



$$I^G(J^P) = 1^-(0^-)$$

We have omitted some results that have been superseded by later experiments. The omitted results may be found in our 1988 edition Physics Letters **B204** 1 (1988).

NODE=S008

NODE=S008

$\pi^\pm$ MASS

NODE=S008M

NODE=S008M

The most accurate charged pion mass measurements are based upon x-ray wavelength measurements for transitions in π^- -mesonic atoms. The observed line is the blend of three components, corresponding to different K-shell occupancies. JECKELMANN 94 revisits the occupancy question, with the conclusion that two sets of occupancy ratios, resulting in two different pion masses (Solutions A and B), are equally probable. We choose the higher Solution B since only this solution is consistent with a positive mass-squared for the muon neutrino, given the precise muon momentum measurements now available (DAUM 91, ASSAMAGAN 94, and ASSAM-AGAN 96) for the decay of pions at rest. Earlier mass determinations with pi-mesonic atoms may have used incorrect K-shell screening corrections.

Measurements with an error of > 0.005 MeV have been omitted from this Listing.

VALUE (MeV)	DOCUMENT ID	TECN	CHG	COMMENT
139.57039±0.00018 OUR FIT	Error includes scale factor of 1.8.			
139.57039±0.00017 OUR AVERAGE	Error includes scale factor of 1.6. See the ideogram below.			
139.57021±0.00014	¹ DAUM 19	SPEC		$\pi^+ \rightarrow \mu^+ \nu_\mu$
139.57077±0.00018	² TRASSINELLI 16	CNTR		X-ray transitions in pionic N ₂
139.57071±0.00053	³ LENZ 98	CNTR	—	pionic N ₂ -atoms gas target
139.56995±0.00035	⁴ JECKELMANN 94	CNTR	—	π^- atom, Soln. B
• • • We do not use the following data for averages, fits, limits, etc. • • •				
139.57022±0.00014	⁵ ASSAMAGAN 96	SPEC	+	$\pi^+ \rightarrow \mu^+ \nu_\mu$
139.56782±0.00037	⁶ JECKELMANN 94	CNTR	—	π^- atom, Soln. A
139.56996±0.00067	⁷ DAUM 91	SPEC	+	$\pi^+ \rightarrow \mu^+ \nu$
139.56752±0.00037	⁸ JECKELMANN 86B	CNTR	—	Mesonic atoms
139.5704 ±0.0011	⁷ ABELA 84	SPEC	+	See DAUM 91
139.5664 ±0.0009	⁹ LU 80	CNTR	—	Mesonic atoms
139.5686 ±0.0020	CARTER 76	CNTR	—	Mesonic atoms
139.5660 ±0.0024	^{9,10} MARUSHEN... 76	CNTR	—	Mesonic atoms

NODE=S008M

OCCUR=2

OCCUR=2

¹ DAUM 19 value is based on their previous (1991+1996) measurements of the μ^+ momentum of 29.79200 ± 0.00011 MeV for π^+ decay at rest. It also uses $m_\mu = 105.6583745 \pm 0.0000024$ MeV, and assumes conservatively $m_{\nu_\mu} = 2.0 \pm 2.0$ MeV. It is the most precise charged pion mass determination.

NODE=S008M;LINKAGE=G

² TRASSINELLI 16 use the muonic oxygen line for online energy calibration of the pionic line.

NODE=S008M;LINKAGE=F

³ LENZ 98 result does not suffer K-electron configuration uncertainties as does JECKELMANN 94.

NODE=S008M;LINKAGE=LL

⁴ JECKELMANN 94 Solution B (dominant 2-electron K-shell occupancy), chosen for consistency with positive $m_{\nu_\mu}^2$.

NODE=S008M;LINKAGE=JB

⁵ ASSAMAGAN 96 measures the μ^+ momentum p_μ in $\pi^+ \rightarrow \mu^+ \nu_\mu$ decay at rest to be 29.79200 ± 0.00011 MeV/c. Combined with the μ^+ mass and the assumption $m_{\nu_\mu} = 0$, this gives the π^+ mass above; if $m_{\nu_\mu} > 0$, m_{π^+} given above is a lower limit. Combined instead with m_μ and (assuming *CPT*) the π^- mass of JECKELMANN 94, p_μ gives an upper limit on m_{ν_μ} (see the ν_μ).

NODE=S008M;LINKAGE=E

⁶ JECKELMANN 94 Solution A (small 2-electron K-shell occupancy) in combination with either the DAUM 91 or ASSAMAGAN 94 pion decay muon momentum measurement yields a significantly negative $m_{\nu_\mu}^2$. It is accordingly not used in our fits.

NODE=S008M;LINKAGE=JA

⁷ The DAUM 91 value includes the ABELA 84 result. The value is based on a measurement of the μ^+ momentum for π^+ decay at rest, $p_\mu = 29.79179 \pm 0.00053$ MeV, uses $m_\mu = 105.658389 \pm 0.000034$ MeV, and assumes that $m_{\nu_\mu} = 0$. The last assumption means that in fact the value is a lower limit.

NODE=S008M;LINKAGE=A

⁸ JECKELMANN 86B gives $m_\pi/m_e = 273.12677(71)$. We use $m_e = 0.51099906(15)$ MeV from COHEN 87. The authors note that two solutions for the probability distribution

NODE=S008M;LINKAGE=J

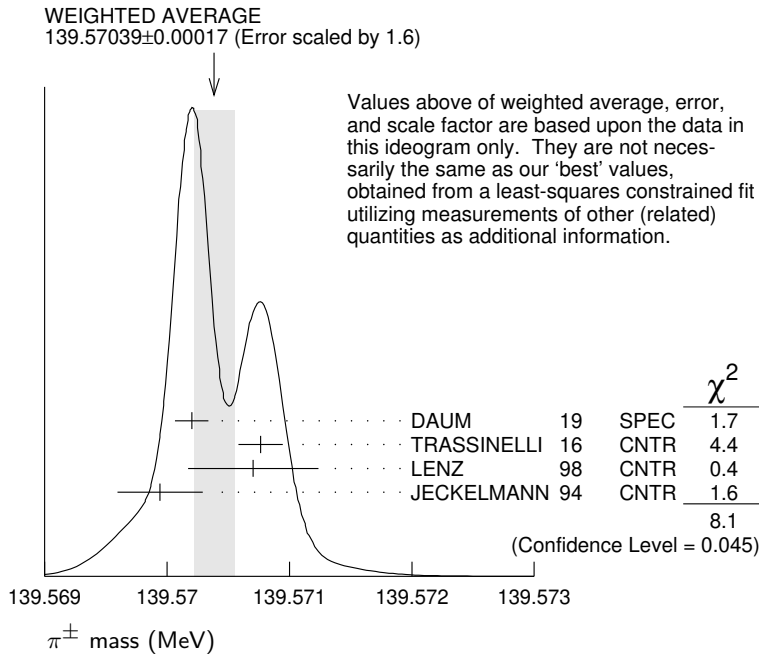
of K-shell occupancy fit equally well, and use other data to choose the lower of the two possible $\pi^\pm$ masses.

⁹ These values are scaled with a new wavelength-energy conversion factor $V\lambda = 1.23984244(37) \times 10^{-6}$ eV m from COHEN 87. The LU 80 screening correction relies upon a theoretical calculation of inner-shell refilling rates.

¹⁰ This MARUSHENKO 76 value used at the authors' request to use the accepted set of calibration γ energies. Error increased from 0.0017 MeV to include QED calculation error of 0.0017 MeV (12 ppm).

NODE=S008M;LINKAGE=C

NODE=S008M;LINKAGE=M



$$m_{\pi^+} = m_{\mu^+}$$

NODE=S008D

Measurements with an error > 0.05 MeV have been omitted from this Listing.

NODE=S008D

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT	
• • • We do not use the following data for averages, fits, limits, etc. • • •						
33.91157±0.00067		¹ DAUM	91	SPEC	+	$\pi^+ \rightarrow \mu^+ \nu$
33.9111 ±0.0011		ABELA	84	SPEC		See DAUM 91
33.925 ±0.025		BOOTH	70	CNTR	+	Magnetic spect.
33.881 ±0.035	145	HYMAN	67	HEBC	+	K^- He

¹ The DAUM 91 value assumes that $m_{\nu_\mu} = 0$ and uses our $m_\mu = 105.658389 \pm 0.000034$ MeV.

NODE=S008D

NODE=S008D;LINKAGE=A

$$(m_{\pi^+} - m_{\pi^-}) / m_{\text{average}}$$

NODE=S008DM

A test of CPT invariance.

NODE=S008DM

VALUE (units 10^{-4})	DOCUMENT ID	TECN
2±5	AYRES	71 CNTR

NODE=S008DM

$\pi^\pm$ MEAN LIFE

NODE=S008T

Measurements with an error > 0.02×10^{-8} s have been omitted.

NODE=S008T

<u>VALUE (10^{-8} s)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>	
2.6033 $\pm$0.0005 OUR AVERAGE	Error includes scale factor of 1.2.				
2.60361 $\pm$ 0.00052	¹ KOPTEV	95	SPEC	+	Surface μ^+ 's
2.60231 $\pm$ 0.00050 $\pm$ 0.00084	NUMAO	95	SPEC	+	Surface μ^+ 's
2.609 $\pm$ 0.008	DUNAITSEV	73	CNTR	+	
2.602 $\pm$ 0.004	AYRES	71	CNTR	$\pm$	
2.604 $\pm$ 0.005	NORDBERG	67	CNTR	+	
2.602 $\pm$ 0.004	ECKHAUSE	65	CNTR	+	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
2.640 $\pm$ 0.008	² KINSEY	66	CNTR	+	

NODE=S008T

¹KOPTEV 95 combines the statistical and systematic errors; the statistical error dominates.

²Systematic errors in the calibration of this experiment are discussed by NORDBERG 67.

$$(\tau_{\pi^+} - \tau_{\pi^-}) / \tau_{\text{average}}$$

A test of *CPT* invariance.

VALUE (units 10^{-4})	DOCUMENT ID	TECN
5.5 ± 7.1	AYRES 71	CNTR
• • • We do not use the following data for averages, fits, limits, etc. • • •		
-14 ± 29	PETRUKHIN 68	CNTR
40 ± 70	BARDON 66	CNTR
23 ± 40	¹ LOBKOWICZ 66	CNTR
¹ This is the most conservative value given by LOBKOWICZ 66.		

π ELECTRIC POLARIZABILITY α_π

See HOLSTEIN 14 for a general review on hadron polarizability.

VALUE (10^{-4} fm^3)	EVTs	DOCUMENT ID	TECN	COMMENT
2.0 ± 0.6 ± 0.7	63k	¹ ADOLPH 15A	SPEC	$\pi^- \gamma \rightarrow \pi^- \gamma$ Compton scatt.
¹ Value is derived assuming $\alpha_\pi = -\beta_\pi$.				

π^+ DECAY MODES

π^- modes are charge conjugates of the modes below.

For decay limits to particles which are not established, see the section on Searches for Axions and Other Very Light Bosons.

Mode	Fraction (Γ_i/Γ)	Confidence level
$\Gamma_1 \quad \mu^+ \nu_\mu$	[a] (99.98770 ± 0.00004) %	
$\Gamma_2 \quad \mu^+ \nu_\mu \gamma$	[b] (2.00 ± 0.25) × 10 ⁻⁴	
$\Gamma_3 \quad e^+ \nu_e$	[a] (1.230 ± 0.004) × 10 ⁻⁴	
$\Gamma_4 \quad e^+ \nu_e \gamma$	[b] (7.39 ± 0.05) × 10 ⁻⁷	
$\Gamma_5 \quad e^+ \nu_e \pi^0$	(1.036 ± 0.006) × 10 ⁻⁸	
$\Gamma_6 \quad e^+ \nu_e e^+ e^-$	(3.2 ± 0.5) × 10 ⁻⁹	
$\Gamma_7 \quad \mu^+ \nu_\mu \nu \bar{\nu}$	< 9	× 10 ⁻⁶ 90%
$\Gamma_8 \quad e^+ \nu_e \nu \bar{\nu}$	< 1.6	× 10 ⁻⁷ 90%

Lepton Family number (*LF*) or Lepton number (*L*) violating modes

$\Gamma_9 \quad \mu^+ \bar{\nu}_e$	<i>L</i> [c] < 1.5	× 10 ⁻³ 90%
$\Gamma_{10} \quad \mu^+ \nu_e$	<i>LF</i> [c] < 8.0	× 10 ⁻³ 90%
$\Gamma_{11} \quad \mu^- e^+ e^+ \nu$	<i>LF</i> < 1.6	× 10 ⁻⁶ 90%

[a] Measurements of $\Gamma(e^+ \nu_e)/\Gamma(\mu^+ \nu_\mu)$ always include decays with γ 's, and measurements of $\Gamma(e^+ \nu_e \gamma)$ and $\Gamma(\mu^+ \nu_\mu \gamma)$ never include low-energy γ 's. Therefore, since no clean separation is possible, we consider the modes with γ 's to be subreactions of the modes without them, and let $[\Gamma(e^+ \nu_e) + \Gamma(\mu^+ \nu_\mu)]/\Gamma_{\text{total}} = 100\%$.

[b] See the Particle Listings below for the energy limits used in this measurement; low-energy γ 's are not included.

[c] Derived from an analysis of neutrino-oscillation experiments.

π^+ BRANCHING RATIOS

$$\Gamma(e^+ \nu_e)/\Gamma_{\text{total}}$$

$$\Gamma_3/\Gamma$$

See note [a] in the list of π^+ decay modes just above, and see also the next block of data. See also the note on "Decay Constants of Charged Pseudoscalar Mesons" in the D_s^+ Listings.

VALUE (units 10^{-4})	DOCUMENT ID
1.230 ± 0.004 OUR EVALUATION	

NODE=S008T;LINKAGE=A

NODE=S008T;LINKAGE=N

NODE=S008DT

NODE=S008DT

NODE=S008DT

NODE=S008DT;LINKAGE=L

NODE=S008EPL

NODE=S008EPL

NODE=S008EPL

NODE=S008EPL;LINKAGE=A

NODE=S008230;NODE=S008

NODE=S008

DESIG=1;OUR EVAL;→ UNCHECKED ←

DESIG=3

DESIG=2

DESIG=5

DESIG=4

DESIG=6

DESIG=11

DESIG=10

NODE=S008;CLUMP=A

DESIG=7

DESIG=8

DESIG=9

LINKAGE=LX2

LINKAGE=LX1

LINKAGE=CL

NODE=S008235

NODE=S008R2

NODE=S008R2

NODE=S008R2

→ UNCHECKED ←

$$[\Gamma(e^+\nu_e) + \Gamma(e^+\nu_e\gamma)] / [\Gamma(\mu^+\nu_\mu) + \Gamma(\mu^+\nu_\mu\gamma)] \quad (\Gamma_3 + \Gamma_4) / (\Gamma_1 + \Gamma_2)$$

See note [a] in the list of π^+ decay modes above. See NUMAO 92 for a discussion of e - μ universality. See also the note on “Decay Constants of Charged Pseudoscalar Mesons” in the D_s^+ Listings.

VALUE (units 10^{-4})	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
1.2327 ± 0.0023 OUR AVERAGE					
1.2344 ± 0.0023 ± 0.0019	400k	AGUILAR-AR... 15	CNTR	+	Stopping π^+
1.2346 ± 0.0035 ± 0.0036	120k	CZAPEK 93	CALO		Stopping π^+
1.2265 ± 0.0034 ± 0.0044	190k	BRITTON 92	CNTR		Stopping π^+
1.218 ± 0.014	32k	BRYMAN 86	CNTR		Stopping π^+
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.273 ± 0.028	11k	¹ DICAPUA 64	CNTR		
1.21 ± 0.07		ANDERSON 60	SPEC		

¹ DICAPUA 64 has been updated using the current mean life.

$$\Gamma(\mu^+\nu_\mu\gamma) / \Gamma_{\text{total}}$$

$$\Gamma_2 / \Gamma$$

Note that measurements here do not cover the full kinematic range.

VALUE (units 10^{-4})	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
2.0 ± 0.24 ± 0.08					
		¹ BRESSI 98	CALO	+	Stopping π^+
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.24 ± 0.25	26	CASTAGNOLI 58	EMUL		$KE_\mu < 3.38 \text{ MeV}$

¹ BRESSI 98 result is given for $E_\gamma > 1 \text{ MeV}$ only. Result agrees with QED expectation, 2.283×10^{-4} and does not confirm discrepancy of earlier experiment CASTAGNOLI 58.

$$\Gamma(e^+\nu_e\gamma) / \Gamma_{\text{total}}$$

$$\Gamma_4 / \Gamma$$

The very different values reflect the very different kinematic ranges covered (bigger range, bigger value). And none of them covers the whole kinematic range.

VALUE (units 10^{-8})	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
73.86 ± 0.54					
	65k	¹ BYCHKOV 09	PIBE		$e^+\nu\gamma$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •					
16.1 ± 2.3		² BOLOTOV 90B	SPEC		17 GeV $\pi^- \rightarrow e^- \bar{\nu}_e \gamma$
5.6 ± 0.7	226	³ STETZ 78	SPEC		$P_e > 56 \text{ MeV}/c$
3.0	143	DEPOMMIER 63B	CNTR		$(KE)_{e^+\gamma} > 48 \text{ MeV}$

¹ This BYCHKOV 09 value is for $E_\gamma > 10 \text{ MeV}$ and $\Theta_{e^+\gamma} > 40^\circ$.

² BOLOTOV 90B is for $E_\gamma > 21 \text{ MeV}$, $E_e > 70 - 0.8 E_\gamma$.

³ STETZ 78 is for an $e^-\gamma$ opening angle $> 132^\circ$. Obtains 3.7 when using same cutoffs as DEPOMMIER 63B.

$$\Gamma(e^+\nu_e\pi^0) / \Gamma_{\text{total}}$$

$$\Gamma_5 / \Gamma$$

VALUE (units 10^{-8})	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
1.036 ± 0.006 OUR AVERAGE					
1.036 ± 0.006	64k	^{1,2} POCANIC 04	PIBE	+	π decay at rest
1.026 ± 0.039	1224	³ MCFARLANE 85	CNTR	+	Decay in flight
1.00 +0.08 -0.10	332	DEPOMMIER 68	CNTR	+	
1.07 ± 0.21	38	⁴ BACASTOW 65	OSPK	+	
1.10 ± 0.26		⁴ BERTRAM 65	OSPK	+	
1.1 ± 0.2	43	⁴ DUNAITSEV 65	CNTR	+	
0.97 ± 0.20	36	⁴ BARTLETT 64	OSPK	+	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
1.15 ± 0.22	52	⁴ DEPOMMIER 63	CNTR	+	See DEPOMMIER 68

¹ POCANIC 04 normalizes to $e^+\nu_e$ decays, using the PDG 2004 value $B(\pi^+ \rightarrow e^+\nu_e) = (1.230 \pm 0.004) \times 10^{-4}$. We add their statistical (0.004×10^{-8}) , systematic (0.004×10^{-8}) and systematic error due to the uncertainty of $B(\pi^+ \rightarrow e^+\nu_e)$ (0.003×10^{-8}) in quadrature.

² This result can be used to calculate V_{ud} from pion beta decay: $V_{ud}^{PIBETA} = 0.9728 \pm 0.0030$.

³ MCFARLANE 85 combines a measured rate $(0.394 \pm 0.015)/s$ with 1982 PDG mean life.

⁴ DEPOMMIER 68 says the result of DEPOMMIER 63 is at least 10% too large because of a systematic error in the π^0 detection efficiency, and that this may be true of all the previous measurements (also V. Soergel, private communication, 1972).

NODE=S008R10

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NODE=S008R10;LINKAGE=D

NODE=S008R1

NODE=S008R1

NODE=S008R1

NODE=S008R1;LINKAGE=RB

NODE=S008R4

NODE=S008R4

NODE=S008R4

NODE=S008R4;LINKAGE=BY

NODE=S008R4;LINKAGE=A

NODE=S008R4;LINKAGE=S

NODE=S008R3

NODE=S008R3

NODE=S008R3;LINKAGE=PC

NODE=S008R3;LINKAGE=PO

NODE=S008R3;LINKAGE=M

NODE=S008R3;LINKAGE=D

$$\Gamma(e^+ \nu_e e^+ e^-) / \Gamma(\mu^+ \nu_\mu)$$

 Γ_6/Γ_1

VALUE (units 10^{-9})	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$3.2 \pm 0.5 \pm 0.2$	98		EGLI	89	SPEC Uses $R_{PCAC} = 0.068 \pm 0.004$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.46 \pm 0.16 \pm 0.07$	7	¹ BARANOV	92	SPEC	Stopped π^+
< 4.8	90	KORENCHE...	76B	SPEC	
< 34	90	KORENCHE...	71	OSPK	

¹ This measurement by BARANOV 92 is of the structure-dependent part of the decay. The value depends on values assumed for ratios of form factors.

NODE=S008R5;LINKAGE=A

$$\Gamma(\mu^+ \nu_\mu \nu \bar{\nu}) / \Gamma(\mu^+ \nu_\mu)$$

 Γ_7/Γ_1

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 8.6 \times 10^{-6}$	90	9.1M	AGUILAR-AR...20A	SPEC	fit E_μ spectrum

NODE=S008R01

NODE=S008R01

OCCUR=2

$$\Gamma(e^+ \nu_e \nu \bar{\nu}) / \Gamma_{\text{total}}$$

 Γ_8/Γ

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 1.6 \times 10^{-7}$	90	1.3M	AGUILAR-AR...20A	SPEC	fit E_e spectrum

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 5 \times 10^{-6}$	90	PICCIOTTO	88	SPEC	
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NODE=S008R9

NODE=S008R9

$$\Gamma(\mu^+ \bar{\nu}_e) / \Gamma_{\text{total}}$$

 Γ_9/Γ

Forbidden by total lepton number conservation. See the note on "Decay Constants of Charged Pseudoscalar Mesons" in the D_s^+ Listings.

VALUE (units 10^{-3})	CL%	DOCUMENT ID	TECN	COMMENT
< 1.5	90	¹ COOPER	82	HLBC Wideband ν beam

¹ COOPER 82 limit on $\bar{\nu}_e$ observation is here interpreted as a limit on lepton number violation.

NODE=S008R6

NODE=S008R6

NODE=S008R6

NODE=S008R6;LINKAGE=L

$$\Gamma(\mu^+ \nu_e) / \Gamma_{\text{total}}$$

 Γ_{10}/Γ

Forbidden by lepton family number conservation.

VALUE (units 10^{-3})	CL%	DOCUMENT ID	TECN	COMMENT
< 8.0	90	¹ COOPER	82	HLBC Wideband ν beam

¹ COOPER 82 limit on ν_e observation is here interpreted as a limit on lepton family number violation.

NODE=S008R7

NODE=S008R7

NODE=S008R7

NODE=S008R7;LINKAGE=L

$$\Gamma(\mu^- e^+ e^+ \nu) / \Gamma_{\text{total}}$$

 Γ_{11}/Γ

Forbidden by lepton family number conservation.

VALUE (units 10^{-6})	CL%	DOCUMENT ID	TECN	CHG
< 1.6	90	BARANOV	91B	SPEC +

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 7.7	90	KORENCHE...	87	SPEC +
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NODE=S008R8

NODE=S008R8

NODE=S008R8

π^+ — POLARIZATION OF EMITTED μ^+

NODE=S008240

$$\pi^+ \rightarrow \mu^+ \nu$$

Tests the Lorentz structure of leptonic charged weak interactions.

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

< (-0.9959)	90	¹ FETSCHER	84	RVUE +	
-0.99 ± 0.16		² ABELA	83	SPEC -	μ X-rays

¹ FETSCHER 84 uses only the measurement of CARR 83.

² Sign of measurement reversed in ABELA 83 to compare with μ^+ measurements.

NODE=S008POL

NODE=S008POL

NODE=S008POL

NODE=S008POL;LINKAGE=F

NODE=S008POL;LINKAGE=A

See the related review(s):

Form Factors for Semileptonic Kaon ($K_{\ell 3}$), Radiative Pion ($\pi_{\ell 2\gamma}$) and Kaon ($K_{\ell 2\gamma}$) Decays

$\pi^\pm$ FORM FACTORS F_V , VECTOR FORM FACTOR

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.0254 ± 0.0017 OUR AVERAGE				
0.0258 ± 0.0017	65k	¹ BYCHKOV	09 PIPE	$e^+ \nu \gamma$ at rest
0.014 ± 0.009		² BOLOTOV	90B SPEC	17 GeV $\pi^- \rightarrow e^- \bar{\nu}_e \gamma$
0.023 $\begin{smallmatrix} +0.015 \\ -0.013 \end{smallmatrix}$	98	EGLI	89 SPEC	$\pi^+ \rightarrow e^+ \nu_e e^+ e^-$

¹ The BYCHKOV 09 F_A and F_V results are highly (anti-)correlated: $F_A + 1.0286 F_V = 0.03853 \pm 0.00014$.

² BOLOTOV 90B only determines the absolute value.

 F_A , AXIAL-VECTOR FORM FACTOR

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.0119 ± 0.0001	65k	^{1,2} BYCHKOV	09 PIPE	$e^+ \nu \gamma$ at rest
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0115 ± 0.0004	41k	^{1,3} FRLEZ	04 PIPE	$\pi^+ \rightarrow e^+ \nu \gamma$ at rest
0.0106 ± 0.0060		^{1,4} BOLOTOV	90B SPEC	17 GeV $\pi^- \rightarrow e^- \bar{\nu}_e \gamma$
0.021 $\begin{smallmatrix} +0.011 \\ -0.013 \end{smallmatrix}$	98	EGLI	89 SPEC	$\pi^+ \rightarrow e^+ \nu_e e^+ e^-$
0.0135 ± 0.0016		^{1,4} BAY	86 SPEC	$\pi^+ \rightarrow e^+ \nu \gamma$
0.006 ± 0.003		^{1,4} PIILONEN	86 SPEC	$\pi^+ \rightarrow e^+ \nu \gamma$
0.011 ± 0.003		^{1,4,5} STETZ	78 SPEC	$\pi^+ \rightarrow e^+ \nu \gamma$

¹ These values come from fixing the vector form factor at the CVC prediction, $F_V = 0.0259 \pm 0.0005$.

² When F_V is released, the BYCHKOV 09 F_A is 0.0117 ± 0.0017 , and F_A and F_V results are highly (anti-)correlated: $F_A + 1.0286 F_V = 0.03853 \pm 0.00014$.

³ The sign of $\gamma = F_A / F_V$ is determined to be positive.

⁴ Only the absolute value of F_A is determined.

⁵ The result of STETZ 78 has a two-fold ambiguity. We take the solution compatible with later determinations.

VECTOR FORM FACTOR SLOPE PARAMETER a

This is a in $F_V(q^2) = F_V(0) (1 + a q^2)$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.10 ± 0.06	65k	BYCHKOV	09 PIPE	$e^+ \nu \gamma$ at rest

 R , SECOND AXIAL-VECTOR FORM FACTOR

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
0.059 $\begin{smallmatrix} +0.009 \\ -0.008 \end{smallmatrix}$	98	EGLI	89 SPEC	$\pi^+ \rightarrow e^+ \nu_e e^+ e^-$

 $\pi^\pm$ CHARGE RADIUS

The charge radius of the pion $\sqrt{\langle r_\pi^2 \rangle}$ is defined in relation to the form factor of the pion electromagnetic vertex, called vector form factor VFF, F_V^π . The VFF is a function of the squared four-momentum transfer t , or of the squared c.m. energy s , depending on the channel in which the photon exchange takes place. In both cases, it is related to the slope of the VFF at zero, namely

$$\langle r_\pi^2 \rangle = 6 \frac{dF_V^\pi(q)}{dq} (q=0) \text{ where } q = t, s.$$

The quantity cannot be measured directly. It can be extracted from the cross sections of three processes: pion electroproduction, $eN \rightarrow eN\pi$, and pion electron scattering $e\pi \rightarrow e\pi$, for the t channel, and positron electron annihilation into two charged pions, $e^+e^- \rightarrow \pi^+\pi^-$, for the s channel. We encode all measurements, but we do not use electroproduction data in averaging because the extraction of the pion radius involves, in this case, theoretical uncertainties that cannot be controlled at the needed level of accuracy. In case of analyses based on the same data set, as ANANTHANARAYAN 17 and COLANGELO 19, which cannot be averaged, we combine the results into a common value, with the uncertainty range chosen to cover both analyses. Note that for consistency the form factor needs to be defined in both channels with the vacuum polarisation removed. For details see COLANGELO 19 or Appendix B of ANANTHANARAYAN 16A.

VALUE (fm)	DOCUMENT ID	TECN	COMMENT
0.659 ± 0.004 OUR AVERAGE			
0.656 ± 0.005	¹ PDG	19 FIT	
0.65 ± 0.05 ± 0.06	ESCHRICH	01 CNTR	$\pi e \rightarrow \pi e$
0.663 ± 0.006	AMENDOLIA	86 CNTR	$\pi e \rightarrow \pi e$
0.663 ± 0.023	DALLY	82 CNTR	$\pi e \rightarrow \pi e$

NODE=S008245

NODE=S008FV
NODE=S008FV

NODE=S008FV;LINKAGE=BY

NODE=S008FV;LINKAGE=A

NODE=S008FA
NODE=S008FA

NODE=S008FA;LINKAGE=FL

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NODE=S008FA;LINKAGE=B

NODE=S008FSV

NODE=S008FSV
NODE=S008FSVNODE=S008FR
NODE=S008FR

NODE=S008CR

NODE=S008CR

NODE=S008CR

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.640 ± 0.007	² CUI	21A	FIT	Fit existing data
0.655 ± 0.004	³ COLANGELO	19	FIT	Fit existing data
0.657 ± 0.003	⁴ ANANTHANA..17	FIT	Fit existing data	
0.6603 ± 0.0005 ± 0.0004	⁵ HANHART	17	FIT	Fit existing data
0.740 ± 0.031	⁶ LIESENFELD	99	CNTR	$ep \rightarrow e\pi^+ n$
0.661 ± 0.012	⁷ BIJNENS	98	CNTR	χ PT extraction
0.660 ± 0.024	AMENDOLIA	84	CNTR	$\pi e \rightarrow \pi e$
0.711 ± 0.009 ± 0.016	⁶ BEBEK	78	CNTR	$eN \rightarrow e\pi N$
0.678 ± 0.004 ± 0.008	⁸ QUENZER	78	CNTR	$e^+e^- \rightarrow \pi^+\pi^-$
0.78 ^{+0.09} _{-0.10}	ADYLOV	77	CNTR	$\pi e \rightarrow \pi e$
0.74 ^{+0.11} _{-0.13}	BARDIN	77	CNTR	$ep \rightarrow e\pi^+ n$
0.56 ± 0.04	DALLY	77	CNTR	$\pi e \rightarrow \pi e$

¹ This value combines the measurements of ANANTHANARAYAN 17 and COLANGELO 19 which are based on the same data set. The uncertainty range is chosen to cover both results.

² CUI 21A perform a fit including spacelike data only. Employ a new mathematical procedure based on interpolation via continued fractions augmented by statistical sampling. Also do not impose the charge conserving normalization condition $F(0) = 1$.

³ COLANGELO 19 fit existing F_V data, using an extended Omnes dispersive representation. This analysis is based on the same data set of ANANTHANARAYAN 17. Accordingly, they cannot be averaged. We combine the results into a common value, with the uncertainty range chosen to cover the uncertainty ranges of both analyses.

⁴ ANANTHANARAYAN 17 fit existing F_V data, using a mixed phase-modulus dispersive representation. This analysis is based on the same data set of COLANGELO 19. Accordingly, they cannot be averaged. We combine the results into a common value, with the uncertainty range chosen to cover the uncertainty ranges of both analyses.

⁵ According to the authors the uncertainty could be underestimated. The value quoted omits the BaBar data AUBERT 09.

⁶ The extractions could contain an additional theoretical uncertainty which cannot be sufficiently quantified.

⁷ BIJNENS 98 fits existing data.

⁸ The extraction is based on a parametrization that does not have correct analytic properties.

NODE=S008CR;LINKAGE=E

NODE=S008CR;LINKAGE=H

NODE=S008CR;LINKAGE=F

NODE=S008CR;LINKAGE=A

NODE=S008CR;LINKAGE=D

NODE=S008CR;LINKAGE=B

NODE=S008CR;LINKAGE=RK

NODE=S008CR;LINKAGE=C

$\pi^\pm$ REFERENCES

NODE=S008

We have omitted some papers that have been superseded by later experiments. The omitted papers may be found in our 1988 edition Physics Letters **B204** 1 (1988).

NODE=S008

CUI	21A	PL B822 136631	Z.-F. Cui <i>et al.</i>	(NJU, ECT, HZDR)	REFID=61548
AGUILAR-AR...	20A	PR D102 012001	A. Aguilar-Arevalo <i>et al.</i>	(PIENU Collab.)	REFID=60332
COLANGELO	19	JHEP 1902 006	G. Colangelo, M. Hoferichter, P. Stoffer	(BERN+)	REFID=59585
DAUM	19	PL B796 11	M. Daum, R. Frosch, P.-R. Kettle	(PSI)	REFID=59977
PDG	19	RPP 2019 at pdg.lbl.gov	M. Tanabashi <i>et al.</i>	(PDG Collab.)	REFID=59587
ANANTHANA...	17	PRL 119 132002	B. Ananthanarayan, I. Caprini, D. Das		REFID=58205
HANHART	17	EPJ C77 98	C. Hanhart <i>et al.</i>		REFID=57839
ANANTHANA...	16A	PR D93 116007	B. Ananthanarayan <i>et al.</i>		REFID=59586
TRASSINELLI	16	PL B759 583	M. Trassinelli <i>et al.</i>		REFID=57155
ADOLPH	15A	PRL 114 062002	C. Adolph <i>et al.</i>	(COMPASS Collab.)	REFID=56443
AGUILAR-AR...	15	PRL 115 071801	A. Aguilar-Arevalo <i>et al.</i>	(PIENU Collab.)	REFID=56490
HOLSTEIN	14	ARNPS 64 51	B. Holstein, S. Scherer	(MASA, MAINZ)	REFID=57150
AUBERT	09	PR D79 011102	B. Aubert <i>et al.</i>	(BABAR Collab.)	REFID=52639
BYCHKOV	09	PRL 103 051802	M. Bychkov <i>et al.</i>	(PSI PIBETA Collab.)	REFID=52934
FRLEZ	04	PRL 93 181804	E. Frlez <i>et al.</i>	(PSI PIBETA Collab.)	REFID=50262
POCANIC	04	PRL 93 181803	D. Pocanic <i>et al.</i>	(PSI PIBETA Collab.)	REFID=50261
ESCHRICH	01	PL B522 233	I. Eschrich <i>et al.</i>	(FNAL SELEX Collab.)	REFID=48455
LIESENFELD	99	PL B468 20	A. Liesenfeld <i>et al.</i>		REFID=48590
BIJNENS	98	JHEP 9805 014	J. Bijnens <i>et al.</i>		REFID=48591
BRESSI	98	NP B513 555	G. Bressi <i>et al.</i>		REFID=46136
LENZ	98	PL B416 50	S. Lenz <i>et al.</i>		REFID=45832
ASSAMAGAN	96	PR D53 6065	K.A. Assamagan <i>et al.</i>	(PSI, ZURI, VILL+)	REFID=44719
KOPTEV	95	JETPL 61 877	V.P. Koptev <i>et al.</i>	(PNPI)	REFID=44395
		Translated from ZETFP 61 865.			
NUMAO	95	PR D52 4855	T. Numao <i>et al.</i>	(TRIU, BRICO)	REFID=44480
ASSAMAGAN	94	PL B335 231	K.A. Assamagan <i>et al.</i>	(PSI, ZURI, VILL+)	REFID=43832
JECKELMANN	94	PL B335 326	B. Jeckelmann, P.F.A. Goudsmit, H.J. Leisi	(WABRN+)	REFID=43827
CZAPEK	93	PRL 70 17	G. Czapek <i>et al.</i>	(BERN, VILL)	REFID=43126
BARANOV	92	SJNP 55 1644	V.A. Baranov <i>et al.</i>	(JINR)	REFID=43145
		Translated from YAF 55 2940.			
BRITTON	92	PRL 68 3000	D.I. Britton <i>et al.</i>	(TRIU, CARL)	REFID=41994
Also		PR D49 28	D.I. Britton <i>et al.</i>	(TRIU, CARL)	REFID=43753
NUMAO	92	MPL A7 3357	T. Numao	(TRIU)	REFID=43155
BARANOV	91B	SJNP 54 790	V.A. Baranov <i>et al.</i>	(JINR)	REFID=41949
		Translated from YAF 54 1298.			

DAUM	91	PL B265 425	M. Daum <i>et al.</i>	(VILL)	REFID=41608
BOLOTOV	90B	PL B243 308	V.N. Bolotov <i>et al.</i>	(INRM)	REFID=41319
EGLI	89	PL B222 533	S. Egli <i>et al.</i>	(SINDRUM Collab.)	REFID=40818
Also		PL B175 97	S. Egli <i>et al.</i>	(AACH3, ETH, SIN, ZURI)	REFID=10686
PDG	88	PL B204 1	G.P. Yost <i>et al.</i>	(LBL+)	REFID=40022
PICCIOTTO	88	PR D37 1131	C.E. Picciotto <i>et al.</i>	(TRIU, CNRC)	REFID=40667
COHEN	87	RMP 59 1121	E.R. Cohen, B.N. Taylor	(RISC, NBS)	REFID=11616
KORENCHE...	87	SJNP 46 192	S.M. Korenchenko <i>et al.</i>	(JINR)	REFID=40461
		Translated from YAF 46 313.			
AMENDOLIA	86	NP B277 168	S.R. Amendolia <i>et al.</i>	(CERN NA7 Collab.)	REFID=48592
BAY	86	PL B174 445	A. Bay <i>et al.</i>	(LAUS, ZURI)	REFID=10684
BRYMAN	86	PR D33 1211	D.A. Bryman <i>et al.</i>	(TRIU, CNRC)	REFID=10685
Also		PRL 50 7	D.A. Bryman <i>et al.</i>	(TRIU, CNRC)	REFID=10418
JECKELMANN	86B	NP A457 709	B. Jeckelmann <i>et al.</i>	(ETH, FRIB)	REFID=43813
Also		PRL 56 1444	B. Jeckelmann <i>et al.</i>	(ETH, FRIB)	REFID=10140
PIILONEN	86	PRL 57 1402	L.E. Piilonen <i>et al.</i>	(LANL, TEMP, CHIC)	REFID=10688
McFARLANE	85	PR D32 547	W.K. McFarlane <i>et al.</i>	(TEMP, LANL)	REFID=10683
ABELA	84	PL 146B 431	R. Abela <i>et al.</i>	(SIN)	REFID=10139
Also		PL 74B 126	M. Daum <i>et al.</i>	(SIN)	REFID=10681
Also		PR D20 2692	M. Daum <i>et al.</i>	(SIN)	REFID=10682
AMENDOLIA	84	PL 146B 116	S.R. Amendolia <i>et al.</i>	(CERN NA7 Collab.)	REFID=48593
FETSCHER	84	PL 140B 117	W. Fetscher	(ETH)	REFID=10679
ABELA	83	NP A395 413	R. Abela <i>et al.</i>	(BASL, KARLK, KARLE)	REFID=10675
CARR	83	PRL 51 627	J. Carr <i>et al.</i>	(LBL, NWES, TRIU)	REFID=10240
COOPER	82	PL 112B 97	A.M. Cooper <i>et al.</i>	(RL)	REFID=10401
DALLY	82	PRL 48 375	E.B. Dally <i>et al.</i>		REFID=48594
LU	80	PRL 45 1066	D.C. Lu <i>et al.</i>	(YALE, COLU, JHU)	REFID=10136
BEBEK	78	PR D17 1693	C.J. Bebek <i>et al.</i>		REFID=48595
QUENZER	78	PL 76B 512	A. Quenzer <i>et al.</i>	(LALO)	REFID=20123
STETZ	78	NP B138 285	A.W. Stetz <i>et al.</i>	(LBL, UCLA)	REFID=10672
ADYLOV	77	NP B128 461	G.T. Adylov <i>et al.</i>		REFID=48596
BARDIN	77	NP B120 45	G. Bardin <i>et al.</i>		REFID=48597
DALLY	77	PRL 39 1176	E.B. Dally <i>et al.</i>		REFID=48598
CARTER	76	PRL 37 1380	A.L. Carter <i>et al.</i>	(CARL, CNRC, CHIC+)	REFID=10667
KORENCHE...	76B	JETP 44 35	S.M. Korenchenko <i>et al.</i>	(JINR)	REFID=10668
		Translated from ZETF 71 69.			
MARUSHEN...	76	JETPL 23 72	V.I. Marushenko <i>et al.</i>	(PNPI)	REFID=10669
		Translated from ZETFP 23 80.			
Also		Private Comm.	R.E. Shafer	(FNAL)	REFID=10670
Also		Private Comm.	A. Smirnov	(PNPI)	REFID=10671
DUNAITSEV	73	SJNP 16 292	A.F. Dunaitsev <i>et al.</i>	(SERP)	REFID=10664
		Translated from YAF 16 524.			
AYRES	71	PR D3 1051	D.S. Ayres <i>et al.</i>	(LRL, UCSB)	REFID=10653
Also		PR 157 1288	D.S. Ayres <i>et al.</i>	(LRL)	REFID=10654
Also		PRL 21 261	D.S. Ayres <i>et al.</i>	(LRL, UCSB)	REFID=10655
Also		Thesis UCRL 18369	D.S. Ayres	(LRL)	REFID=10656
Also		PRL 23 1267	A.J. Greenberg <i>et al.</i>	(LRL, UCSB)	REFID=10657
KORENCHE...	71	SJNP 13 189	S.M. Korenchenko <i>et al.</i>	(JINR)	REFID=10660
		Translated from YAF 13 339.			
BOOTH	70	PL 32B 723	P.S.L. Booth <i>et al.</i>	(LIVP)	REFID=10652
DEPOMMIER	68	NP B4 189	P. Depommier <i>et al.</i>	(CERN)	REFID=10650
PETRUKHIN	68	JINR P1 3862	V.I. Petrukhin <i>et al.</i>	(JINR)	REFID=10651
HYMAN	67	PL 25B 376	L.G. Hyman <i>et al.</i>	(ANL, CMU, NWES)	REFID=10122
NORDBERG	67	PL 24B 594	M.E. Nordberg, F. Lobkowicz, R.L. Burman	(ROCH)	REFID=10647
BARDON	66	PRL 16 775	M. Bardon <i>et al.</i>	(COLU)	REFID=10642
KINSEY	66	PR 144 1132	K.F. Kinsey, F. Lobkowicz, M.E. Nordberg	(ROCH)	REFID=10644
LOBKOWICZ	66	PRL 17 548	F. Lobkowicz <i>et al.</i>	(ROCH, BNL)	REFID=10645
BACASTOW	65	PR 139 B407	R.B. Bacastow <i>et al.</i>	(LRL, SLAC)	REFID=10638
BERTRAM	65	PR 139 B617	W.K. Bertram <i>et al.</i>	(MICH, CMU)	REFID=10639
DUNAITSEV	65	JETP 20 58	A.F. Dunaitsev <i>et al.</i>	(JINR)	REFID=10640
		Translated from ZETF 47 84.			
ECKHAUSE	65	PL 19 348	M. Eckhause <i>et al.</i>	(WILL)	REFID=10641
BARTLETT	64	PR 136 B1452	D. Bartlett <i>et al.</i>	(COLU)	REFID=10635
DICAPUA	64	PR 133 B1333	M. di Capua <i>et al.</i>	(COLU)	REFID=10636
Also		Private Comm.	L. Pondrom	(WISC)	REFID=10637
DEPOMMIER	63	PL 5 61	P. Depommier <i>et al.</i>	(CERN)	REFID=10633
DEPOMMIER	63B	PL 7 285	P. Depommier <i>et al.</i>	(CERN)	REFID=10634
ANDERSON	60	PR 119 2050	H.L. Anderson <i>et al.</i>	(EFI)	REFID=10631
CASTAGNOLI	58	PR 112 1779	C. Castagnoli, M. Muchnik	(ROMA)	REFID=10630