pi^+\pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and  $\rho(1450)$  components.  
<sup>7</sup> Using relativistic Breit-Wigner and taking into account  $\rho$ - $\omega$  interference.  
<sup>8</sup> Systematic errors not evaluated.  
<sup>9</sup> From fit of 3-parameter relativistic Breit-Wigner to helicity-zero part of  $P$ -wave intensity.  
 CHABAUD 83 includes data of GRAYER 74.  
<sup>10</sup> HEYN 81 includes all spacelike and timelike  $F_\pi$  values until 1978.  
<sup>11</sup> Of HYAMS 68 six parametrizations this is theoretically soundest. MR  
<sup>12</sup> Includes MALAMUD 69, ARMENISE 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, GOLDHABER 64, ABOLINS 63.
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 **$\Gamma_{\rho(770)^0} - \Gamma_{\rho(770)^\pm}$** 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3 ±1.3 OUR AVERAGE</b>		Error includes scale factor of 1.4.		
-0.2 ±1.0		<sup>1</sup> SCHAEL 05C	ALEP $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$	
3.6 ±1.8 ±1.7 1.98M		<sup>2</sup> ALOISIO 03	KLOE 1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.66±0.85		<sup>3</sup> BARTOS 17A	RVUE $e^+e^- \rightarrow \pi^+\pi^-, \tau^- \rightarrow \pi^-\pi^0\nu_\tau$	

<sup>1</sup> From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEL 05C and  $e^+e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05. Supersedes BARATE 97M.

<sup>2</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>3</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, AMBROSINO 11A, and FUJIKAWA 08.

 **$\Gamma_{\rho(770)^+} - \Gamma_{\rho(770)^-}$** 

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.8±2.0±0.5</b>	1.98M	<sup>1</sup> ALOISIO 03	KLOE 1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$	

<sup>1</sup> Without limitations on masses and widths.

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 **$\rho(770)$  DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1 \pi\pi$	~ 100	%
<b><math>\rho(770)^\pm</math> decays</b>		
$\Gamma_2 \pi^\pm\pi^0$	~ 100	%
$\Gamma_3 \pi^\pm\gamma$	( 4.5 ±0.5 ) × 10 <sup>-4</sup>	S=2.2
$\Gamma_4 \pi^\pm\eta$	< 6 × 10 <sup>-3</sup>	CL=84%
$\Gamma_5 \pi^\pm\pi^+\pi^-\pi^0$	< 2.0 × 10 <sup>-3</sup>	CL=84%

**$\rho(770)^0$  decays**

$\Gamma_6$	$\pi^+ \pi^-$	$\sim 100$	%	
$\Gamma_7$	$\pi^+ \pi^- \gamma$	( 9.9 $\pm$ 1.6 )	$\times 10^{-3}$	
$\Gamma_8$	$\pi^0 \gamma$	( 4.7 $\pm$ 0.8 )	$\times 10^{-4}$	S=1.7
$\Gamma_9$	$\eta \gamma$	( 3.00 $\pm$ 0.21 )	$\times 10^{-4}$	
$\Gamma_{10}$	$\pi^0 \pi^0 \gamma$	( 4.5 $\pm$ 0.8 )	$\times 10^{-5}$	
$\Gamma_{11}$	$\mu^+ \mu^-$	[a] ( 4.55 $\pm$ 0.28 )	$\times 10^{-5}$	
$\Gamma_{12}$	$e^+ e^-$	[a] ( 4.72 $\pm$ 0.05 )	$\times 10^{-5}$	
$\Gamma_{13}$	$\pi^+ \pi^- \pi^0$	( 1.01 $\pm$ 0.54 )	$\times 10^{-4}$	
$\Gamma_{14}$	$\pi^+ \pi^- \pi^+ \pi^-$	( 1.8 $\pm$ 0.9 )	$\times 10^{-5}$	
$\Gamma_{15}$	$\pi^+ \pi^- \pi^0 \pi^0$	( 1.6 $\pm$ 0.8 )	$\times 10^{-5}$	
$\Gamma_{16}$	$\pi^0 e^+ e^-$	< 1.2	$\times 10^{-5}$	CL=90%
$\Gamma_{17}$	$\eta e^+ e^-$			

[a] The  $\omega \rho$  interference is then due to  $\omega \rho$  mixing only, and is expected to be small. If  $e\mu$  universality holds,  $\Gamma(\rho^0 \rightarrow \mu^+ \mu^-) = \Gamma(\rho^0 \rightarrow e^+ e^-) \times 0.99785$ .

**CONSTRAINED FIT INFORMATION**

An overall fit to the total width and a partial width uses 10 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 10.7$  for 8 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$$\begin{matrix} & & -100 \\ & \Gamma & \begin{bmatrix} & -100 \\ & 15 & -15 \\ -100 & & \end{bmatrix} \\ x_3 & & \begin{matrix} x_2 & x_3 \end{matrix} \end{matrix}$$

	Mode	Rate (MeV)	Scale factor
$\Gamma_2$	$\pi^\pm \pi^0$	150.2 $\pm$ 2.4	
$\Gamma_3$	$\pi^\pm \gamma$	0.068 $\pm$ 0.007	2.3

**CONSTRAINED FIT INFORMATION**

An overall fit to the total width, a partial width, and 7 branching ratios uses 21 measurements and one constraint to determine 9 parameters. The overall fit has a  $\chi^2 = 9.5$  for 13 degrees of freedom.

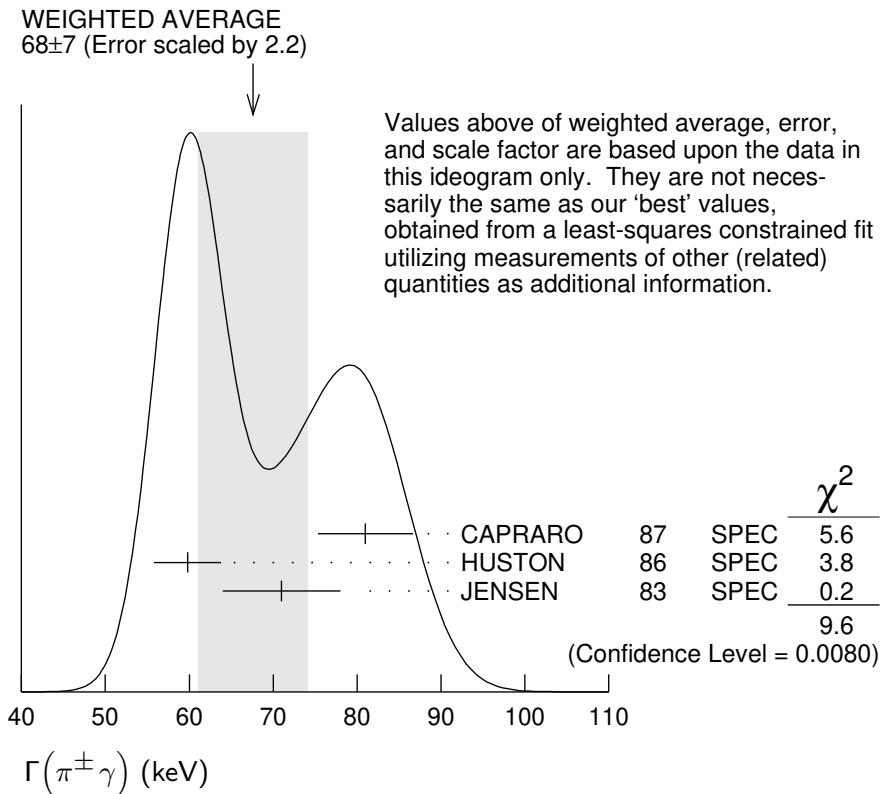
The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \cdot \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_7$	-100							
$x_8$	-5 0							
$x_9$	-1 0 1							
$x_{10}$	-1 0 0 0							
$x_{11}$	2 -3 0 0 0							
$x_{12}$	0 0 -6 -9 0 0							
$x_{14}$	-1 0 0 0 0 0 0							
$\Gamma$	0 0 3 5 0 0 -54 0							
	$x_6$	$x_7$	$x_8$	$x_9$	$x_{10}$	$x_{11}$	$x_{12}$	$x_{14}$

	Mode	Rate (MeV)		Scale factor
$\Gamma_6$	$\pi^+ \pi^-$	147.5	$\pm 0.9$	
$\Gamma_7$	$\pi^+ \pi^- \gamma$	1.48	$\pm 0.24$	
$\Gamma_8$	$\pi^0 \gamma$	0.070	$\pm 0.012$	1.7
$\Gamma_9$	$\eta \gamma$	0.0447	$\pm 0.0032$	
$\Gamma_{10}$	$\pi^0 \pi^0 \gamma$	0.0066	$\pm 0.0012$	
$\Gamma_{11}$	$\mu^+ \mu^-$	[a]	0.0068 $\pm 0.0004$	
$\Gamma_{12}$	$e^+ e^-$	[a]	0.00704 $\pm 0.00006$	
$\Gamma_{14}$	$\pi^+ \pi^- \pi^+ \pi^-$	0.0027	$\pm 0.0014$	

### $\rho(770)$ PARTIAL WIDTHS

$\Gamma(\pi^\pm \gamma)$		$\Gamma_3$
<i>VALUE (keV)</i>	<i>DOCUMENT ID</i>	<i>TECN</i> <i>CHG</i> <i>COMMENT</i>
<b>68 <math>\pm 7</math> OUR FIT</b>	Error includes scale factor of 2.3.	
<b>68 <math>\pm 7</math> OUR AVERAGE</b>	Error includes scale factor of 2.2. See the ideogram below.	
81 $\pm 4$ $\pm 4$	CAPRARO 87 SPEC -	200 $\pi^- A \rightarrow \pi^- \pi^0 A$
59.8 $\pm 4.0$	HUSTON 86 SPEC +	202 $\pi^+ A \rightarrow \pi^+ \pi^0 A$
71 $\pm 7$	JENSEN 83 SPEC -	156–260 $\pi^- A \rightarrow \pi^- \pi^0 A$



### $\Gamma(\pi^0\gamma)$

$\Gamma_8$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
77±17±11	36500	1 ACHASOV	03	SND $0.60\text{--}0.97 e^+e^- \rightarrow \pi^0\gamma$
121±31		DOLINSKY	89	ND $e^+e^- \rightarrow \pi^0\gamma$

<sup>1</sup> Using  $\Gamma_{\text{total}} = 147.9 \pm 1.3$  MeV and  $B(\rho \rightarrow \pi^0\gamma)$  from ACHASOV 03.

### $\Gamma(\eta\gamma)$

$\Gamma_9$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
62±17		1 DOLINSKY	89	ND $e^+e^- \rightarrow \eta\gamma$

<sup>1</sup> Solution corresponding to constructive  $\omega$ - $\rho$  interference.

### $\Gamma(e^+e^-)$

$\Gamma_{12}$

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>7.04 ±0.06 OUR FIT</b>				
<b>7.04 ±0.06 OUR AVERAGE</b>				
7.048±0.057±0.050	900k	1 AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
7.06 ±0.11 ±0.05	114k	2,3 AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
6.77 ±0.10 ±0.30		BARKOV	85	OLYA $e^+e^- \rightarrow \pi^+\pi^-$
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				
7.12 ±0.02 ±0.11	800k	4 ACHASOV	06	SND $e^+e^- \rightarrow \pi^+\pi^-$
6.3 ±0.1		5 BENAYOUN	98	RVUE $e^+e^- \rightarrow \pi^+\pi^-$ , $\mu^+\mu^-$

- <sup>1</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.  
<sup>2</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference.  
<sup>3</sup> From a fit in the energy range 0.61 to 0.96 GeV. Update of AKHMETSHIN 02.  
<sup>4</sup> Supersedes ACHASOV 05A.  
<sup>5</sup> Using the data of BARKOV 85 in the hidden local symmetry model.

$\Gamma(\pi^+\pi^-\pi^+\pi^-)$				$\Gamma_{14}$
VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$2.8 \pm 1.4 \pm 0.5$	153	AKHMETSHIN 00	CMD2	$0.6-0.97 e^+ e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

$$\rho(770) \Gamma(e^+e^-)\Gamma(i)/\Gamma^2(\text{total})$$

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma \times \Gamma_6/\Gamma$
VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ OUR AVERAGE				
$4.89 \pm 0.04$				
$4.889 \pm 0.015 \pm 0.039$		<sup>1</sup> ACHASOV 21	SND	$e^+e^- \rightarrow \pi^+\pi^-$
$4.876 \pm 0.023 \pm 0.064$	800k	<sup>2,3</sup> ACHASOV 06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$4.72 \pm 0.02$		<sup>4</sup> BENAYOUN 10	RVUE	$0.4-1.05 e^+e^-$

- <sup>1</sup> From a fit of the cross section in the energy range  $0.525 < \sqrt{s} < 0.883$  GeV parameterized by the sum of the Breit-Wigner amplitudes for the  $\rho(770)$ ,  $\omega$  and  $\rho(1450)$  resonances.

<sup>2</sup> Supersedes ACHASOV 05A.

<sup>3</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.

<sup>4</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.

$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}$				$\Gamma_{12}/\Gamma \times \Gamma_9/\Gamma$
VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$\bullet \bullet \bullet$ OUR FIT				
$1.42 \pm 0.10$				
$\bullet \bullet \bullet$ OUR AVERAGE				
$1.32 \pm 0.14 \pm 0.08$	33k	<sup>1</sup> ACHASOV 07B	SND	$0.6-1.38 e^+e^- \rightarrow \eta\gamma$
$1.50 \pm 0.65 \pm 0.09$	17.4k	<sup>2</sup> AKHMETSHIN 05	CMD2	$0.60-1.38 e^+e^- \rightarrow \eta\gamma$
$1.61 \pm 0.20 \pm 0.11$	23k	<sup>3,4</sup> AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
$1.85 \pm 0.49$		<sup>5</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \eta\gamma$
$\bullet \bullet \bullet$ We do not use the following data for averages, fits, limits, etc. $\bullet \bullet \bullet$				
$1.05 \pm 0.02$		<sup>6</sup> BENAYOUN 10	RVUE	$0.4-1.05 e^+e^-$

<sup>1</sup> From a combined fit of  $\sigma(e^+e^- \rightarrow \eta\gamma)$  with  $\eta \rightarrow 3\pi^0$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$ , and fixing  $B(\eta \rightarrow 3\pi^0)/B(\eta \rightarrow \pi^+\pi^-\pi^0) = 1.44 \pm 0.04$ . Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>2</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>3</sup> From the  $\eta \rightarrow 3\pi^0$  decay and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .

<sup>4</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).

<sup>5</sup> Recalculated by us from the cross section in the peak.

<sup>6</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.

$\Gamma(e^+ e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}$        $\Gamma_{12}/\Gamma \times \Gamma_8/\Gamma$ 

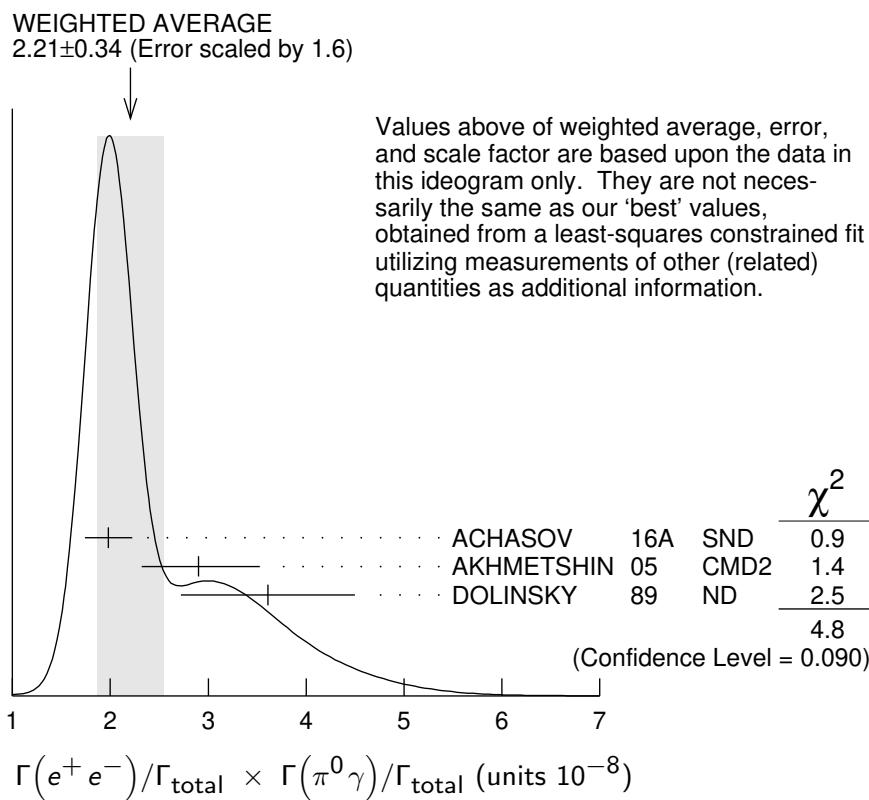
VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.2 ± 0.4 OUR FIT</b>				Error includes scale factor of 1.7.
<b>2.21 ± 0.34 OUR AVERAGE</b>				Error includes scale factor of 1.6. See the ideogram below.
1.98 ± 0.22 ± 0.10	1 ACHASOV	16A SND	0.60-1.38 $e^+ e^- \rightarrow \pi^0 \gamma$	
2.90 ± 0.60 ± 0.55	18k AKHMETSHIN	05 CMD2	0.60-1.38 $e^+ e^- \rightarrow \pi^0 \gamma$	
3.61 ± 0.74 ± 0.49	10k DOLINSKY	89 ND	$e^+ e^- \rightarrow \pi^0 \gamma$	
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.875 ± 0.026	3 BENAYOUN	10 RVUE	0.4-1.05 $e^+ e^-$	
2.37 ± 0.53 ± 0.33	36k ACHASOV	03 SND	0.60-0.97 $e^+ e^- \rightarrow \pi^0 \gamma$	

<sup>1</sup> From the VMD model with the rho(770), omega(782), phi(1020) resonances, and an additional resonance describing the total contribution of the rho(1450) and omega(1420) states. Supersedes ACHASOV 03.

<sup>2</sup> Recalculated by us from the cross section in the peak.

<sup>3</sup> A simultaneous fit of  $e^+ e^- \rightarrow \pi^+ \pi^-$ ,  $\pi^+ \pi^- \pi^0$ ,  $\pi^0 \gamma$ ,  $\eta \gamma$  data.

<sup>4</sup> Using  $\sigma_{\phi \rightarrow \pi^0 \gamma}$  from ACHASOV 00 and  $m_\rho = 775.97$  MeV in the model with the energy-independent phase of rho-omega interference equal to  $(-10.2 \pm 7.0)^\circ$ .


 $\Gamma(e^+ e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$        $\Gamma_{12}/\Gamma \times \Gamma_{13}/\Gamma$ 

VALUE (units $10^{-9}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.903 ± 0.076	1 BENAYOUN	10 RVUE	0.4-1.05 $e^+ e^-$	
4.58 ± 2.46 ± 1.56	1.2M ACHASOV	03D RVUE	0.44-2.00 $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$	

<sup>1</sup> A simultaneous fit of  $e^+ e^- \rightarrow \pi^+ \pi^-$ ,  $\pi^+ \pi^- \pi^0$ ,  $\pi^0 \gamma$ ,  $\eta \gamma$  data.<sup>2</sup> Statistical significance is less than  $3\sigma$ .

## $\rho(770)$ BRANCHING RATIOS

### $\Gamma(\pi^\pm \eta)/\Gamma(\pi\pi)$

<u>VALUE</u> (units $10^{-4}$ )	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<60	84	FERBEL	66	HBC	$\pm$ $\pi^\pm p$ above 2.5

### $\Gamma(\pi^\pm \pi^+ \pi^- \pi^0)/\Gamma(\pi\pi)$

<u>VALUE</u> (units $10^{-4}$ )	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<20	84	FERBEL	66	HBC	$\pm$ $\pi^\pm p$ above 2.5

• • • We do not use the following data for averages, fits, limits, etc. • • •

35  $\pm$  40 JAMES 66 HBC + 2.1  $\pi^+ p$

### $\Gamma(\pi^+ \pi^- \gamma)/\Gamma_{\text{total}}$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0099 <math>\pm</math> 0.0016 OUR FIT</b>				

**0.0099  $\pm$  0.0016**

<sup>1</sup> DOLINSKY 91 ND  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.0111  $\pm$  0.0014

<sup>2</sup> VASSERMAN 88 ND  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma$

<0.005

<sup>3</sup> VASSERMAN 88 ND  $e^+ e^- \rightarrow \pi^+ \pi^- \gamma$

<sup>1</sup> Bremsstrahlung from a decay pion and for photon energy above 50 MeV.

<sup>2</sup> Superseded by DOLINSKY 91.

<sup>3</sup> Structure radiation due to quark rearrangement in the decay.

### $\Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}$

<u>VALUE</u> (units $10^{-4}$ )	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>• • • We do not use the following data for averages, fits, limits, etc. • • •</b>				

4.20  $\pm$  0.52

<sup>1</sup> ACHASOV 16A SND 0.60–1.38  $e^+ e^- \rightarrow \pi^0 \gamma$

6.21  $^{+1.28}_{-1.18} \pm 0.39$

<sup>2,3</sup> AKHMETSHIN 05 CMD2 0.60–1.38  $e^+ e^- \rightarrow \pi^0 \gamma$

5.22  $\pm 1.17 \pm 0.75$

<sup>3,4</sup> ACHASOV 03 SND 0.60–0.97  $e^+ e^- \rightarrow \pi^0 \gamma$

6.8  $\pm 1.7$

<sup>5</sup> BENAYOUN 96 RVUE 0.54–1.04  $e^+ e^- \rightarrow \pi^0 \gamma$

7.9  $\pm 2.0$

<sup>3</sup> DOLINSKY 89 ND  $e^+ e^- \rightarrow \pi^0 \gamma$

<sup>1</sup> Using  $B(\rho \rightarrow e^+ e^-)$  from PDG 15. Supersedes ACHASOV 03.

<sup>2</sup> Using  $B(\rho \rightarrow e^+ e^-) = (4.67 \pm 0.09) \times 10^{-5}$ .

<sup>3</sup> Not independent of the corresponding  $\Gamma(e^+ e^-) \times \Gamma(\pi^0 \gamma)/\Gamma_{\text{total}}^2$ .

<sup>4</sup> Using  $B(\rho \rightarrow e^+ e^-) = (4.54 \pm 0.10) \times 10^{-5}$ .

<sup>5</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution.

$\Gamma(\eta\gamma)/\Gamma_{\text{total}}$	$\Gamma_9/\Gamma$			
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG COMMENT</u>
<b>3.00±0.21 OUR FIT</b>				
<b>2.90±0.32 OUR AVERAGE</b>				
2.79±0.34±0.03	33k	1 ACHASOV 07B	SND	0.6–1.38 $e^+e^- \rightarrow \eta\gamma$
3.6 ± 0.9		2 ANDREWS 77	CNTR 0	6.7–10 $\gamma\text{Cu}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
3.21±1.39±0.20	17.4k	3,4 AKHMETSHIN 05	CMD2	0.60–1.38 $e^+e^- \rightarrow \eta\gamma$
3.39±0.42±0.23		2,5,6 AKHMETSHIN 01B	CMD2	$e^+e^- \rightarrow \eta\gamma$
1.9 $^{+0.6}_{-0.8}$		7 BENAYOUN 96	RVUE	0.54–1.04 $e^+e^- \rightarrow \eta\gamma$
4.0 ± 1.1		2,4 DOLINSKY 89	ND	$e^+e^- \rightarrow \eta\gamma$
<sup>1</sup> ACHASOV 07B reports $[\Gamma(\rho(770) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\rho(770) \rightarrow e^+e^-)] = (1.32 \pm 0.14 \pm 0.08) \times 10^{-8}$ which we divide by our best value $B(\rho(770) \rightarrow e^+e^-) = (4.72 \pm 0.05) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00D and ACHASOV 06A. <sup>2</sup> Solution corresponding to constructive $\omega$ - $\rho$ interference. <sup>3</sup> Using $B(\rho \rightarrow e^+e^-) = (4.67 \pm 0.09) \times 10^{-5}$ and $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ . <sup>4</sup> Not independent of the corresponding $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ . <sup>5</sup> The combined fit from 600 to 1380 MeV taking into account $\rho(770)$ , $\omega(782)$ , $\phi(1020)$ , and $\rho(1450)$ (mass and width fixed at 1450 MeV and 310 MeV respectively). <sup>6</sup> Using $B(\rho \rightarrow e^+e^-) = (4.75 \pm 0.10) \times 10^{-5}$ from AKHMETSHIN 02 and $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ . <sup>7</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution. Constructive $\rho$ - $\omega$ interference solution.				

$\Gamma(\pi^0\pi^0\gamma)/\Gamma_{\text{total}}$	$\Gamma_{10}/\Gamma$			
<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>4.5±0.8 OUR FIT</b>				
<b>4.5<math>^{+0.9}_{-0.8}</math> OUR AVERAGE</b>				
5.2 $^{+1.5}_{-1.3}$ ±0.6	190	1 AKHMETSHIN 04B	CMD2	0.6–0.97 $e^+e^- \rightarrow \pi^0\pi^0\gamma$
4.1 $^{+1.0}_{-0.9}$ ±0.3	295	2 ACHASOV 02F	SND	0.36–0.97 $e^+e^- \rightarrow \pi^0\pi^0\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.8 $^{+3.4}_{-1.8}$ ±0.5	63	3 ACHASOV 00G	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
<sup>1</sup> This branching ratio includes the conventional VMD mechanism $\rho \rightarrow \omega\pi^0$ , $\omega \rightarrow \pi^0\gamma$ , and the new decay mode $\rho \rightarrow f_0(500)\gamma$ , $f_0(500) \rightarrow \pi^0\pi^0$ with a branching ratio $(2.0^{+1.1}_{-0.9}) \times 10^{-5}$ differing from zero by 2.0 standard deviations. <sup>2</sup> This branching ratio includes the conventional VMD mechanism $\rho \rightarrow \omega\pi^0$ , $\omega \rightarrow \pi^0\gamma$ , and the new decay mode $\rho \rightarrow f_0(500)\gamma$ , $f_0(500) \rightarrow \pi^0\pi^0$ with a branching ratio $(1.9^{+0.9}_{-0.8} \pm 0.4) \times 10^{-5}$ differing from zero by 2.4 standard deviations. Supersedes ACHASOV 00G. <sup>3</sup> Superseded by ACHASOV 02F.				

$\Gamma(\mu^+\mu^-)/\Gamma(\pi^+\pi^-)$	$\Gamma_{11}/\Gamma_6$
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<u>VALUE</u> (units $10^{-5}$ )	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>4.60±0.28 OUR FIT</b>			
<b>4.6 ±0.2 ±0.2</b>	ANTIPOV 89	SIGM	$\pi^- \text{Cu} \rightarrow \mu^+ \mu^- \pi^- \text{Cu}$
• • • We do not use the following data for averages, fits, limits, etc. • • •			
8.2 +1.6 -3.6	1 ROTHWELL 69	CNTR	Photoproduction
5.6 ±1.5	2 WEHMANN 69	OSPK	12 $\pi^- \text{C, Fe}$
9.7 +3.1 -3.3	3,4 HYAMS 67	OSPK	11 $\pi^- \text{Li, H}$

<sup>1</sup> Possibly large  $\rho$ - $\omega$  interference leads us to increase the minus error.

<sup>2</sup> Result contains  $11 \pm 11\%$  correction using SU(3) for central value. The error on the correction takes account of possible  $\rho$ - $\omega$  interference and the upper limit agrees with the upper limit of  $\omega \rightarrow \mu^+ \mu^-$  from this experiment.

<sup>3</sup> But he even enlarges his error to take residual  $\omega$  contamination into account. Since his value is high, seems the other experiments also can't have too many  $\omega$ 's. But maybe Hyams has additional  $\mu$ 's from  $\rho \rightarrow \pi\pi$ , decaying  $\pi$ 's.

<sup>4</sup> HYAMS 67's mass resolution is 20 MeV. The  $\omega$  region was excluded.

$\Gamma(e^+e^-)/\Gamma(\pi\pi)$	$\Gamma_{12}/\Gamma_1$
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<u>VALUE</u> (units $10^{-4}$ )	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.40±0.05	1,2 BENAKSAS 72	OSPK	$e^+ e^- \rightarrow \pi^+ \pi^-$
<sup>1</sup> The $\rho'$ contribution is not taken into account.			
<sup>2</sup> Barkov excludes Auslender and Benaksas for large statistical and systematic errors.			

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$	$\Gamma_{13}/\Gamma$
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<u>VALUE</u> (units $10^{-4}$ )	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.88±0.23±0.30</b>			1 LEES	21B BABR	$10.5 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.01^{+0.54}_{-0.36} \pm 0.34$	1.2M	2 ACHASOV	03D RVUE	$0.44-2.00 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
<1.2	90	VASSERMAN	88B ND	$e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> From the cross section for  $e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$  with contributions from  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ ,  $\omega(1420)$ , and  $\omega(1650)$ . Statistical evidence is more than  $6\sigma$ .

<sup>2</sup> Statistical significance is less than  $3\sigma$ .

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\pi\pi)$	$\Gamma_{13}/\Gamma_1$
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<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •					
~0.01		BRAMON 86	RVUE	0	$J/\psi \rightarrow \omega \pi^0$
<0.01	84	1 ABRAMS 71	HBC	0	$3.7 \pi^+ p$

<sup>1</sup> Model dependent, assumes  $l = 1, 2$ , or  $3$  for the  $3\pi$  system.

$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.8 \pm 0.9</math> OUR FIT</b>					
<b><math>1.8 \pm 0.9 \pm 0.3</math></b>	153		AKHMETSHIN 00	CMD2	$0.6-0.97 e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<20	90	KURDADZE	88	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
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$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma(\pi\pi)$   $\Gamma_{14}/\Gamma_1$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<15	90	ERBE	69	HBC	$0 \quad 2.5-5.8 \gamma p$
<20		CHUNG	68	HBC	$0 \quad 3.2, 4.2 \pi^- p$
<20	90	HUSON	68	HLBC	$0 \quad 16.0 \pi^- p$
<80		JAMES	66	HBC	$0 \quad 2.1 \pi^+ p$

$\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>1.60 \pm 0.74 \pm 0.18</math></b>		<sup>1</sup> ACHASOV 09A	SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 4	90	AULCHENKO 87c	ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
<20	90	KURDADZE 86	OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

<sup>1</sup> Assuming no interference between the  $\rho$  and  $\omega$  contributions.

$\Gamma(\pi^0e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{16}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;1.2</b>	90	ACHASOV 08	SND	$0.36-0.97 e^+e^- \rightarrow \pi^0e^+e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.6		AKHMETSHIN 05A	CMD2	$0.72-0.84 e^+e^-$
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$\Gamma(\eta e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{17}/\Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
• • • We do not use the following data for averages, fits, limits, etc. • • •			
<0.7	AKHMETSHIN 05A	CMD2	$0.72-0.84 e^+e^-$

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AMBROSINO	11A	PL B700 102	F. Ambrosino <i>et al.</i> (KLOE Collab.)
GARCIA-MAR...	11	PRL 107 072001	R. Garcia-Martin <i>et al.</i> (MADR, CRAC)
BATLEY	10C	EPJ C70 635	J.R. Batley <i>et al.</i> (CERN NA48/2 Collab.)
BENAYOUN	10	EPJ C65 211	M. Benayoun <i>et al.</i>
DUBNICKA	10	APS 60 1	S. Dubnicka, A.Z. Dubnickova
ACHASOV	09A	JETP 109 379	M.N. Achasov <i>et al.</i> (SND Collab.)
		Translated from ZETF 136 442.	
AUBERT	09AS	PRL 103 231801	B. Aubert <i>et al.</i> (BABAR Collab.)
ACHASOV	08	JETP 107 61	M.N. Achasov <i>et al.</i> (SND Collab.)
		Translated from ZETF 134 80.	
FUJIKAWA	08	PR D78 072006	M. Fujikawa <i>et al.</i> (BELLE Collab.)
ACHASOV	07B	PR D76 077101	M.N. Achasov <i>et al.</i> (SND Collab.)
AKHMETSHIN	07	PL B648 28	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
ACHASOV	06	JETP 103 380	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
		Translated from ZETF 130 437.	
ACHASOV	06A	PR D74 014016	M.N. Achasov <i>et al.</i> (SND Collab.)
AULCHENKO	06	JETPL 84 413	V.M. Aulchenko <i>et al.</i> (Novosibirsk CMD-2 Collab.)
		Translated from ZETFP 84 491.	
ACHASOV	05A	JETP 101 1053	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
		Translated from ZETF 128 1201.	
AKHMETSHIN	05	PL B605 26	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
AKHMETSHIN	05A	PL B613 29	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
ALOISIO	05	PL B606 12	A. Aloisio <i>et al.</i> (KLOE Collab.)
AULCHENKO	05	JETPL 82 743	V.M. Aulchenko <i>et al.</i> (Novosibirsk CMD-2 Collab.)
		Translated from ZETFP 82 841.	
SCHAEL	05C	PRPL 421 191	S. Schael <i>et al.</i> (ALEPH Collab.)
AKHMETSHIN	04	PL B578 285	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
AKHMETSHIN	04B	PL B580 119	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
PELAEZ	04A	MPL A19 2879	J.R. Pelaez (MADU)
ACHASOV	03	PL B559 171	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
ACHASOV	03D	PR D68 052006	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
ALOISIO	03	PL B561 55	A. Aloisio <i>et al.</i> (KLOE Collab.)
SANZ-CILLERO	03	EPJ C27 587	J.J. Sanz-Cillero, A. Pich
ACHASOV	02	PR D65 032002	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
ACHASOV	02F	PL B537 201	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
AKHMETSHIN	02	PL B527 161	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
AKHMETSHIN	01B	PL B509 217	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
COLANGELO	01	NP B603 125	G. Colangelo, J. Gasser, H. Leytwyler
PICH	01	PR D63 093005	A. Pich, J. Portoles
ACHASOV	00	EPJ C12 25	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
ACHASOV	00D	JETPL 72 282	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
		Translated from ZETFP 72 411.	
ACHASOV	00G	JETPL 71 355	M.N. Achasov <i>et al.</i> (Novosibirsk SND Collab.)
		Translated from ZETFP 71 519.	
AKHMETSHIN	00	PL B475 190	R.R. Akhmetshin <i>et al.</i> (Novosibirsk CMD-2 Collab.)
ANDERSON	00A	PR D61 112002	S. Anderson <i>et al.</i> (CLEO Collab.)
ABELE	99E	PL B469 270	A. Abele <i>et al.</i> (Crystal Barrel Collab.)
BENAYOUN	98	EPJ C2 269	M. Benayoun <i>et al.</i> (IPNP, NOVO, ADLD+)
BREITWEG	98B	EPJ C2 247	J. Breitweg <i>et al.</i> (ZEUS Collab.)
GARDNER	98	PR D57 2716	S. Gardner, H.B. O'Connell
Also		PR D62 019903 (errat.)	S. Gardner, H.B. O'Connell
ABELE	97	PL B391 191	A. Abele <i>et al.</i> (Crystal Barrel Collab.)
ADAMS	97	ZPHY C74 237	M.R. Adams <i>et al.</i> (E665 Collab.)
BARATE	97M	ZPHY C76 15	R. Barate <i>et al.</i> (ALEPH Collab.)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i> (OBELIX Collab.)
BOGOLYUB...	97	PAN 60 46	M.Y. Bogolyubsky <i>et al.</i> (MOSU, SERP)
		Translated from YAF 60 53.	
O'CONNELL	97	NP A623 559	H.B. O'Connell <i>et al.</i> (ADLD)
BENAYOUN	96	ZPHY C72 221	M. Benayoun <i>et al.</i> (IPNP, NOVO)
BERNICA	94	PR D50 4454	A. Bernicha, G. Lopez Castro, J. Pestieau (LOUV+)
WEIDENAUER	93	ZPHY C59 387	P. Weidenauer <i>et al.</i> (ASTERIX Collab.)
AGUILAR-...	91	ZPHY C50 405	M. Aguilar-Benitez <i>et al.</i> (LEBC-EHS Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i> (NOVO)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i> (MPIM)
ANTIROV	89	ZPHY C42 185	Y.M. Antipov <i>et al.</i> (SERP, JINR, BGNA+)
BISELLO	89	PL B220 321	D. Bisello <i>et al.</i> (DM2 Collab.)
DOLINSKY	89	ZPHY C42 511	S.I. Dolinsky <i>et al.</i> (NOVO)
DUBNICKA	89	JP G15 1349	S. Dubnicka <i>et al.</i> (JINR, SLOV)
GESHKEN...	89	ZPHY C45 351	B.V. Geshkenbein (ITEP)
KURDADZE	88	JETPL 47 512	L.M. Kurdadze <i>et al.</i> (NOVO)
		Translated from ZETFP 47 432.	

VASSERMAN	88	SJNP 47 1035 Translated from YAF 47 1635.	I.B. Vasserman <i>et al.</i>	(NOVO)
VASSERMAN	88B	SJNP 48 480 Translated from YAF 48 753.	I.B. Vasserman <i>et al.</i>	(NOVO)
AULCHENKO	87C	IYF 87-90 Preprint	V.M. Aulchenko <i>et al.</i>	(NOVO)
CAPRARO	87	NP B288 659	L. Capraro <i>et al.</i>	(CLER, FRAS, MILA+)
BRAMON	86	PL B173 97	A. Bramon, J. Casulleras	(BARC)
HUSTON	86	PR D33 3199	J. Huston <i>et al.</i>	(ROCH, FNAL, MINN)
KURDADZE	86	JETPL 43 643 Translated from ZETFP 43 497.	L.M. Kurdadze <i>et al.</i>	(NOVO)
BARKOV	85	NP B256 365	L.M. Barkov <i>et al.</i>	(NOVO)
DRUZHININ	84	PL 144B 136	V.P. Druzhinin <i>et al.</i>	(NOVO)
CHABAUD	83	NP B223 1	V. Chabaud <i>et al.</i>	(CERN, CRAC, MPIM)
JENSEN	83	PR D27 26	T. Jensen <i>et al.</i>	(ROCH, FNAL, MINN)
HEYN	81	ZPHY C7 169	M.F. Heyn, C.B. Lang	(GRAZ)
BOHACIK	80	PR D21 1342	J. Bohacik, H. Kuhnelt	(SLOV, WIEN)
COHEN	80	PR D22 2595	D. Cohen <i>et al.</i>	(ANL)
LANG	79	PR D19 956	C.B. Lang, A. Mas-Parareda	(GRAZ)
BARTALUCCI	78	NC 44A 587	S. Bartalucci <i>et al.</i>	(DESY, FRAS)
WICKLUND	78	PR D17 1197	A.B. Wicklund <i>et al.</i>	(ANL)
ANDREWS	77	PRL 38 198	D.E. Andrews <i>et al.</i>	(ROCH)
DEUTSCH...	76	NP B103 426	M. Deutschmann <i>et al.</i>	(AACH3, BERL, BONN+)
ENGLER	74	PR D10 2070	A. Engler <i>et al.</i>	(CMU, CASE)
ESTABROOKS	74	NP B79 301	P.G. Estabrooks, A.D. Martin	(DURH)
GRAYER	74	NP B75 189	G. Grayer <i>et al.</i>	(CERN, MPIM)
RATCLIFF	74	Private Comm.		
BYERLY	73	PR D7 637	W.L. Byerly <i>et al.</i>	(MICH)
GLADDING	73	PR D8 3721	G.E. Gladding <i>et al.</i>	(HARV)
HYAMS	73	NP B64 134	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
PROTOPOP...	73	PR D7 1279	S.D. Protopopescu <i>et al.</i>	(LBL)
BALLAM	72	PR D5 545	J. Ballam <i>et al.</i>	(SLAC, LBL, TUFTS)
BENAKSAS	72	PL 39B 289	D. Benaksas <i>et al.</i>	(ORSAY)
JACOBS	72	PR D6 1291	L.D. Jacobs	(SACL)
RATCLIFF	72	PL 38B 345	B.N. Ratcliff <i>et al.</i>	(SLAC)
ABRAMS	71	PR D4 653	G.S. Abrams <i>et al.</i>	(LBL)
ALVENSLEB...	70	PRL 24 786	H. Alvensleben <i>et al.</i>	(DESY)
BIGGS	70	PRL 24 1197	P.J. Biggs <i>et al.</i>	(DARE)
ERBE	69	PR 188 2060	R. Erbe <i>et al.</i>	(German Bubble Chamber Collab.)
MALAMUD	69	Argonne Conf. 93	E.I. Malamud, P.E. Schlein	(UCLA)
REYNOLDS	69	PR 184 1424	B.G. Reynolds <i>et al.</i>	(FSU)
ROTHWELL	69	PRL 23 1521	P.L. Rothwell <i>et al.</i>	(NEAS)
WEHMANN	69	PR 178 2095	A.A. Wehmann <i>et al.</i>	(HARV, CASE, SLAC+)
ARMENISE	68	NC 54A 999	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ+)
BATON	68	PR 176 1574	J.P. Baton, G. Laurens	(SACL)
CHUNG	68	PR 165 1491	S.U. Chung <i>et al.</i>	(LRL)
FOSTER	68	NP B6 107	M. Foster <i>et al.</i>	(CERN, CDEF)
GOUNARIS	68	PRL 21 244	G.J. Gounaris, J.J. Sakurai	
HUSON	68	PL 28B 208	R. Huson <i>et al.</i>	(ORSAY, MILA, UCLA)
HYAMS	68	NP B7 1	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
LANZEROTTI	68	PR 166 1365	L.J. Lanzerotti <i>et al.</i>	(HARV)
PISUT	68	NP B6 325	J. Pisut, M. Roos	(CERN)
ASBURY	67B	PRL 19 865	J.G. Asbury <i>et al.</i>	(DESY, COLU)
BACON	67	PR 157 1263	T.C. Bacon <i>et al.</i>	(BNL)
EISNER	67	PR 164 1699	R.L. Eisner <i>et al.</i>	(PURD)
HUWE	67	PL 24B 252	D.O. Huwe <i>et al.</i>	(COLU)
HYAMS	67	PL 24B 634	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
MILLER	67B	PR 153 1423	D.H. Miller <i>et al.</i>	(PURD)
ALFF-...	66	PR 145 1072	C. Alff-Steinberger <i>et al.</i>	(COLU, RUTG)
FERBEL	66	PL 21 111	T. Ferbel	(ROCH)
HAGOPIAN	66	PR 145 1128	V. Hagopian <i>et al.</i>	(PENN, SACL)
HAGOPIAN	66B	PR 152 1183	V. Hagopian, Y.L. Pan	(PENN, LRL)
JACOBS	66B	UCRL 16877	L.D. Jacobs	(LRL)
JAMES	66	PR 142 896	F.E. James, H.L. Kraybill	(YALE, BNL)
ROSS	66	PR 149 1172	M. Ross, L. Stodolsky	
SOEDING	66	PL B19 702	P. Soeding	
WEST	66	PR 149 1089	E. West <i>et al.</i>	(WISC)
BLIEDEN	65	PL 19 444	H.R. Blieden <i>et al.</i>	(CERN MMS Collab.)
CARMONY	64	PRL 12 254	D.D. Carmony <i>et al.</i>	(UCB)
GOLDHABER	64	PRL 12 336	G. Goldhaber <i>et al.</i>	(LRL, UCB)
ABOLINS	63	PRL 11 381	M.A. Abolins <i>et al.</i>	(UCSD)